



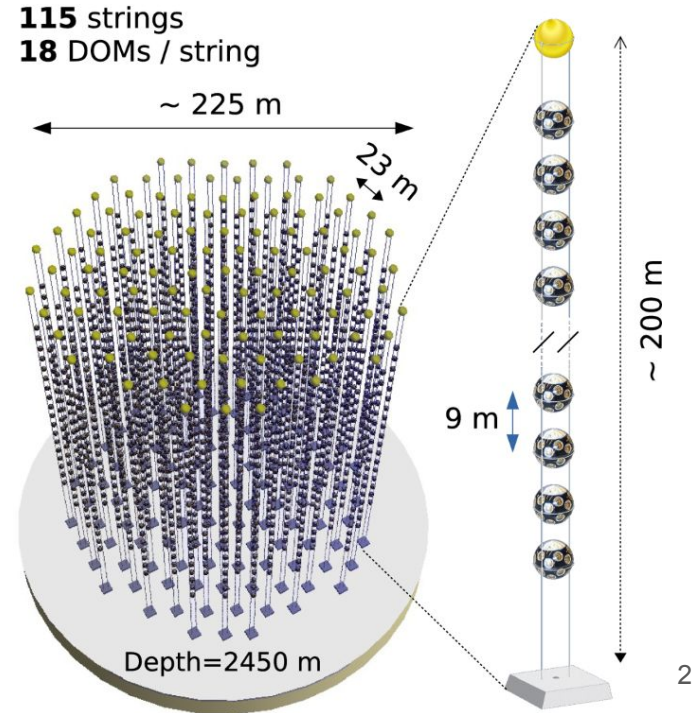
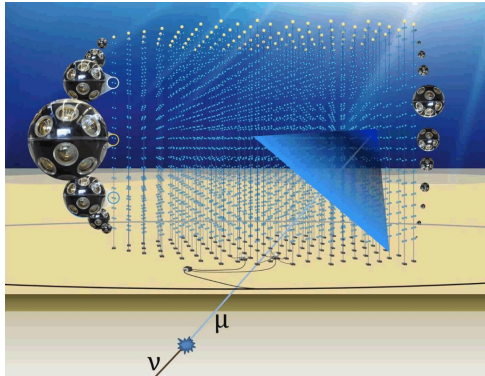
Sensitivities for DM search in the Galactic Center with KM3NeT/ORCA-6

Adriana Bariego-Quintana
Instituto de Física Corpuscular (CSIC-UV)

