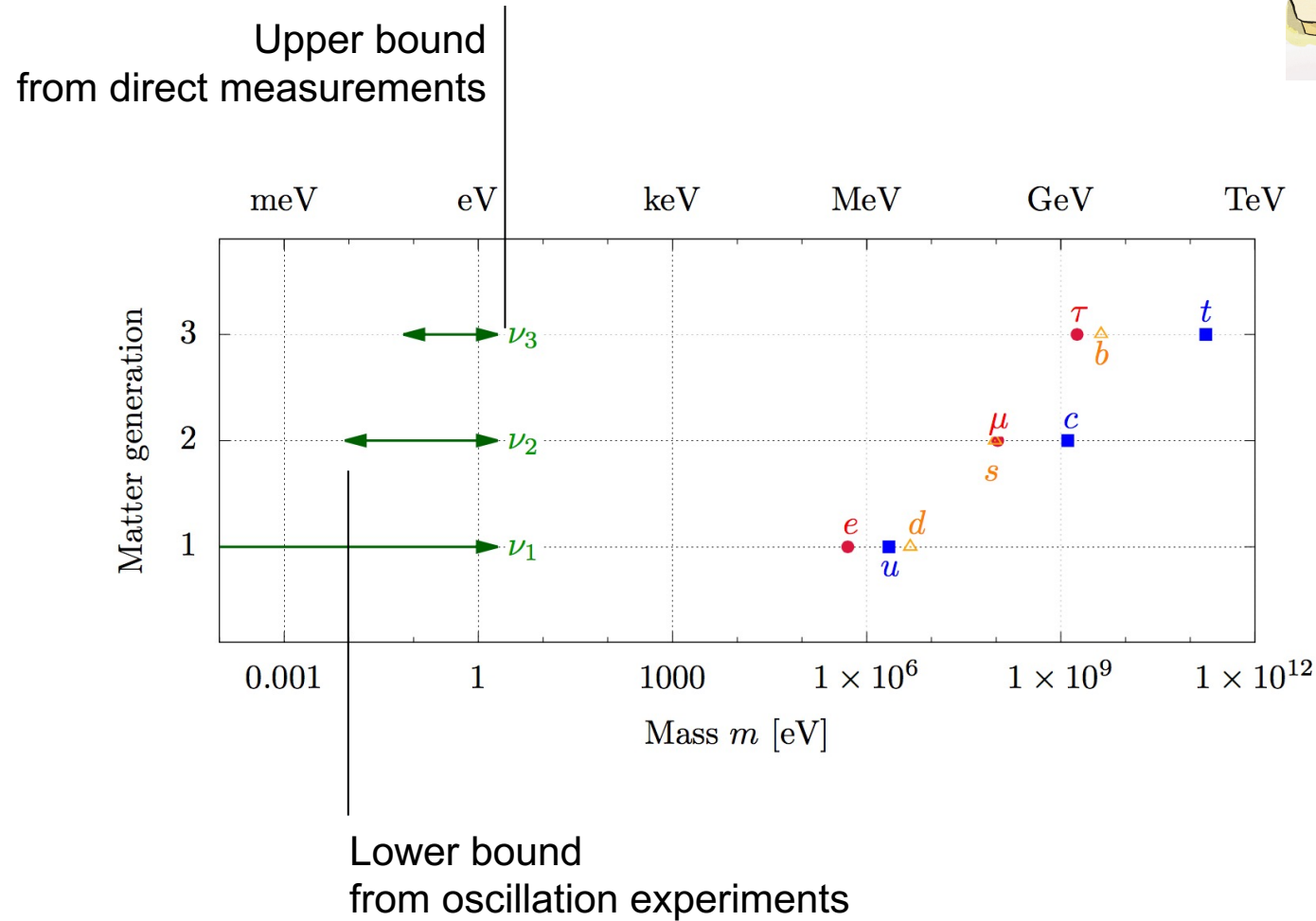
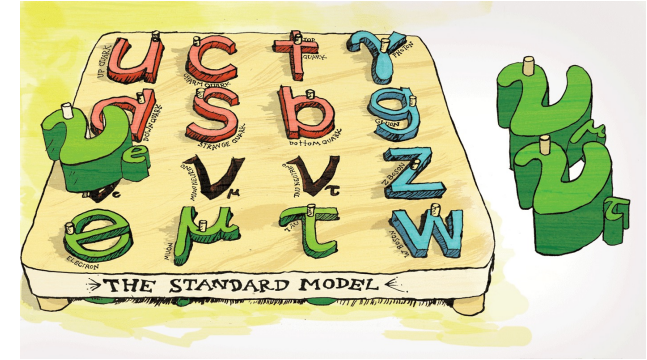


# First direct neutrino-mass measurement with sub-eV sensitivity



Susanne Mertens  
Max Planck Institute for Physics & Technical University Munich

# Neutrino mass



# Neutrino mass

## Cosmology

potential:  $m_\nu = 10 - 50 \text{ meV}$

$$m_\nu = \sum_i m_i$$



## Search for $0\nu\beta\beta$

potential:  $m_{\beta\beta} = 7 - 17 \text{ meV}$

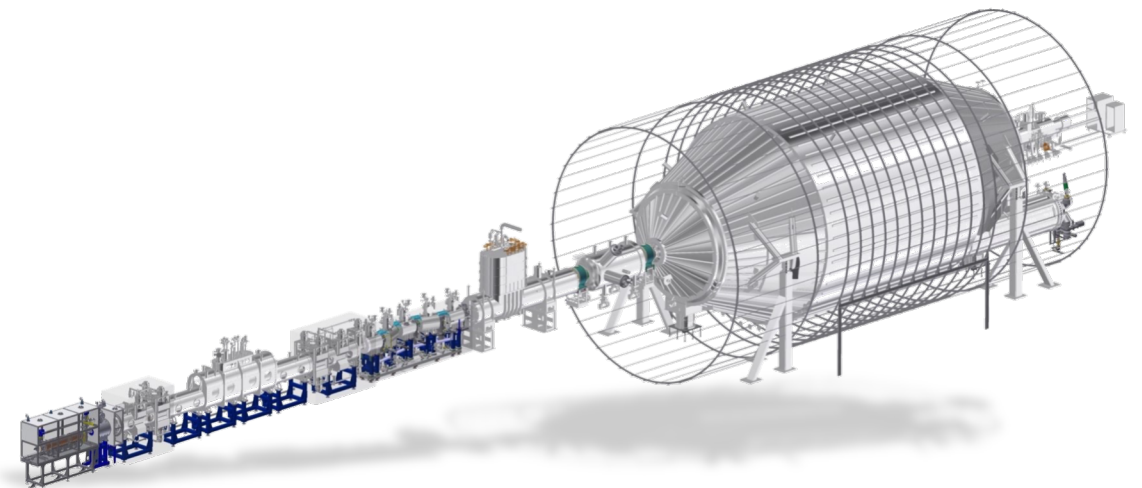
$$m_{\beta\beta} = \left| \sum_i U_{ei}^2 m_i \right|$$



## Kinematics of $\beta$ -decay

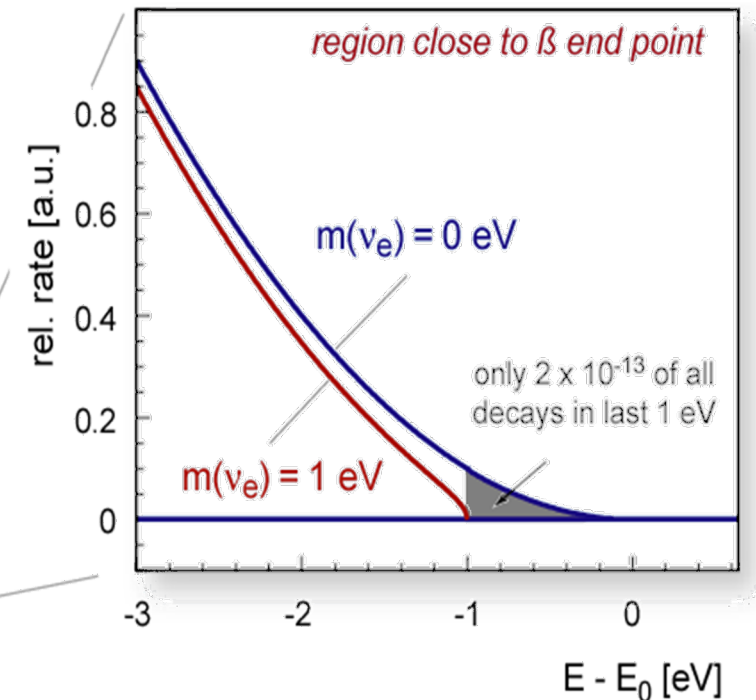
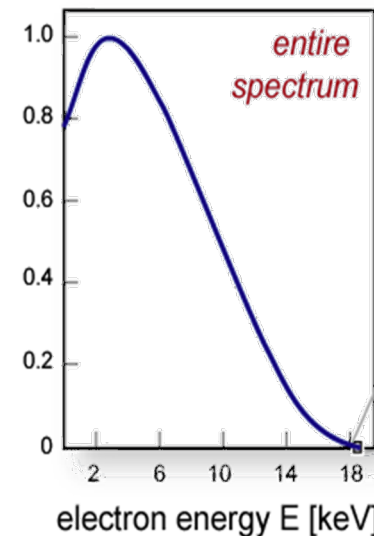
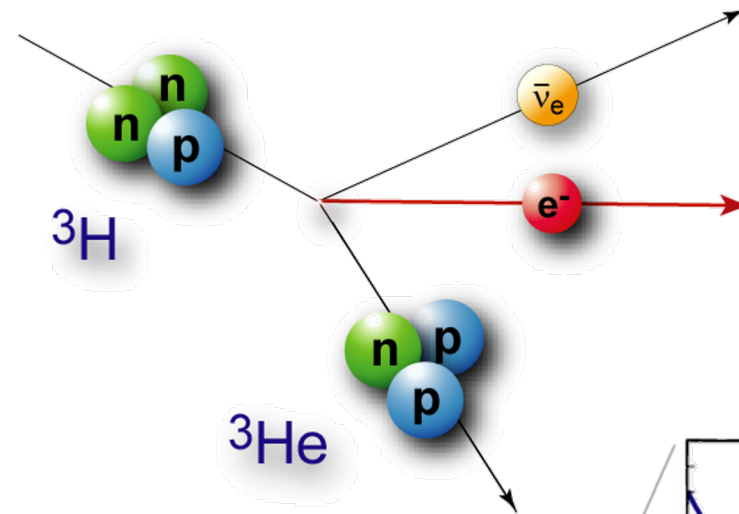
potential:  $m_\beta = 50 - 200 \text{ meV}$

$$m_\beta^2 = \sum_i |U_{ei}|^2 \cdot m_i^2$$



# General Idea

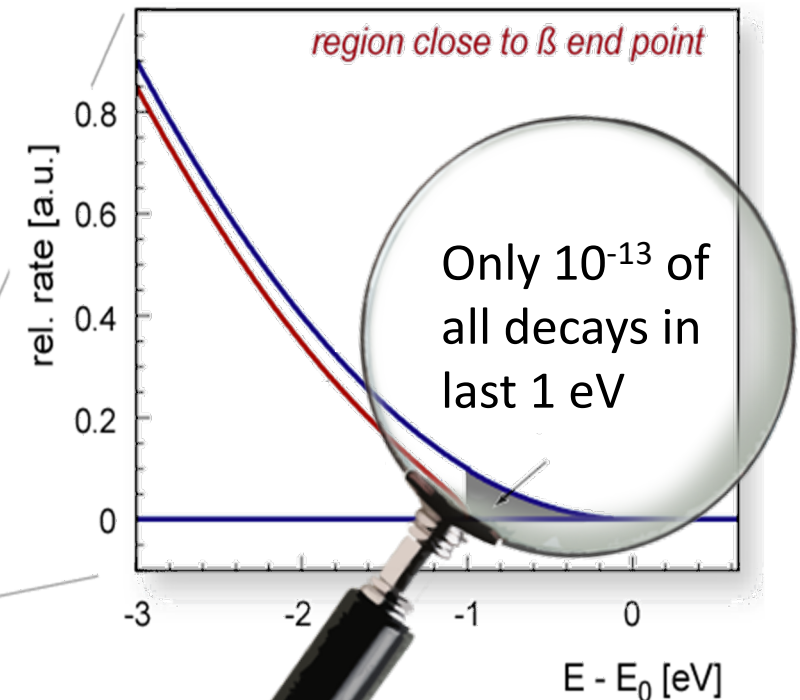
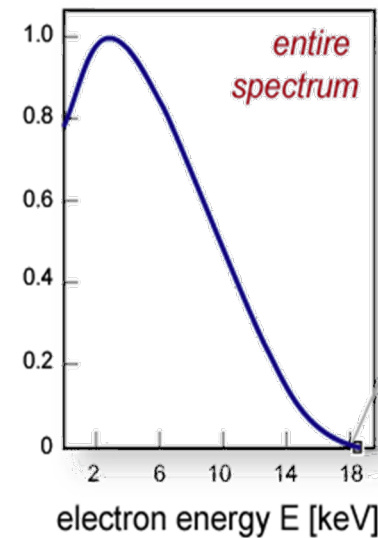
- Non-zero neutrino mass distorts the spectrum close to the endpoint
- ✓ Independent of cosmology
- ✓ Independent of neutrino nature
- Observable:  $m_\nu^2 = \sum_i |U_{ei}|^2 \cdot m_i^2$



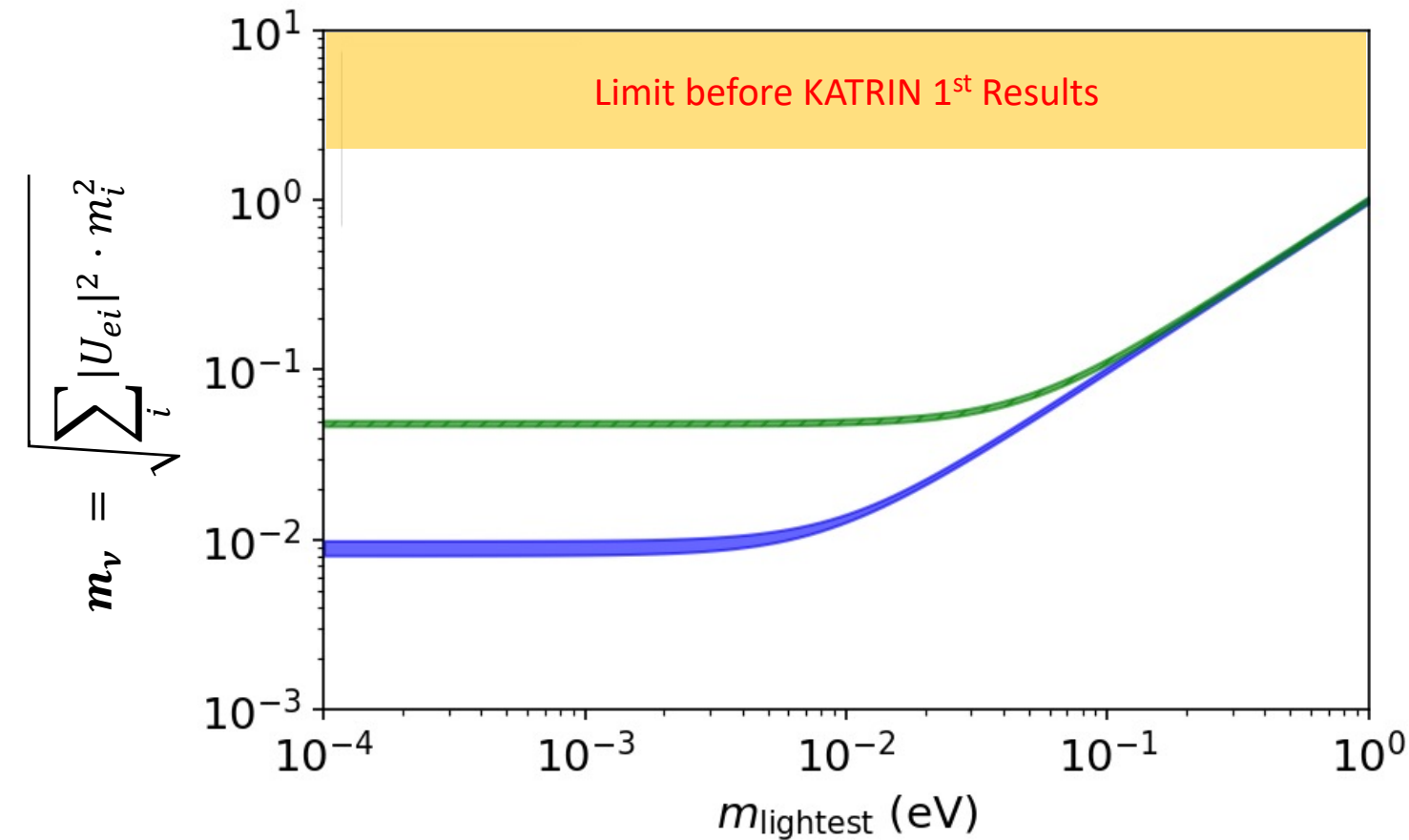
# The challenge

## Key requirements:

- Ultra-strong radioactive source
  - Tritium (12.3 years,  $E_0 = 18.6$  keV)
  - Holmium (4500 years,  $E_0 = 2.8$  keV)
- Excellent energy resolution ( $\sim 1$  eV)
- Low background ( $< 100$  mcps)

