# Probing Dark Sectors with EDMs

Based on MA, M. Rahat, N. Valori, O. Vives <u>2407.21100</u>

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 $\mathscr{L}_{\rm SM} = \mathscr{L}_{\rm Gauge} + \mathscr{L}_{\rm Yukawa} + \mathscr{L}_{\rm Higgs}$ 



- The Standard Model of Particle Physics is extremely successful in explaining a wide variety of phenomena
- No evidence of new states

### • Yet we know that it cannot be the full story...

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# The need for Beyond SM physics

- Strong CP Problem
- Hierarchy Problem
- Flavour puzzle
- • •



# The need for Beyond SM physics

- Strong CP Problem
- Hierarchy Problem
- Flavour puzzle

- Neutrino masses
- Dark matter
- Baryon asymmetry of the Universe

 $\bullet$ 

. . .



normal hierarchy (NH) inverted hierarchy (IH)  $m^2 \uparrow$  $\uparrow m^2$  $\nu_3$  $\left[\Delta m^2_{
m sol}\right]$  $\Delta m_{\rm atm}^2$ dark matte **25%**  $\Delta m^2_{\rm atm}$ dark energy **70%**  $\Delta m_{\rm sol}^2$  $\nu_3$  $\nu_1$  $\nu_{\mu} \ \nu_{\tau}$  $\nu_e$  $\frac{n_B - n_{\bar{B}}}{\sim} \sim 10^{-10}$  $n_{\gamma}$ 



## Dark Sector (DS)

- Dark sectors (=set of particles that interact feebly with the SM) are attractive because
  - They can provide natural **Dark Matter candidates**
  - violation)

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### SM

 $SU(3) \times SU(2) \times U(1)$ 

 $\ell, q, u, d, \nu, H$ 

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## Vector portal to the DS: simplest case

• The SM and DS gauge bosons could mix kinetically



 $U(1)_Y \quad U(1)_D$ 

## Vector portal to the DS: simplest case

The SM and DS gauge bosons could mix kinetically





- The kinetic mixing for two U(1) is renormalizable ( $\epsilon$  is dimensionless)
- To make the dark sector feebly interacting with the SM we have to take  $\epsilon \ll 1$



• If the U(1) is not embedded in a larger group at higher energies,  $\epsilon$  is just another tree-level Lagrangian coupling



### Vector portal to the DS: Non-abelian dark sectors



• The vector bosons of the dark sector carry an adjoint index A:  $X^{A}_{\mu\nu}$ .

 $\Rightarrow$  Vector mixing is only possible at the non-renormalizable level!



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• • •

 $T^A \equiv \text{generator}$  $\Sigma^A \equiv adjoint scalar$  $\phi \equiv$  scalar fundamental rep



### Vector portal to the DS: Non-abelian dark sectors

 $G_D = SU(N)_D$ 

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 $\bullet$ 

$$\frac{\langle \Sigma^A \rangle}{\Lambda} B^{\mu\nu} X^A_{\mu\nu} \equiv \epsilon^A B^{\mu\nu} X^A_{\mu\nu}$$

- $T^A \equiv \text{generator}$  $\Sigma^A \equiv adjoint scalar$  $\phi \equiv$  scalar fundamental rep
- Kinetic mixing possible if scalar gets vevs (or group confines) and is naturally suppressed if  $\Lambda$  is large: [Alonso-Alvarez et al '23]



### Intermezzo: CP violation

- CP distinguishes matter and anti-matter  $\Rightarrow$  CP violation necessary to explain baryon and anti-baryon asymmetry!  $\hat{C}\hat{P} \Rightarrow$  $i \xrightarrow{\Delta B} f$ 
  - All CP violation in the SM is parametrised by the Jarlskog invariant  $J \simeq 10^{-3}$  [Jarlskog '85] [PDG RPP '22]
    - (Although we do not know if CP is violated in the lepton mixing sector)
    - We know that CP violation is absent or ridicolously small in the strong interactions)

[Sakharov '6  
$$\overline{i} \longrightarrow \overline{f}$$
 It must be  $\Gamma(i \to f) \neq \Gamma(\overline{i} \to \overline{f})$   
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 $\Rightarrow$  CP violation is both needed and a clean probe of beyond Standard Model physics

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### $G_D = SU(N)_D$

### No way to define a C charge for $\Sigma$ that make both CP even: CP is violated

$$\frac{\Sigma^A}{\Lambda} B^{\mu\nu} X^A_{\mu\nu}$$

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• Consider non-abelian dark sectors (quite popular for DM model building and more...)