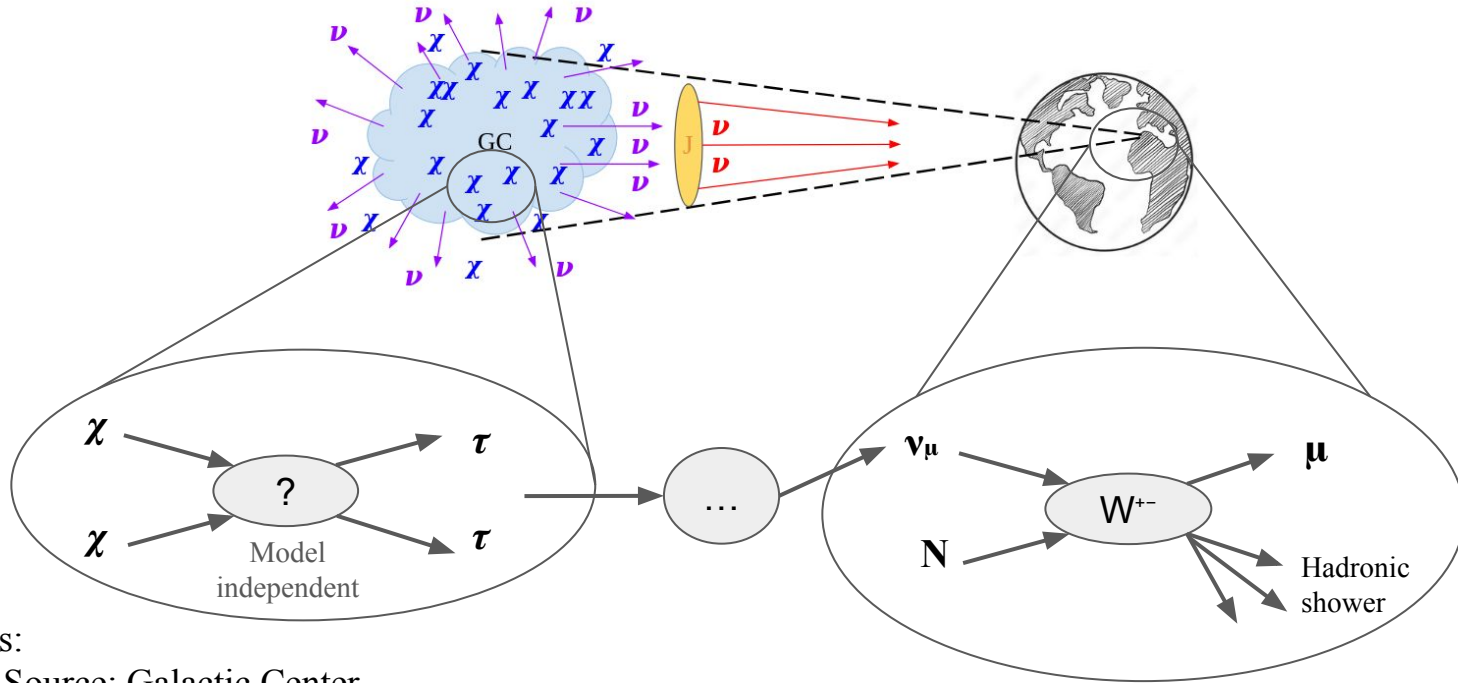
KM3NeT: ARCA and ORCA detectors.



- The KM3NeT infrastructure, located in abyssal sites of the Mediterranean Sea, is composed of two undersea Cherenkov neutrino telescopes: ORCA, a dense detector optimised for the measurement of low energy neutrinos, and ARCA, a cubic kilometer detector, intended for low fluxes of astrophysical neutrinos.
- It is composed of Digital Optical Modules with a segmented photo-sensitive area provided by 31 photomultiplier tubes (PMTs), arranged in vertical strings called Detection Units.
- The Cherenkov light is induced by ultra-relativistic charged particles that are produced in neutrino interactions near the detector, this light is measured in the PMTs.
- 23 DUs have been already deployed at the ORCA site and 28 at ARCA.



Theoretical framework.

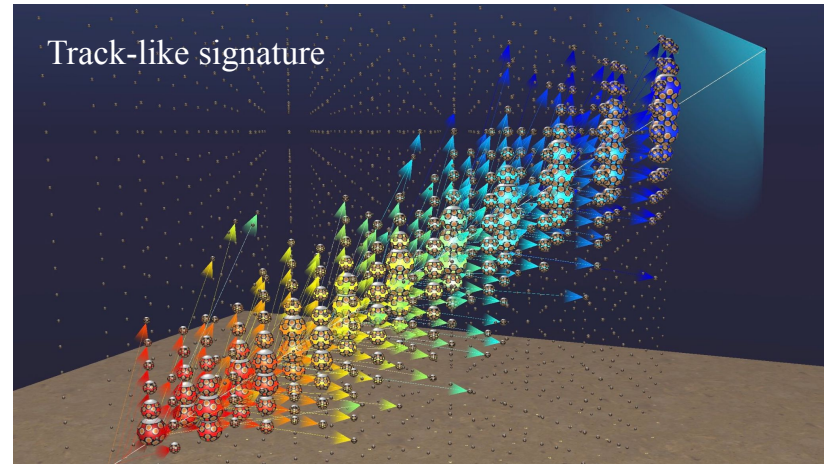


Details:

- Source: Galactic Center
- Density Profile: Navarro-Frenk-White (cuspy)
- Halo shape: spherical
- Model independent approach
- DM mass from 50 to 10 000 GeV/c²
- Annihilation channels $\mu^+ \mu^-$, $\tau^+ \tau^-$, $b\bar{b}$, $W^+ W^-$, HH , $\nu\bar{\nu}$

Data Set.