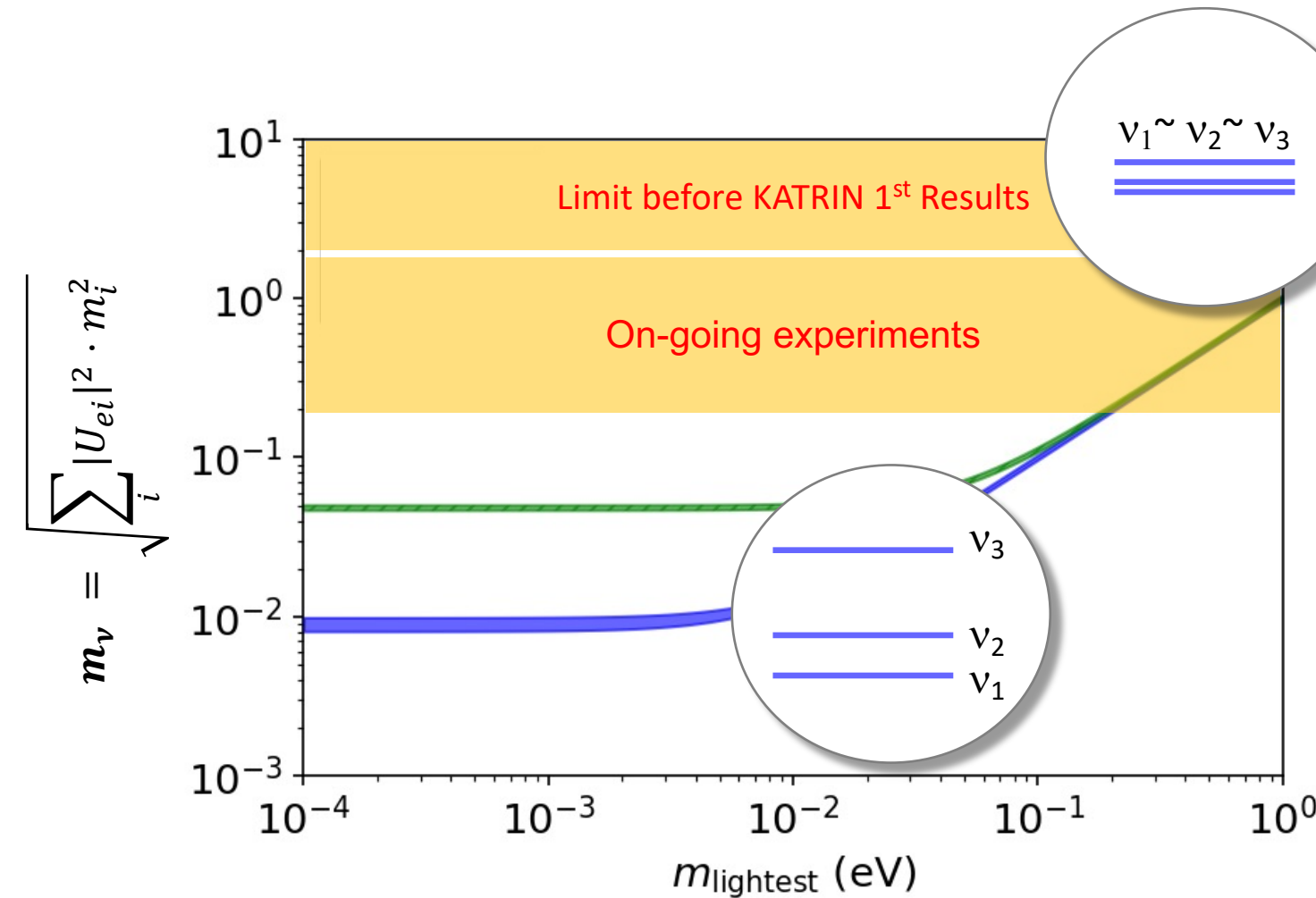


# Where do we stand?



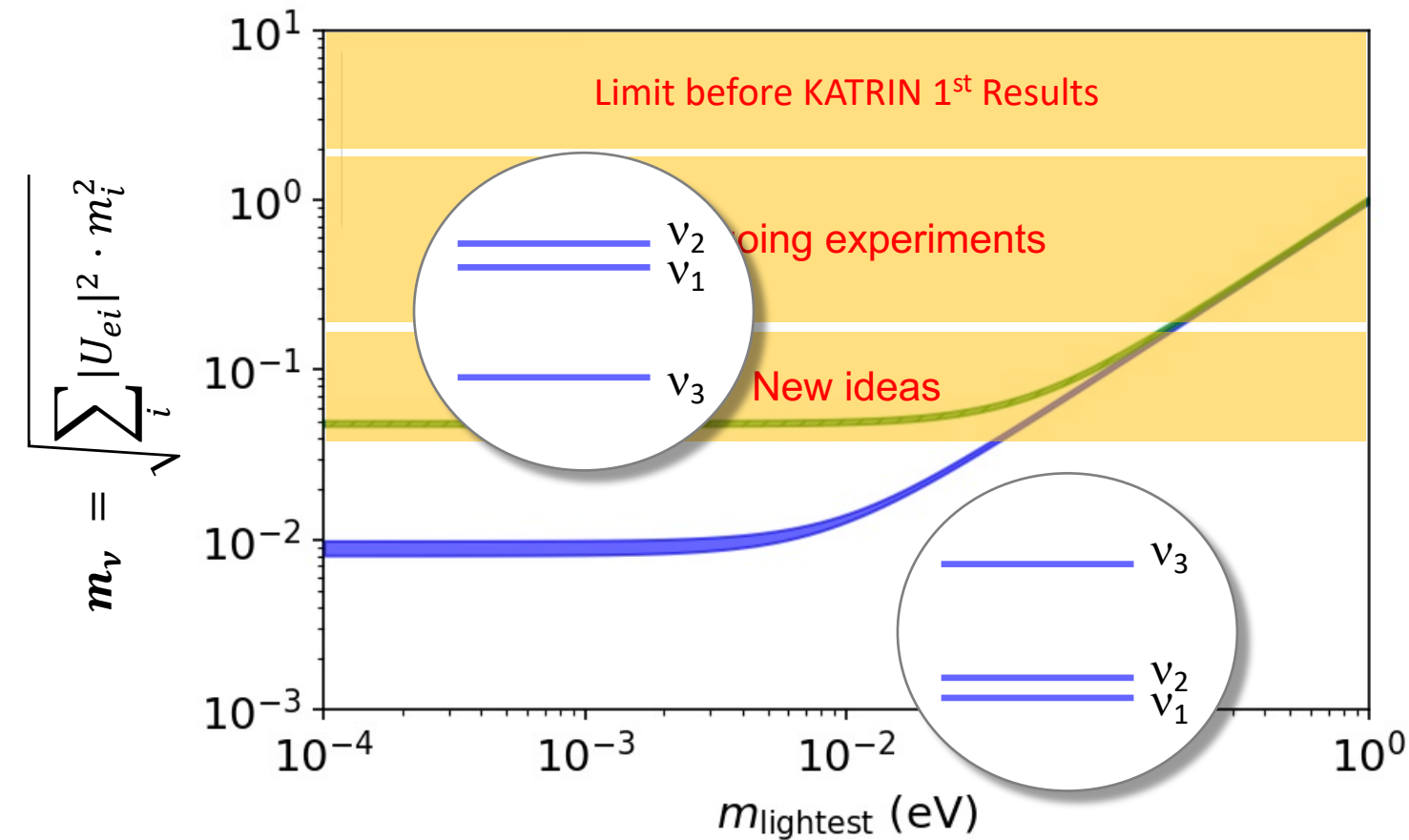
- Limit before KATRIN 1<sup>st</sup> Results:  
Mainz and Troitsk Experiment  
V. N. Aseev et al., Phys. Rev. D 84 (2011) 112003  
Kraus, C., Bornschein, B., Bornschein, L. et al. Eur. Phys. J. C (2005)

# Where do we stand?



- Limit before KATRIN 1<sup>st</sup> Results: Mainz and Troitsk Experiment  
V. N. Aseev et al., Phys. Rev. D 84 (2011) 112003  
Kraus, C., Bornschein, B., Bornschein, L. et al. Eur. Phys. J. C (2005)
- KATRIN goal: Distinguish between **degenerate** and **hierarchical** scenario

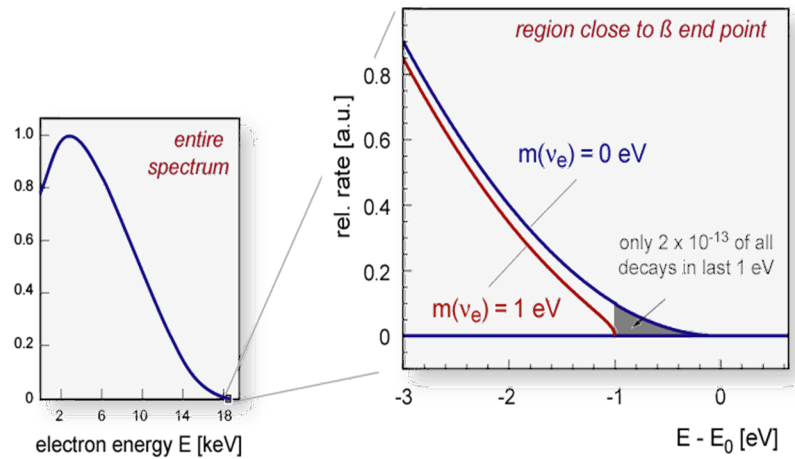
# Where do we stand?



- Limit before KATRIN 1<sup>st</sup> Results:  
Mainz and Troitsk Experiment  
V. N. Aseev et al., Phys. Rev. D 84 (2011) 112003  
Kraus, C., Bornschein, B., Bornschein, L. et al. Eur. Phys. J. C (2005)
- KATRIN goal:  
Distinguish between **degenerate** and **hierarchical** scenario
- Future:  
Resolve **normal** vs **inverted** neutrino mass ordering



# Experimental efforts



Electrostatic  
filter (MAC-E)

counting  
above  
threshold

frequency



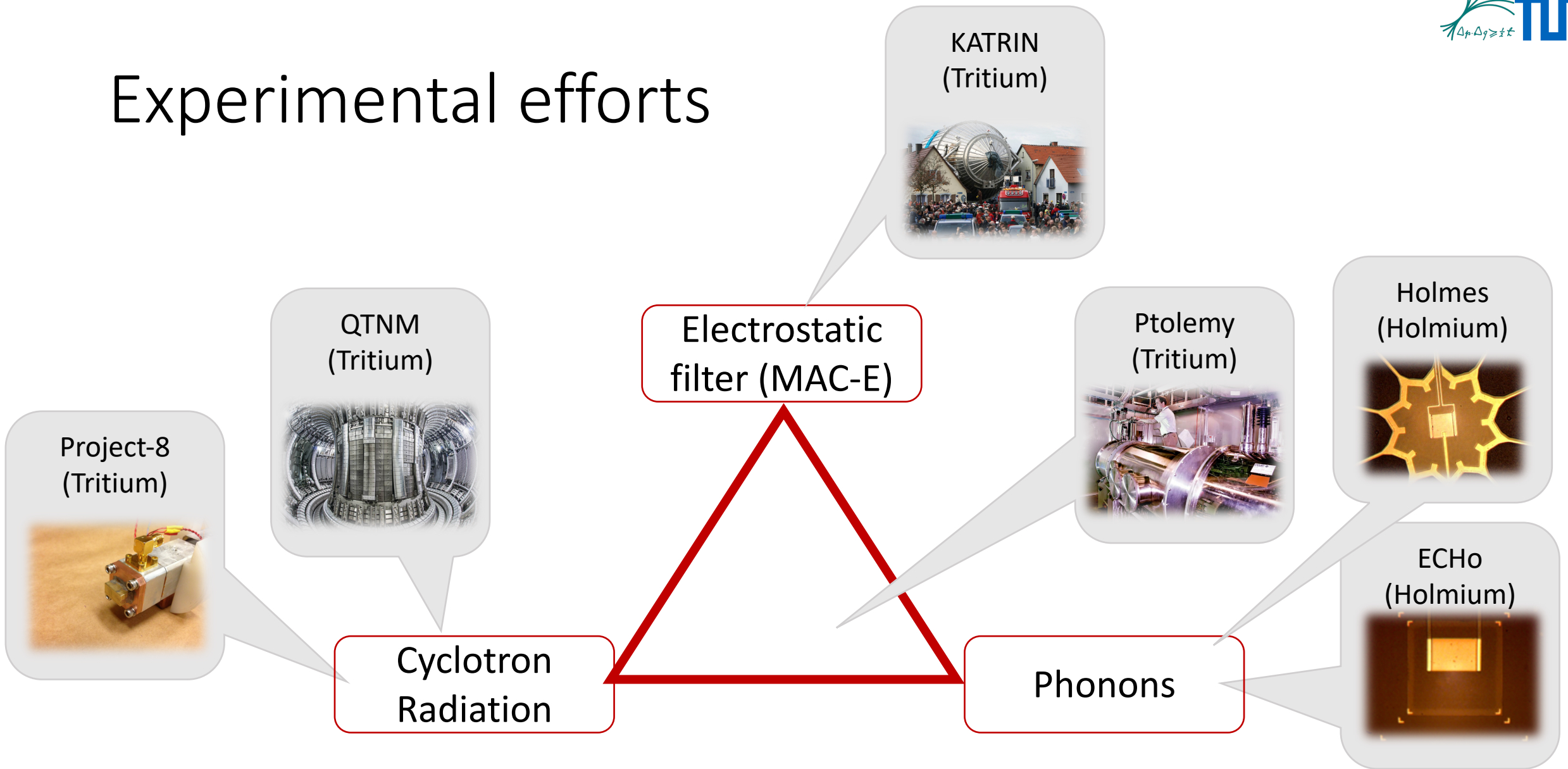
Cyclotron  
Radiation

heat



Phonons

# Experimental efforts



Karlsruhe  
Tritium  
Neutrino  
Experiment



# KATRIN

- Experimental site: Karlsruhe Institute of Technology (KIT)
- International Collaboration (150 members)
- Design sensitivity: 0.2 eV (90% CL)  
(1000 days of measurement time)



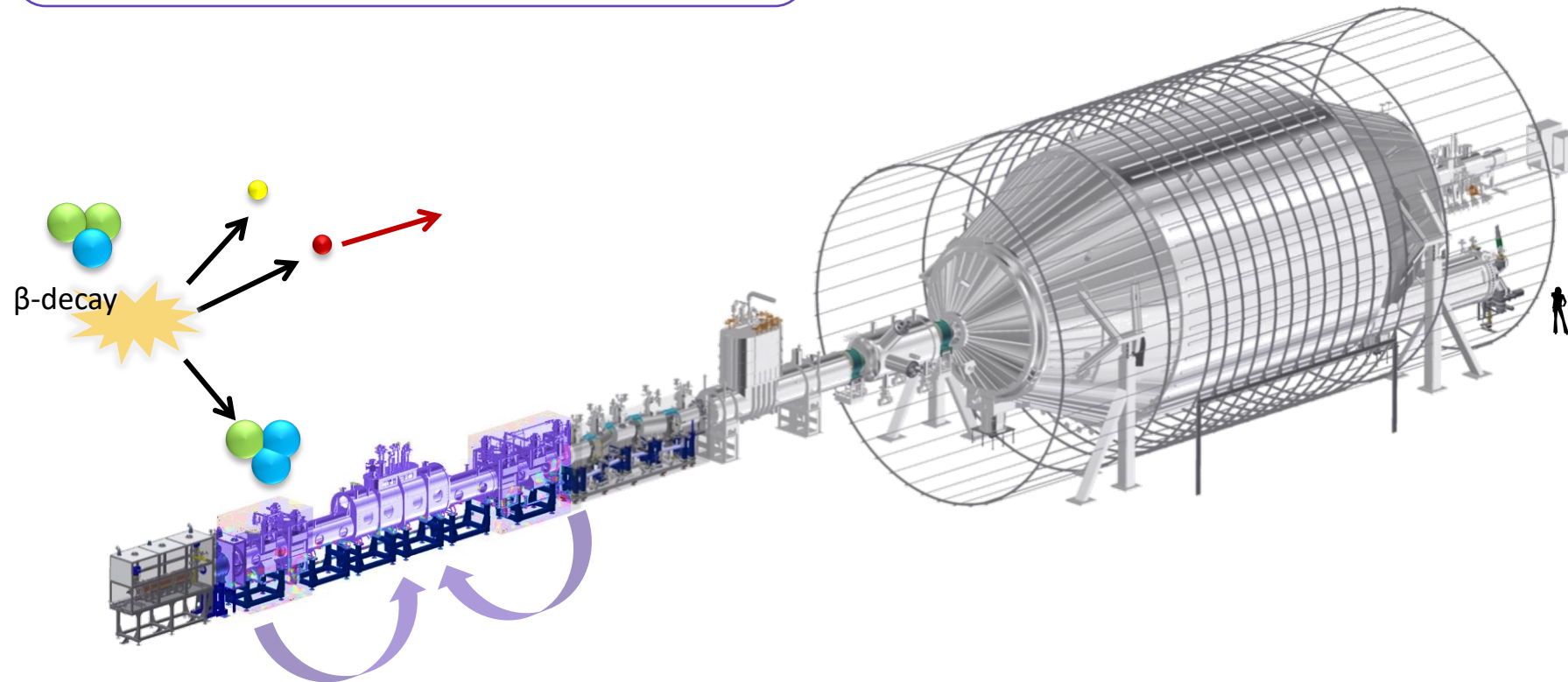
# Overview

- How does KATRIN work ?
- What are the latest results?
- What's next?
- What else can we do with the data?

# KATRIN Working Principle

## Windowless gaseous tritium source

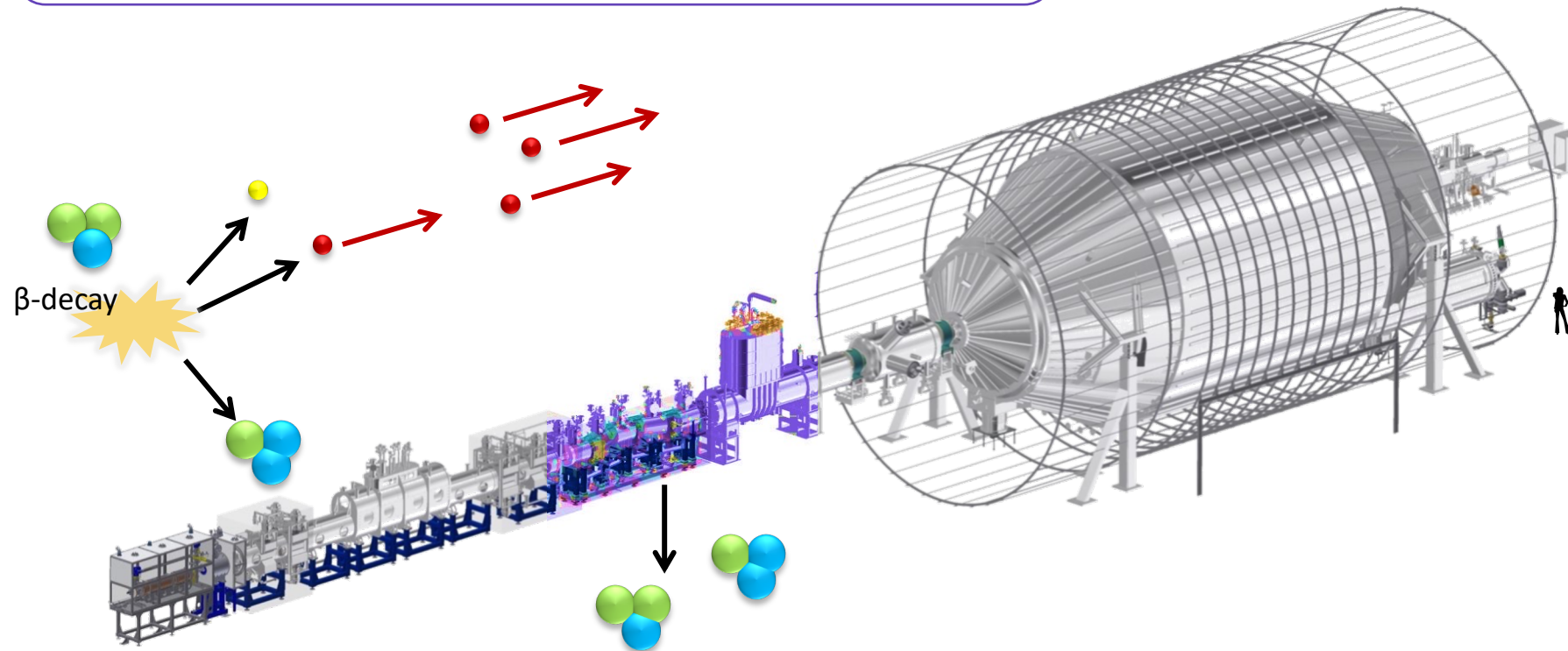
- molecular tritium in closed loop system
- $10^{11}$  decays/s



# KATRIN Working Principle

## Transport section

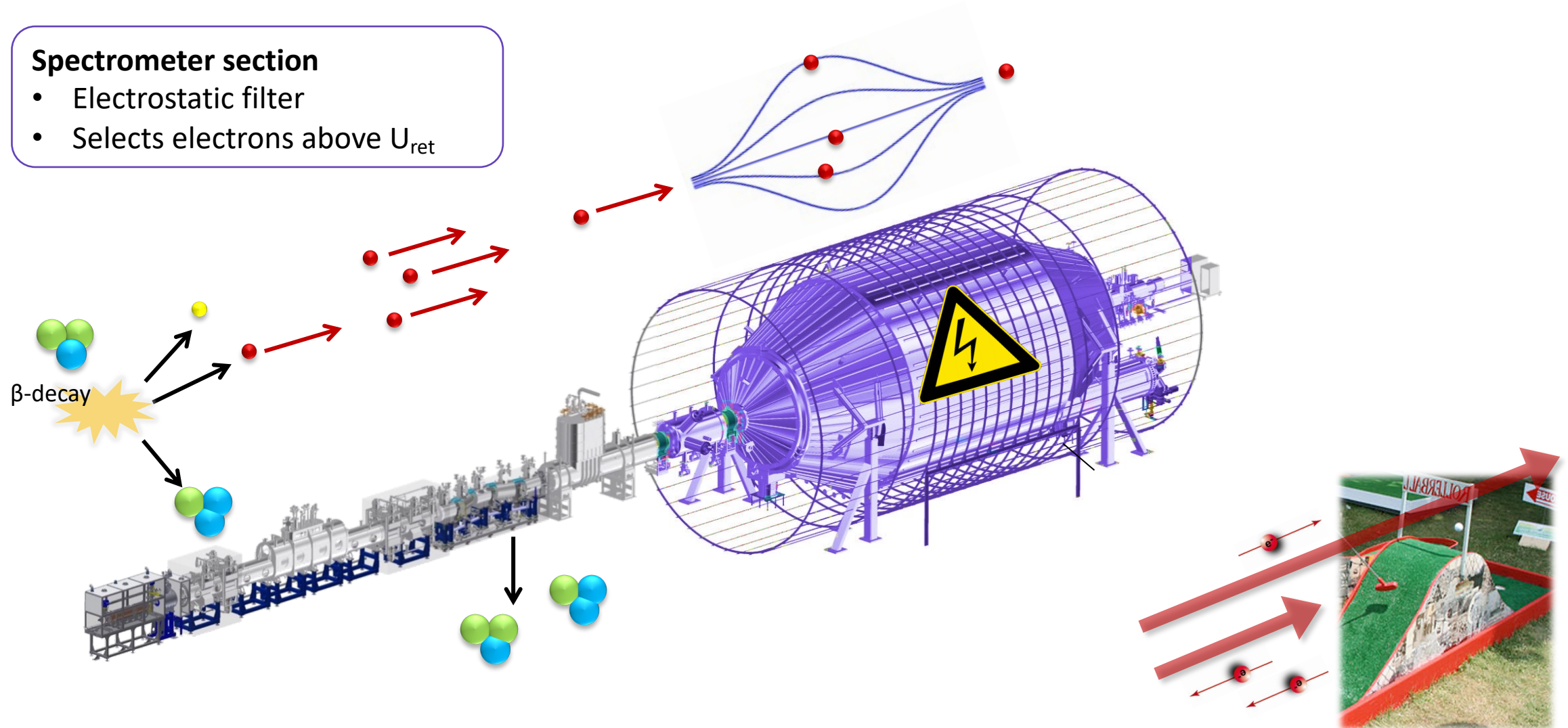
- magnetic guidance of electrons (@ 4 T)
- tritium flow reduction by  $> 10^{14}$  + tritium ion removal



