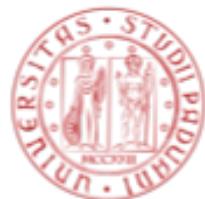


May 28th, 2025

Chern-Simons interaction with a non-canonical axion -paving the way for polarized tensor modes from inflation-

@ Planck 2025



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Jun'ya Kume (University of Padova)

Collaborators: Marco Peloso, Nicola Bartolo

Based on: arXiv:2501.02890 [astro-ph.CO]

Outline

- Inflationary phenomenology with Chern-Simons term
- Production of U(1) gauge field and sourced perturbations
- Non-canonical case: suppression of scalar mode
- Summary & Discussion

- **inflation**: accelerating expansion before Hot Big-Bang (Starobinsky, Guth, Sato, ...)

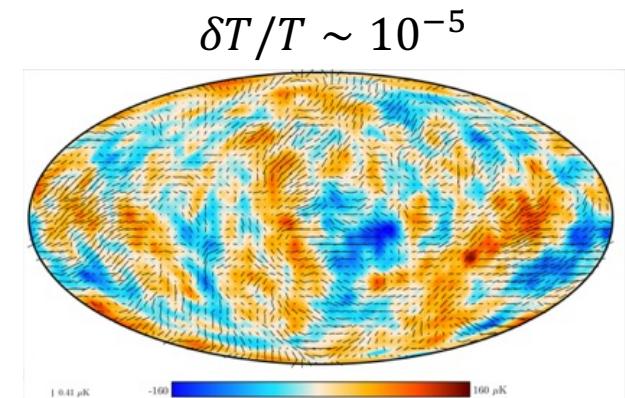
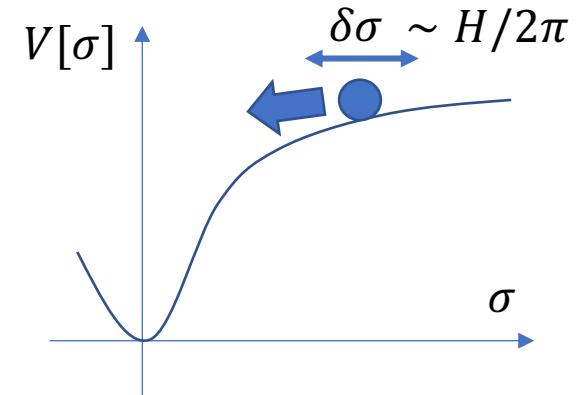
solves classical problems in Big-Bang cosmology

→ **origin of CMB fluctuation!**

ex.) a scalar field (inflaton) $\sigma(t, \vec{x}) = \bar{\sigma}(t) + \delta\sigma(t, \vec{x})$

→ curvature perturbation $\zeta \cong H\delta\sigma/\dot{\bar{\sigma}}$

power spectrum & spectral index: $P_\zeta(k)$, $n_s - 1 = d\ln P_\zeta / d\ln k$



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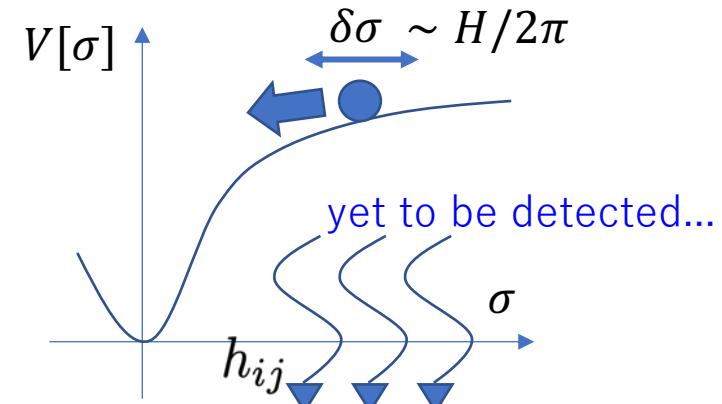
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Tensor perturbation as well! → B-mode polarization

tensor-to-scalar ratio: $r \sim P_h/P_\zeta \lesssim 0.03$

vacuum fluctuation:

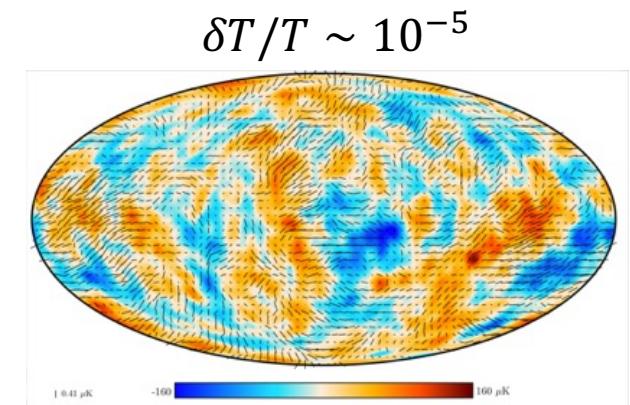
$$P_h = (2/\pi^2)(H/M_{Pl})^2$$



$$H \lesssim 10^{13} \text{ GeV} \dots !?$$

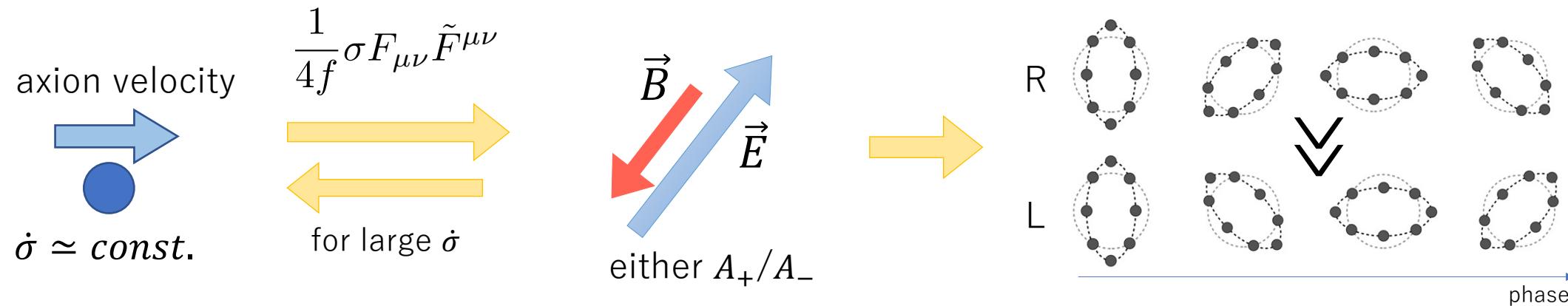
not true if sourced component exists

“What can we learn from the tensor signal?”



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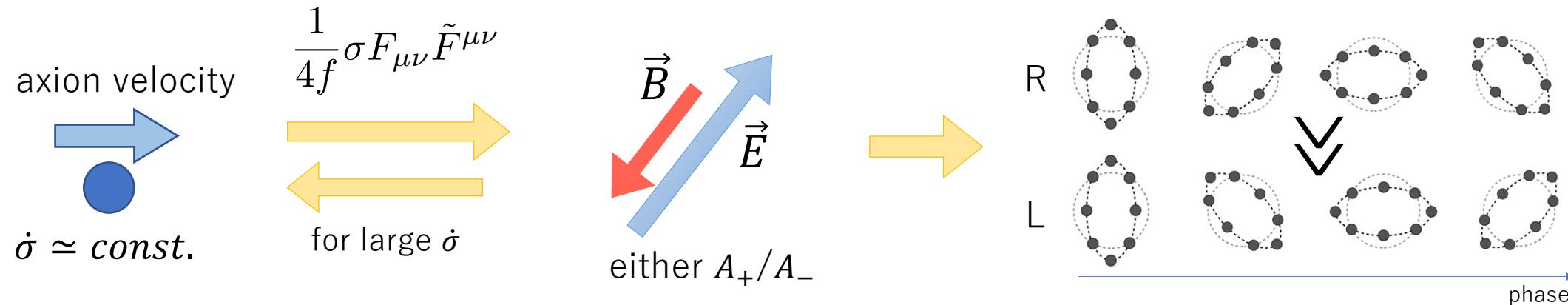
- **Polarized tensor modes** from U(1) CS term (Anber & Sorbo 2006, Sorbo 2011, ...)



※ fermion production: Domcke & Mukaida 2018, Gorbar+2021, Fujita, JK+ 2022, ... also talks by Oleksandr & Richard

Circularly polarized tensor pert. → a smoking gun for identifying the inflationary model!?