- Data set: ICRC 2023 ORCA-6 data set
- Detector configuration: 6 detection units of ORCA
- Lifetime: 510 days of data taking
- Background: atmospheric muons
- Event selection:
 - Track events are considered $\nu_\mu + X \xrightarrow{CC} \mu + Y$
 - Selection cuts to reject atmospheric muons
 - Anti-noise cuts: upgoing events, likelihood of the reconstruction, number of hits of events.
 - Particle identity cuts for tracks: probability of event being muon
 - 4445 events in the data sample survive the cuts.



Analysis Method.

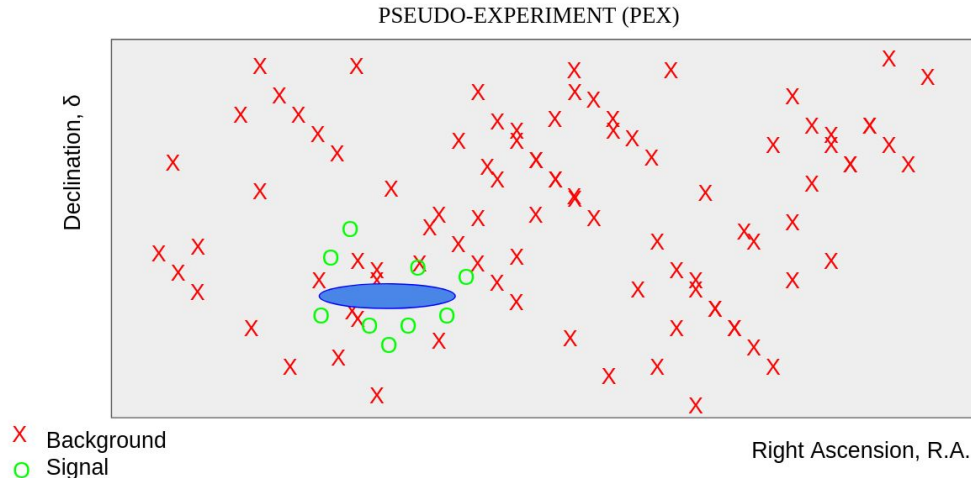


$$-\log \mathcal{L} = - \prod_{i=1}^{N_{events}} \log [n_{sg} P_{sg}(\psi_i, E_i) + (N_{events} - n_{sg}) P_{bg}(\delta_i, E_i)]$$

Probability of an event to be signal
events characterised by angular distance and energy

Probability of an event to be background
events characterised by declination and energy

We create signal and background event coordinates and energies in pseudo-experiments. Signal and background probability functions are used to determine the amount of signal in each pseudo-experiment. We maximise the likelihood to fit the number of signal events that we are sensitive to