# KATRIN Working Principle

## Spectrometer section

- Electrostatic filter
- Selects electrons above  $U_{ret}$

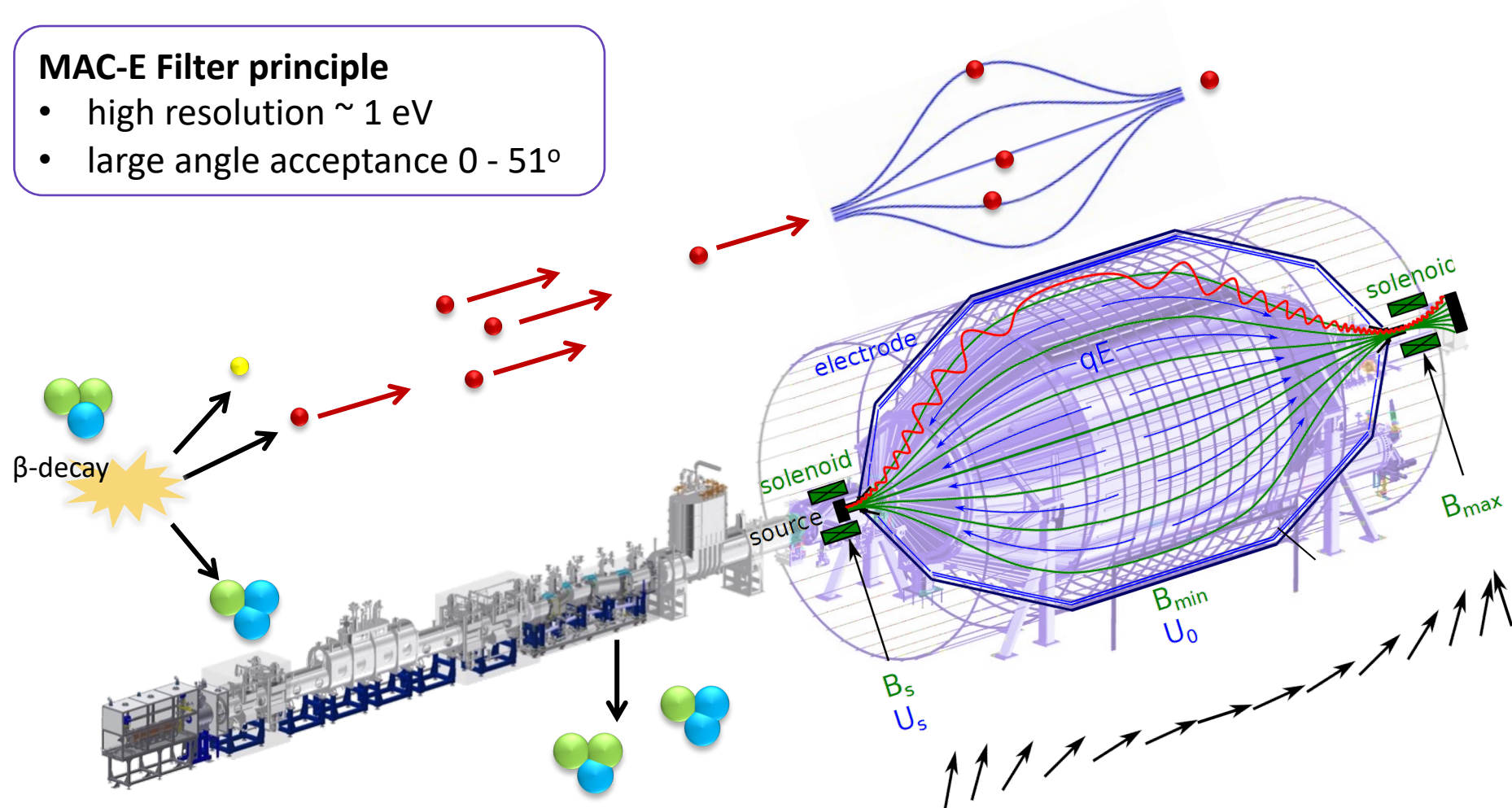




# KATRIN Working Principle

## MAC-E Filter principle

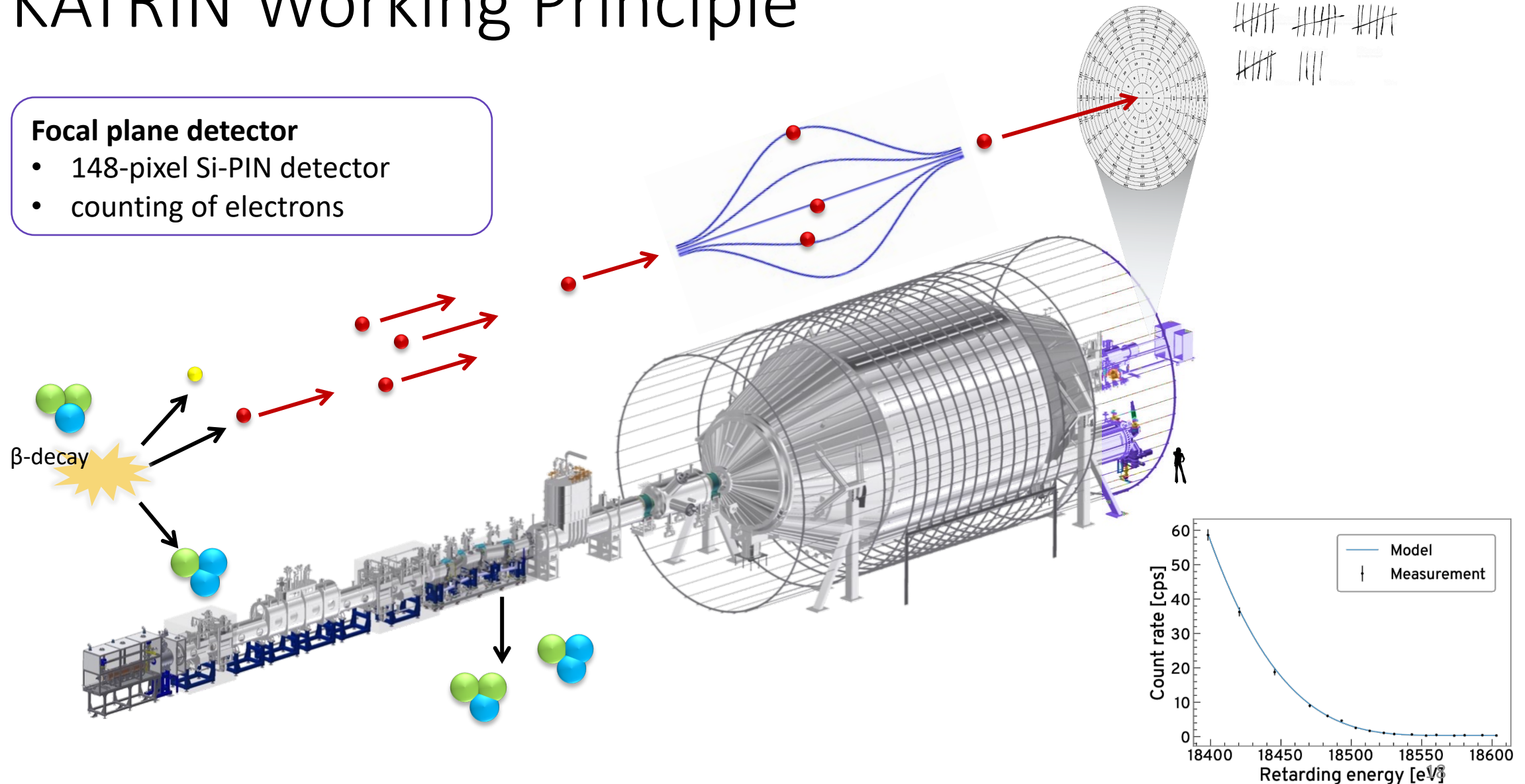
- high resolution  $\sim 1$  eV
- large angle acceptance  $0 - 51^\circ$



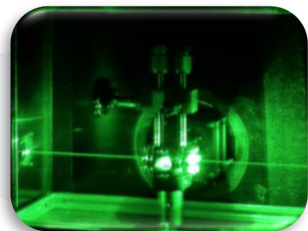
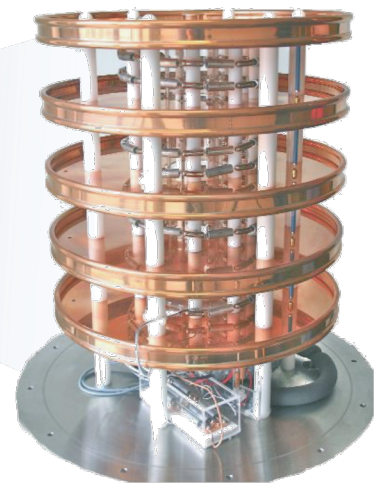
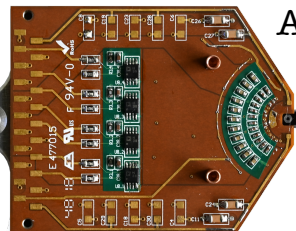
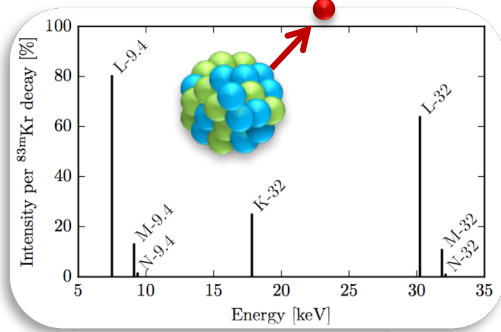
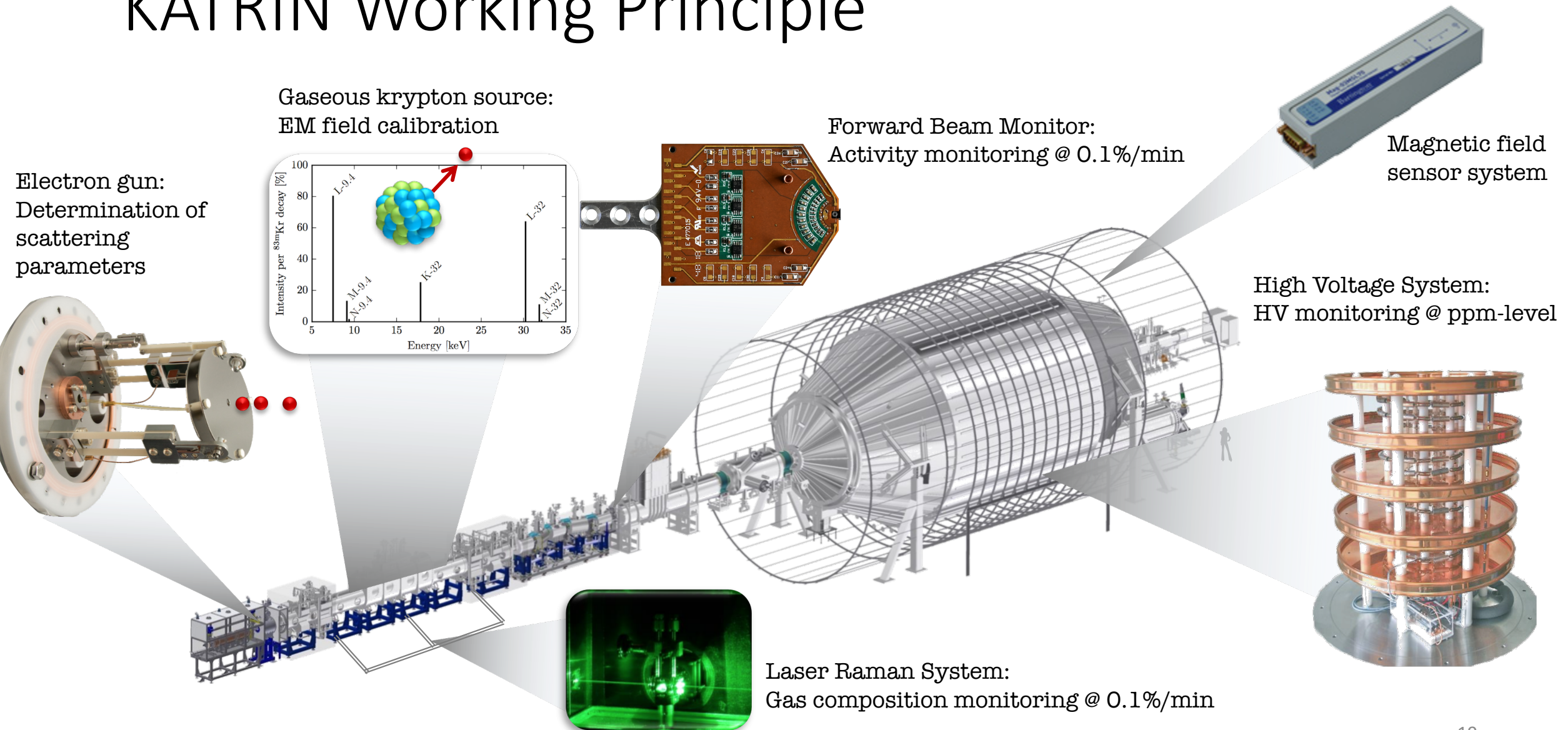
# KATRIN Working Principle

## Focal plane detector

- 148-pixel Si-PIN detector
- counting of electrons

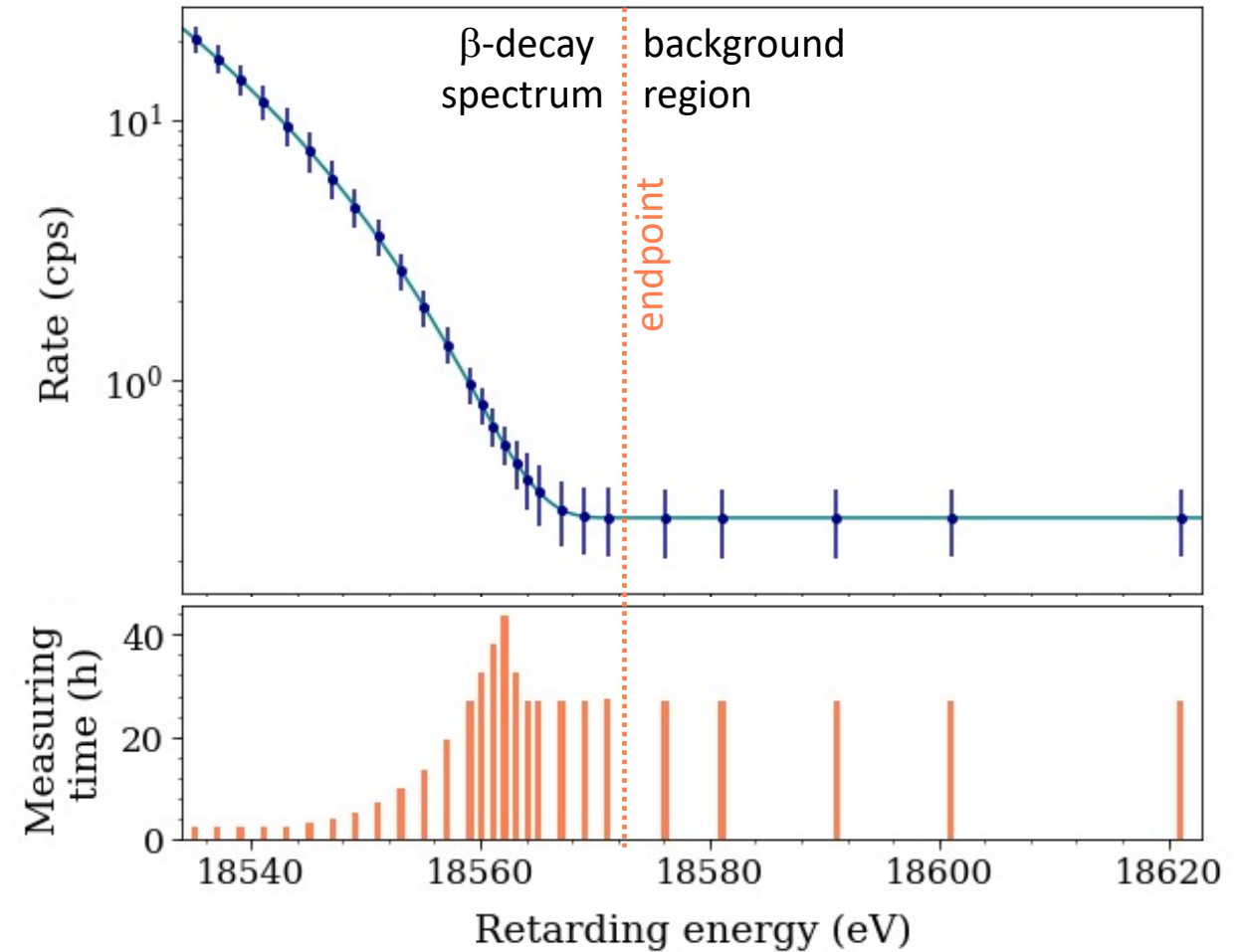
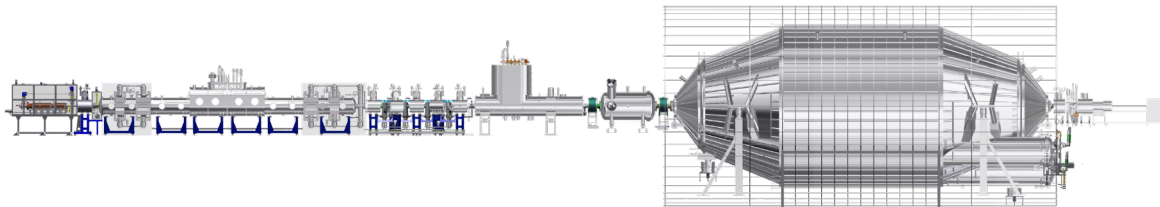


# KATRIN Working Principle



# Measurement strategy

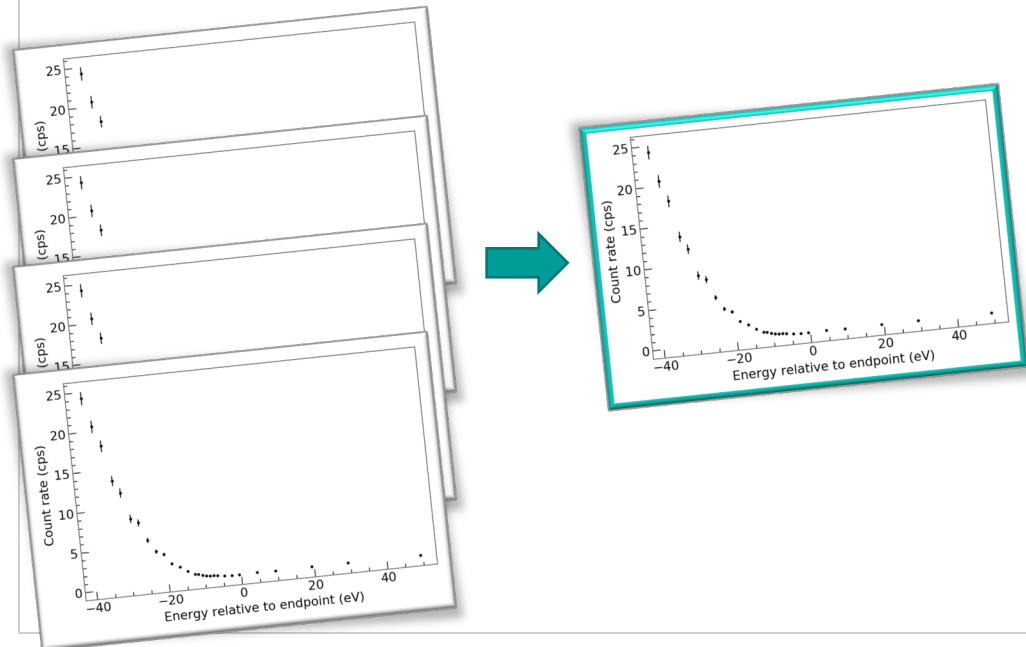
- Scan: **~ 30 HV set points**
- Scanning time: **~ 2 hours**
- Analysis interval:  **$E_0 - 40 \text{ eV}$  ,  $E_0 + 130 \text{ eV}$**
- Campaign: **several hundreds of scans**