[Buttazzo et al '19,'20] [Landini, Wang '20] [MA et al '20] [Frigerio et al '23] [Borah et al '22]

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- Assume that there is at least one scalar  $\Sigma^A$  trasforming in the adjoint that acquire a VEV  $\langle \Sigma^A \rangle \neq 0$

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- Relevant interactions:

$$rac{\langle \Sigma^A 
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u}$$

$$\overset{\gamma}{\Longrightarrow}\overset{\epsilon}{\Longrightarrow}\overset{A'}{\longrightarrow}$$

**Kinetic mixing** 

[Buttazzo et al '19,'20] [Landini, Wang '20] [MA et al '20] [Frigerio et al '23] [Borah et al '22]



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$$H)(\Sigma^A \Sigma^A)$$

h -----  $\Sigma^A$ 

### **Higgs-Dark scalar mixing**



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**Higgs-Dark scalar mixing** 

• Contribution to the EDMs:







$$\mathcal{L}_{\text{EDM}} = -\frac{i}{2} d_f \left( \bar{\psi}_f \sigma_{\mu\nu} \gamma_5 \psi_f \right) F^{\mu\nu}$$

- Experiments are sensitive to EDMs much larger than what the SM predicts
  - $\Rightarrow$  They still provide stringent constraint on CP violating New Physics

[Chupp et al, 1710.0250]

SUSY CPV Two Higgs Doublet Model Left-Right Symmetric Model

Next generation of experiments expect to improve their sensitivities of orders of magnitude



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Experiment	Current bound/Upcoming sensitivity
JILA eEDM	$<4.1\times10^{-30}$ e cm
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 $d_{\rho}$ 

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ACME III	$\sim 1 \times 10^{-30}~{\rm e~cm}$
YBF	$\sim 1 \times 10^{-31}$ e cm
BaF	$\sim 1 \times 10^{-33}$ e cm

 $d_{\mu} < 1.8 \times 10^{-19} \ e \cdot \text{cm} \rightarrow 6 \times 10^{-23} \ e \cdot \text{cm}$ 

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[muEDM, 2201.06561]





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Next generation of experiments expect to improve their sensitivities of orders of magnitude

• In flavour blind models, the electron EDM is the most constraining





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$$\Sigma B_{\mu\nu} X_{\rho\sigma}$$

$$\langle \Sigma \rangle = v_D, \ \epsilon = \tilde{\epsilon}$$

### eEDM vs Dark Photon searches

Assume that CP-even and CP-odd operators have a similar coefficient





- Assume that the dark photon acquire a mass  $M_X \sim v_D$  and the scalar mass is also  $m_{\Sigma} \sim v_D$
- $\beta$  is the mixing angle between Higgs and the dark scalar



$$\langle \Sigma^3 \rangle = v_D$$
  $M_{X_3} = 0, M$ 

 $SU(2)_D \rightarrow U(1)_D$ 

 $_{X_3} = 0, M_{X_{1,2}} = g_D v_D$ 

$$\langle \Sigma^3 \rangle = v_D$$
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• The two massless field mix kinetically

$$\frac{\epsilon}{v_D} \Sigma^A B^{\mu\nu} X^A_{\mu\nu} \Rightarrow \epsilon B^{\mu\nu} X^3_{\mu\nu}$$

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• Define the dark photon as the field that does not couple with SM particles  $\Rightarrow U(1)_D$  charged particles acquire a millicharge



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$$\Leftarrow \frac{\tilde{\epsilon}}{v_D} \ \epsilon^{\mu\nu\rho\sigma} \ \Sigma^A B_{\mu\nu} X^A_{\rho\sigma}$$



## EDMs vs millicharged particles searches

• Assuming  $\epsilon \sim \tilde{\epsilon}$  and  $g_D \sim 1$ , can compare the sensitivity of the EDM with other searches for millicharged particles

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- The massive gauge boson  $X_1, X_2$  combine into  $U(1)_D$  charged states  $W_D^{\pm}$  which are also **millicharged under**  $U(1)_{\rm em}$ . The *x* axis correspond to the vector masses
- $\beta$  is the mixing angle between Higgs and the dark scalar

### • Assuming $\epsilon \sim \tilde{\epsilon}$ and $g_D \sim 1$ , can compare the sensitivity of the EDM with other searches for



- With two adjoints  $\Sigma^A_1, \ \Sigma^A_2$  we can break completely  $SU(2)_D$ 

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### Conclusions

- The Standard Model must be extended to address various unresolved puzzles
- CP violating New Physics is motivated by the observed baryon-antibaryon asymmetry
- EDM are exceptionally sensitive observables to CP violation

The EDM could indirectly probe Dark Sectors that respect a non-abelian gauge symmetry

If CP violation is large in the Dark Sector, EDM can have a better sensitivity to kinetic mixing parameters than other experimental probes (dark photon searches, millicharged particles)



