

**Università di Padova** 7-10 June 2021



### Contents

### 1. Inter-nucleon forces

- Brief introduction to the nuclear many-body problem
- Properties and modelling of nuclear forces
- The modern view: chiral effective field theory

### 2. Ab initio techniques for the nuclear many-body problem

- Configuration-interaction approaches
- Techniques to mitigate the "curse of dimensionality" (SRG, NO2B, IT)
- $\circ$  Mean field and correlations
- Expansion methods for closed-shell nuclei
- Symmetry breaking
- Expansion methods for open-shell nuclei
- State of the art and open problems

#### 3. Equation of state of nuclear matter & connections to astrophysics

- Neutron stars & Tolman-Oppenheimer-Volkoff equations
- Equation of state of neutron-star matter
- Astrophysical constraints on the nuclear EoS

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### Correlation expansion methods: the idea



• Must be simple enough (such that it can be computed easily and exactly)

• Must be rich enough (such that it is a suitable starting point for the expansion)

 $\circ$  Obtained by

1) Splitting  $H = H_0 + H_1$  2) Solving for  $H_0$  (one-body operator)  $H_0 |\phi_k\rangle = \epsilon_k |\phi_k\rangle$ 

# Concept of mean field

• Slater determinant as reference state

$$\circ \text{ One-body potential: } H_0 = \sum_{i=1}^A h_0(i) \rightarrow H_0 |\phi_k\rangle = \epsilon_k |\phi_k\rangle \implies h_0 |\alpha\rangle = \varepsilon_\alpha |\alpha\rangle \quad \forall i$$
  

$$A \text{ one-body problems}$$
  

$$\circ \text{ Build Slater } |\phi_0\rangle = \prod_{i=1}^A a_{\alpha_i}^{\dagger} |0\rangle$$

• Independent particles: nucleons move inside a (one-body) potential well or *mean field* 

#### • Does an independent-particle picture make any sense at all?

- $\,\circ\,$  Inter-particle distance in nuclei ~ 2 fm
- $\circ$  Range of nuclear interaction ~ 2 fm



- ✓ Fermi statistics helps out
- ✓ Large mean free path  $\lambda$ 
  - → Quantum liquid

[Rios & Somà 2012; Lopez *et al.* 2014]



# Liquid drop model & semi-empirical mass formula

• Picture the nucleus as a (suspended) drop of (incompressible) liquid with surface tension



# Non-interacting shell model



### • What creates regular patterns?

- Nucleon shells? (cf. electrons in the atom)
- $\circ$  Yet, no obvious common potential



### Empirical mean-field potentials



### Hartree-Fock with realistic potentials



### Dynamical vs. static correlations

• **Dynamical**/short-range/weak/single-particle correlations

- **Dynamical**: characterised by short time scales & high excitation energies
- **Short-range**: governed by short-distance dynamics
- $\circ$  **Weak**: perturbative (at least with low-resolution interactions)
- **Single-particle**: can be described in terms of excitations of a few nucleons
  - Short-range correlations
  - Sociated with short-range repulsion of NN interactions
- **Static**/long-range/strong/collective correlations
- Static: characterised by long time scales & low excitation energies
- Long-range: governed by long-distance dynamics
- **Strong**: strongly non-perturbative
- **Collective**: description involves coherent excitations of many particles
  - Section State Control State Control State St
  - Solution → Associated with presence of bound *np* and virtual *nn* pairs (at least pairing correlations)

# Closed- vs. open-shell systems

• Mean-field configurations have different features as *N* and *Z* vary

Nucleons **entirely** fill levels below a magic number **Closed-shell** systems Larger energy gap, excitations hindered, enhanced stability

Dynamical correlations drive bulk properties\* Nucleons partially fill levels below a magic number  $\underbrace{\phantom{}}$ 

Smaller (→ 0) energy gap, excitations enabled, lesser stability

Both static & dynamical correlations are important

\*Each observables is impacted differently

# Correlationsiappartitlellalateticitations

● In configuration-space methods, correlations are accounted for by means of **ph excitations** 



• **Configuration-interaction techniques** (e.g. no-core shell model)

- Few-p-few-h and many-p-many-h excitations treated on an equal footing
- Efficient treatment of both dynamical and static correlations

### Expansion techniques

- Expansion in the rank *k* of *kp-k*h excitations
- Efficient treatment of **dynamical** correlations, difficult to treat **static** correlations

# Correlations via particle-hole excitations

• The two different strategies reflect into two different truncations of the Hilbert space