# Data combination

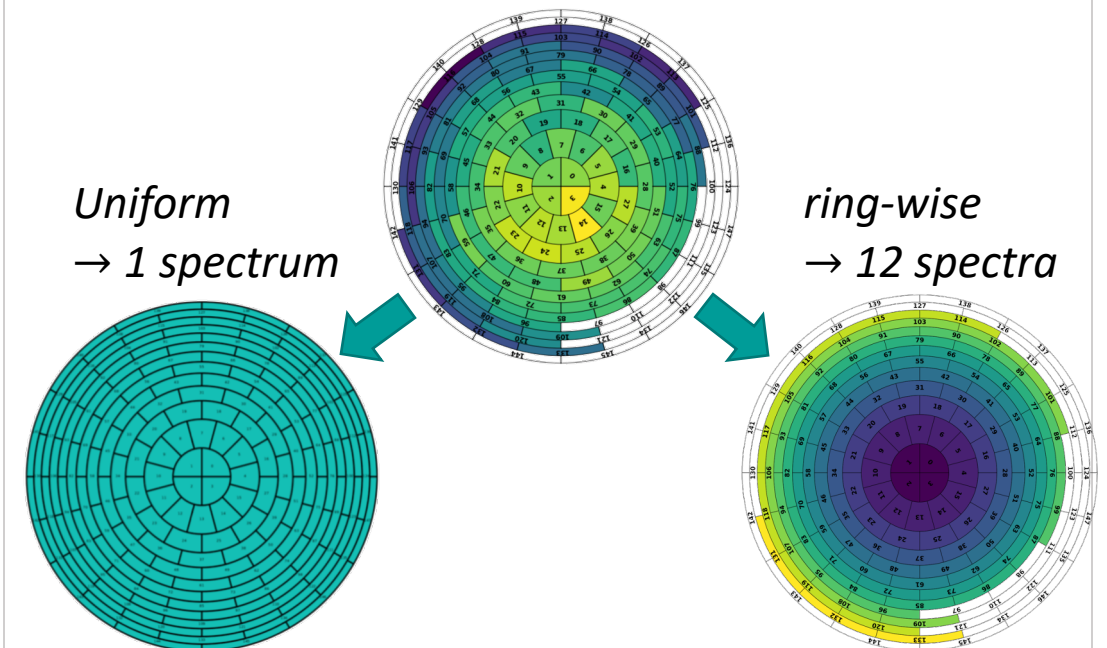
## Scan combination

- sum the counts at the same HV set point
- use average HV ( $\sigma_{HV} < 34$  mV)



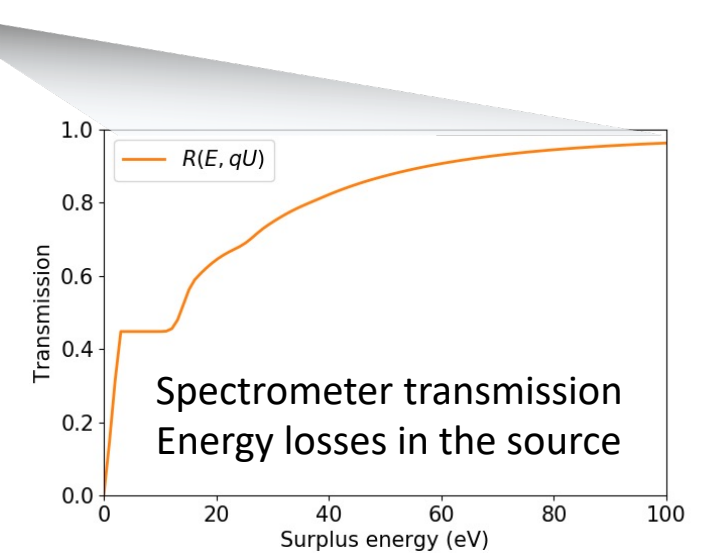
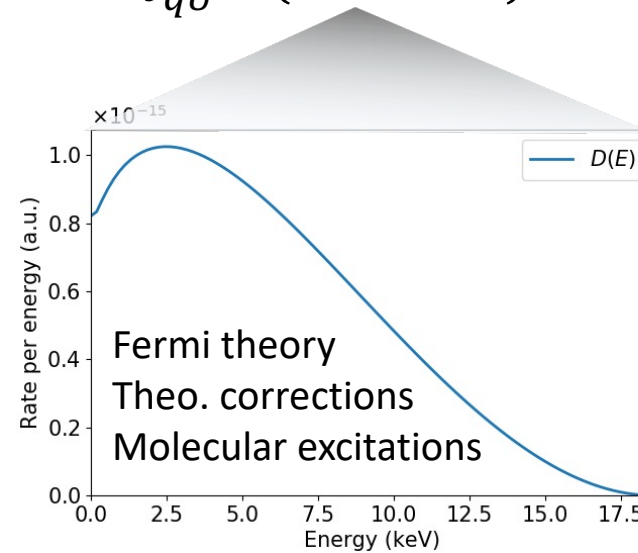
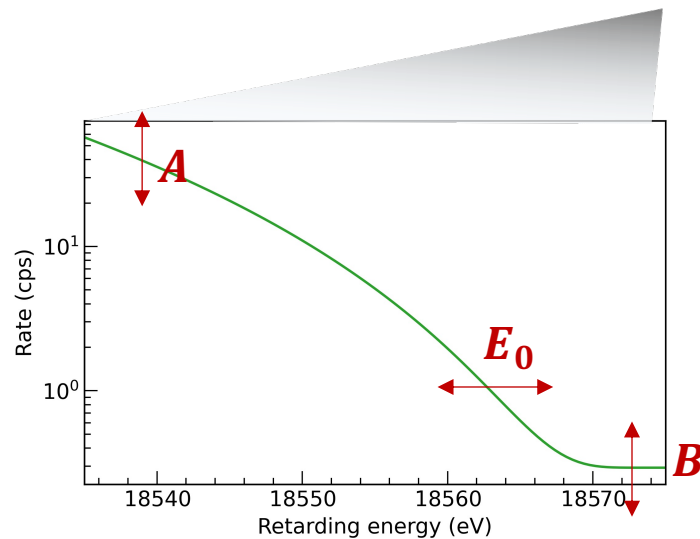
## Pixel combination

- sum the counts of ALL pixels or in a ring
- use average response function



# Data analysis

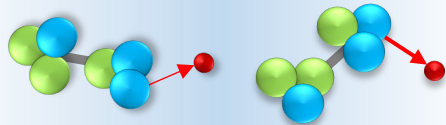
- Fit of theoretical prediction:  $\Gamma(qU) \propto A \cdot \int_{qU}^{E_0} D(E; m_v^2, E_0) \cdot R(qU, E) dE + B$



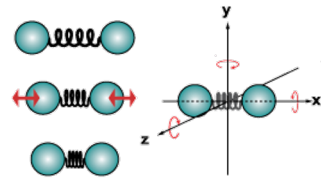
- Free parameters:  $m_v^2, E_0, B, A$
- Fit model informed by **theoretical** and **experimental** inputs (e-gun, krypton, monitoring, ...)

# Systematic uncertainties

## Source electric potential

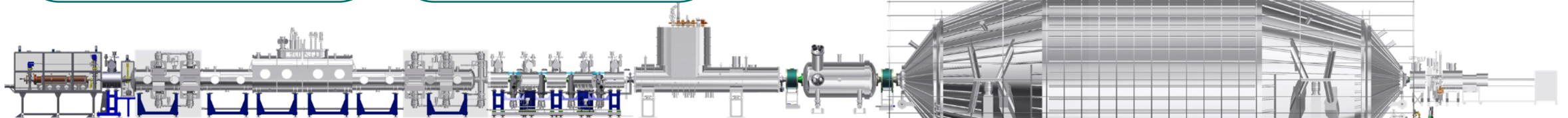
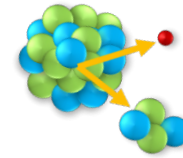


## Molecular Final States



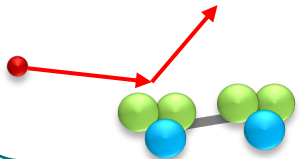
## Background:

- time correlation (radon, penning trap)
- retarding potential dependence

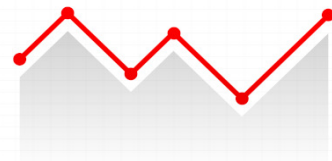


## Scattering

- energy loss
- column density



## Activity fluctuations

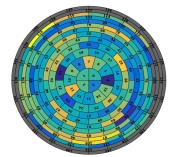


## Magnetic fields

- source
- spectrometer
- detector



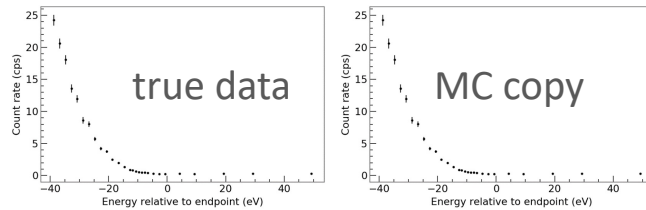
## Detection efficiency



# Blinded analysis

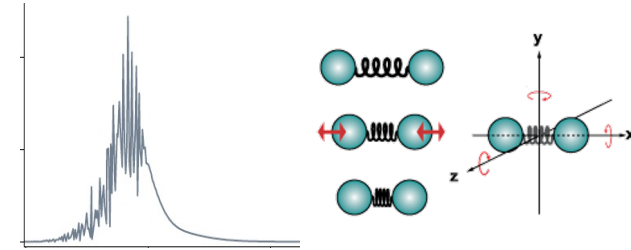
## Freeze analysis on MC-twin data

- MC-copy of each scan (with  $m_\nu = 0$  eV)



## Blinded model

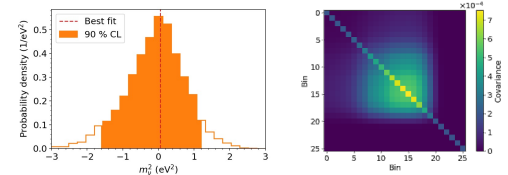
- Modified molecular final state dist.



$m_\nu^2$

## Independent analysis strategies

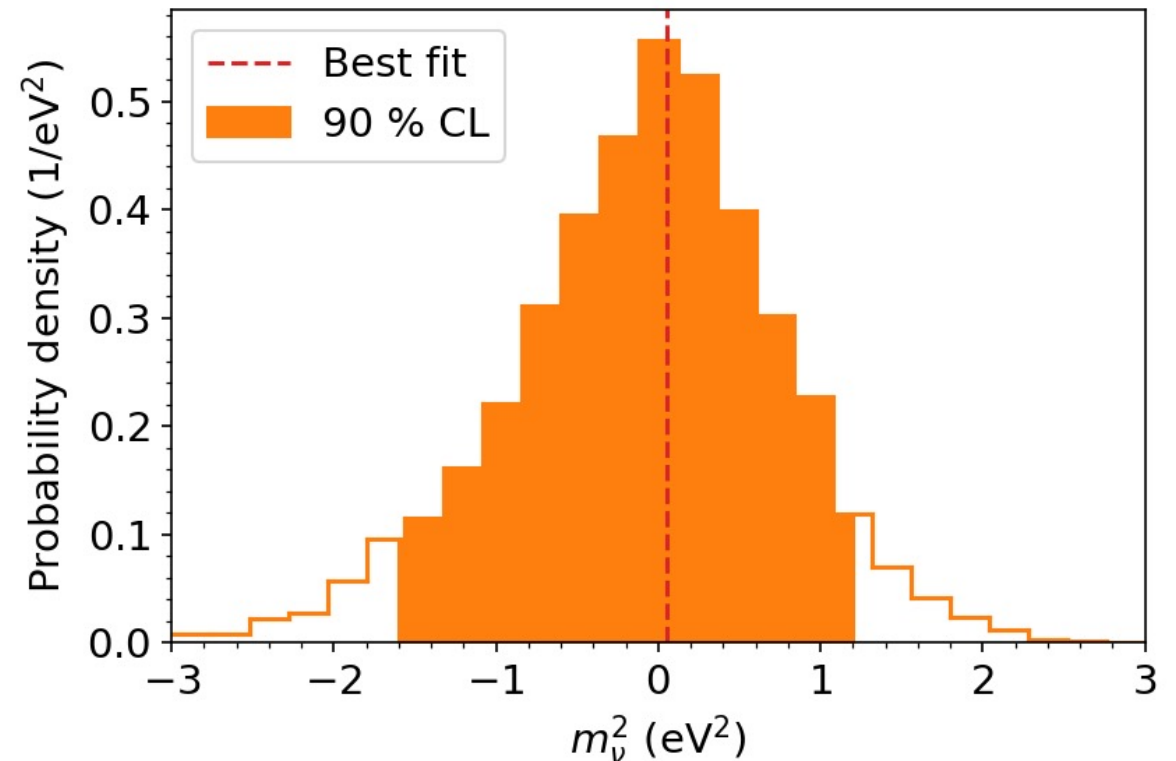
- Covariance matrix
- Monte Carlo propagation
- Pull term





# MC propagation of uncertainties

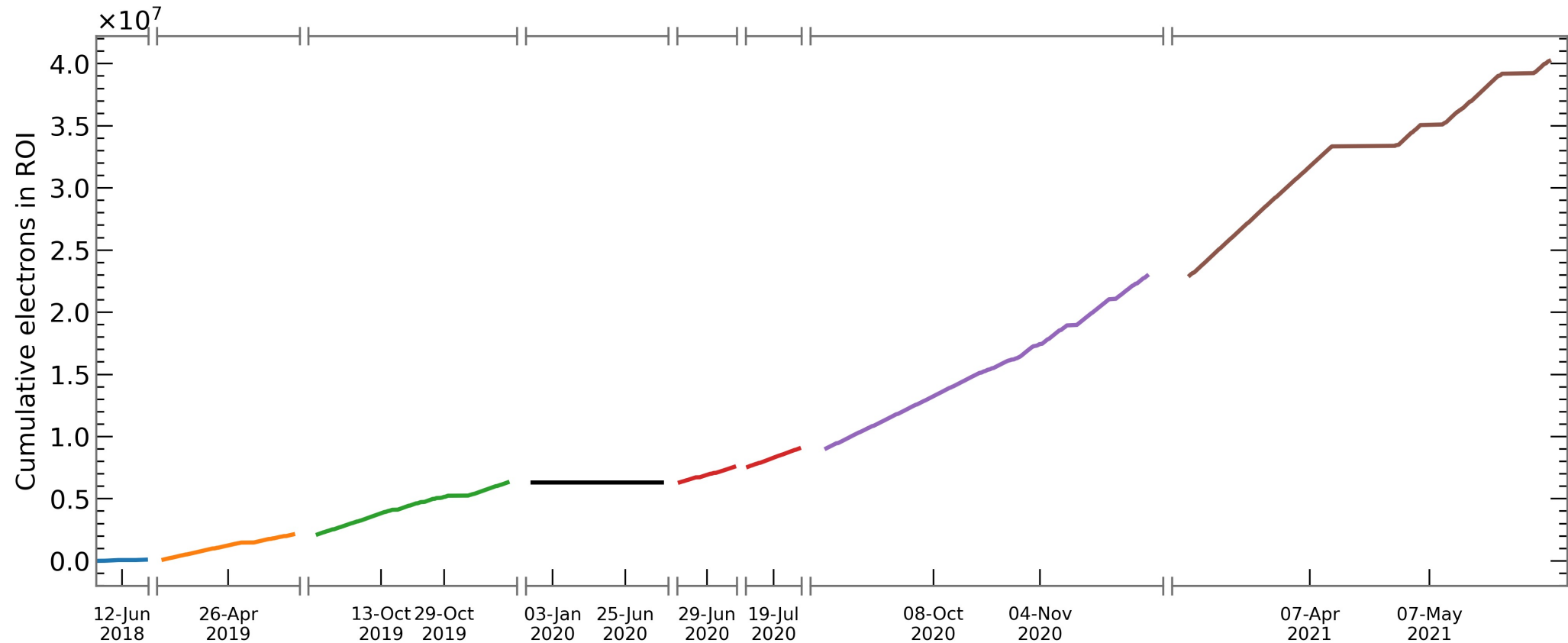
- Fit performed  $10^5$  times
- Each time the systematic parameter is varied according to its uncertainty
- Width of  $m_\nu^2$  distribution reflects systematic uncertainty from this effect



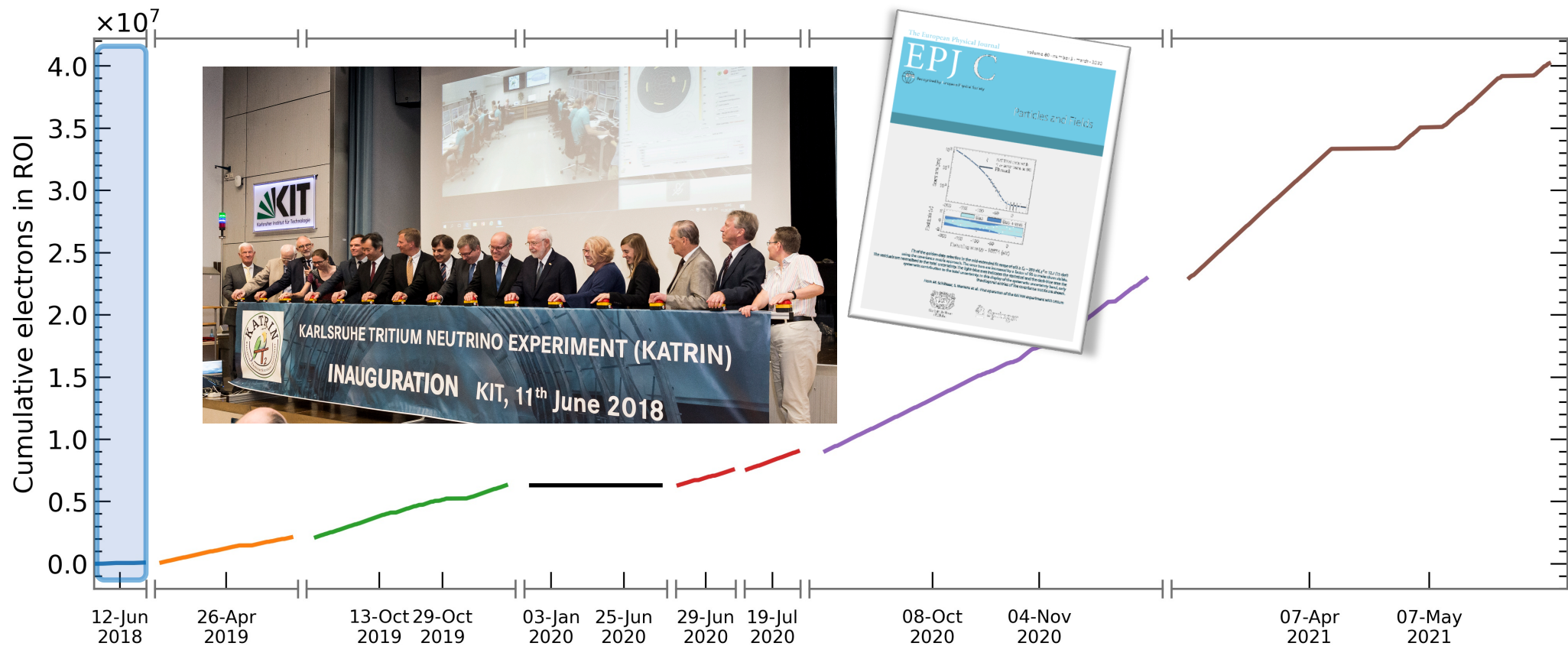
# Overview

- How does KATRIN work ?
- What are the latest results?
- What's next?
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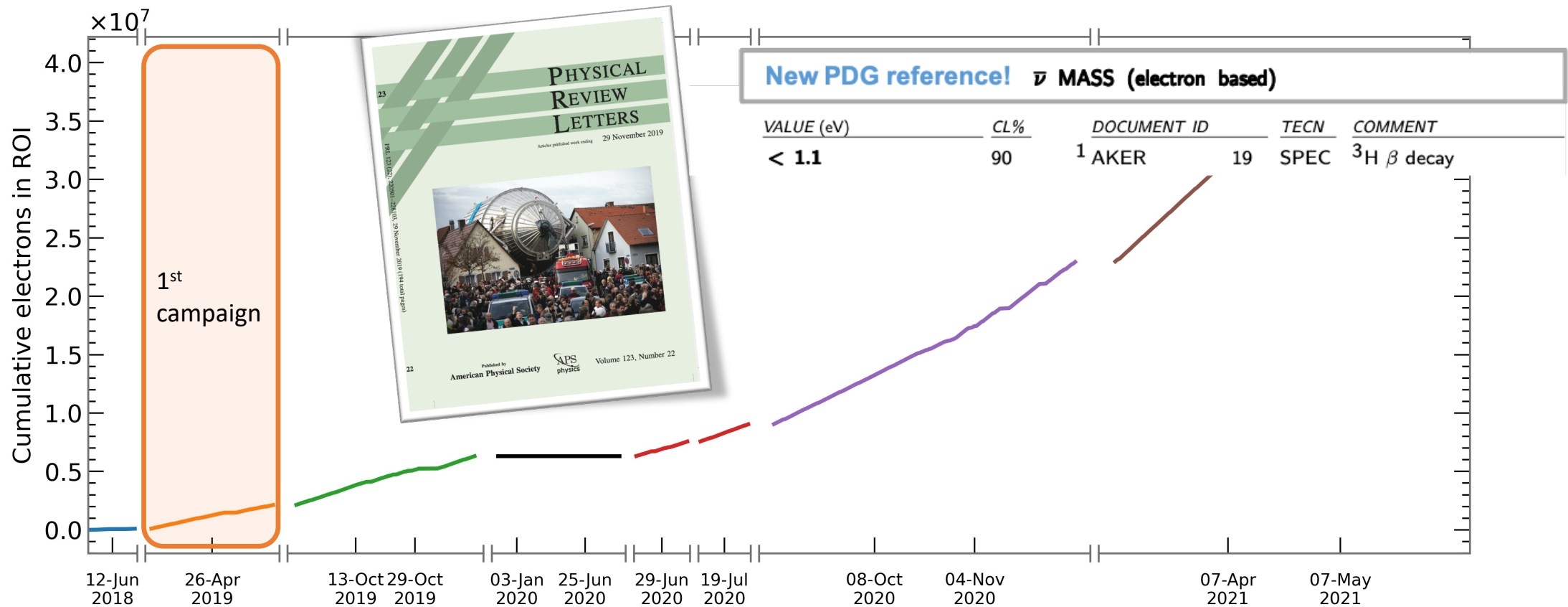
# KATRIN Data Taking Overview