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Circularly polarized tensor pert. → a smoking gun for identifying the inflationary model!?

- also **sourcing scalar mode**: total tensor-to-scalar not increased...
- $A + A \rightarrow \delta\sigma$: **highly non-Gaussian** → $\dot{\sigma}$ cannot be large... (Barnaby & Peloso 2010, ...)

💡 axion as spectator: CS term → axion pert. → curvature pert. (indirect sourcing)

- CMB non-Gaussianity bound → axion can roll only a few e-folds

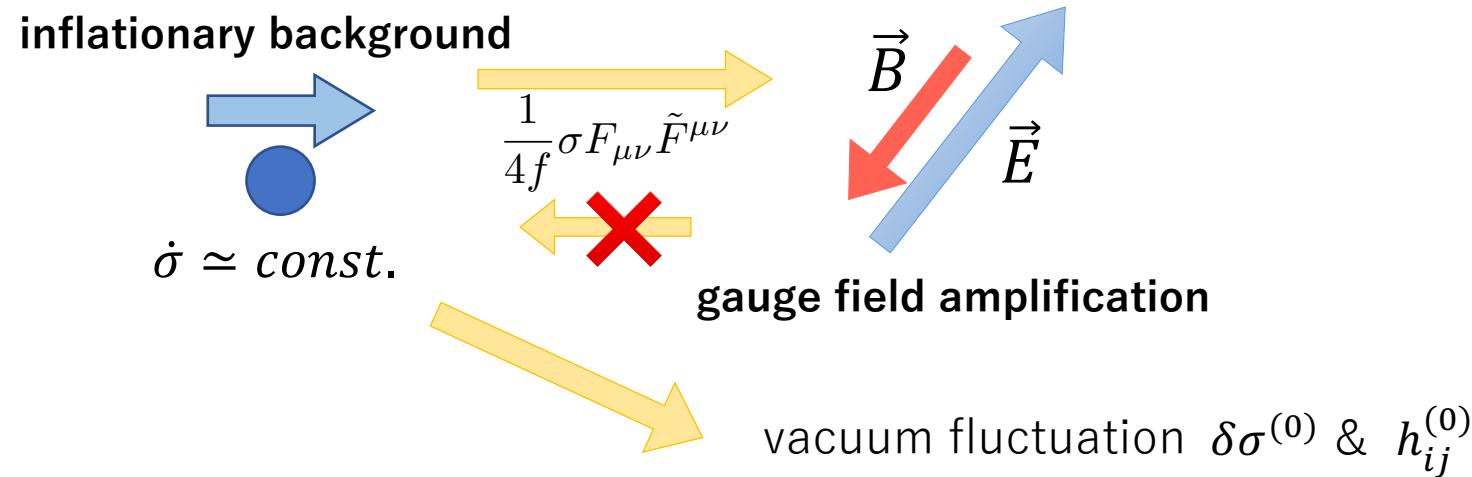
(Ferreira & Sloth 2014, Namba+ 2015)

- Any other possibilities? 🤔

Outline

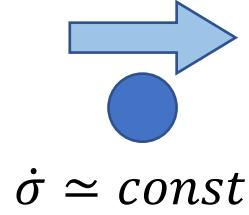
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- The system & working assumption

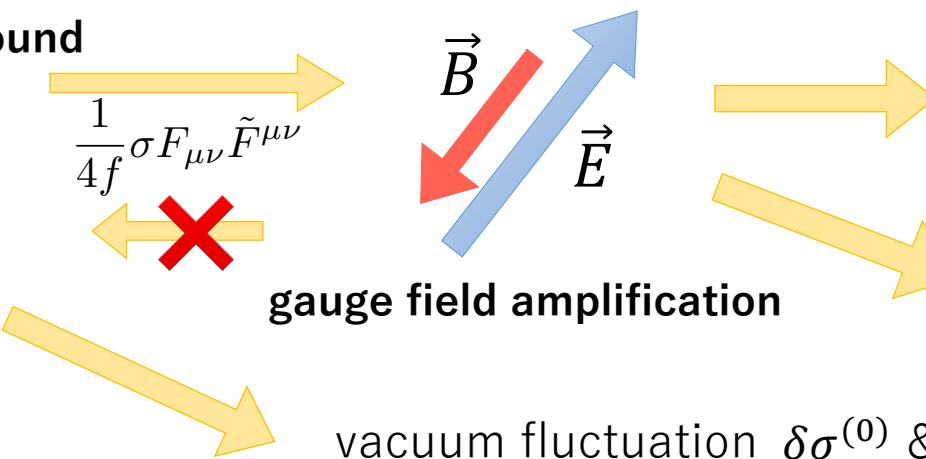


- The system & working assumption

inflationary background



$$\frac{1}{4f}\sigma F_{\mu\nu}\tilde{F}^{\mu\nu}$$



vacuum fluctuation $\delta\sigma^{(0)}$ & $h_{ij}^{(0)}$

sourced components:

inverse decay
& gravitational coupling $\delta\sigma^{(1)}$

polarized tensor pert. $h_{ij}^{(1)}$

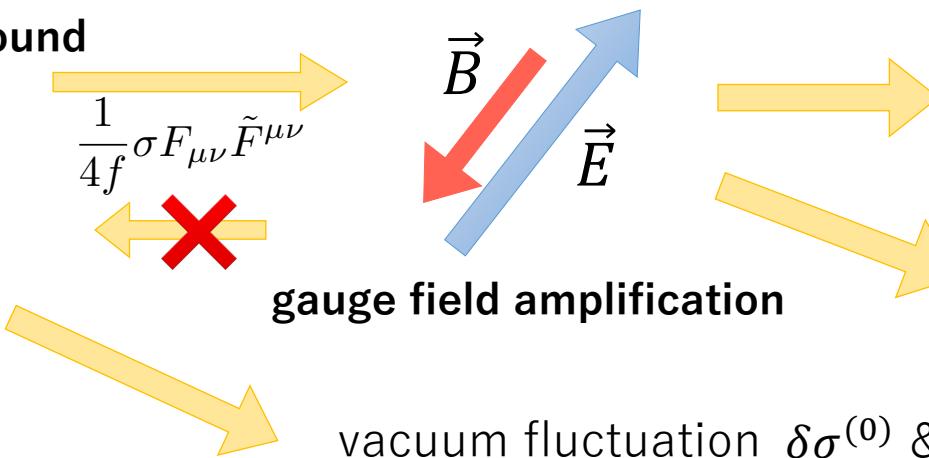
uncorrelated

- The system & working assumption

inflationary background

$$\dot{\sigma} \simeq \text{const.}$$

$$\frac{1}{4f}\sigma F_{\mu\nu}\tilde{F}^{\mu\nu}$$



sourced components:

inverse decay
& gravitational coupling $\delta\sigma^{(1)}$

polarized tensor pert. $h_{ij}^{(1)}$

uncorrelated

observables of our interest:

$$P_\zeta(k) = P_\zeta^{(0)}(k) + P_\zeta^{(1)}(k) \quad \text{with vacuum pert.: } 8\pi^2\epsilon_I P_\zeta^{(0)} = H^2/M_{Pl}^2$$

$$f_{NL} = \frac{10}{9(2\pi)^{5/2}} \frac{k^6}{P_\zeta^2} F|_{|\vec{k}_1|=|\vec{k}_2|=|\vec{k}_3|=k} \quad \langle \hat{\zeta}(0^-, \vec{k}_1) \hat{\zeta}(0^-, \vec{k}_2) \hat{\zeta}(0^-, \vec{k}_3) \rangle \equiv \delta^{(3)}(\vec{k}_1 + \vec{k}_2 + \vec{k}_3) F(\vec{k}_1, \vec{k}_2, \vec{k}_3)$$

$$r = (P_+ + P_-)/P_\zeta$$

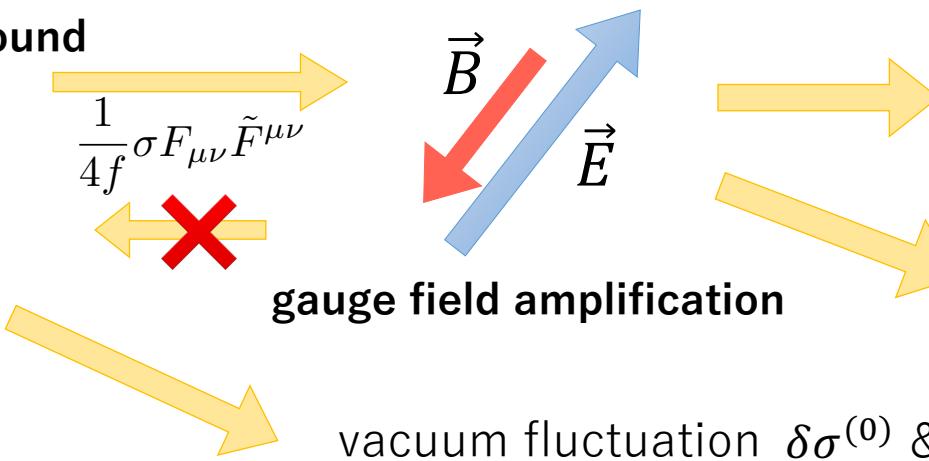
P_\pm : tensor power spectrum for +/- polarization