### Configuration-interaction techniques

• Truncation in terms of the total number of HO excitation quanta N<sub>max</sub> of the many-body states

E.g.  $N_{max} = 6$  calculation No-core shell model





### • Expansion techniques

- $\circ$  Truncation in terms of the energy of the states included in the one-body basis
- (On top of this there's the truncation on the rank *k* of *k*p-*k*h excitations that are included)







### Closed vs. open shells



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### Formal perturbation theory

• Very popular in description of **many-electron systems** (since early days)

- Coulomb interaction sufficiently weak to allow perturbative treatment
- Applications in many-nucleon systems traditionally hindered by strong short-range repulsion
   SRG techniques have completely changed this view → renaissance of MBPT in nuclear physics
   Starting point: splitting of the many-body Hamiltonian ("partitioning")

$$U = U + U$$

$$H \equiv H_0 + H_1$$
  
unperturbed perturbation

• Eigenvalue equation for *H*<sup>0</sup> must be **numerically accessible** 

 $H_0|\Phi_k\rangle = E_k^{(0)}|\Phi_k\rangle \rightarrow \text{set of unperturbed eigenstates and eigenenergies} \quad \{|\Phi_k\rangle, E_k^{(0)}; k \in \mathbb{N}\}$ 

### form orthonormal basis of Hilbert space

• In the following  $|\Psi_0^A\rangle = |\Psi^A\rangle$   $|\Phi_0\rangle = |\Phi\rangle$   $E_0^{(0)} = E^{(0)}$ 

 $\odot$  Goal: approach exact (g.s.) wave function and energy by systematically including effects of  $H_1$ 

• Define **projectors** associated with the partitioning of H

$$P \equiv |\Phi\rangle\langle\Phi|$$
  $Q \equiv 1 - P$  with  $P|\Phi\rangle = |\Phi\rangle$   $Q|\Phi\rangle = 0$ 

$$\rightarrow Q \text{ can be written as} \qquad Q \equiv \sum_{k}' |\Phi_k\rangle \langle \Phi_k| \equiv \sum_{|\Phi_k\rangle \neq |\Phi\rangle} |\Phi_k\rangle \langle \Phi_k|$$

• Express the exact wave function as

$$|\Psi^{A}\rangle = P|\Psi^{A}\rangle + Q|\Psi^{A}\rangle$$
  
=  $|\Phi\rangle + |\chi\rangle$   
unperturbed wave function correlated wave function

• It follows that the exact energy can be written as

$$E^{A} = \langle \Phi | H | \Psi^{A} \rangle$$
  
=  $\underbrace{\langle \Phi | H_{0} | \Phi \rangle + \langle \Phi | H_{1} | \Phi \rangle}_{\equiv E_{ref}} + \underbrace{\langle \Phi | H_{1} | \chi \rangle}_{\equiv \Delta E}$   
reference energy correlation energy

• Introduce resolvent operator

$$R^{\rm RS} \equiv \sum_{k}' \frac{|\Phi_k\rangle \langle \Phi_k|}{E^{(0)} - E_k^{(0)}}$$
  
Rayleigh-Schrödinger

which has the property  $R^{RS}|\Phi\rangle = 0$ 

• Finally, one gets

$$|\chi\rangle = \sum_{k=1}^{\infty} (RH_1)^k |\Phi\rangle_c \qquad \text{correlated wave function}$$
$$\Delta E = \langle \Phi | H_1 \sum_{k=1}^{\infty} (RH_1)^k |\Phi\rangle_c \qquad \text{correlation energy}$$

and can write exact wave function and g.s. energy as **power series in**  $H_1$ 

$$|\Psi^{A}\rangle \equiv \sum_{p=0}^{\infty} |\Psi^{(p)}\rangle \qquad E^{A} \equiv \sum_{p=0}^{\infty} E^{(p)} \qquad \underbrace{\text{e.g.}}_{k} \qquad E^{(2)} = \sum_{k}' \frac{\langle \Phi | H_{1} | \Phi_{k} \rangle \langle \Phi_{k} | H_{1} | \Phi \rangle}{E^{(0)} - E^{(0)}_{k}}$$

#### • Application to many-nucleon systems

- Characterise unperturbed Hamiltonian
- Define basis, derive working expressions (many-body matrix elements & unperturbed energies)
- $\circ$  Use Wick's theorem  $\rightarrow$  many-body diagrams