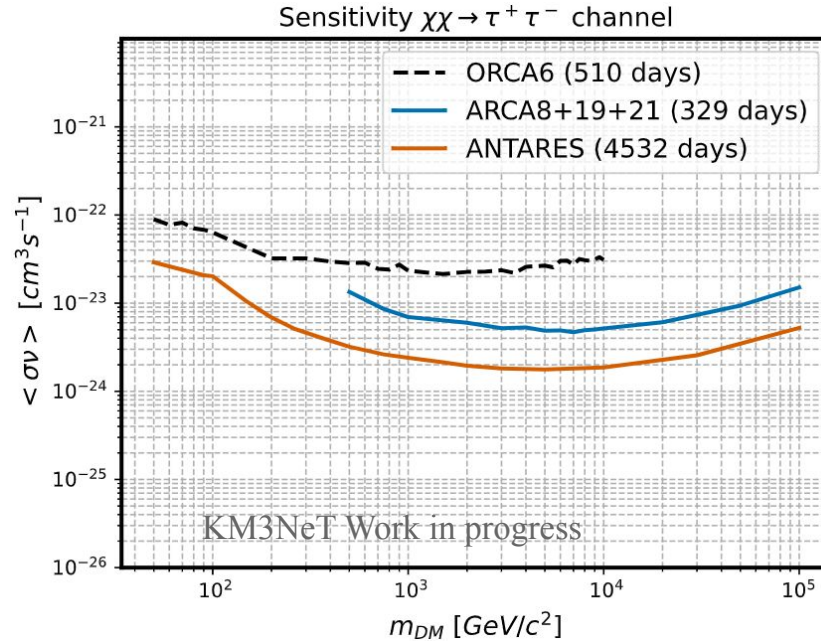
Sensitivity.



For the first time the ORCA-6 data set has been used to search for DM annihilating in the GC. We have obtained the sensitivity of this detector configuration considering WIMP masses in the range 50 GeV/c² to 10 000 GeV/c²

$$\langle\sigma v\rangle = \frac{n_{90}}{Acc \cdot T} \cdot \frac{8\pi m_{DM}^2}{J}$$

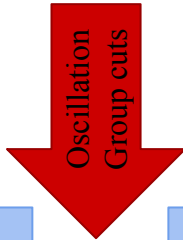
In addition, the combination of ARCA-8+19+21 has been used to obtain the 90% upper limit of the thermal cross section for the DM annihilating at the GC. This result is compatible with background.



Thank you for your time.
-Adriana-

BACKUP SLIDES

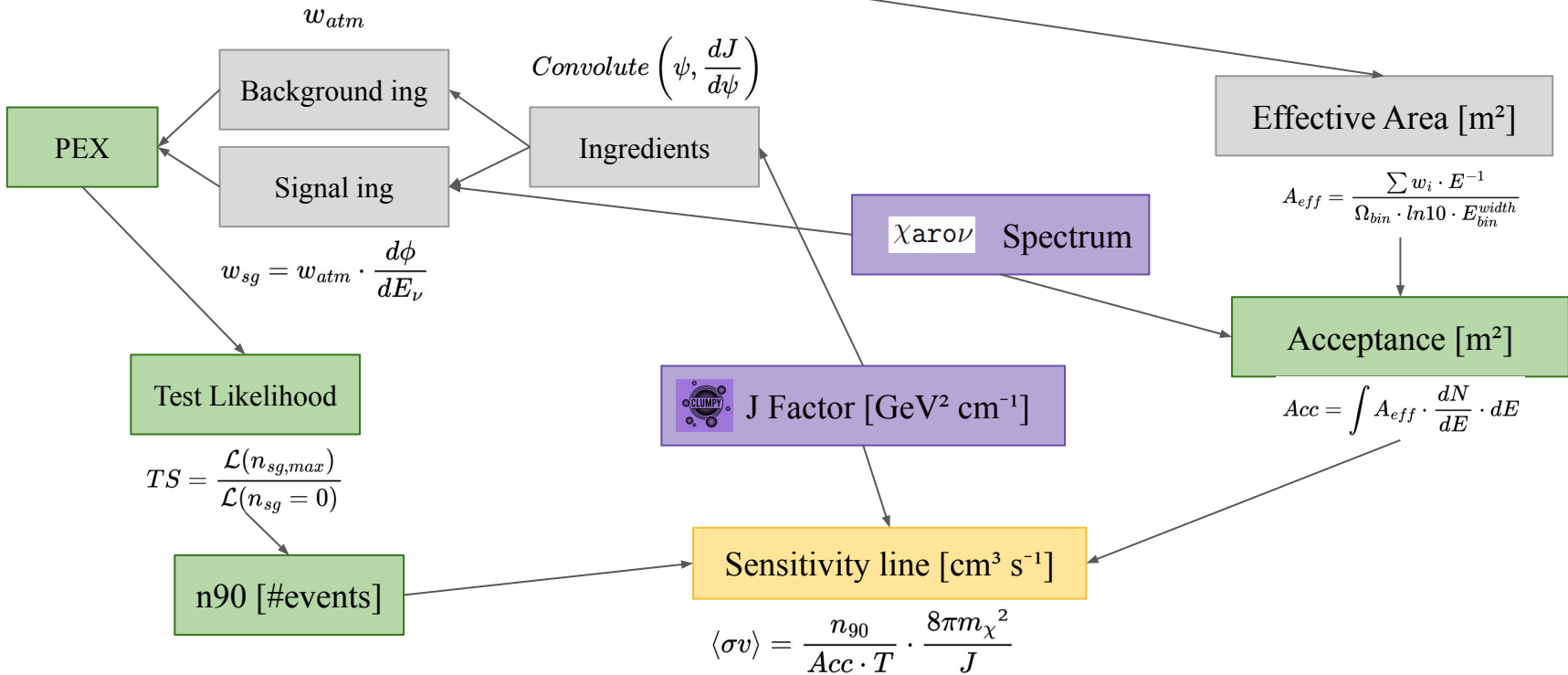
Analysis mindmap.



