

PLANCK 2025

# Thermal Effects in Freeze-In Dark Matter Production

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María José Fernández Lozano

arXiv:2506.xxxxx

JGU Mainz

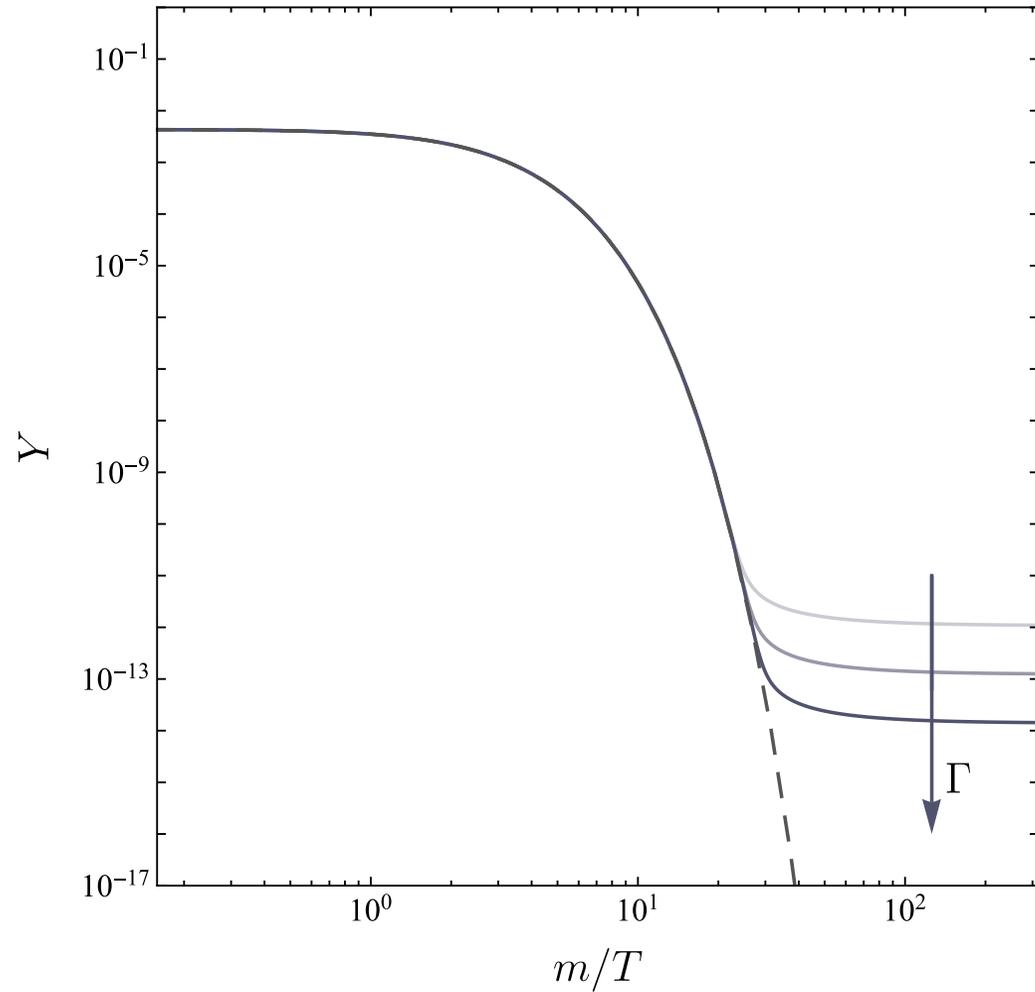
in collaboration with

MATHIAS BECKER

JULIA HARZ

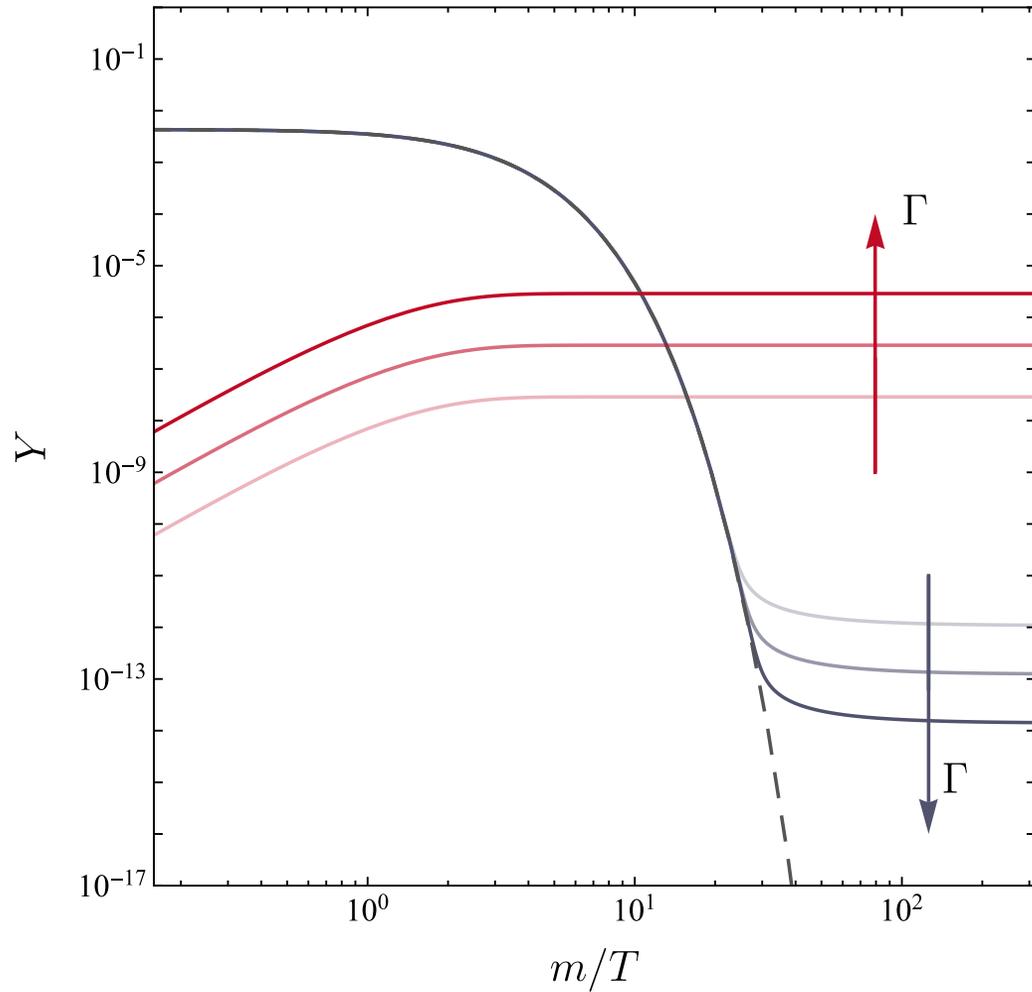
CARLOS TAMARIT

# DM PRODUCTION MECHANISMS



Freeze-out  
(WIMPs)