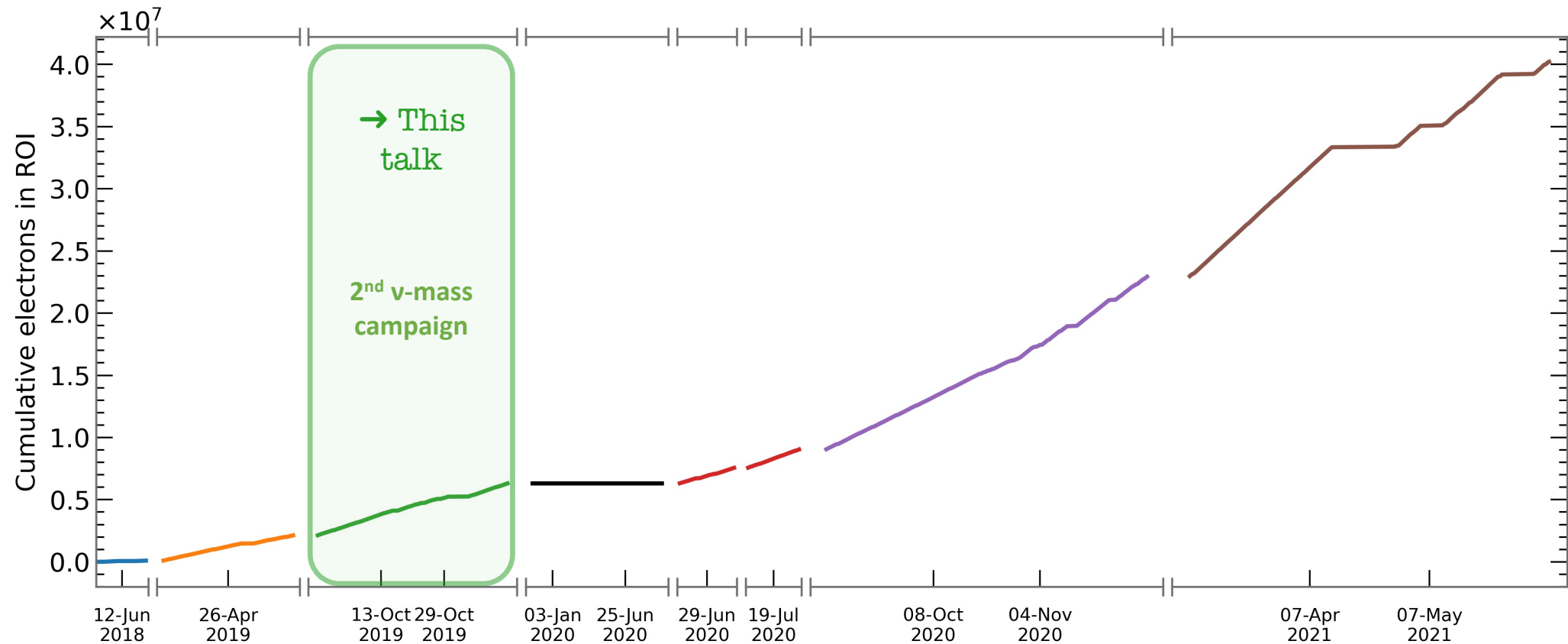
# KATRIN Data Taking Overview



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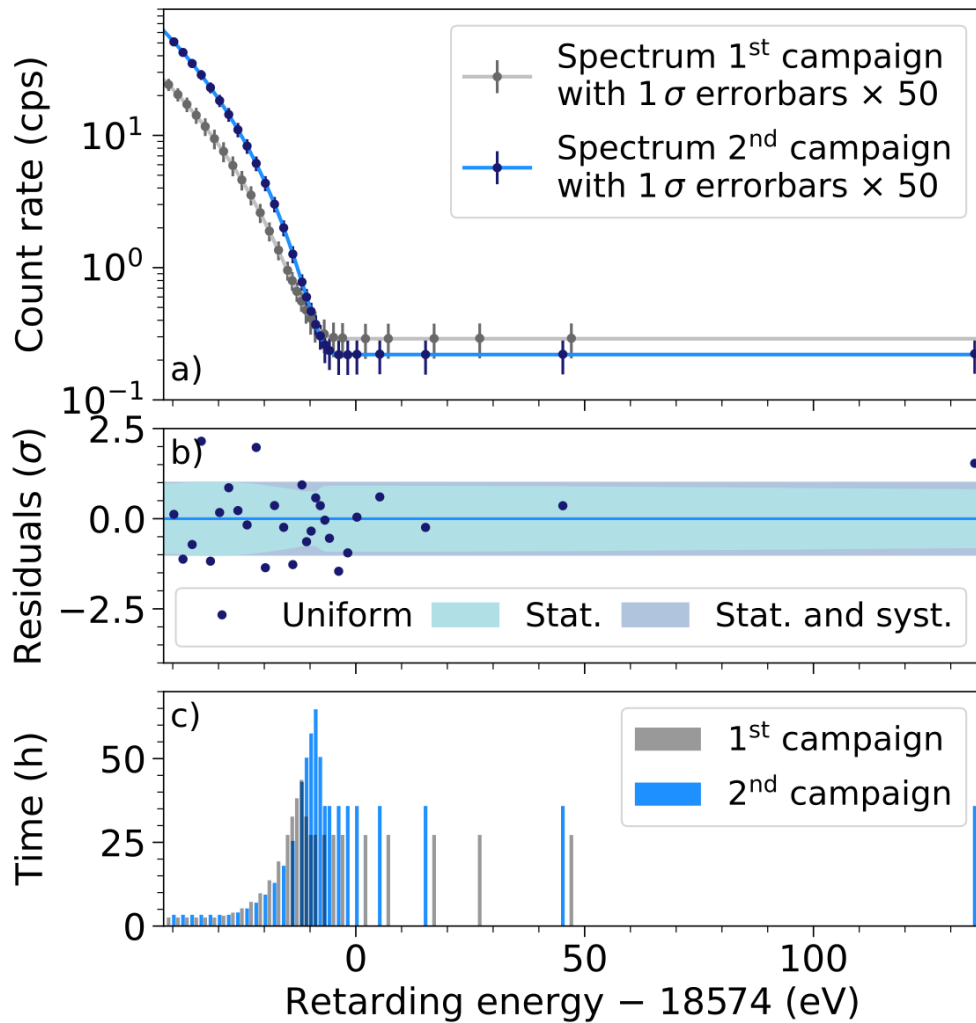
# KATRIN Data Taking Overview



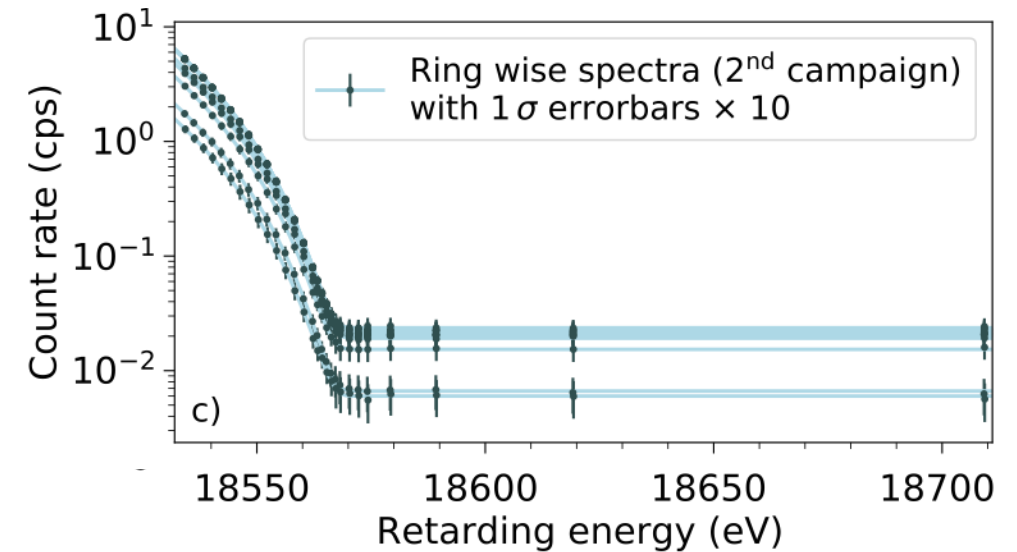
# Improvements wrt 1<sup>st</sup> campaign

	1 <sup>st</sup> campaign PRL 123 (2019)		2 <sup>nd</sup> campaign This talk
<b>Campaign date</b>	April-May 2019		Sept-Nov 2019
<b>Total scan time</b>	522 h (274 scans)		744 h (361 scans)
<b>Source activity</b>	25 GBq	<b>nominal activity</b> →	98 GBq
<b>Background</b>	290 mcps	<b>reduction -25%</b> →	220 mcps
<b>Tritium purity</b>	97.6%	<b>raised purity</b> →	98.7%
<b>Electrons in RoI</b>	2 Mio	<b>stats doubled</b> →	4.3 Mio

# New data release

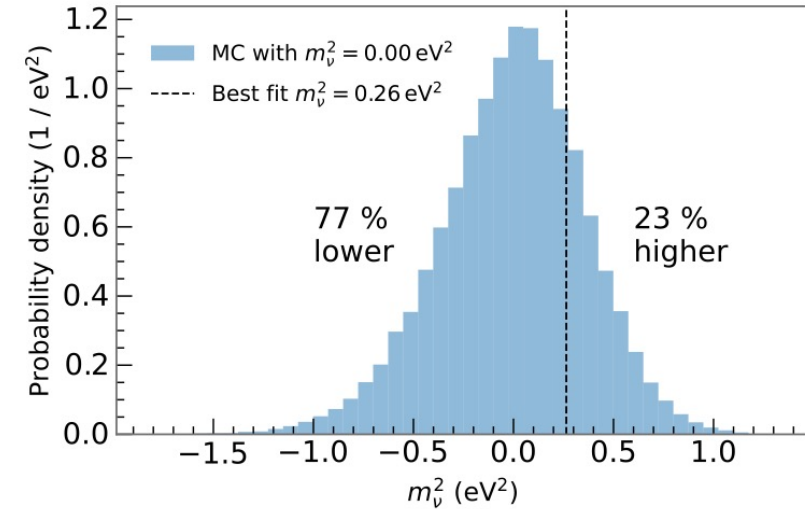
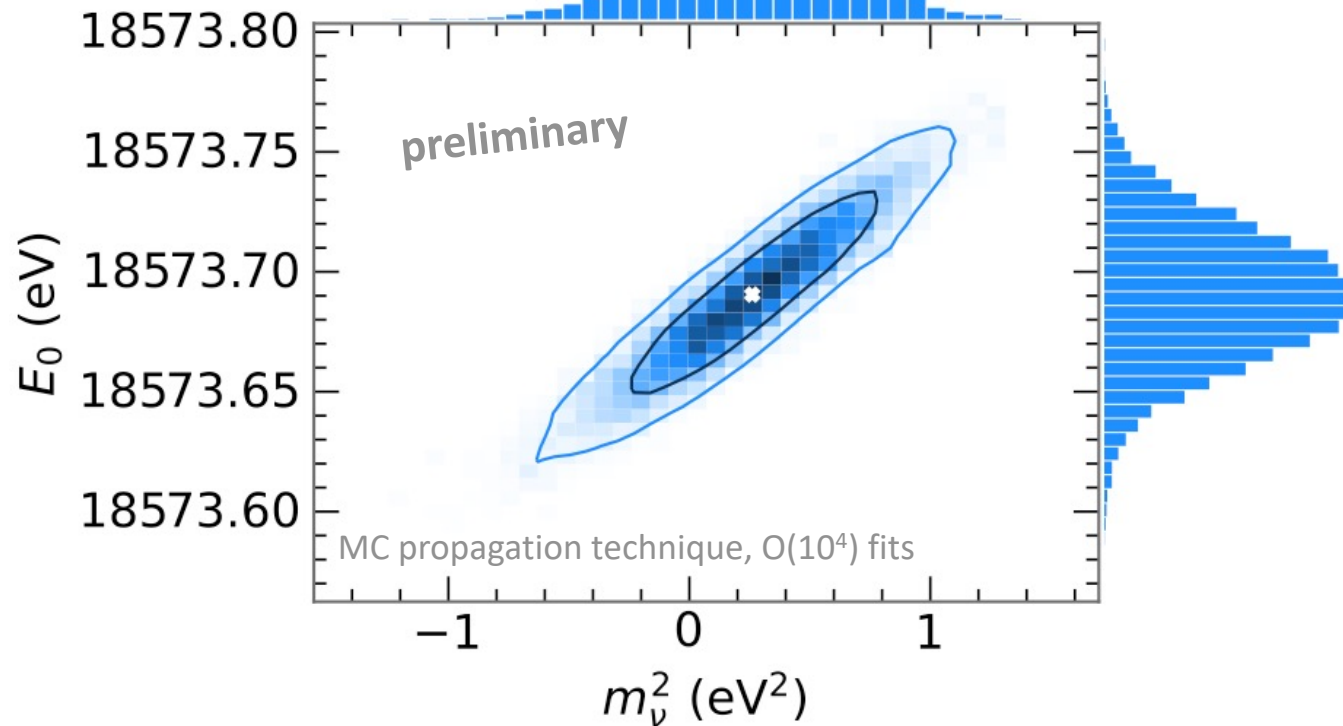


- total statistics: 4 million events
- excellent goodness-of-fit: p-value = 0.8
- Uncertainties are statistics dominated
- *Uniform* and *ring-wise* fit lead to consistent results





# Best Fit



$$m_\nu^2 = (0.26^{+0.34}_{-0.34}) \text{ eV}^2$$

✓ compatible with zero

$$E_0 = 18573.69 \pm 0.03 \text{ eV}$$

✓ Q-value :  $18575.2 \pm 0.5 \text{ eV}$

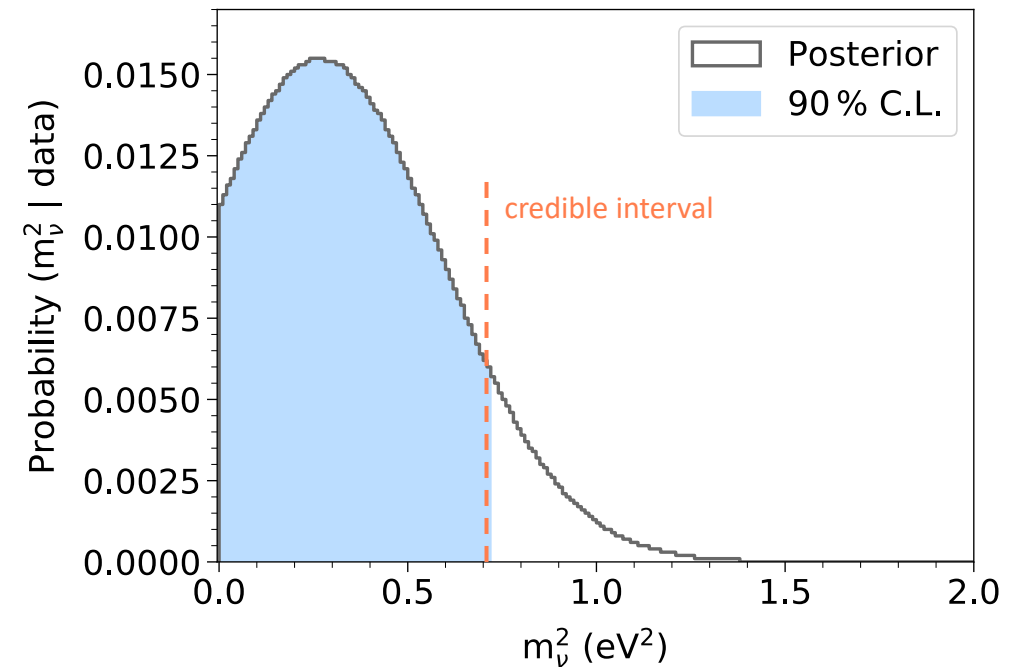
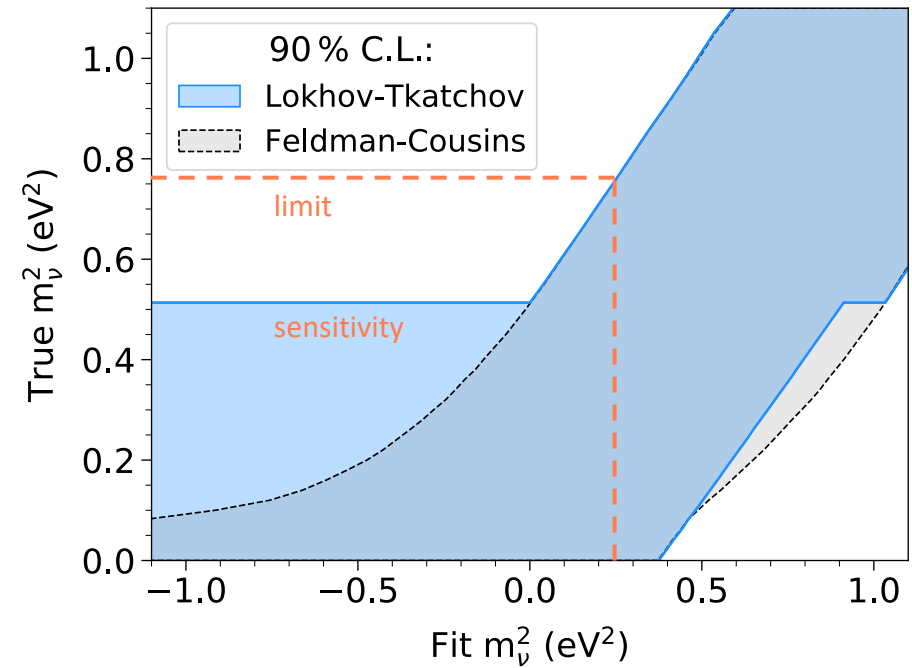
✓ good agreement with literature

Q =  $18575.72 \pm 0.07 \text{ eV}$

E. Myers et al. Phys. Rev. Lett. 114, 013003 (2015)

# New upper limit

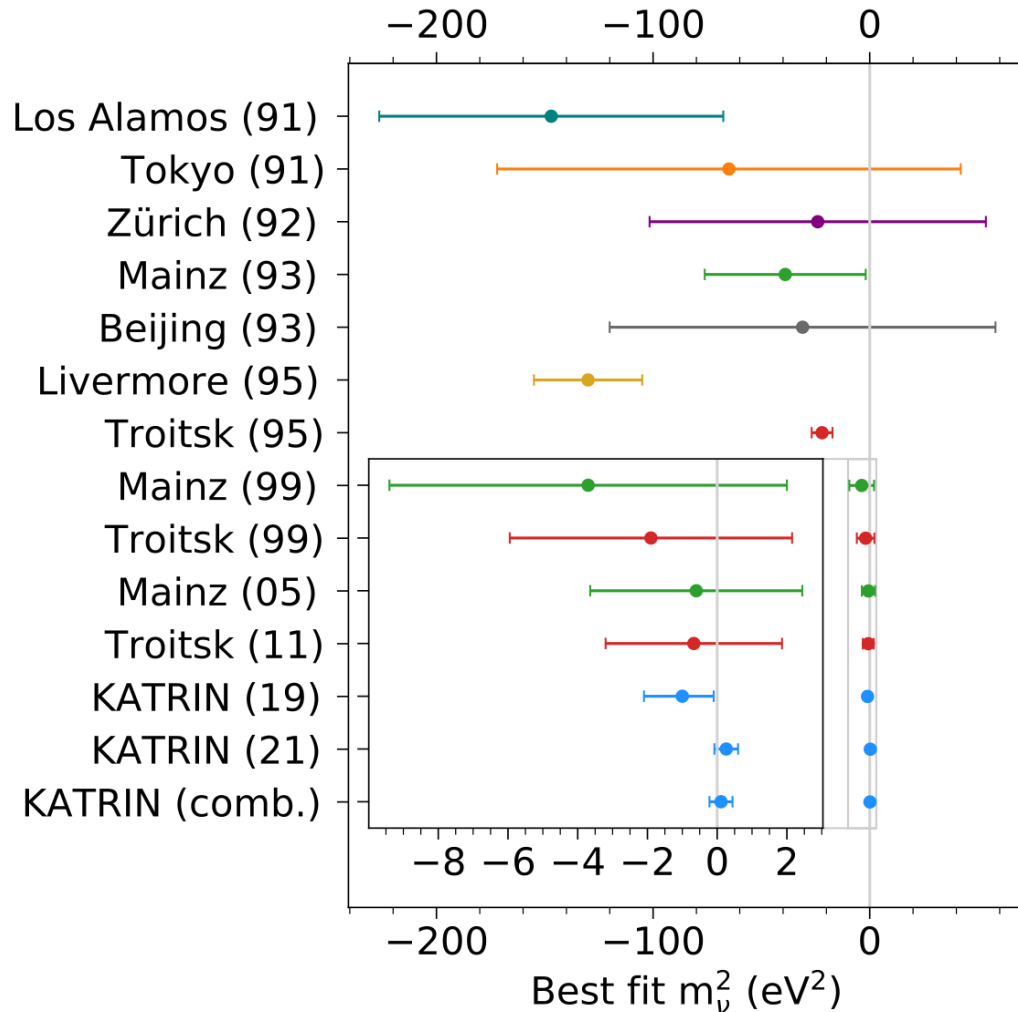
- Frequentist limit:  $m_\nu < 0.9$  eV (90% CL)
- Bayesian:  $m_\nu < 0.85$  eV (90% CI)
- Sensitivity:  $m_\nu < 0.7$  eV (90% CL)