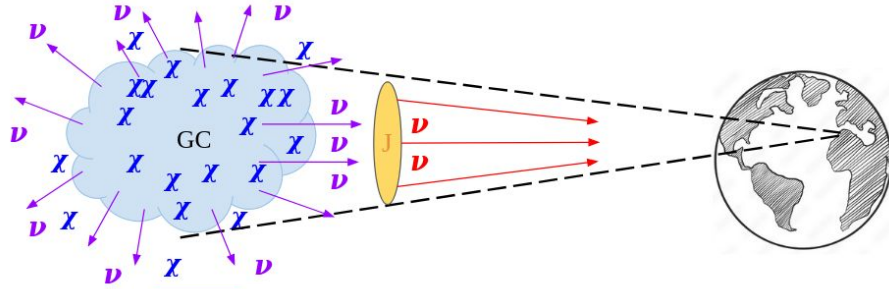
DST Files

Cut DST Files

Analysis on cut variables



Theoretical framework.

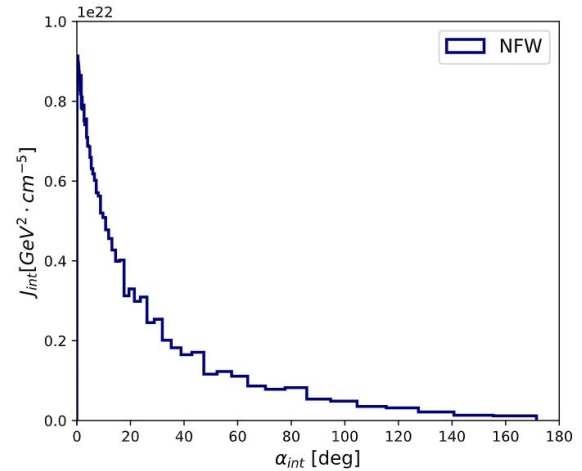
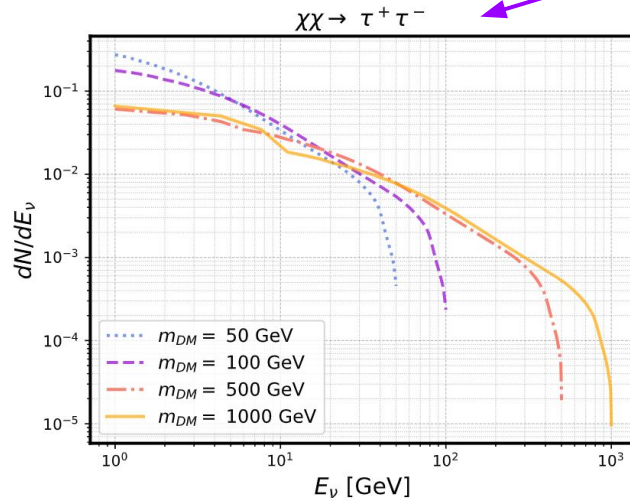