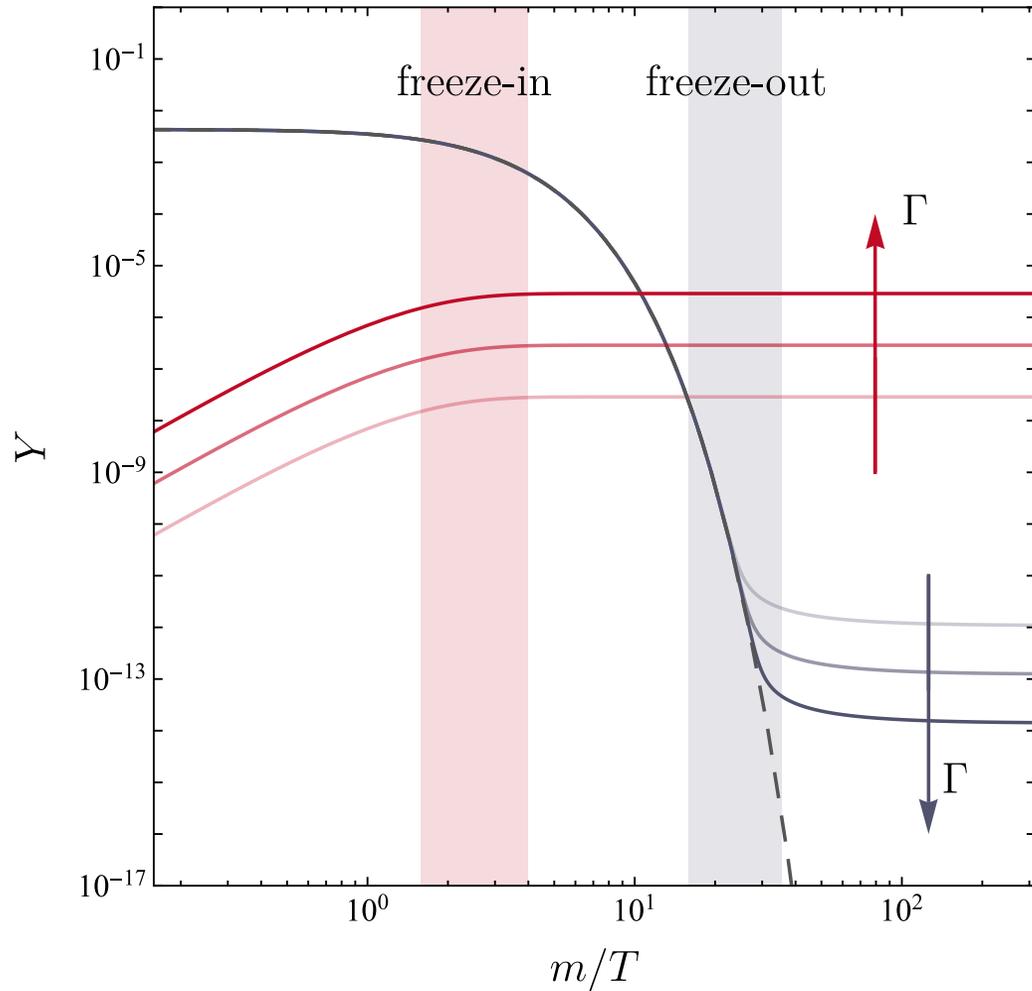
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Freeze-In  
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Decoupling at  
 $T \sim m_\chi/25$

Decoupling at  
 $T \sim m/5$

*Freeze-in models are sensitive to thermal corrections*

How does the thermal plasma affect the freeze-in mechanism?

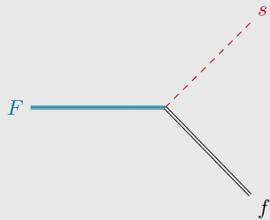
How does the **thermal plasma** affect the freeze-in mechanism?

Thermal masses, modified distribution functions, off-shell effects, thermal widths, **multiple soft scatterings (LPM)**, ...

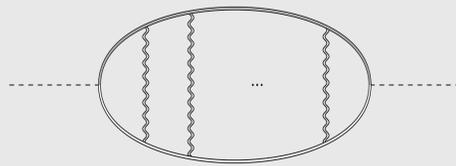
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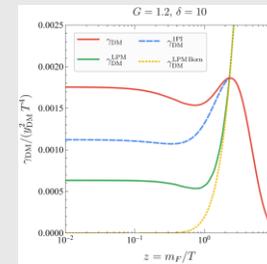
## Framework