#### • Reference state: Slater determinant

$$|\Phi
angle\equiv\prod_{i=1}^{A}c_{i}^{\dagger}|0
angle$$

#### Normal-ordered Hamiltonian

• Apply Wick's theorem with respect to the reference Slater to *H* 

$$H = H^{[0]} + \sum_{pq} H^{[2]}_{pq} : c_p^{\dagger} c_q : + \frac{1}{4} \sum_{pqrs} H^{[4]}_{pqrs} : c_p^{\dagger} c_q^{\dagger} c_s c_r : + \dots$$

NO2B approximation

 $\circ$  Effective (normal-ordered) operators

◦ Each NO operator of rank  $k_{eff}$  receives contributions from original operators with  $k_{eff} ≤ k ≤ k_{max}$ 

#### • Partitioning

• Add and subtract a diagonal (NO) one-body operator

• Recall: an **orthonormal basis of the** *A***-body Hilbert space** can be built via Slater determinants

$$\mathcal{H}^{A} = \{ |\Phi\rangle, |\Phi_{i}^{a}\rangle, |\Phi_{ij}^{ab}\rangle, |\Phi_{ijk}^{abc}\rangle, ... \} \qquad \text{where} \quad |\Phi_{ij\cdots}^{ab\cdots}\rangle \equiv c_{a}^{\dagger}c_{b}^{\dagger}\cdots c_{j}c_{i}|\Phi\rangle$$

 $\rightarrow$  eigenbasis of  $H_0$ 

$$H_0|\Phi^{ab\cdots}_{ij\cdots}\rangle = (H^{[0]} + \epsilon^{ab\cdots}_{ij\cdots})|\Phi^{ab\cdots}_{ij\cdots}\rangle$$

with 
$$\epsilon_{ij\cdots}^{ab\cdots} \equiv (e_a + e_b + \cdots) - (e_i + e_j + \cdots)$$

 $H_0|\Phi\rangle = H^{[0]}|\Phi\rangle$ 

• **Convention**: one-body states occupied (unoccupied) in the reference determinant are labeled by i,j,k, ... (a,b,c, ...) and are referred to as **hole** (**particle**) states

### • Choice of partitioning

• Simplest choice in nuclear physics: HO Hamiltonian  $H_0 \equiv \frac{\vec{p}^2}{2m} + \frac{1}{2}m\omega^2 \vec{r}^2$ 

• Common (more refined) choice: Hartree-Fock reference

- $\rightarrow$  Solve variational HF problem and build  $\bar{H}^{[2]}$  from one-body HF Hamiltonian
- $\rightarrow$  Møller-Plesset partitioning  $\breve{H}^{[2]} = 0 \rightarrow H_1 = H^{[4]}$
- Resolvent operator

$$R = -\sum_{ai} \frac{|\Phi_i^a\rangle\langle\Phi_i^a|}{\epsilon_i^a} - \left(\frac{1}{2!}\right)^2 \sum_{abij} \frac{|\Phi_{ij}^{ab}\rangle\langle\Phi_{ij}^{ab}|}{\epsilon_{ij}^{ab}} - \left(\frac{1}{3!}\right)^2 \sum_{abcijk} \frac{|\Phi_{ijk}^{abc}\rangle\langle\Phi_{ijk}^{abc}|}{\epsilon_{ijk}^{abc}} + \dots$$

• Second-order energy correction

$$E^{(2)} = -\sum_{ai} \frac{H^{[2]}_{ai} H^{[2]}_{ia}}{\epsilon^a_i} - \frac{1}{4} \sum_{abij} \frac{H^{[4]}_{abij} H^{[4]}_{ijab}}{\epsilon^{ab}_{ij}}$$

#### **Computational advantages**

- No computation of Hamiltonian matrix
- $\circ$  Non-iterative calculation
- $\circ$  Polynomial scaling  $O(N^4)$







#### **- - 1**

### Non-perturbative methods

#### • Expansion of the exact wave function



Perturbative methods: expansion coefficients computed independently

- ➡ Non-perturbative methods: expansion coefficients computed self-consistently
- Truncated CI: expansion coefficients computed via a diagonalisation
- Examples of non-perturbative approaches
  - Coupled-cluster theory (CC)

 $\Rightarrow$  Exponential ansatz for the wave function  $|\Psi_{CC}\rangle = e^T |\Phi\rangle$ 

• In-medium similarity renormalisation group (IMSRG)