- The system & working assumption

inflationary background

$$\dot{\sigma} \simeq \text{const.}$$

$$\frac{1}{4f}\sigma F_{\mu\nu}\tilde{F}^{\mu\nu}$$



gauge field amplification

sourced components:

inverse decay
& gravitational coupling $\delta\sigma^{(1)}$

polarized tensor pert. $h_{ij}^{(1)}$

observables of our interest:

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$$r = (P_+ + P_-)/P_\zeta$$

P_\pm : tensor power spectrum for +/- polarization

$\hat{\zeta}\sigma$ can be a spectator $\rightarrow I = \{\phi, \sigma\}, \quad \varphi_I = \{\phi, \sigma\}, \quad \epsilon_I, \eta_I \ll 1.$

- Amplification of U(1) gauge field

$$\mathcal{L} \supset \frac{1}{4f} \sigma F_{\mu\nu} \tilde{F}^{\mu\nu} \rightarrow \text{circular pol.: } A_+/A_-$$

Quantization: $\hat{A}_\lambda(\tau, \mathbf{k}) = \hat{a}_{\mathbf{k}}^{(\lambda)} \mathcal{A}_\lambda(\tau, k) + \hat{a}_{-\mathbf{k}}^{(\lambda)\dagger} \mathcal{A}_\lambda^*(\tau, k)$

$$\Rightarrow \left[\partial_\tau^2 + k^2 \pm 2k \frac{\xi}{\tau} \right] \mathcal{A}_\pm(\tau, k) = 0 \quad \text{with} \quad \xi \equiv \frac{\dot{\sigma}}{2fH}$$

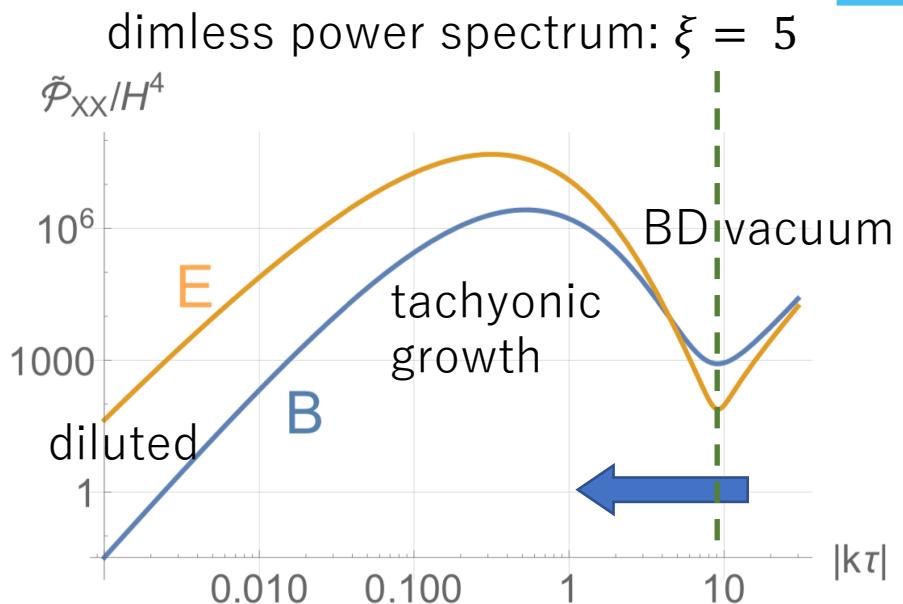
Only + mode experiences **exponential growth!** ($\xi \lesssim 5$ for negligible backreaction)

Imposing Bunch-Davies condition:

$$\mathcal{A}_+(\tau, k) = \frac{1}{\sqrt{2k}} e^{\pi\xi/2} W_{-i\xi, 1/2}(2ik\tau) \cong \frac{1}{\sqrt{2k}} \left(\frac{k}{2\xi aH} \right)^{1/4} e^{\pi\xi - 2\sqrt{2\xi k/(aH)}}$$

← around the peak

Electromagnetic variables: $\hat{E}_i \equiv -\frac{1}{a^2} \hat{A}'_i$, $\hat{B}_i \equiv \frac{1}{a^2} \epsilon_{ijk} \partial_j \hat{A}_k$



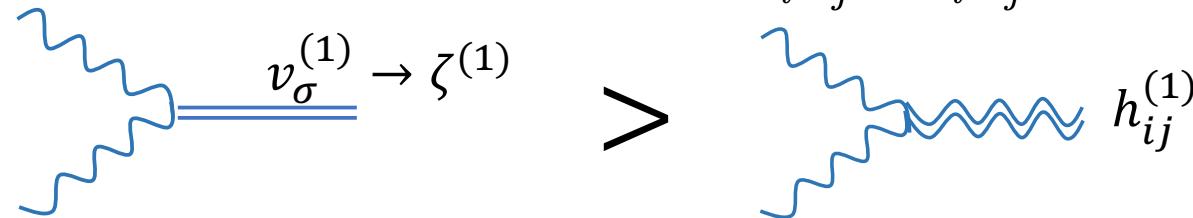
- Sourced perturbation

$$\Delta S_{sAA} = \frac{1}{2} \int d\tau d^3k \left[-\frac{\dot{\varphi}_I}{2M_p^2 H} \frac{1}{a} \left(\underline{a^4 Q_1} - \underline{\frac{ik_k}{k^2} (a^4 Q_{2,k})'} \right) v_I^\dagger + a^3 \frac{Q_3}{f} v_\sigma^\dagger + \text{h.c.} \right]$$

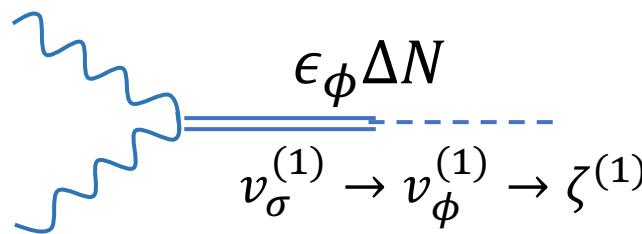
v_I : canonical variables
 $I = \{\phi, \sigma\}$

with $Q_1(x) \equiv \frac{1}{2} (\hat{E}_i \hat{E}_i + \hat{B}_i \hat{B}_i)$ $Q_{2,k}(x) \equiv \epsilon_{ijk} \hat{E}_i \hat{B}_j \ll Q_3(x) \equiv \hat{E}_i \hat{B}_i$

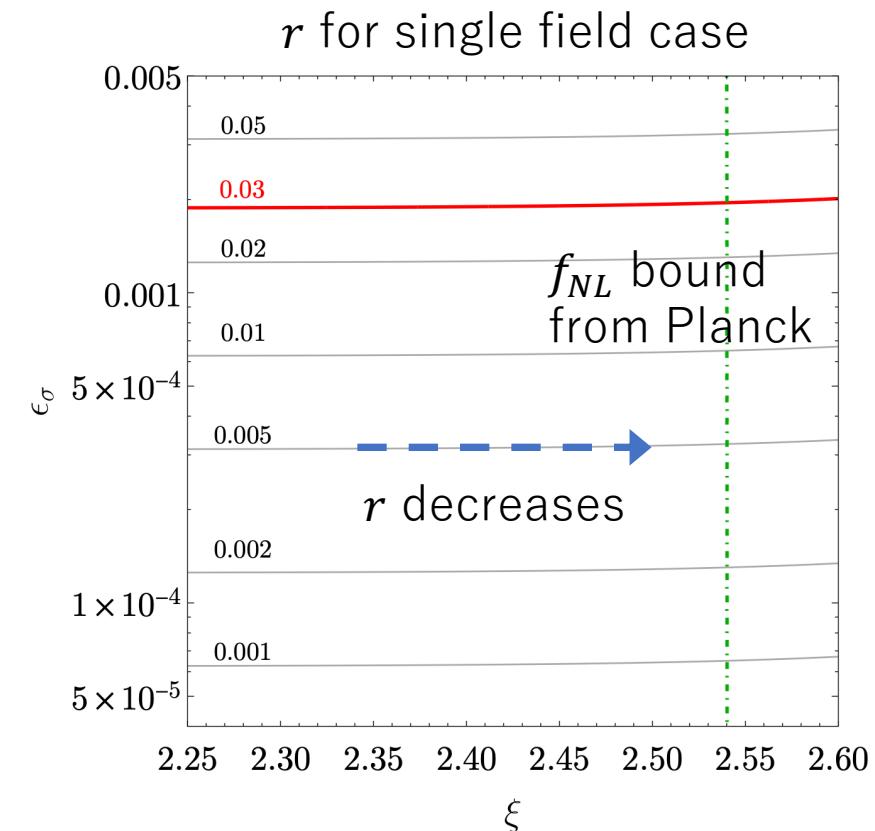
- single field case: $\{I, \varphi_I\} \rightarrow \sigma$