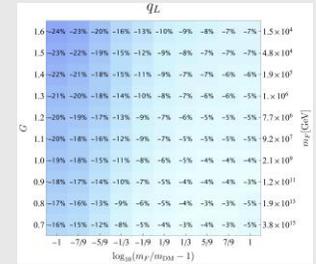
## The LPM Effect



## Complete Rate



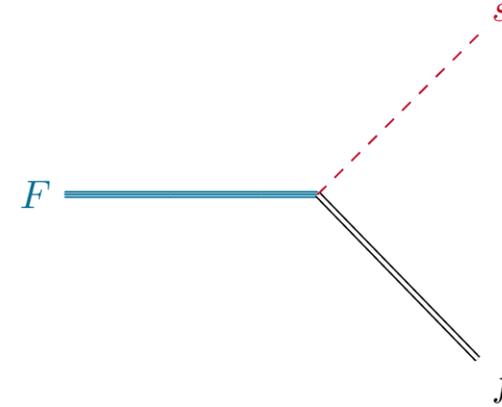
## Impact on $\Omega h^2$



# VECTORLIKE SCALAR FIMP MODEL

$$\mathcal{L}_{\text{int}} = - [y_{\text{DM}} \bar{F} f s + \text{h.c.}]$$

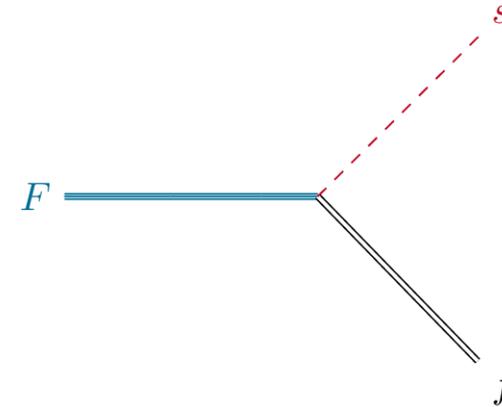
portal coupling      dark mediator      scalar dark matter particle



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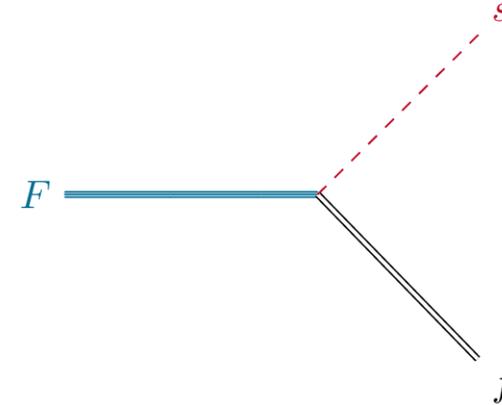
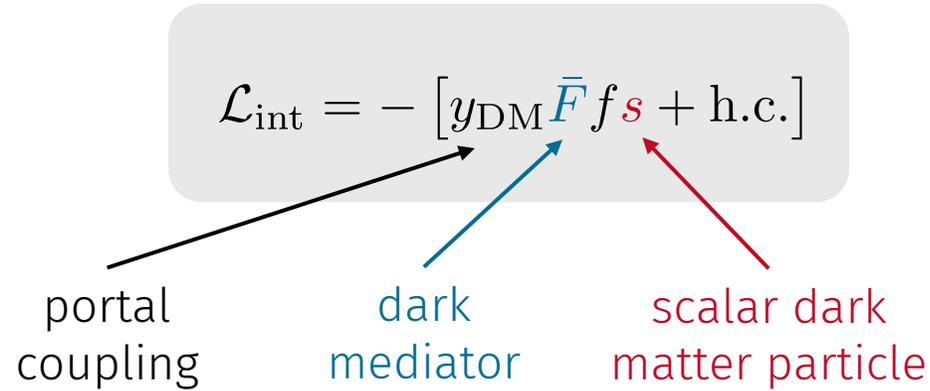
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5 possible realizations of this model:

$$e_L, q_L, e_R, d_R, u_R$$



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Parametrized with 4 quantities:

$$y_{\text{DM}}$$

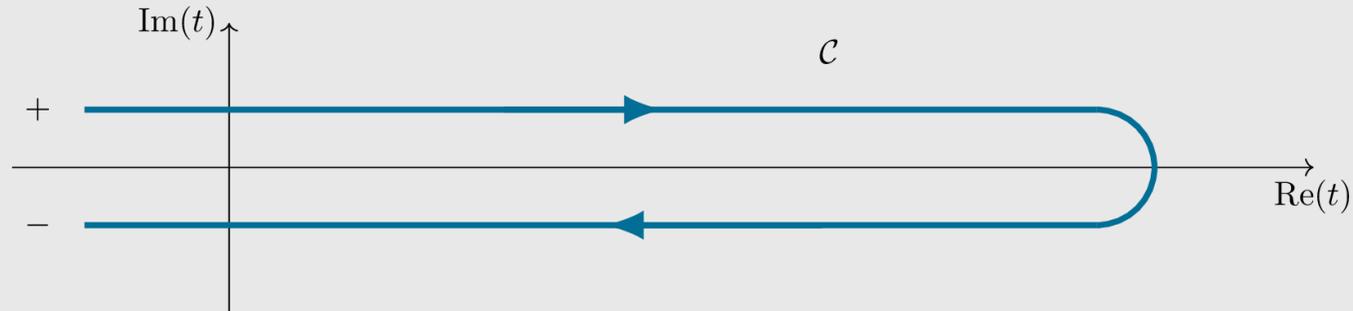
$$G = Y^2 g_1^2 + C_2(\mathcal{R}_2) g_2^2 + C_2(\mathcal{R}_3) g_3^2$$

$$\delta = \frac{m_F - m_{\text{DM}}}{m_{\text{DM}}}$$

$$m_F$$

## CTP Formalism

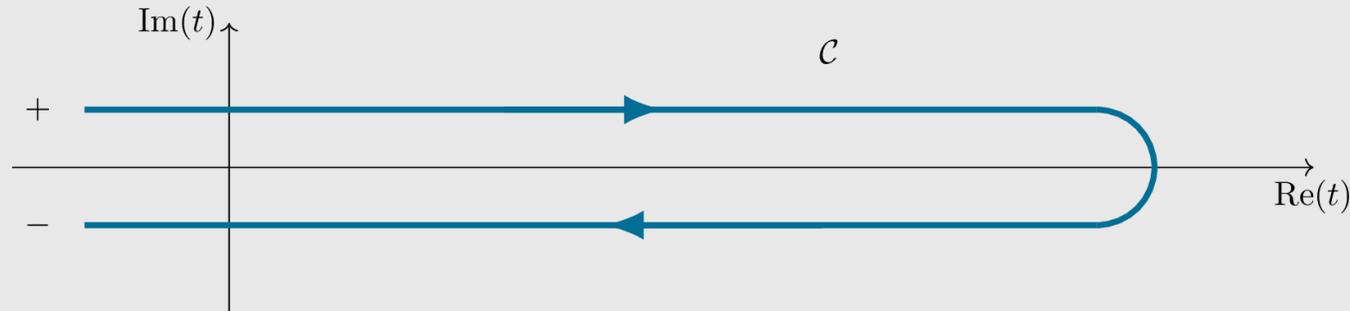
Fields are defined on a complex time contour that doubly transverses the real-time axis