SRG evolution for H normal-ordered w.r.t. to a reference Slater determinant

### Green's function techniques

• The goal is to solve the *A*-body Schrödinger equation

$$H|\Psi_k^A\rangle = E_k^A|\Psi_k^A\rangle$$

• Instead of working with the full *A*-body wave function  $|\Psi_k^A\rangle$ , rewrite the Schrödinger equation in terms of **1-**, **2-**, .... *A*-body objects  $G_1=G$ ,  $G_{2,...}G_A$  (Green's functions)

→ *A*-1 coupled equations

- 1-, 2-, …. *A*-body Green's functions yield **expectation values of 1-, 2-,** …. *A*-body operators
  - → In practice, one usually needs 1- and/or 2-body GFs (~ 1- & 2-body density matrices)

• One-body Green's function obtained by solving **Dyson equation** (derived from Schrödinger eq.)

$$G = G^{(0)} + G^{(0)} \Sigma G$$
  
unperturbed Green's function many-body effects contained in the self-energy  $\Sigma$ 

• Bonus: one-body Green's function contains information about *A***±1** excitation energy spectra

## Benchmarks



Convergence of many-body results

 $\circ$  Different strategies to solve HP=EP

• Same input Hamiltonian (except lattice EFT)

 $\circ$  All methods agree within 5%

 $\odot$  Physics of oxygen isotopes

 $\circ$  Energy trend reproduced by 2N+3N results

 $\circ$  Correct drip line only with 3N forces

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# Degeneracy of open-shell systems

### ● Configuration-space methods expand the solution in ph excitations on top of a reference state

- Standard formulation: reference state has the same symmetries of the exact wave function
- Very efficient to account for dynamical (weak) correlations → closed-shell nuclei
- $\circ$  What about static (strong) correlations?  $\rightarrow$  open-shell nuclei

• Open-shell nuclei are (near-)degenerate with respect to ph excitations

 $\circ$  Gap at the Fermi surface decreases ( $\rightarrow 0$  in the limit)

 $\circ$  ph hierarchy becomes ill-defined



# Breakdown of ph expansion



a b



# Single- vs multi-reference strategy

### • Multi-reference strategy

- Reopens the gap via **IR diagonalisation**
- Ref: linear combination of Slater dets.
- $\circ$  UV correlations via ph excitations



- Reopens the gap via **symmetry breaking**
- Ref: single Slater determinant
- UV correlations via ph excitations



$$|\Psi_0^{J=0A}\rangle = \Omega_0^{UV} |\Theta_{0IR}^{J=0A}\rangle$$

✓ Symmetries are automatically preserved× ph expansion: complicated formalism



$$|\Psi_0^{J=0\,A}\rangle = \Omega_0^{SB} |\Phi_0^{SB}\rangle$$

✓ ph expansion: simpler formalismX Symmetries must be restored

# Symmetry breaking

### • Enlarge the variational space

- Requiring the w.f. ansatz to have the same symmetries of *H* is too restrictive
- $\circ$  In most cases, mean-field solution spontaneously breaks symmetries if allowed

### • Lift the degeneracy

 $\circ$  Trade the ph degeneracy for one in the transformations of the associated symmetry group



#### • Symmetry restoration

• Symmetry breaking is **fictitious** in finite systems

 $\circ$  Symmetry breaking is an intermediate step  $\rightarrow$  symmetries must be **restored** at the end

# Symmetry breaking

• Which symmetries **can** be broken?

Physical symmetry	Group	Casimir	Correlations
Rotational inv.	SU(2)	$\hat{ec{J}}^2$	Deformation
Particle-number inv.	$U(1)_{\rm N} \ge U(1)_{\rm Z}$	$\hat{N},\hat{Z}$	Pairing

• Which symmetries **should** be broken?

- In principle, the more the better (provided calculations are feasible)
- In practice, there are different formal & computational consequences
  - ⇒ Breaking U(1) requires a modification of the (bases of the) formalism, breaking SU(2) doesn't
  - Computationally, breaking U(1) is cheaper than breaking SU(2) (symmetry reduction)

U(1)-breaking methods were first developed

- Ultimately, it depends on the system
  - Singly open-shell nuclei □ Sufficient to break U(1)
  - Doubly open-shell nuclei ⇒ Necessary to break SU(2)

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- Generalises GFs to U(1) breaking
  - Normal + anomalous propagators
  - First symmetry-breaking ab initio method
- Application to isotopic chains (Z=18-24)
  - $\circ$  *Z* = 20 (calcium): magic number
  - $\circ$  Above and below: doubly open shells
- Binding energies
  - $\circ$  Systematic calculations possible
  - Underbinding corrected at ADC(3) level