- two-field case:



σ decays $\Delta N \lesssim 5$ e-folds
after the CMB mode crossing
→ visible sourced tensor
(Ferreira & Sloth 2014, Namba+ 2015)



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- Non-canonical scalar and perturbation

$$\mathcal{L} = K_\phi(X) - V(\hat{\phi}) \quad X \equiv -\frac{1}{2}g^{\mu\nu}\partial_\mu\hat{\phi}\partial_\nu\hat{\phi}$$

Scalar field (b.g. & pert.) acquires large inertia:

$$\left(K_{\phi,1} + \dot{\phi}^2 K_{\phi,2}\right) \ddot{\phi} + 3K_{\phi,1} \frac{\dot{a}}{a} \dot{\phi} + \frac{\partial V}{\partial \phi} = 0$$

- examples, but not for a pseudo-scalar...
- k-inflation:
Armendariz-Picon 1999, Garriga & Mukhanov 1999
DBI-model: Silverstein & Tong 2003, Alishahiha+ 2004

$$K_{\phi,n} \equiv (\partial/\partial X)^n K_\phi$$

sound speed: $c_{s,I}^2 \equiv \frac{K_{I,1}}{K_{I,1} + \dot{\varphi}_I^2 K_{I,2}}$



parametrical changes $\hat{\zeta} \simeq \frac{c_{s,I} H \tau}{\sqrt{2\epsilon_I} M_{\text{Pl}}} v_I$

- power spectrum

$$\mathcal{P}_\zeta^{(0)} = \frac{H^2}{8\pi^2 c_{s,I} \epsilon_I M_{\text{Pl}}^2} \rightarrow r = 16\epsilon_I c_{s,I}$$

- non-Gaussianity

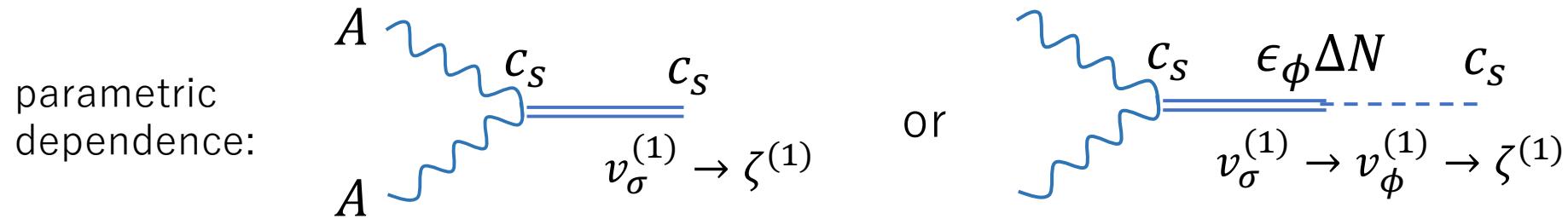
$$f_{\text{NL}} = \frac{\langle \zeta \zeta \zeta \rangle}{\langle \zeta \zeta \rangle^2} \propto \frac{1}{c_s^2} \rightarrow \text{CMB bound on single field EFT: } c_s < 0.02$$

※ standard scenario recovered by taking $c_s \rightarrow 1$

※ smaller $c_s \leftrightarrow$ larger inertia → persistence of the vacuum mode, enhancing the non-linearity(?)

- Suppressing the interaction with gauge field

$$\hat{\zeta} \simeq \frac{c_{s,I} H \tau}{\sqrt{2\epsilon_I} M_{\text{Pl}}} v_I \quad \& \quad \frac{1}{2} \int d\tau d^3k \left[-\frac{\sqrt{K_{I,1}} \dot{\phi}_I}{2M_p^2 H} \frac{1}{a c_{s,I}} \left(c_{s,I}^2 a^4 Q_1 - \frac{i k_k}{k^2} (a^4 Q_{2,k})' \right) v_I^\dagger + \frac{c_{s,\sigma}}{\sqrt{K_{\sigma,1}}} a^3 \frac{Q_3}{f} v_\sigma^\dagger + \text{h.c.} \right]$$

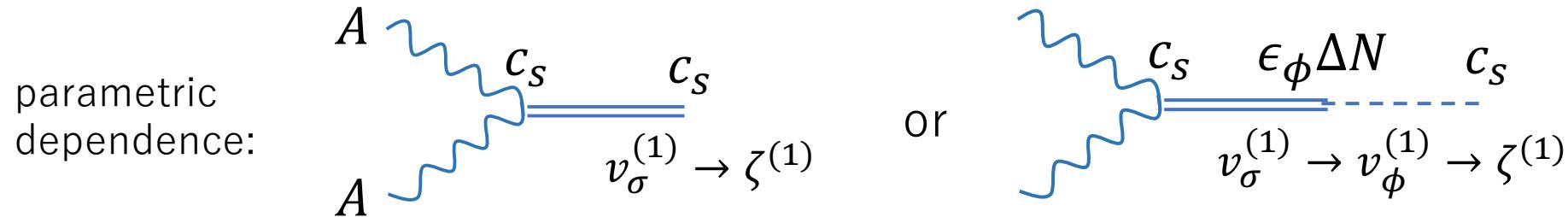


single field case: $\hat{\zeta}^{(1)}(0^-, \vec{k}) \simeq \frac{H^2}{M_p^2} \int_{-\infty}^{0^-} d\tau' k \left(-\frac{k^3 \tau'^3}{3} \right) \left[\frac{1}{2} J_{1,2}(\tau', \vec{k}) - \frac{c_{s,\sigma}^2 \xi}{\epsilon_\sigma} J_3(\tau', \vec{k}) \right]$

※ Q_2 not suppressed by $c_s \rightarrow$ must be included for smaller c_s

- Suppressing the interaction with gauge field

$$\hat{\zeta} \simeq \frac{c_{s,I} H \tau}{\sqrt{2\epsilon_I} M_{\text{Pl}}} v_I \quad \& \quad \frac{1}{2} \int d\tau d^3k \left[-\frac{\sqrt{K_{I,1}} \dot{\phi}_I}{2M_p^2 H} \frac{1}{a c_{s,I}} \left(c_{s,I}^2 a^4 Q_1 - \frac{i k_k}{k^2} (a^4 Q_{2,k})' \right) v_I^\dagger + \frac{c_{s,\sigma}}{\sqrt{K_{\sigma,1}}} a^3 \frac{Q_3}{f} v_\sigma^\dagger + \text{h.c.} \right]$$