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**Back-up Slides** 

## **CP** violation in the SM

Where does CP violation in the SM come from? lacksquare

$$-\mathscr{L}_{W^{\pm}} = \frac{g}{\sqrt{2}} \overline{u_{Li}} \gamma^{\mu} \left( V_{\text{CKM}} \right)_{ij} d_{Lj} W_{\mu}^{+} + h \cdot c .$$

$$\Rightarrow If V_{CKM} \neq V_{CKM}^*, CP \text{ is violated} \qquad V_{CKM} = \begin{pmatrix} V_{ud} \\ V_{cd} \\ V_{td} \end{pmatrix}$$
Possible only with at least three flavours
[Kobayashi, Maskawa '73]

$$\operatorname{Im}\left(V_{ij}V_{kl}V_{il}^*V_{kj}^*\right) = J \times \sum_{m,n} \varepsilon_{ikm} \varepsilon_{jln}$$

Repeated indices are not summed



• But observables don't care about the basis/parametrization, so CP violation must be proportional to an invariant

Jarlskog invariant [Jarlskog '85]

 $J = s_{12}s_{23}s_{31}c_{12}c_{23}c_{31}^2 \sin \delta = (3.08^{+0.15}_{-0.13}) \times 10^{-5}$ 

[PDG RPP '22]







## **Electric dipole moments (EDMs)**

• A particle with spin  $\vec{S}$  can have a magnetic and an electric dipole moment

$$H = -(\mu \vec{S} \cdot \vec{B} + d \vec{S} \cdot \vec{E})/S$$

$$\mathscr{L}_{\text{EDM}} = -\frac{i}{2} d_f \left( \bar{\psi}_f \sigma_{\mu\nu} \gamma_5 \psi_f \right) F^{\mu\nu} \qquad [\psi] = 3/2, \ [F_{\mu\nu}] = 2, \ [\mathscr{L}] = 4 \Rightarrow [d_f] = 4$$

•  $d_f$  can only be non-zero if CP is not conserved

•  $\vec{S}$  and  $\vec{B}$  are axial vectors, while  $\vec{E}$  is a vector  $\Rightarrow \vec{S} \cdot \vec{B}$  and  $\vec{S} \cdot \vec{E}$  are P and CP even and odd respectively

• In a quantum field theory, the electric dipole moment of a fermion is given by the CP-odd operator

## EDM in the SM vs exp. searches

• EDMs are predicted to be very small in the SM



Experiments are sensitive to much larger values  $\bullet$ 

$$d_e^{(\exp)} \lesssim 4 \times 10^{-30} \ e \cdot \text{cm}$$
$$d_n^{(\exp)} \lesssim 2 \times 10^{-26} \ e \cdot \text{cm}$$

Upper limit on neutron EDM severely constraint CP violation in strong interactions (why so small? = Strong CP problem)

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### Lepton EDMs appear at four or higher-loops

Back-of-the-envelope:

$$d_l \lesssim \frac{m_l}{m_e} \times 10^{-44} \ e \cdot \mathrm{cm}$$

**Quark EDMs appear at three or higher-loops** 

Back-of-the-envelope:

 $d_{u,d} \lesssim 10^{-34} \ e \cdot \mathrm{cm}$ 

For an up-to-date estimate see: [Yamanaka, Yamaguchi '20]

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[JILAeEDM - Roussy et al '23]

[Abel et al '20, PRL 124-081803]

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[Pospelov, Ritz '14]

## EDM prediction dark sector



 $d_l\simeq {1\over 8\pi^2}$ 



 $d_l \simeq e$ 

$$\frac{Y_l}{x^2 v_D} \epsilon \tilde{\epsilon} \beta c_{\theta}^2 e \left( \frac{\log(x_{X\phi})}{x_{X\phi} - 1} - \frac{\log(x_{Xh})}{x_{Xh} - 1} \right)$$

 $x_{ij} = m_i^2 / m_j^2$ 

$$\tan \chi = \tilde{\epsilon}/\epsilon$$

$$\left(rac{Y_l}{8\pi^2 v_D}
ight)\epsilon^2 an\chi \ eta c_{ heta}^2 \log\left(rac{m_h^2}{m_{\phi}^2}
ight)$$

## Why inelastic DM?

• Correct relic abundance for thermal freeze-out require  $\langle \sigma v_{\rm rel} \rangle \sim 1.7 \times 10^{-9} \, {\rm GeV}^{-2}$ 



- $\bullet$ detection too
- In the inelastic DM scenario, there are two DM states  $\chi_H$ ,  $\chi_S$  with a small mass splitting

• Requires  $\epsilon \sim 10^{-4}$ ,  $M_X \sim \text{few} \times m_\chi$  and  $m_\chi \lesssim 1$  GeV to have the correct DM abundance and respect laboratory constraints

But then it conflicts with indirect detection that place constraints on swave annihilation  $m_{\gamma}\gtrsim 30$  GeV, a region severely constrained by direct

• If only inelastic scattering is allowed  $\chi_H \chi_S \to SM$  (which sets the relic abundance), one can avoid the indirect detection bound because  $\chi_H$  would have decayed in  $\chi_S$  by then, and  $\chi_S \chi_S \to SM$  is forbidden