Lokhov & Tkachov, Phys. Part. Nucl. 46 (2015) 347

Feldman & Cousins, Phys. Rev. D57 (1998) 3873

# Historical context



- KATRIN (2021):

**first direct neutrino-mass experiment to reach sub-eV sensitivity and limit**

- 1<sup>st</sup> and 2<sup>nd</sup> campaign combined result:

$$m_\nu^2 = (0.11^{+0.33}_{-0.33}) \text{eV}^2$$

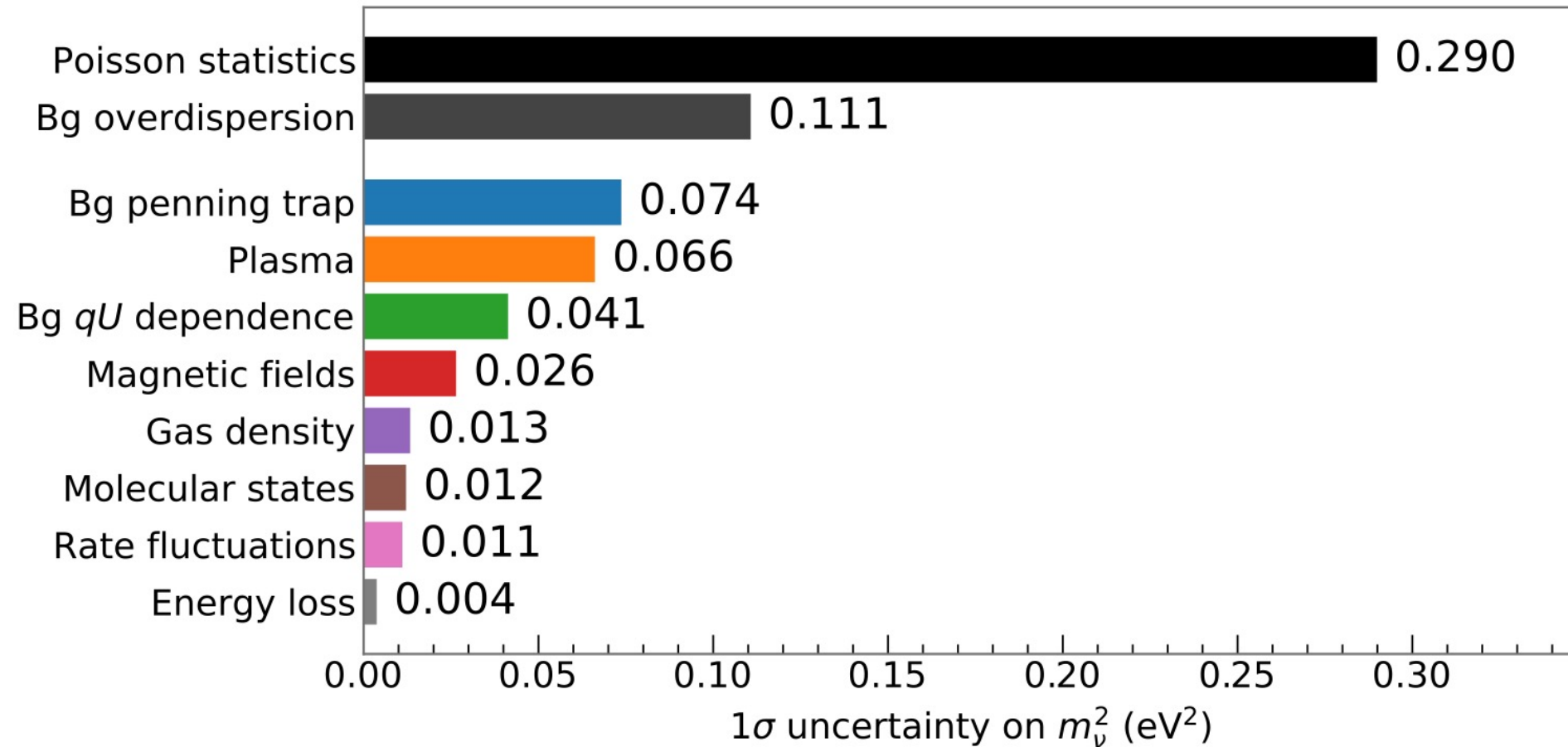
- 1<sup>st</sup> and 2<sup>nd</sup> campaign combined limit:

$$m_\nu < 0.8 \text{ eV (90\% CL)}$$

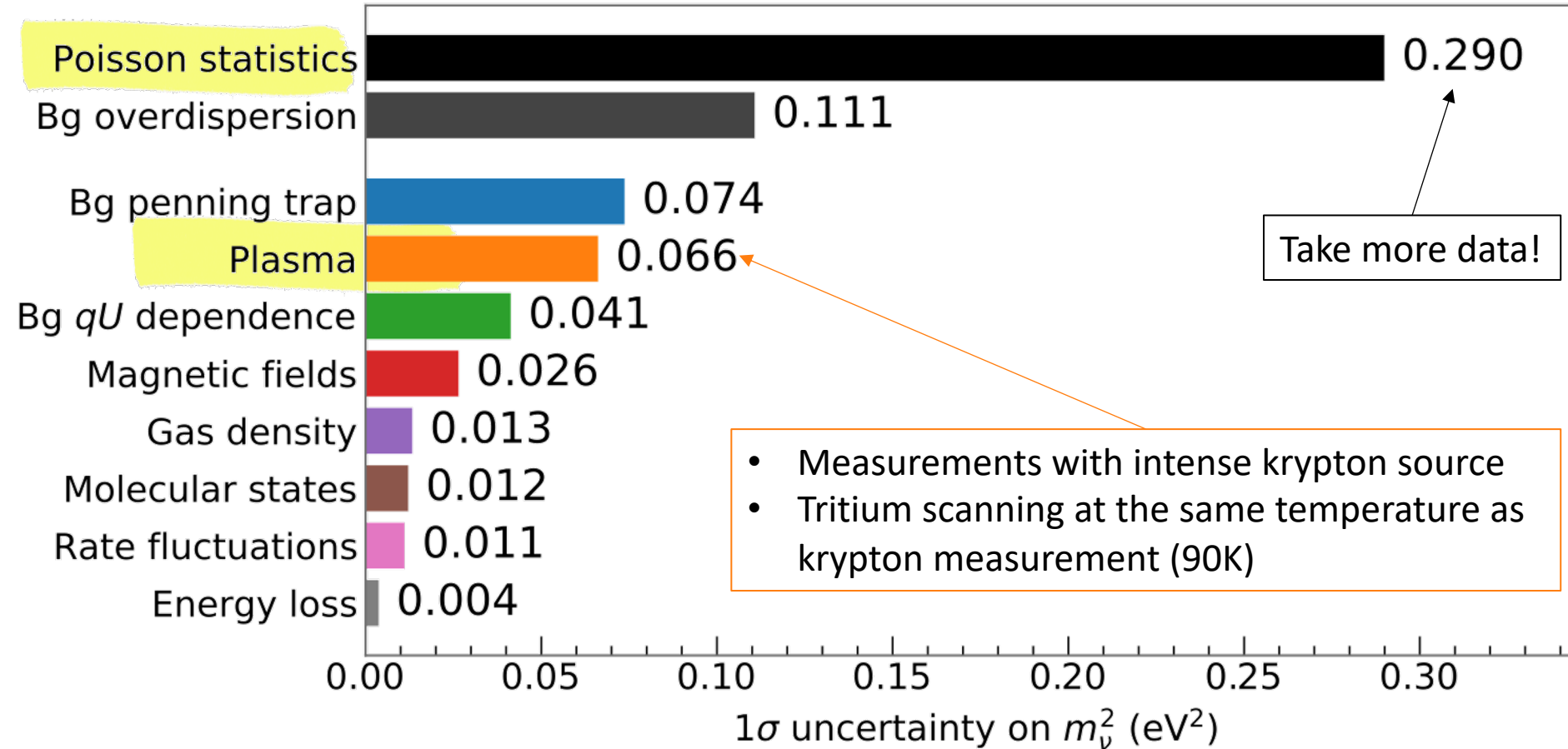
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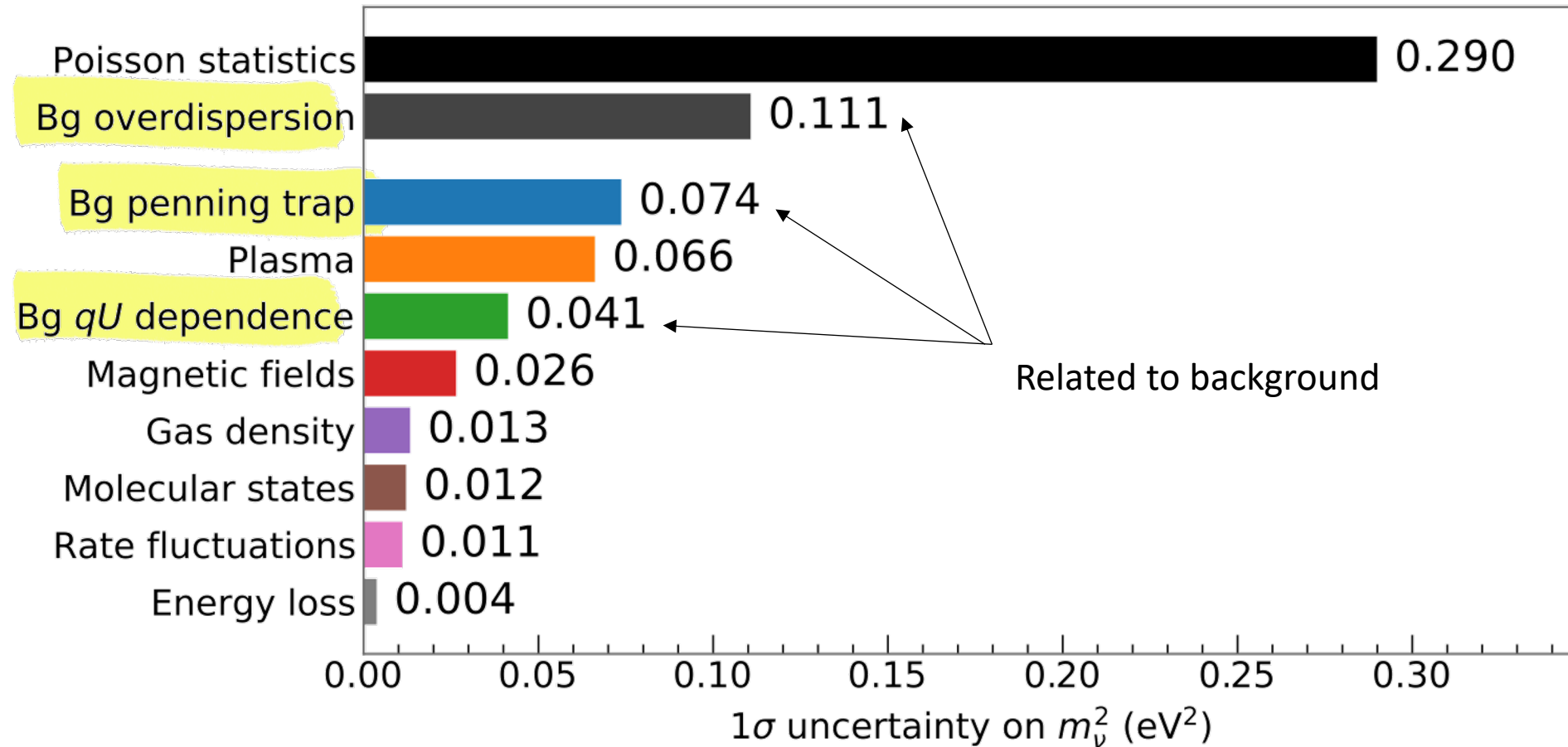
# Uncertainty budget of 2<sup>nd</sup> campaign



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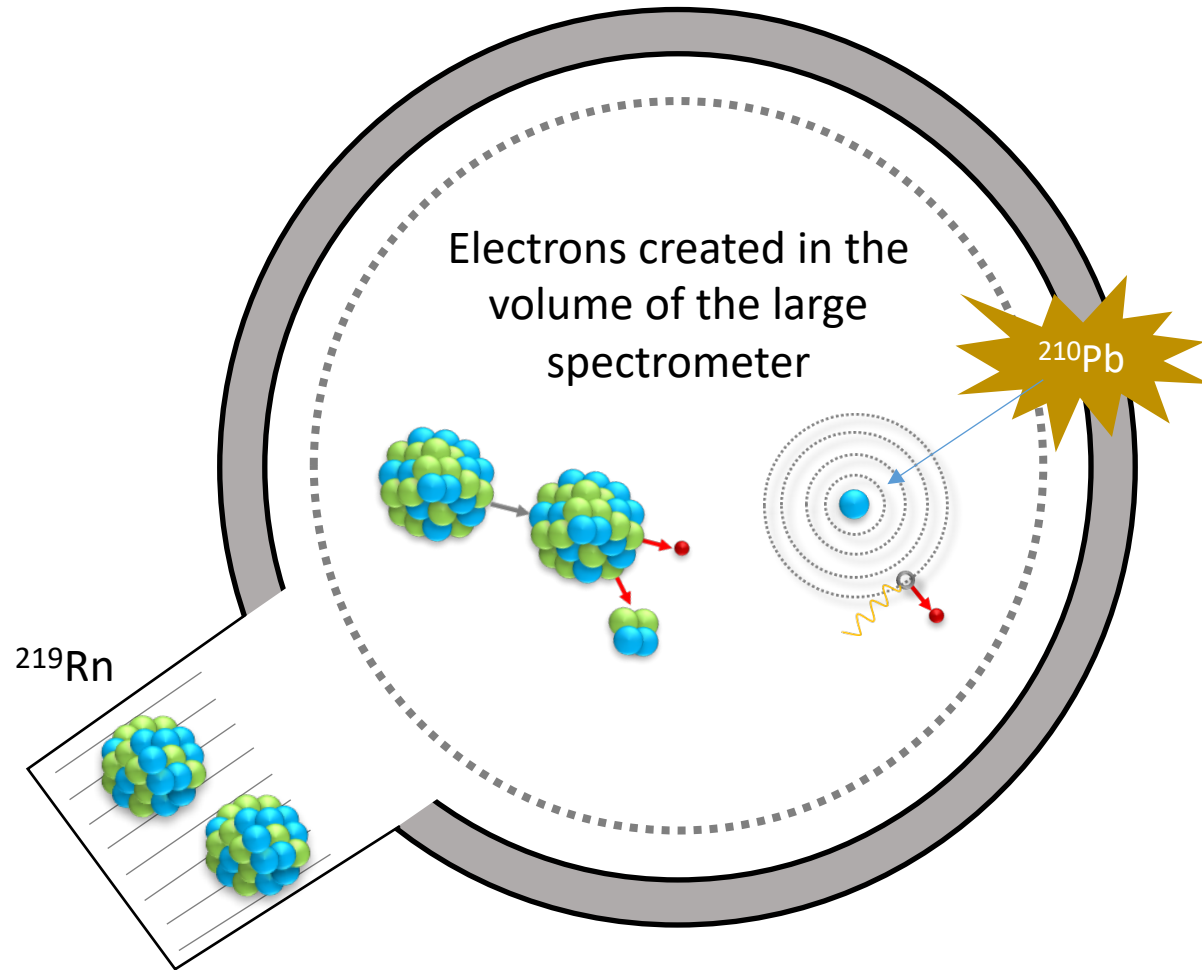