- Generalises GFs to U(1) breaking
  - Normal + anomalous propagators
  - First symmetry-breaking ab initio method
- Application to isotopic chains (Z=18-24)
  - $\circ$  *Z* = 20 (calcium): magic number
  - $\circ$  Above and below: doubly open shells
- Two-neutron separation energies
  - $S_{2n}(N, Z) \equiv |E(N, Z)| |E(N 2, Z)|$
  - Error cancellation in relative quantities
  - $\circ$  Drops correspond to magic gaps



 $\angle \mathbf{I}$ 

LL

Ζ

20



# Deformation across the nuclear chart







➡ Majority of nuclei display a non-spherical mean field (when allowed to do so)

# Breaking SU(2)

### **⊙** Ongoing efforts to extend state-of-the-art techniques to SU(2)-breaking

 $\circ$  Coupled cluster: only SU(2)  $\rightarrow$  deformed CC

 $\circ$  MBPT: on top of U(1) → deformed BMBPT



Example from deformed (= unrestricted) HFB

[Frosini et al. in preparation]



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# Benchmarking many-body approaches

- Different many-body calculations yield very consistent results
  - $\circ$  All within few %
  - MBPT not shown but wouldn't be too far
  - Discrepancies w.r.t. data to be **attributed to the input Hamiltonian**



### Proliferation of nuclear Hamiltonians

### • State of the art (until very recently)

- No routine / consistent account of systematic uncertainties coming from input Hamiltonian
- Precision ultimately depends on the chosen input
- Proliferation of Hamiltonians





## Ab initio prediction of the drip lines

• Systematic survey of light and medium-mass nuclei (method: valence-space IMSRG)

• Good description (+ prediction) of proton and neutron drip lines

• **Rms deviation** on total binding energies = **3.3 MeV** (cf. **0.7 MeV** in energy density functionals)



### Mid-mass isotopic chains

• SCGF calculations along mid-mass chains

ADC(2) level: few % on differential quantities
ADC(3) level: 2.5 % on B.E. of closed-shells

• Radii within few % but some features are missing



**Gorkov ADC(2)** 

#### [Somà *et al.* 2021]

	Ar	Ca	Ti	Cr
$\boxed{NN+3N(\ln l)}$				
E [MeV]	14.1	10.3	14.2	19.2
E/A [MeV]	0.34	0.21	0.29	0.35
$S_{2n}$ [MeV]	2.90	1.56	2.05	2.22
NNLO <sub>sat</sub>				
$\langle r_{\rm ch}^2 \rangle^{1/2}$ [fm]	0.008	0.022	0.019	0.010
$\delta \langle r_{\rm ch}^2 \rangle^{1/2}$ [fm]	0.008	0.024	0.023	0.013



### • Correct reproduction of radii crucial when computing cross sections

• Particularly important for nuclear physics applications to other domains



*Elastic electron scattering* @SCRIT

Inelastic electron scattering @JLAB



[Arthuis *et al.* 2020]

<sup>[</sup>Barbieri et al. 2019]

### Systematic fit of low-energy constants

#### ● Example of fit of low-energy constants (LECs) in the three-body sector

### $\circ$ Two LECs (c\_D & c\_E) at N²LO and N³LO

 $\circ$  **c**<sub>D</sub> as a parameter in the calculation of E(<sup>4</sup>He) & r(<sup>4</sup>He), **c**<sub>E</sub> determined to reproduce E(<sup>3</sup>H)

[Hüther et al. 2020]

1.55 N2LO N3LO 1.5 2  $C_{\rm D}$ 1.45 R<sub>p,rms</sub> [fm] 1.4 1.35  $\Lambda = 500 \text{ MeV}$ exp. exp. 1.3 -31 -30 -29 -28 -30 -29 -28 -27 E [MeV] E [MeV]

### Systematic fit of low-energy constants

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[Hüther et al. 2020]





 $\circ$  Large sensitivity to 3N parameters

 $\circ$  "Tension" between optimal values in few-body systems, mid-mass nuclei & nuclear matter

 $\circ$  Radii not much affected by changes in  $c_D \& c_E$  (regulator more important)

#### -230 - NN+3N(lnl) <sup>36</sup>Ca $-\Delta$ - NNLO<sub>sat</sub> -240 E [MeV] -250 -260 -270 -280 = (a)3.5 $\langle r^2 \rangle^{1/2} \, [fm]$ 3.4 3.3 3.2 3.1 = (c)20 12 14 16 18 22 $\hbar\Omega$ [MeV]