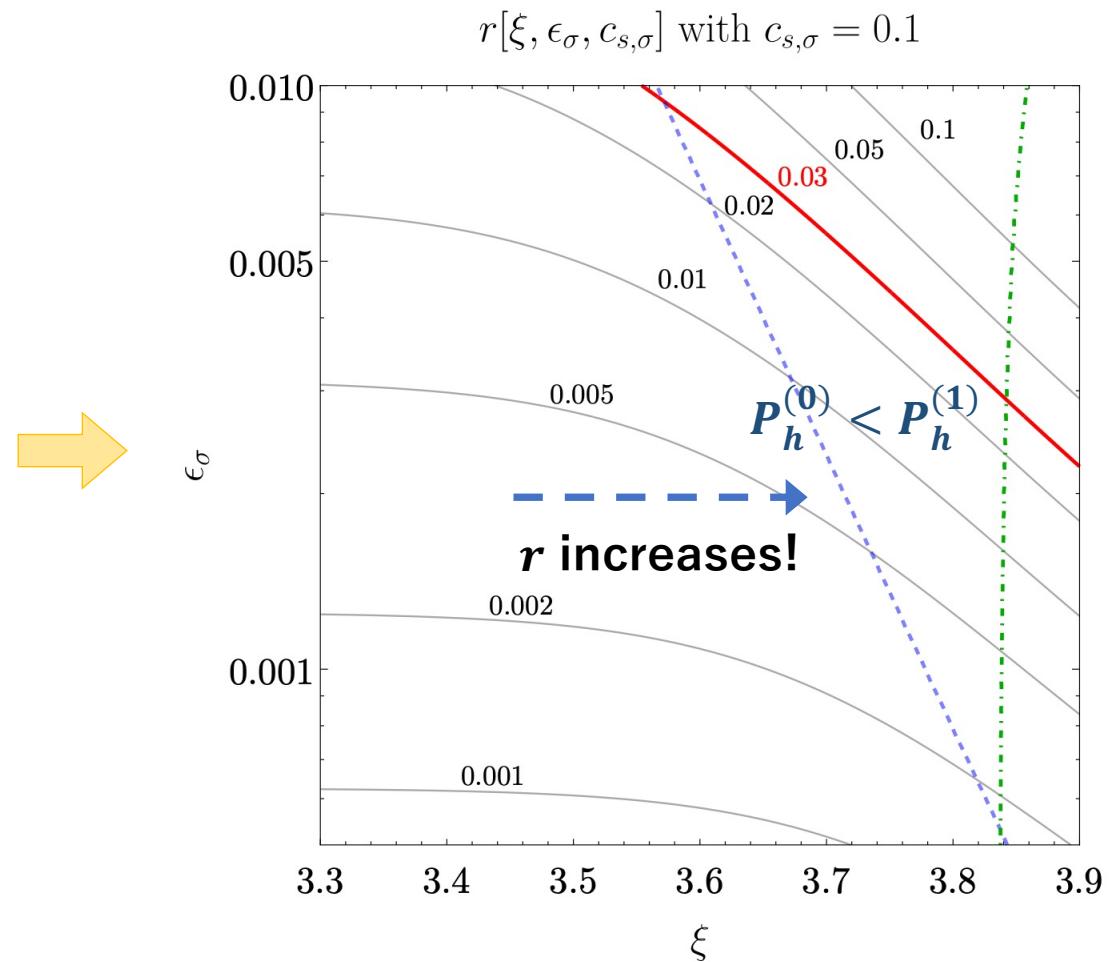
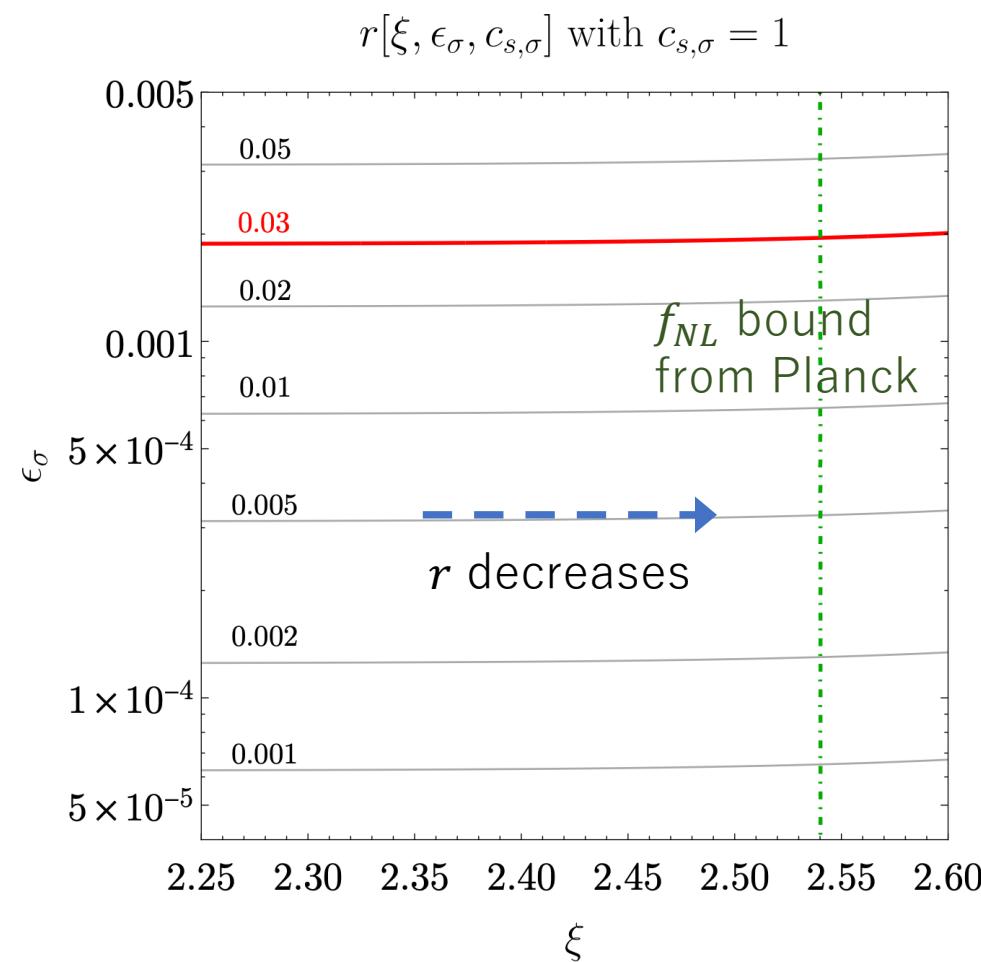
single field case: $\hat{\zeta}^{(1)}(0^-, \vec{k}) \simeq \frac{H^2}{M_p^2} \int_{-\infty}^{0^-} d\tau' k \left(-\frac{k^3 \tau'^3}{3} \right) \left[\frac{1}{2} J_{1,2}(\tau', \vec{k}) - \frac{c_{s,\sigma}^2 \xi}{\epsilon_\sigma} J_3(\tau', \vec{k}) \right]$