# KATRIN backgrounds

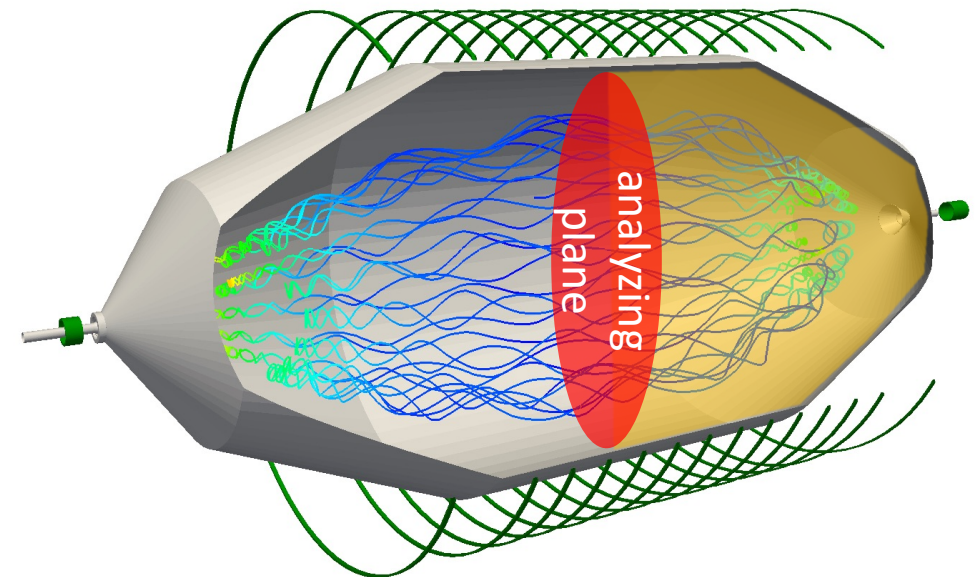




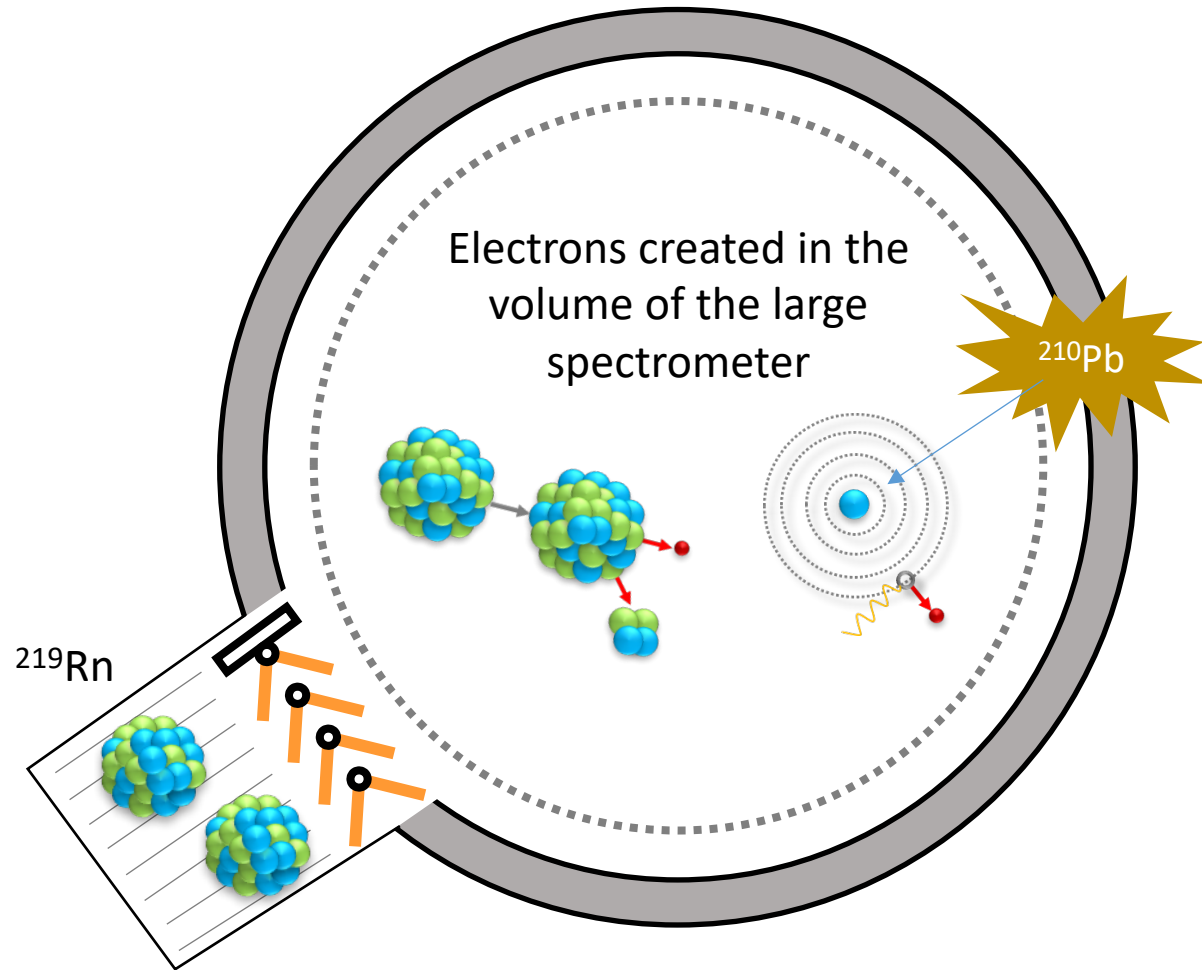
# KATRIN backgrounds



- $^{219}\text{Rn}$ -induced background
- $^{210}\text{Pb}$ -induced background



# KATRIN background mitigation

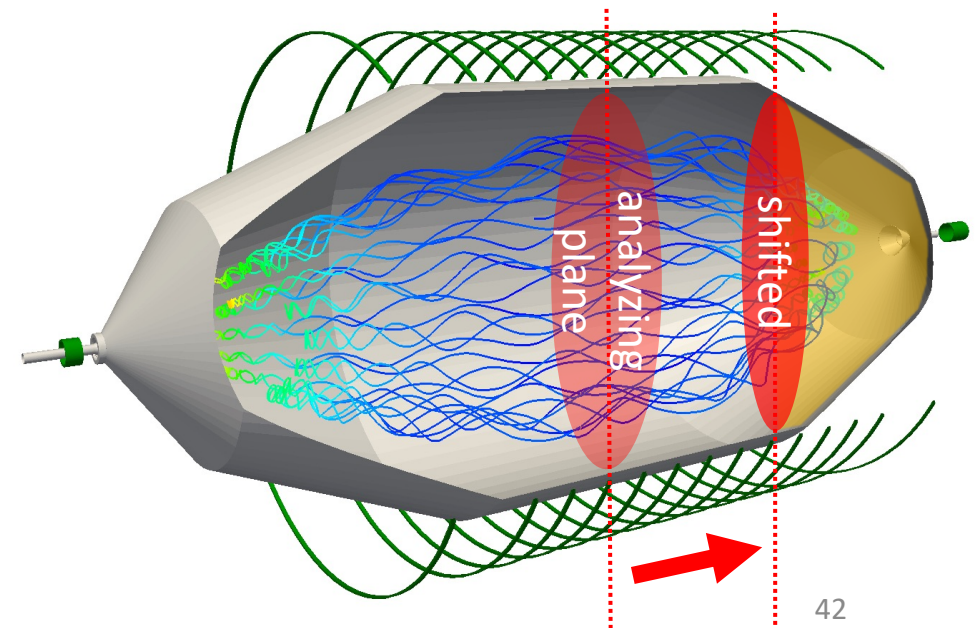


✓ LN cooled baffle + shifted analyzing plane

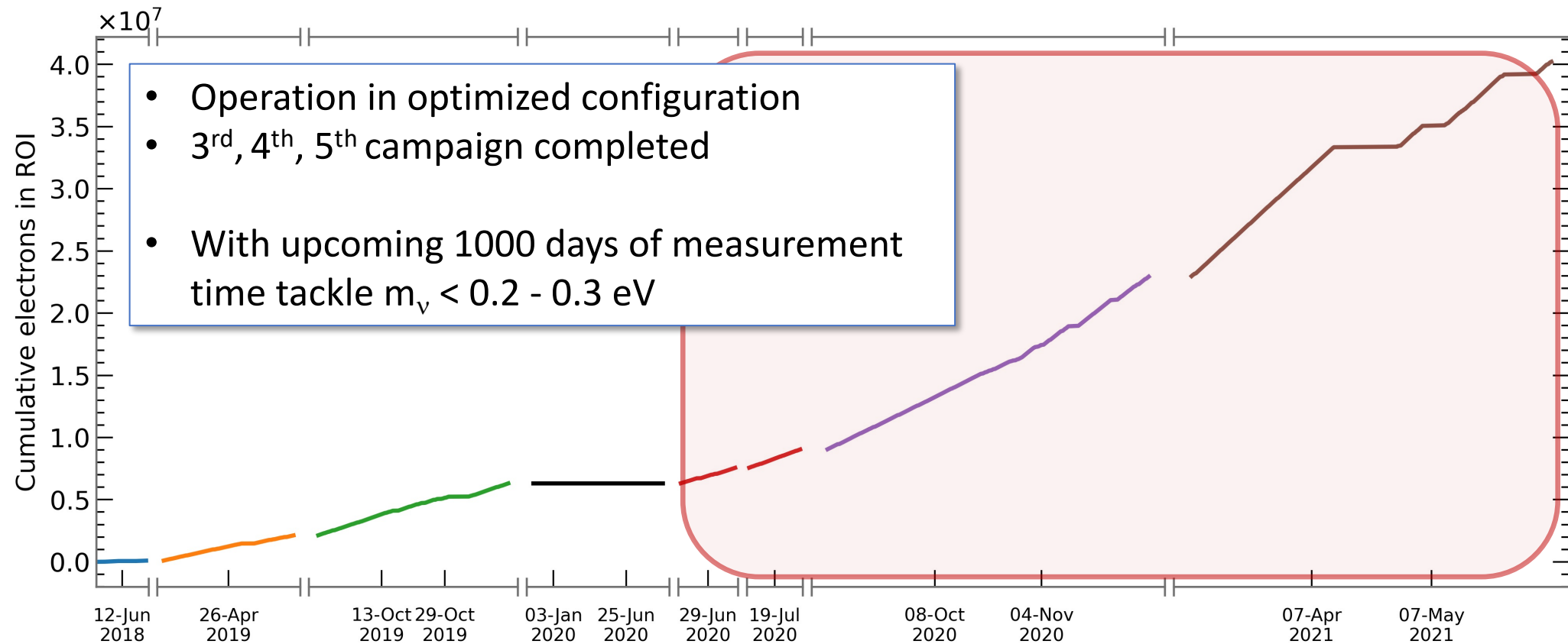
S. Goerhardt, et al., JINST 13 (2018) no.10, T10004

✓ Background reduction by factor of 2.3  
to about 130 mcps (original design: 10 mcps)

✓ Further R&D ongoing



# Outlook



# Overview

- How does KATRIN work ?
- What are the latest results?
- What's next?
- What else can we do with the data?

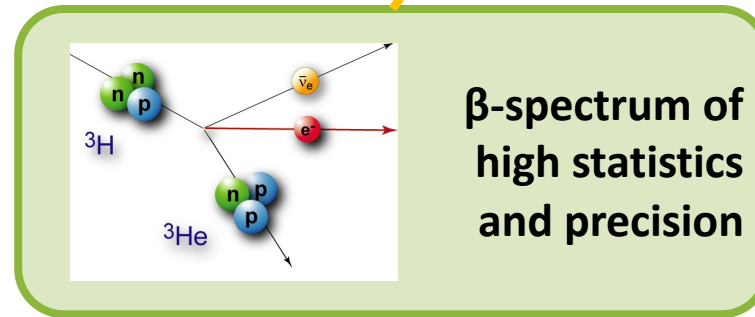
# New Physics with KATRIN



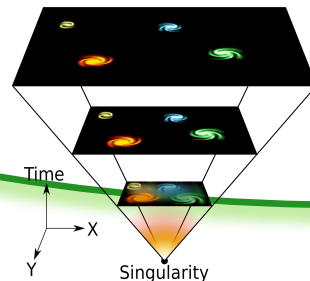
neutrino mass  
(*endpoint shape*)

Search for eV-keV sterile neutrinos (*kink search*)

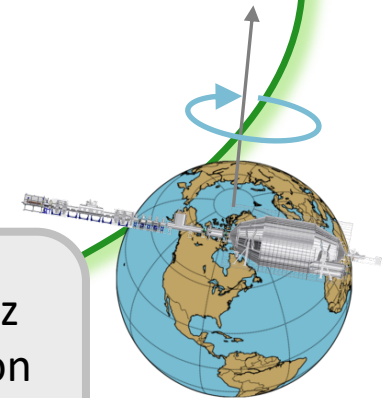
Search for exotic weak interactions  
(*spectrum shape*)



Constrain local overdensity of cosmic relic neutrinos  
(*peak search*)

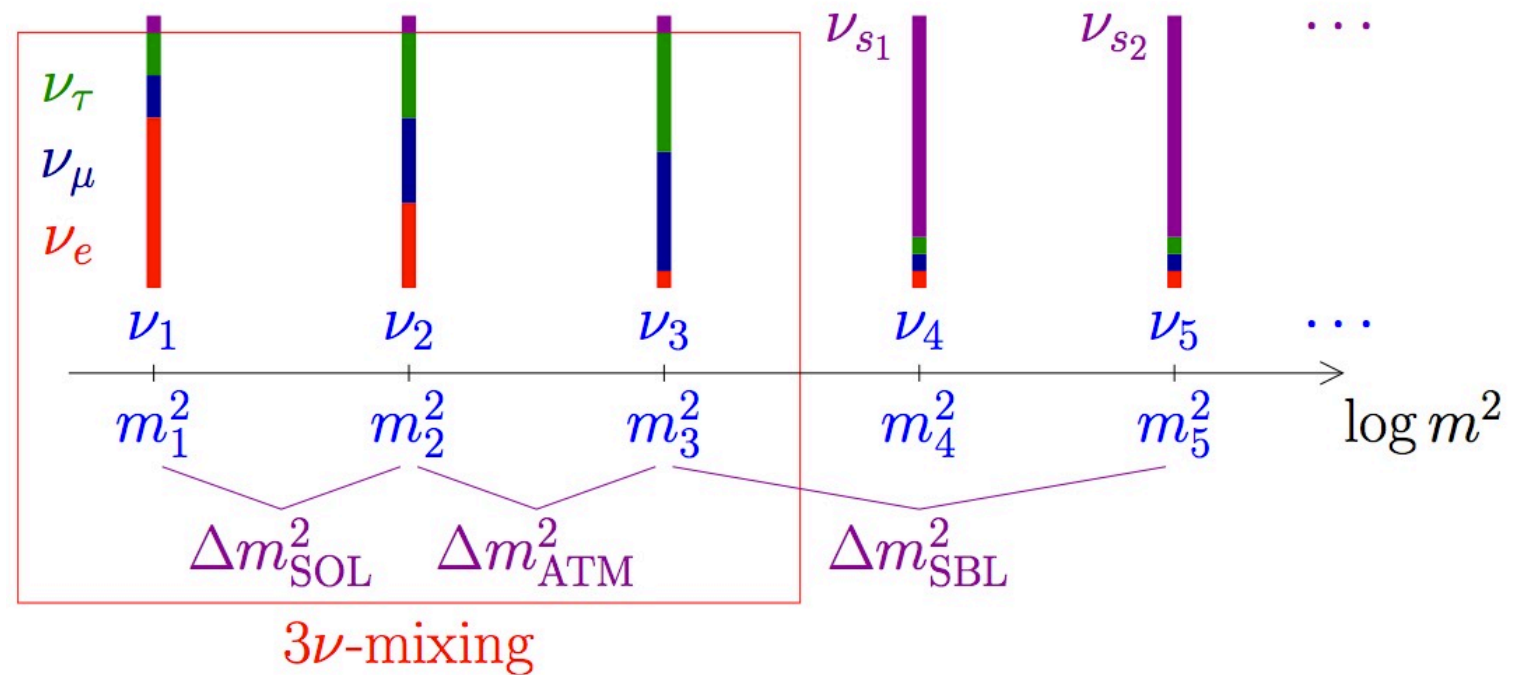


Search for Lorentz invariance violation  
(*sidereal modulation*)



# Sterile neutrinos

- Additional neutrino mass eigenstates or arbitrary scale
- Interaction via their mixing with active states



# Sterile neutrinos

## Heavy sterile neutrinos ( $> \text{GeV}$ )

- Lightness of neutrinos  
+ Matter/Anti-matter asymmetry

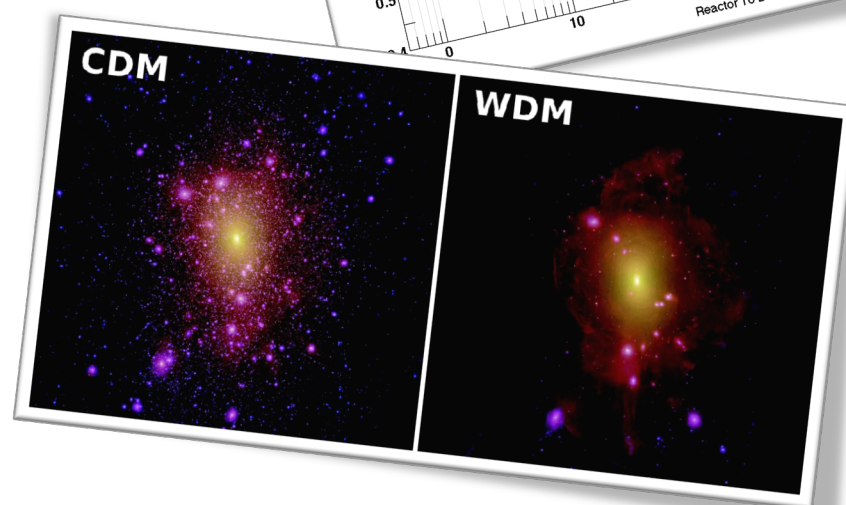
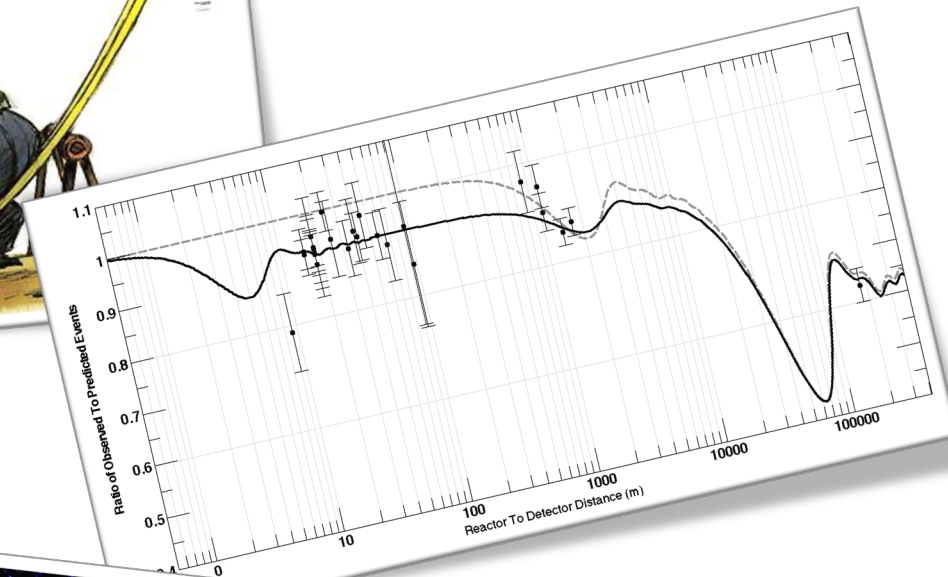
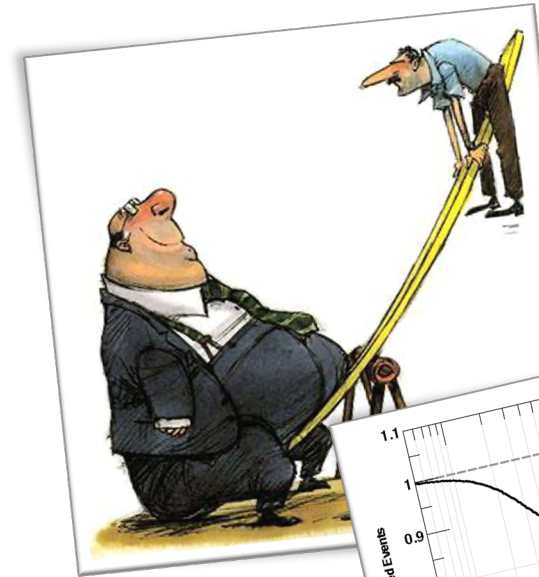
## Light sterile neutrinos ( $\sim 1 \text{ eV}$ )

- Short-baseline neutrino oscillation anomalies

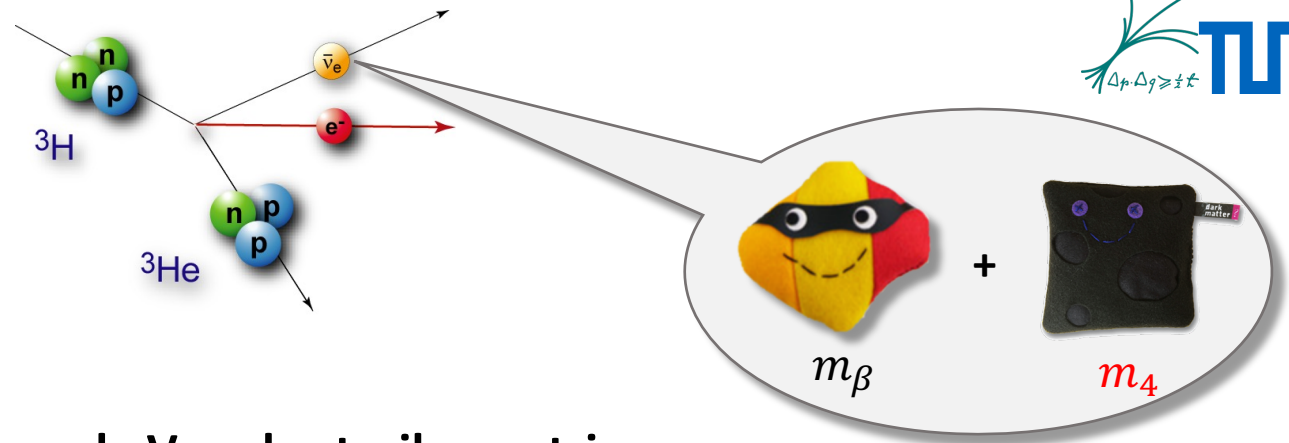
## KeV-scale sterile neutrinos ( $\sim 1 - 50 \text{ keV}$ )