# ħΩ [MeV]

(d)

24 12 14

(b)

 $e_{max} = 9$  $e_{max} = 11$  $e_{max} = 13$ 

-460

-480

-500

-520

-540

-560

-580

4.1

4.0

3.9

3.8

3.7

3.6

3.5

20

18

16

22 24

<sup>68</sup>Ni

[Soma et al. 2020]

#### • Uncertainties from the expansion method

#### $\circ$ Basis truncation





[Soma et al. 2020]

### $\ensuremath{{ \bullet}}$ Uncertainties from the expansion method

- $\circ$  Basis truncation
- Many-body truncation
- Symmetry breaking (if any)
- $\circ$  Neglected induced operators (if any)

0...

	NN+3N	<b>NNLO</b> <sub>sat</sub>		
[Soma <i>et al.</i> 2021]	E	r <sub>ch</sub>	r <sub>ch</sub>	
Model space $(e_{\max})$	0.5%	< 0.1%	0.5%	
Model space ( $e_{3\max}$ )	0.2%	0.2%	0.3%	
ADC truncation	2%	0.5%	< 0.1%	
U(1) breaking	0.2%	< 0.1%	< 0.1%	
Neglected induced op.	2%	1%	_	
Total	2.9%	1.1%	0.6%	

• Uncertainties from the expansion method		NN+3N	$NN+3N(\ln l)$	
$\circ$ Basis truncation	[Soma <i>et al</i> . 2021]	E	r <sub>ch</sub>	r <sub>ch</sub>
<ul> <li>Many-body truncation</li> </ul>	Model space $(e_{\max})$	0.5%	< 0.1%	0.5%
Filling body diamentation	Model space ( $e_{3\max}$ )	0.2% 0.2%		0.3%
<ul> <li>Symmetry breaking (if any)</li> </ul>	ADC truncation	2%	0.5%	< 0.1%
· Naclastad in ducad anomators (if any)	U(1) breaking	0.2%	< 0.1%	< 0.1%
• Neglected induced operators (if any)	Neglected induced op.	2%	1%	_
Ο	Total	2.9%	1.1%	0.6%

#### • Uncertainties from the Hamiltonian

- $\circ$  Ideally, at each order from cutoff variation
- $\circ$  If not possible, use some estimate

[Epelbaum et al. 2015]

$$\Delta X^{\mathrm{N^{3}LO}}(p) = \max \left( Q^{5} \times \left| X^{\mathrm{LO}}(p) \right|, \quad Q^{3} \times \left| X^{\mathrm{LO}}(p) - X^{\mathrm{NLO}}(p) \right|, \quad Q^{2} \times \left| X^{\mathrm{NLO}}(p) - X^{\mathrm{N^{2}LO}}(p) \right|, \quad Q \times \left| X^{\mathrm{N^{2}LO}}(p) - X^{\mathrm{N^{3}LO}}(p) \right| \right),$$

# Towards systematic calculations (with uncertainties)

#### • Evaluation of uncertainties from the Hamiltonian

[Hüther *et al.* 2020]



# Towards systematic calculations (with uncertainties)

#### • Evaluation of uncertainties from the Hamiltonian + the many-body method [Hüther *et al.* 2020] -6 -7 -7 -8-





• Only exact methods available



• Only exact methods available



○ Development of symmetry-conserving methods → doubly closed-shell nuclei
 ○ Development of U(1) symmetry-breaking methods → singly open-shell nuclei



 $\circ$  Development of SU(2) symmetry-breaking methods  $\rightarrow$  **doubly open-shell** nuclei

#### • Towards heavy nuclei

- $\circ$  Size of 3N matrix elements becomes prohibitive
- $\rightarrow$  Techniques from applied maths  $\rightarrow$  Tensor factorisation of the many-body problem

### Oubly open-shell nuclei

- Symmetry breaking? Single- or multi-reference? Scaling?
- → Strategy has to be adapted to the obejective

#### Uncertainty quantification

- Thorough quantification to establish link to QCD & predictive power
- → Development of efficient many-body *emulators* → Towards statistical analyses of LEC fits
- → Issue of renormalisability?

#### • How far can this approach be pushed?

- Not obvious that "chiral EFT in the A-body sector" works all the way up to superheavy nuclei
- → Different types of EFT explored

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