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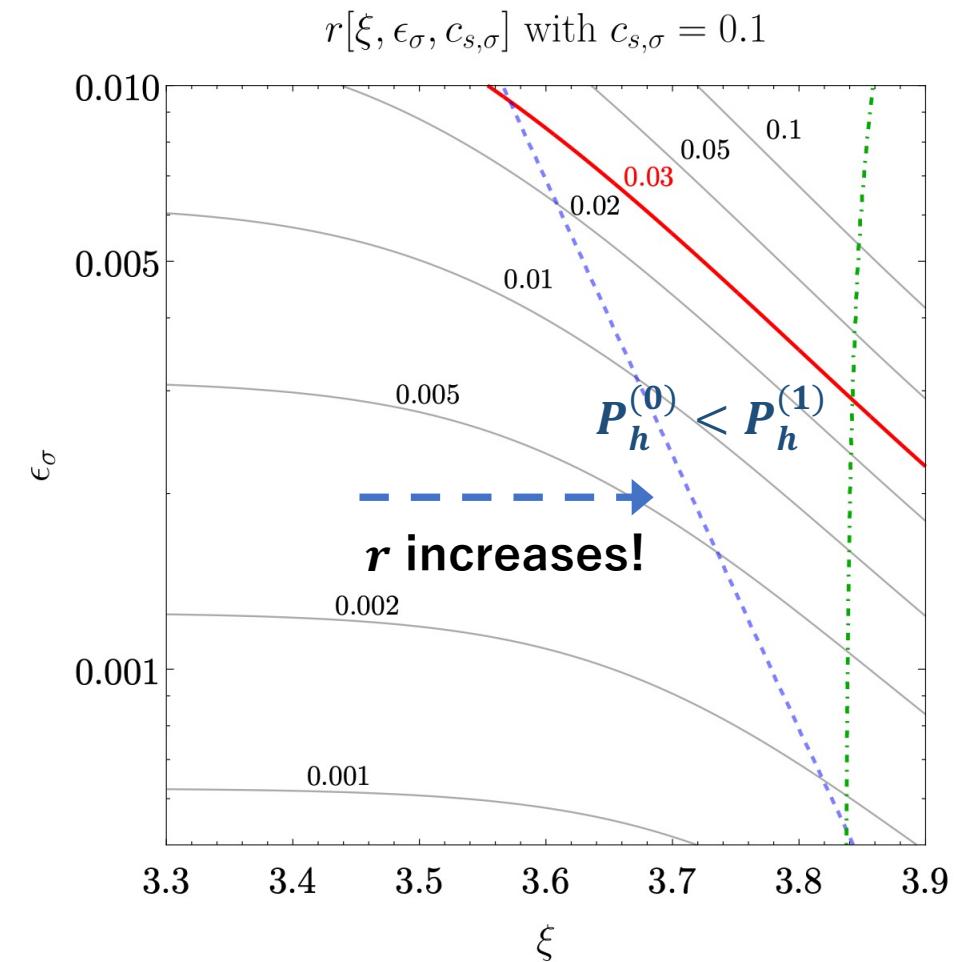
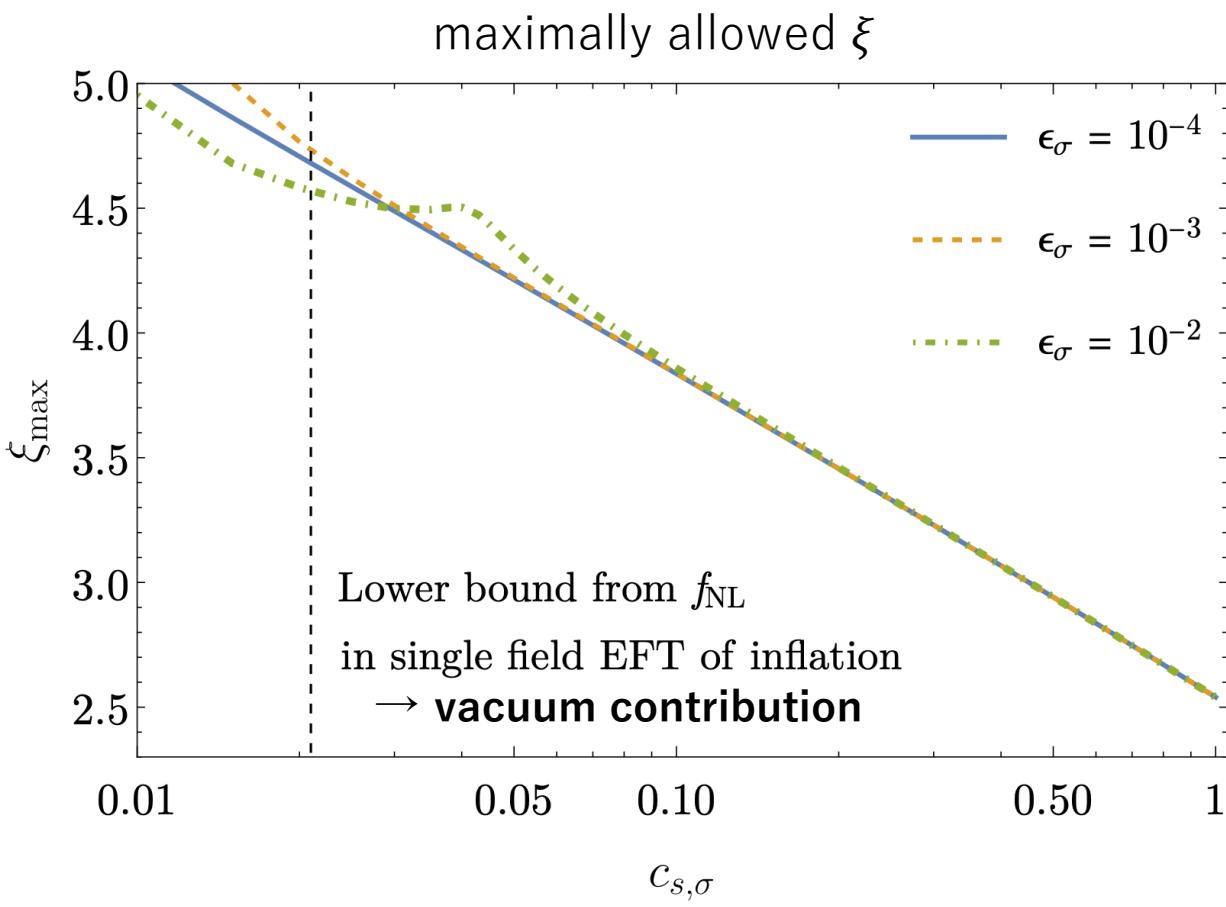
From the standard case, we get the suppression as $P_\zeta^{(1)} \propto c_s^4$ and $f_{NL} \propto c_s^6$!!

※※ Tensor sector not modified!

- Single field case (σ : inflaton)

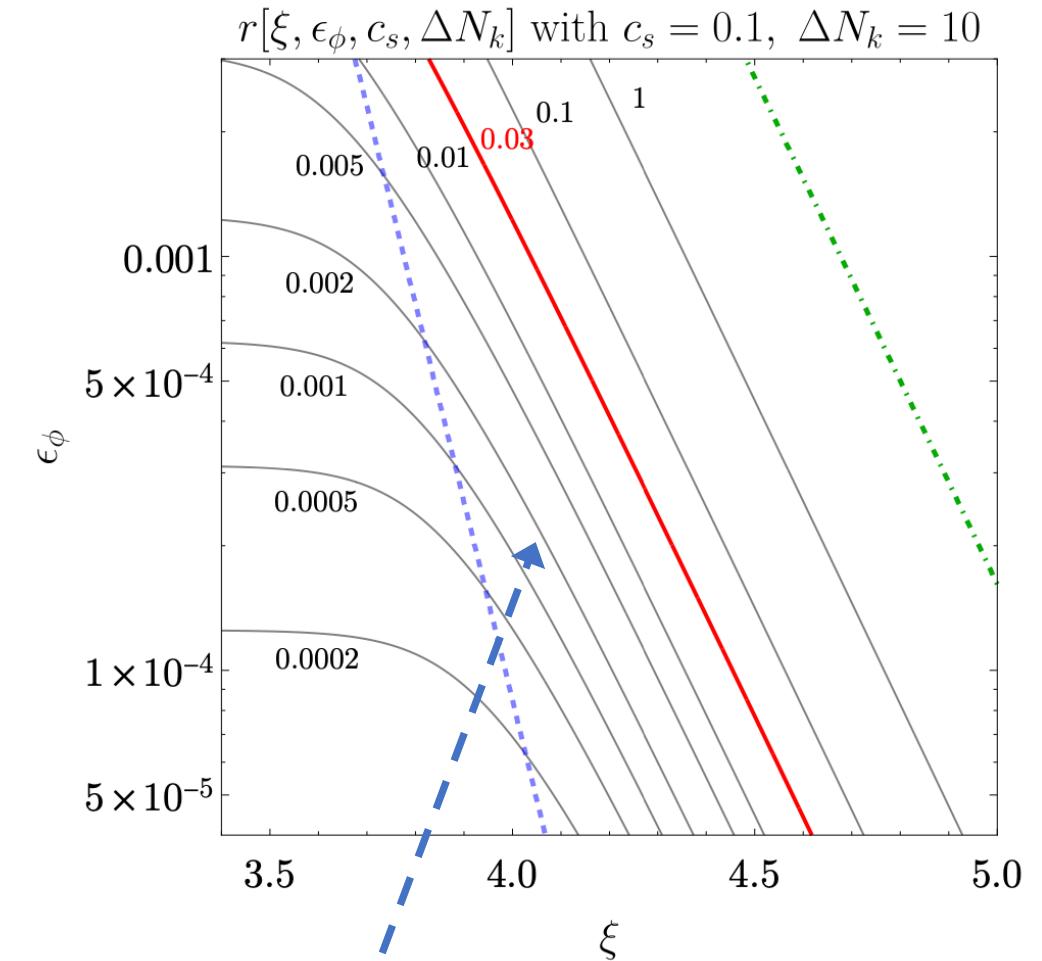
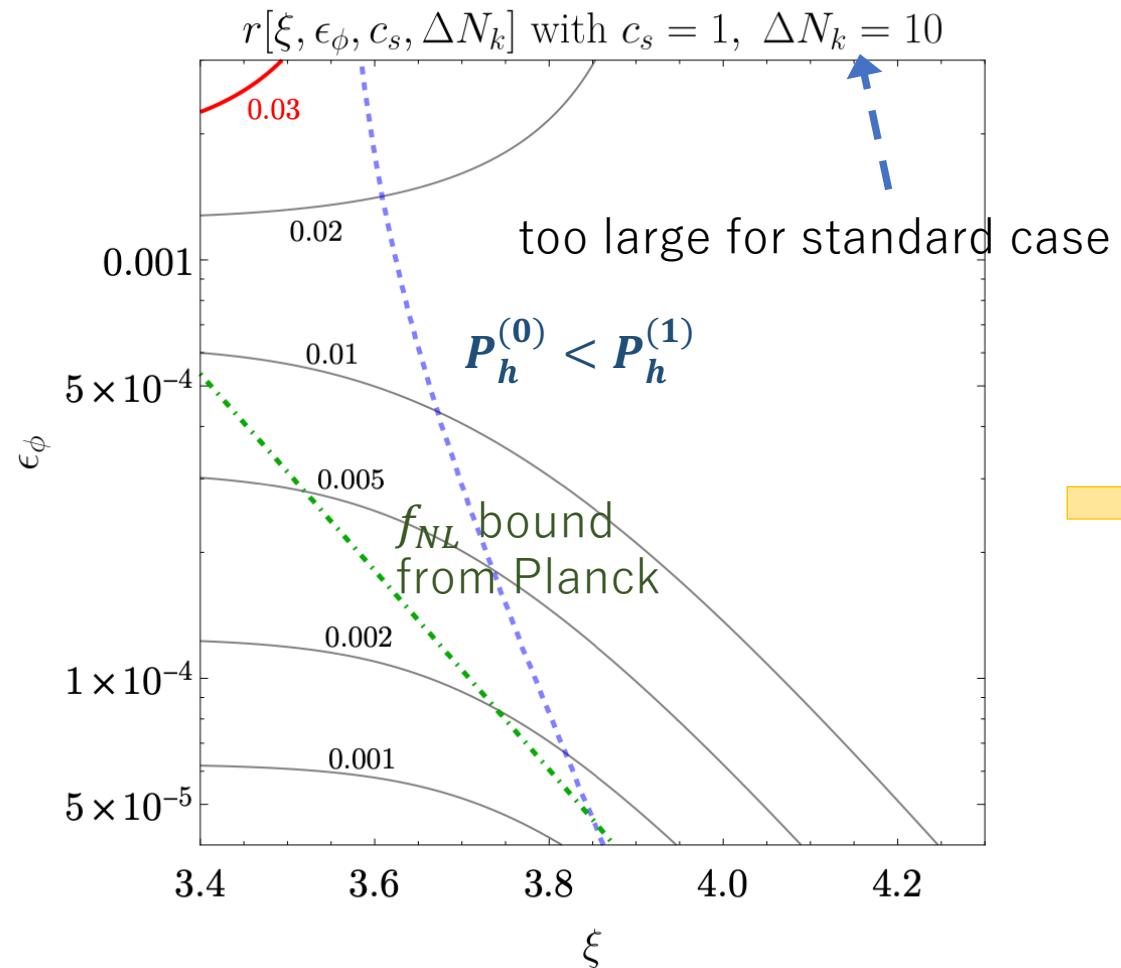