$$\frac{d\Phi_\nu}{dE_\nu} = \frac{1}{4\pi} \frac{\langle\sigma v\rangle}{2m_\chi^2} \frac{dN_\nu}{dE_\nu} \int_{\Delta\Omega} \int_{l.o.s.} \rho^2 dl d\Omega$$

Flux at Earth

Energy Spectrum

J-factor

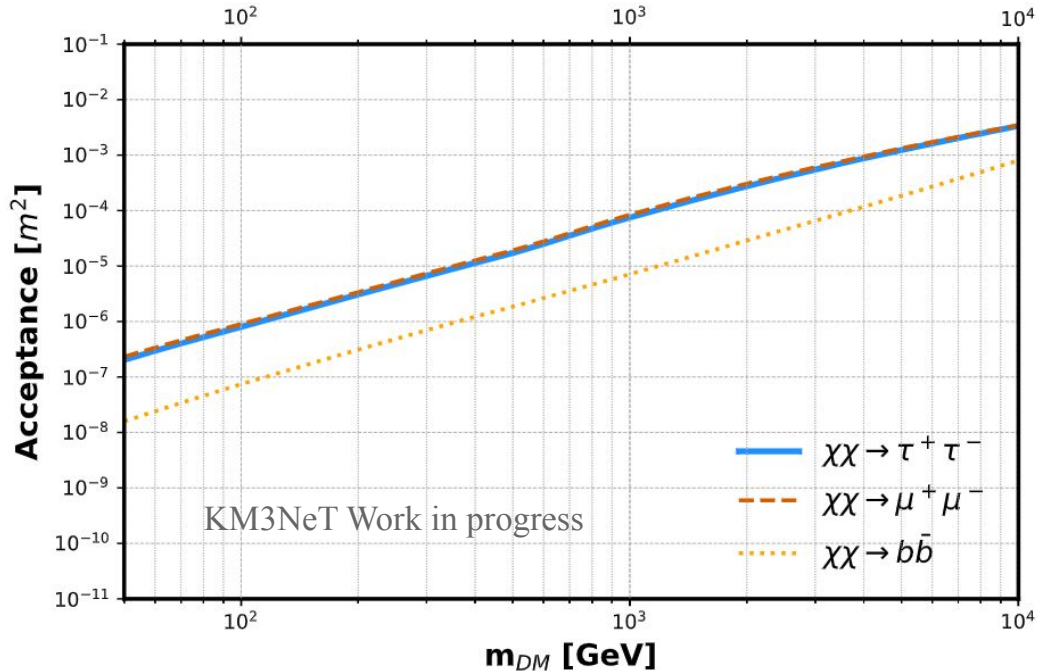


Effective areas and acceptances.

The effective area and the acceptances of the detector are obtained to study the ability of the detector to detect signal events.

$$Acc(m_{WIMP}) = \frac{2 \int^{m_{WIMP}} \bar{A}_{eff}^{\nu} \cdot \frac{dN}{dE_{\nu}} \cdot dE_{\nu}}{\int^{m_{WIMP}} \frac{dN}{dE_{\nu}} \cdot dE_{\nu}}$$

We calculate the average effective areas for tracks using CC muon neutrino files (future implementation of other flavours)



Analysis Method.

The test statistic distributions $TS = \frac{\mathcal{L}(n_{s,max})}{\mathcal{L}(n_s = 0)}$ are used to determine the mean number of signal events which can be identified as background with a 90% confidence level, **n90**.

Number of signal events compatible with background fluctuations on 90% confidence level.