- Dark matter candidate

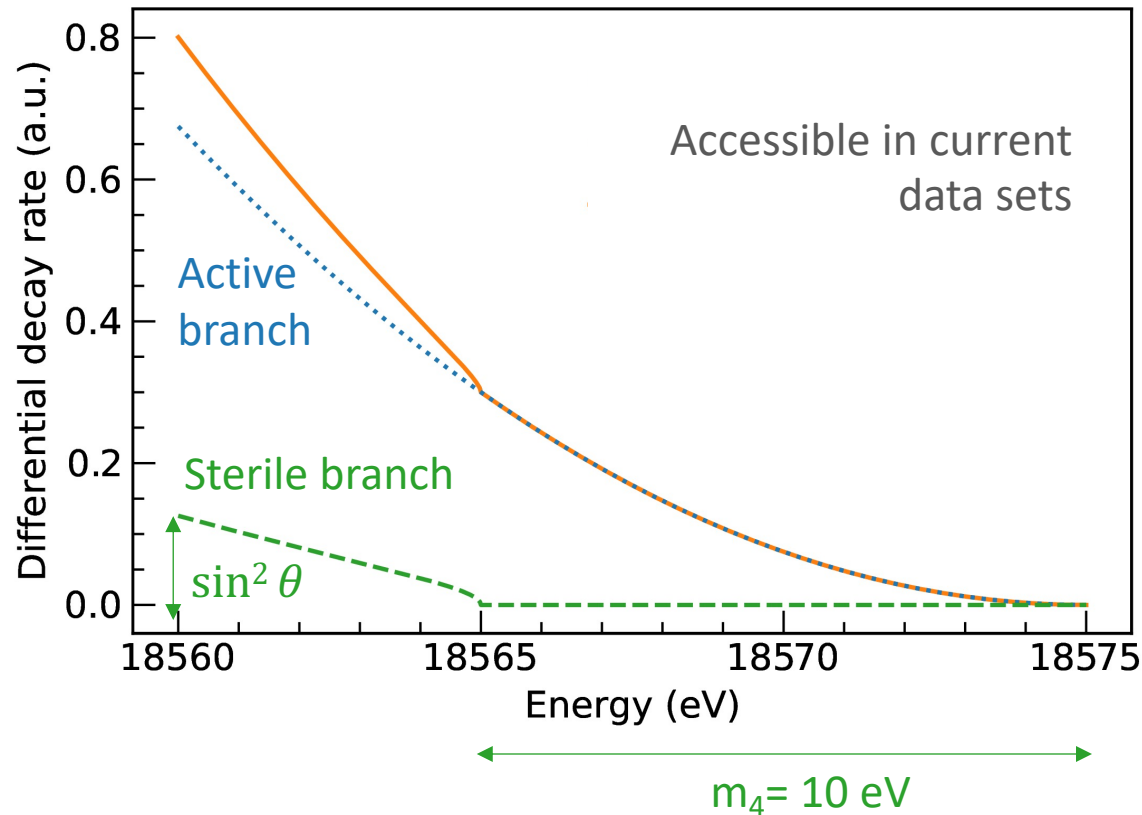
➤ Accessible in beta-decays



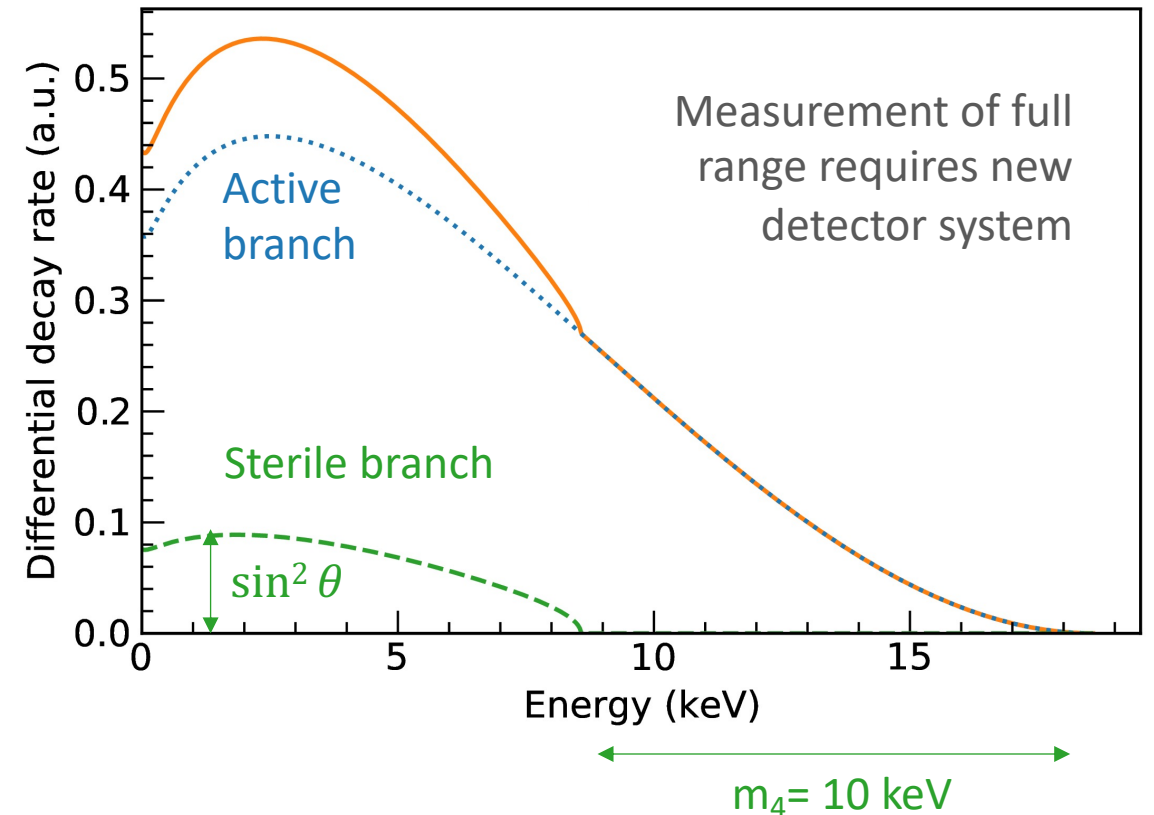
# Sterile neutrinos



## Light sterile neutrinos



## keV-scale sterile neutrinos

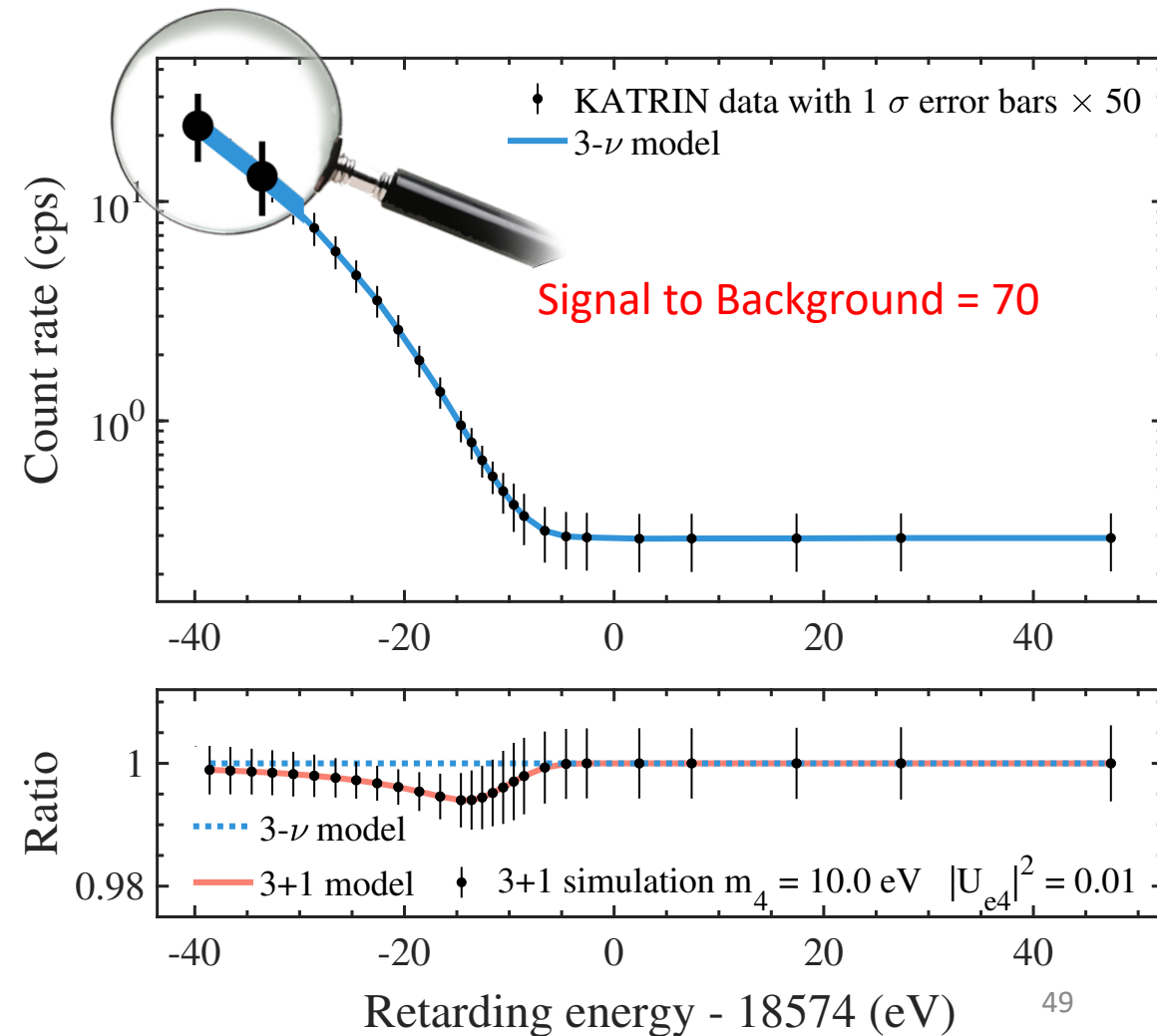




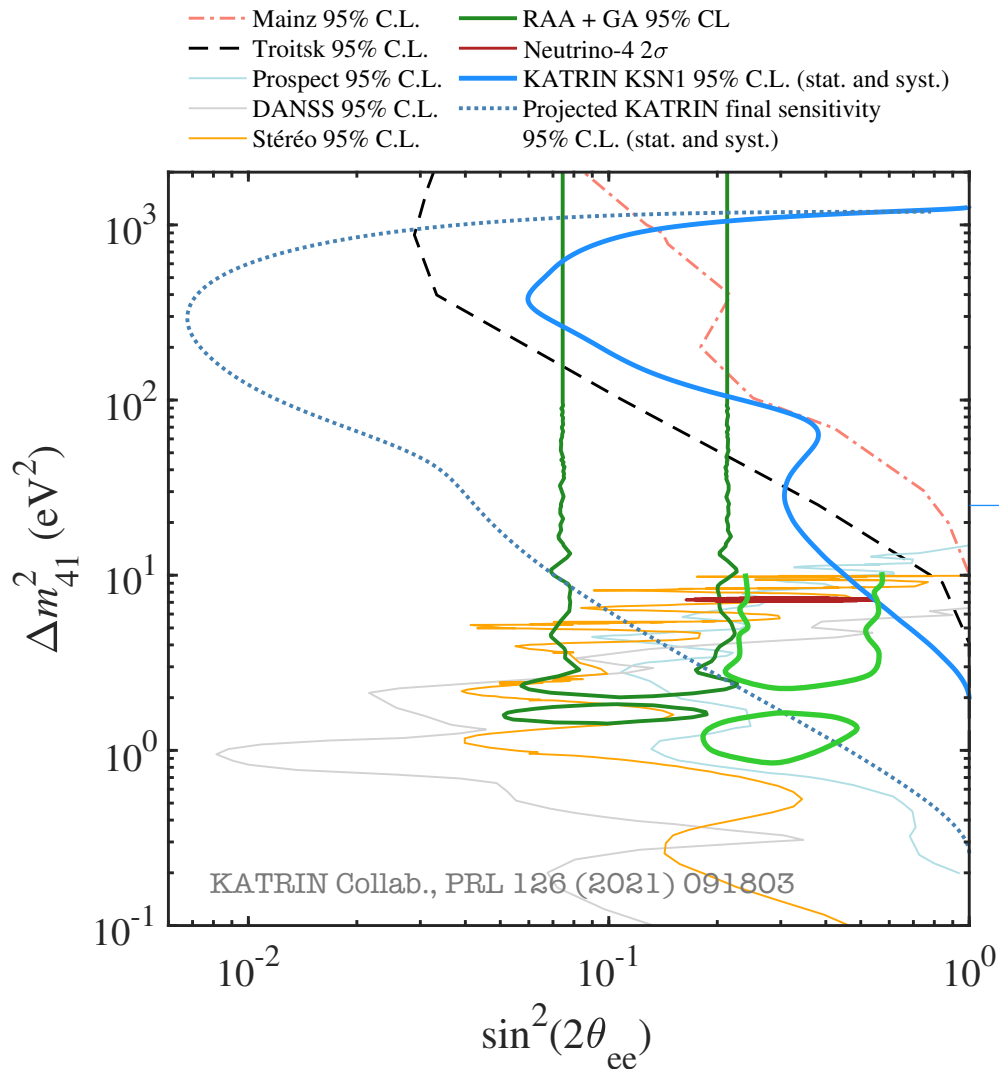
# eV-scale sterile neutrino search

- Search performed on data set of **first** and **second** neutrino mass campaign
- 3+1 sterile neutrino model
- Grid search in  $m_4, |U_{e4}|^2$  plane

$$\frac{d\Gamma}{dE} = \underbrace{(1 - |U_{e4}|^2)}_{\text{light neutrino}} \frac{d\Gamma}{dE}(m_\beta^2) + \underbrace{|U_{e4}|^2}_{\text{heavy neutrino}} \frac{d\Gamma}{dE}(m_4^2)$$



# eV-scale sterile neutrino search (1st campaign)



## High $\Delta m_{41}$ region:

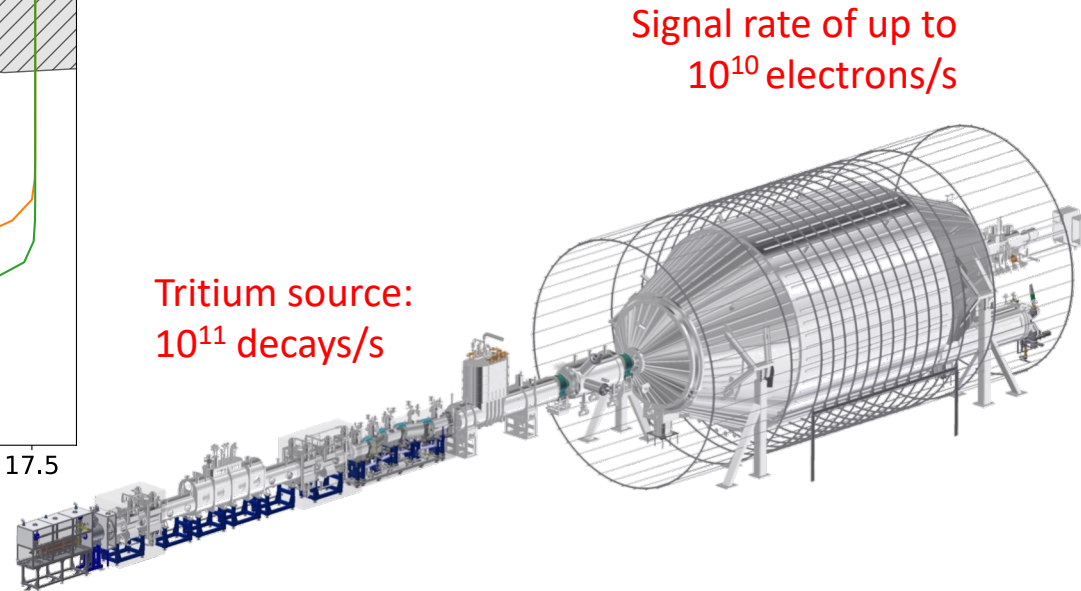
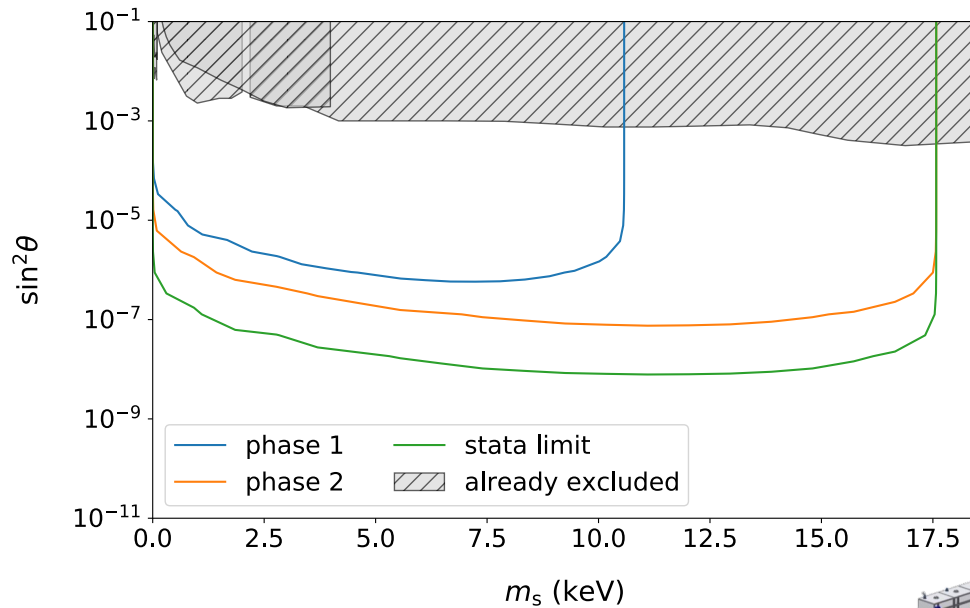
- ✓ Improve exclusion with respect to DANSS, PROSPECT, STEREO
- ✓ Exclude parameter space of Reactor Anomaly (RAA)

## Low $\Delta m_{41}$ region:

- ✓ Improve MAINZ and TROITSK limit
- ✓ The NEUTRINO-4 hint at the edge of exclusion limit
- ✓ Test part of BEST result with future data

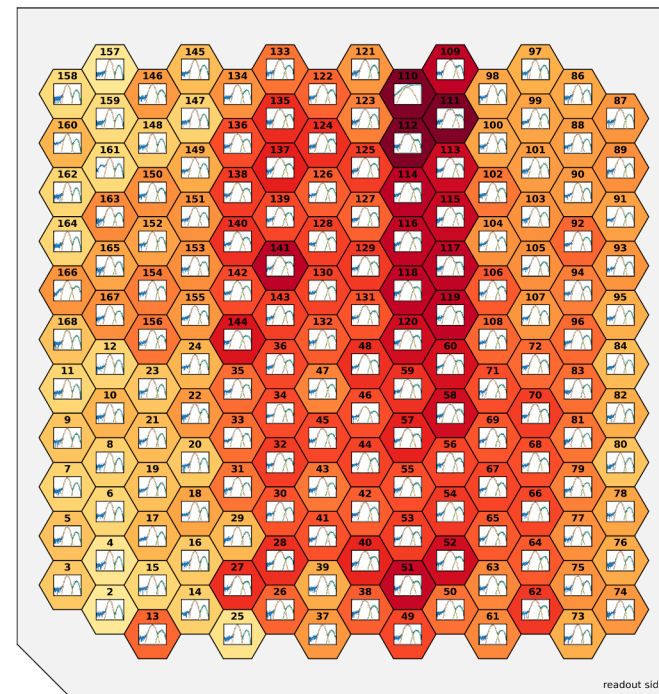
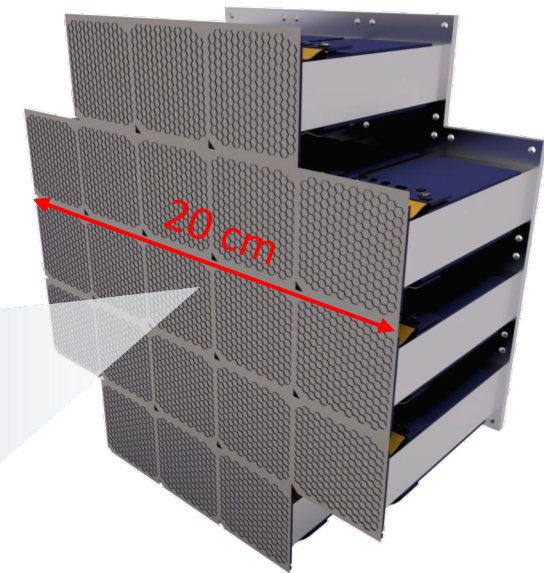
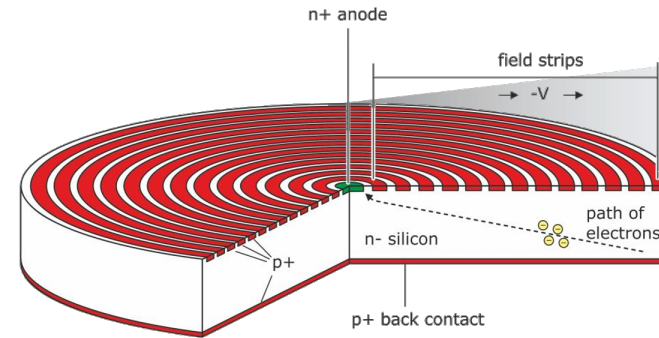
# keV-scale sterile neutrinos

- Idea: make use of the luminous KATRIN source to explore full beta spectrum to search for BSM physics
- Develop a novel detector system

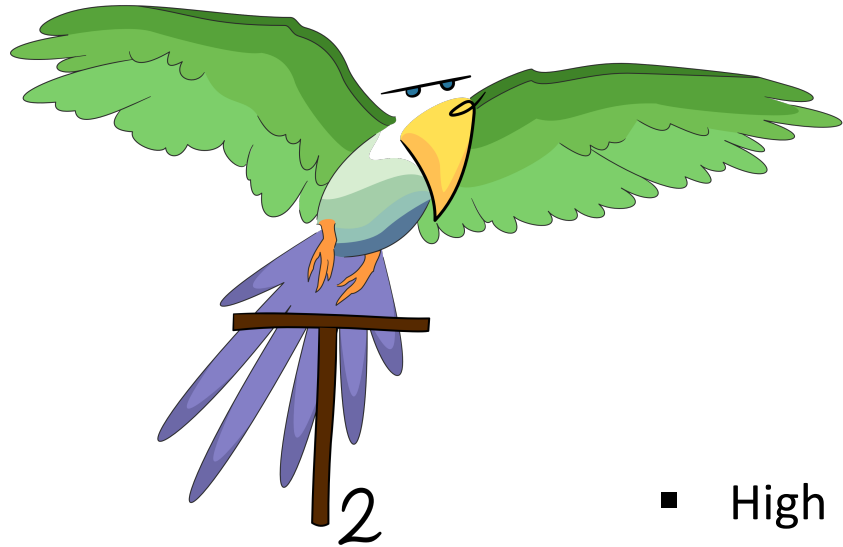


# TRISTAN detector

- Silicon drift detector (SDD) technology
  - ✓ Capability of handling high rates ( $> 10^8$  cps)
  - ✓ Excellent energy resolution (300 eV @ 20 keV)
- Challenge
  - Control of systematics at the ppm-level
  - Operation of 3500 pixel focal plane array
- Status
  - ✓ Excellent performance of prototypes
    - S. Mertens et al, J. Phys. G46 (2019)
    - S. Mertens et al, J. Phys. G48 (2020)
    - M. Gugiatti et al, NIM-A 979 (2020)
    - M. Biassoni et al, Eur. Phys. J. Plus 136, 125 (2021)
    - P. King et al JINST 16 T07007 (2021)
  - ✓ Operation of 166-pixel module (largest SDD module ever operated) – thanks to Polimi and Bicocca



# Conclusion



- First sub-eV neutrino mass limit from a direct experiment
- Various improvements of systematics and background in place
- Sensitivity close to 0.2 eV (90% CL) targeted within next years
- High precision KATRIN data available for interesting new physics searches
- Upgrade of KATRIN beamline with SDD array will allow to extend the measurement interval

Thank you for your attention

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Thanks to  
To my group  
The KATRIN collaboration  
Thierry Lasserre  
And many others  
...

Susanne Mertens

Technical University Munich & Max Planck Institute for Physics