- Single field case (σ : inflaton)



exponential dependence on ξ → sourcing effect can be significantly large!

- two-field case (ϕ : inflaton, σ : spectator)



The new window exists even for $\Delta N \sim 50 - 60 !!$

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Summary & Discussion

✓ axion inflation + U(1) CS term → polarized tensor perturbation?

i) strong f_{NL} bound and ii) decreasing total r ...

✓ Extension: non-canonical kinetic term for axion (and inflaton)

→ Inverse decay contribution is suppressed by c_s !

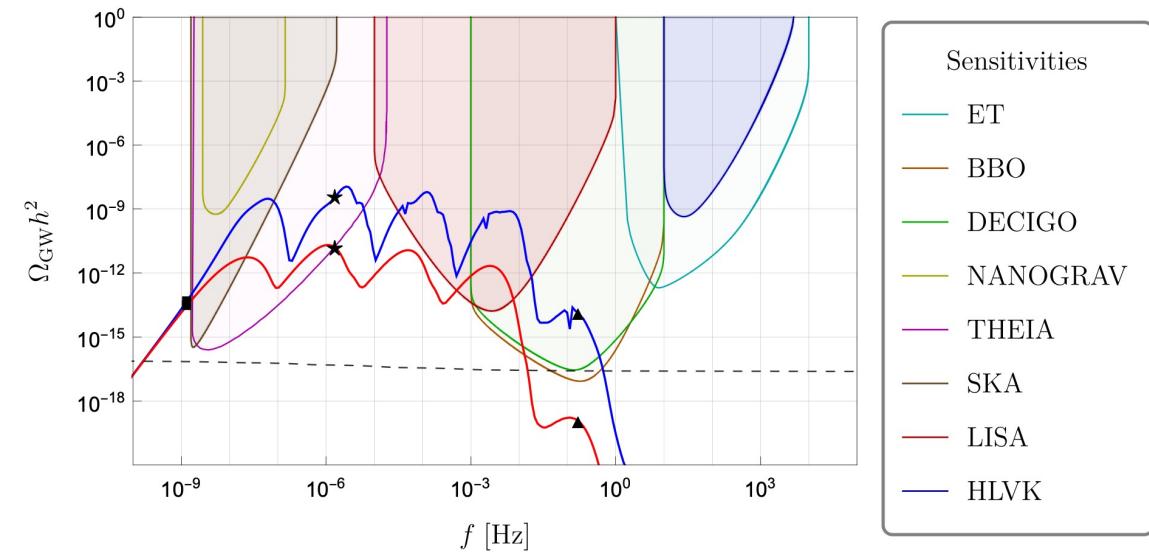
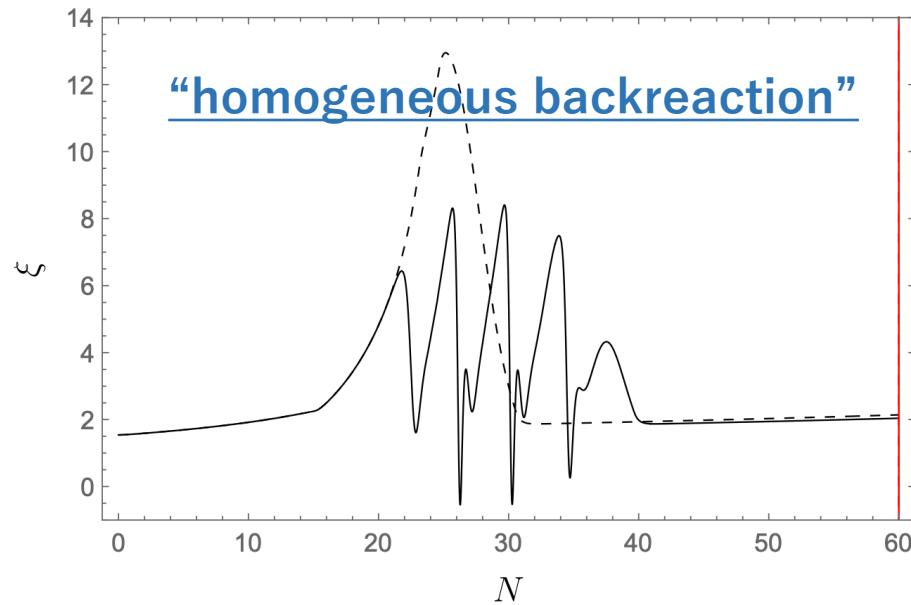
→ Sourced tensor becomes dominant while satisfying f_{NL} bound.

✓ Physics behind: large inertia of scalar

→ generally applicable to other scenarios with “unwanted” scalar

For SU(2): Watanabe & Komatsu 2020, Dimastrogiovanni+ 2023, Murata & Kobayashi 2024...

