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Preliminary studies on RPCs for SWGO

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RPCs for Cosmic Ray Physics

The Yangbajing Cosmic Ray Laboratory





Data taking with Very low maintenance: 2007 - 2012 Angular resolution **0.5° at 1 TeV** (shower front projection with time measurement) Energy Range: 100 GeV - 10 PeV 5600 m² of active area + 1100 m² guard ring







R&D for an RPC project for SWGO

In the framework of PNRR CTAplus, Working Package 1520, INAF

INFN will support the construction and test of an hybrid RPC + Water-Cerenkov of about 100-200 m2

Rome TorVergata group lead the RPC construction project starting from the Argo experience

Timeline for the project:

• Tenders closed by the end of 2023:

1) Front-End Read-Out per RPC production

2) HPL electrodes production for RPC

3) HV, LV e Read Out systems

4) Full assembled RPCs production-> N° 32 RPC total area 73 m²

Note: 1 Cluster of Argo 50 m² produced physics results (Astroparticle Physics 17 (2002) 151–165)

- •Test station preparation before April 2024
- •RPC Production and test: main fraction before end of 2024
- •Site definition for final test end of 2024
- •Installation and operations in 2025

Application for SWGO if approved by the Collaboration

Man Power:

Rome 2 TorVergata INFN: 1.5 FTE for 2024

LNF INFN: interest with the plan to create a small group for middle 2024 (1.5 FTE)



Main differences with Argo: 1) RPC operated in avalanche mode -> Advanced Linearity

Analog calorimetry with RPC operated in saturated avalanche regime



R. Bartoli et al. Intrinsic linearity of bakelite Resistive Plate Chambers operated in streamer mode. NIM Section A, 2019. A. Rocchi, R. Cardarelli, B. Liberti et al. JINST 15 (2020) 12, C12004. 10.1088/1748-0221/15/12/C12004

Main differences with Argo: 2) Gas in closed loop -> Ecogas Mixture and Eco Compatibility



Search for EcoGas Mixture started in 2014, HFO and CO2 based

Search for SF6 substitute, HFCl and other possible candidates

Low cost Closed Loop/ Recirculation System under study



ressure that was always kept by the as been reduced because of the fast current rising when the system was fully recovered. On day 15 an add of 400 cc of fresh gas was needed, to recover the intrinsic leak rate of the system. At day 18 the current rise again, probably due to the Oxygen level increasing, and the voltage was lowered again. After 2 days, surprisingly the chamber recovered, but since the oxygen level in the mixture was too high, one volume of fresh gas was exchanged. From day 20 for more than 5 days the system run very stable with a very low level of oxygen. We suspect that the previous oxygen increase was due to a water residual dissolved air. The last Oxygen increasing trend from day 26 must be further investigated, but it is probably due to a break in the system.



Main differences with Argo: 3) Improved electronics with lower threshold

- New Front-End adequate to avalanche-mode
- 1 Board 8 channels, LVDS Outputs \bullet

Discrete components version of a Full Custom circuit (dedicated to ATLAS activity)



R&D for an **RPC** project for **SWGO**

• **RPC chamber deisgn**





Goal of R&D

- Evaluate performance (efficiency and time resolution) with large strips
- Optimization of Front-end readout
- Mechanical layout evaluation





1 bigger pad to oscilloscope







System layout





System layout



Typical signal waveform ampl < Pedestal Region 0.005 -0.005 Strip signal window for charge estimation Pad signal window for charge estimation –0.01 –120 -60 -100 -80 -40



Efficient Event selection

Max amplitude middle strip in background region (ch2)



Integrated Charge middle strip in background region (ch2)



Cluster size

Cluster size





HV normalized at $P_0 = 1010 \text{ mbar}$ $T_0 = 20^\circ$

Integrated charge HV_{eff} = 9860 V **Strip fired**

Integrated Charge pad (ch3) strip Fired





Time resolution

Threshold time middle strip



Threshold time middle strip

Time resolution, cls, charge scan

HV normalized at $P_0 = 1010$ mbar $T_0 = 20^\circ$

System layout for Front end stdies

Test of Front End Board Developed by R. Cardarelli (TorVergata)

waveform

waveform

Threshold scan @9300 kV (9480 effective) vs 9500 kV (9625 effective)

Characterization of front end board started

- Noise under control (1%-2% of noisy events)
- Threshold studies ongoing

Threshold scan @ HV = 9500 kV (9625 effective)

Characterization of front end board started

- Efficiency sable after 185 mV
- Cluster size and time resolution almost stable

NOTE: **Chamber operated at lower efficiency**

Tor Vergata test Real Size Prototype (RP):

Just Arrived....

We are working!!!

Current-Voltage Characterization

- The Gap current is first QC for the gap
- @ WP (10,0 kV) lgap=2 uA
- Current density = 0,009 uA/m^2

Tor Vergata test Real Size Prototype (RP):

- Two different cabling, for matching strip size with FE pitch 1)
- Two RPC (One strip equivalent active area) Trigger 2)
- 3) VME-TDC based acquisition of LVDS signal

First Results before the end of the year

Conclusions

- Single gap 2 mm
- Strip size 7 cm
- Pad readout

Performance verified and in agreement with expected results with standard gas mixture • Efficiency (driven from geometrical acceptance)

- Strip charge (between 2 and 3 pC @ WP)
- cluster size (average value = 1.1)
- time resolution (better than 1.6 ns @ WP)

Performance of front end board started Study of cabling and layout for real size chamber operations

Next steps:

Characterization of Front End Board Start measurements on real size chamber

• First tests with small RPC segment in the SWGO configuration started few months ago

Backup

total strips Charge vs max amplitude middle strip

Integrated charge HV_{eff} = 10128 V

Integrated Charge pad (ch3) strip Fired

Strip fired