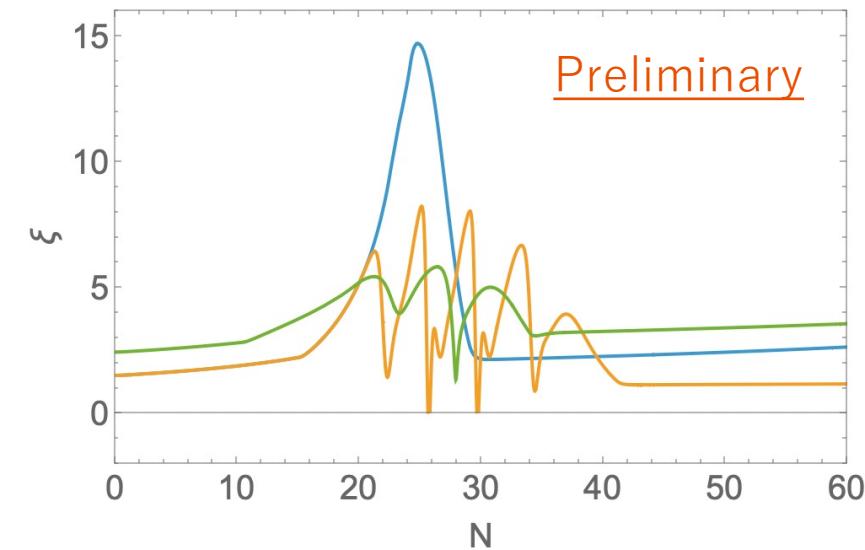
- computation including backreaction (García-Bellido+ 2023, ...) → talk by L. Sorbo



$$\left(K_{\phi,1} + \dot{\phi}^2 K_{\phi,2} \right) \ddot{\phi} + 3K_{\phi,1} \frac{\dot{a}}{a} \dot{\phi} + \frac{\partial V}{\partial \phi} = 0$$

suppressing the oscillatory behavior

broadening peaks in SGWB spectrum...?



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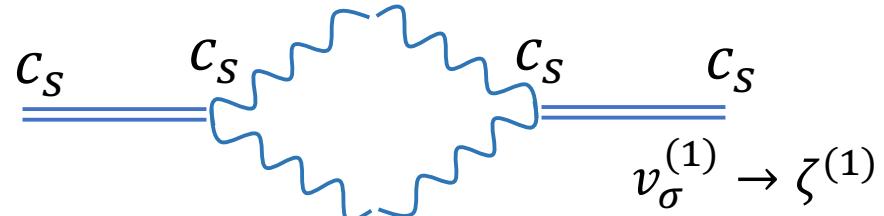


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Based on: arXiv:2501.02890 [astro-ph.CO]

single field



$$\mathcal{P}_\zeta^{(1)}(k) \equiv \left[\mathcal{P}_\zeta^{(0)} \right]^2 e^{4\pi\xi} f_{2,\zeta} [\epsilon_\sigma, c_{s,\sigma}, \xi]$$

$$c_{s,\sigma}^{-2} \quad E \cdot B \text{ contribution} \propto c_{s,\sigma}^6$$

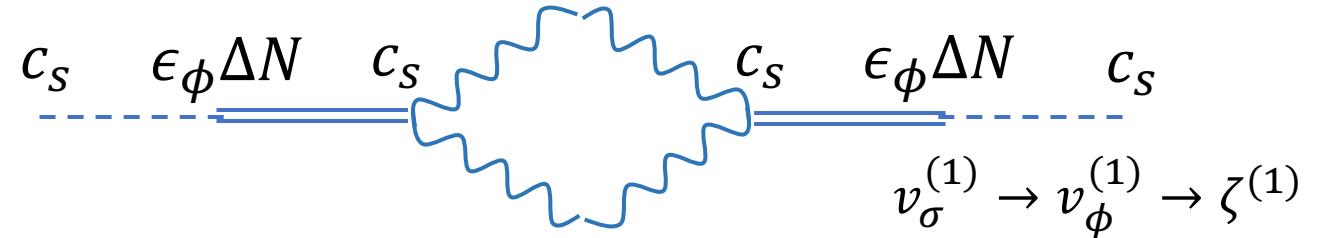
$$f_{2,\zeta} [\epsilon_\sigma, c_{s,\sigma}, \xi] \simeq \frac{7.47 \times 10^{-5}}{\xi^6} c_{s,\sigma}^6$$

$$- \frac{1.92 \times 10^{-5}}{\xi^6} c_{s,\sigma}^4 \epsilon_\sigma + \frac{1.89 \times 10^{-6}}{\xi^6} c_{s,\sigma}^2 \epsilon_\sigma^2 \dots$$

$$F(\vec{k}_1, \vec{k}_2, \vec{k}_3) = \frac{\left[\mathcal{P}_\zeta^{(0)} \right]^3}{k_1^2 k_2^2 k_3^2} e^{6\pi\xi} f_{3,\zeta}$$

\$E \cdot B\$ contribution \$\propto c_{s,\sigma}^9\$

two-field



$$\mathcal{P}_\zeta^{(1)}(k) \equiv \left[\epsilon_\phi \mathcal{P}_\zeta^{(0)} \right]^2 e^{4\pi\xi} f_{2,\zeta} [c_s, \xi, \Delta N_k]$$

$$f_{2,\zeta} \simeq \frac{7.47 \times 10^{-5}}{\xi^6} c_s^6 (1 + c_s^2)^2 \Delta N_k^2 - \frac{1.92 \times 10^{-5}}{\xi^6} c_s^4 (1 + c_s^2) \Delta N_k \dots$$

$$F(\vec{k}_1, \vec{k}_2, \vec{k}_3) = \frac{\left[\epsilon_\phi \mathcal{P}_\zeta^{(0)} \right]^3}{k_1^2 k_2^2 k_3^2} e^{6\pi\xi} f_{3,\zeta}$$

tensor:

$$\mathcal{P}_\lambda^{(1)}(k) \equiv \frac{2}{\pi^2} \frac{H^4}{M_p^4} e^{4\pi\xi} f_{h,\lambda}(\xi) \quad f_{h,+}(\xi) \simeq \frac{4.3 \times 10^{-7}}{\xi^6}$$

- Analytic expression ex.) single field

$$\mathcal{P}_\zeta^{(1)}(k) = \left[\mathcal{P}_\zeta^{(0)} \right]^2 \frac{9\pi^3 \xi^2 e^{4\pi\xi}}{16} \int_1^\infty dx \int_0^1 dy \frac{(x^2 - 1)^2}{\sqrt{x^2 - y^2}} \mathcal{I}_\zeta^2 \left[\epsilon_\sigma, c_{s,\sigma}, \xi, \sqrt{\frac{x+y}{2}}, \sqrt{\frac{x-y}{2}} \right]$$

$$\begin{aligned} \mathcal{I}_\zeta \left[\epsilon_\sigma, c_{s,\sigma}, \xi, \sqrt{\tilde{p}}, \sqrt{\tilde{q}} \right] &\equiv c_{s,\sigma}^3 \left[\tilde{p}^{1/2} + \tilde{q}^{1/2} \right] \mathcal{T}_\zeta^{(E \cdot B)} \left[\xi, \sqrt{\tilde{p}} + \sqrt{\tilde{q}} \right] \\ &+ \epsilon_\sigma c_{s,\sigma} \left[c_{s,\sigma}^2 - (\tilde{p} - \tilde{q})^2 \right] \left\{ \mathcal{T}_\zeta^{(E^2)} \left[\xi, \sqrt{\tilde{p}} + \sqrt{\tilde{q}} \right] + \tilde{p}^{1/2} \tilde{q}^{1/2} \mathcal{T}_\zeta^{(B^2)} \left[\xi, \sqrt{\tilde{p}} + \sqrt{\tilde{q}} \right] \right\} \end{aligned}$$

$$\mathcal{T}_\zeta^{(E^2)} [\xi, Q] \equiv \frac{1}{3\pi^{3/2}\xi^{1/2}} \int_0^\infty dx \frac{-c_s x \cos(c_s x) + \sin(c_s x)}{c_s^3} x^{-1/2} \exp \left[-2\sqrt{2\xi x} Q \right]$$

$$\simeq \frac{1}{3\pi^{3/2}\xi^{1/2}} \int_0^\infty dx \frac{x^{5/2}}{3} \exp \left[-2\sqrt{2\xi x} Q \right] = \frac{5}{32\sqrt{2}\pi^{3/2}\xi^4 Q^7} , \quad \text{---} \quad c_s \text{ does not alter scale}$$

$$\mathcal{T}_\zeta^{(E \cdot B)} [\xi, Q] \equiv \frac{\sqrt{2}}{3\pi^{3/2}} \int_0^\infty dx \frac{-c_s x \cos(c_s x) + \sin(c_s x)}{c_s^3} \exp \left[-2\sqrt{2\xi x} Q \right]$$

$$\simeq \frac{\sqrt{2}}{3\pi^{3/2}} \int_0^\infty dx \frac{x^3}{3} \exp \left[-2\sqrt{2\xi x} Q \right] = \frac{35}{64\sqrt{2}\pi^{3/2}\xi^4 Q^8} , \quad \text{---} \quad \rightarrow \text{amplitude suppression}$$