Advantage: allows for non-equilibrium phenomena

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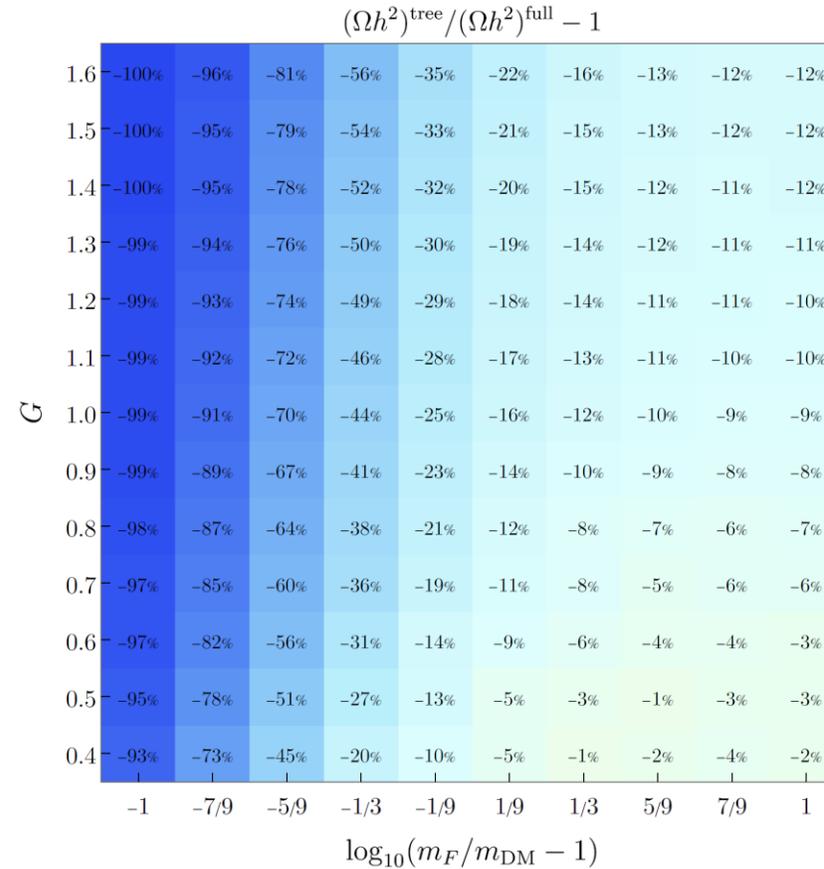
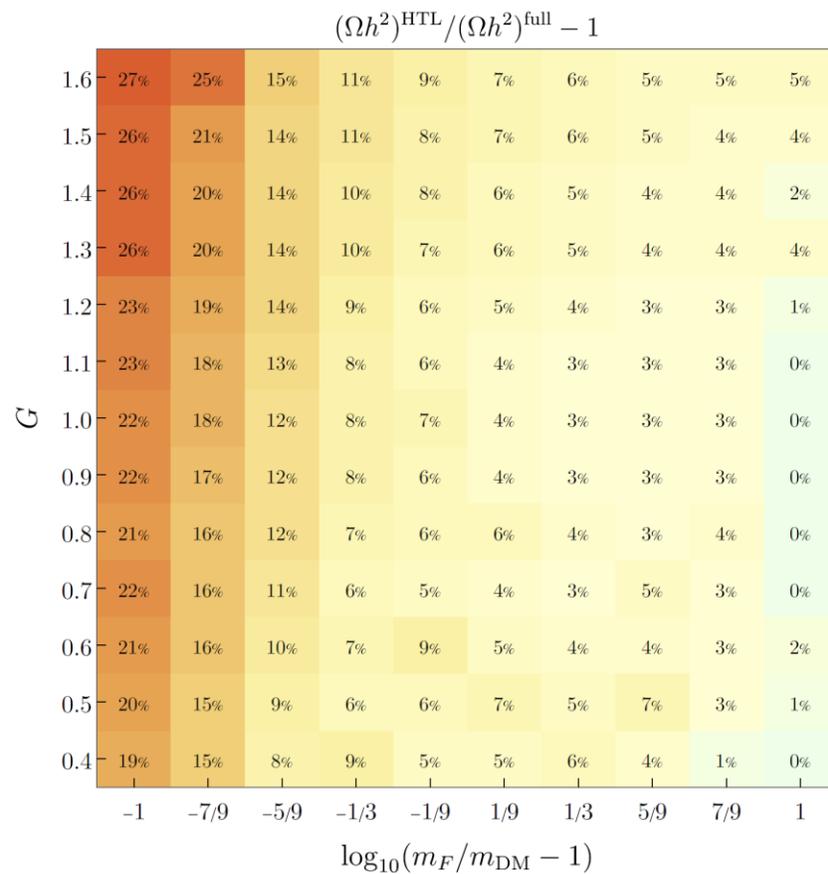
Advantage: allows for **non-equilibrium phenomena**

Equation for the DM production rate from the Kadanoff-Baym and CTP formalism

[Becker, Copello, Harz, Tamarit, 2312.17246]

$$\gamma_{\text{DM}} \equiv \int \frac{d^3 \vec{p}}{(2\pi)^3} \frac{\Pi_s^{\mathcal{A}}(\omega_p, |\vec{p}|)}{\omega_p} f_{\text{B}}(\omega_p)$$

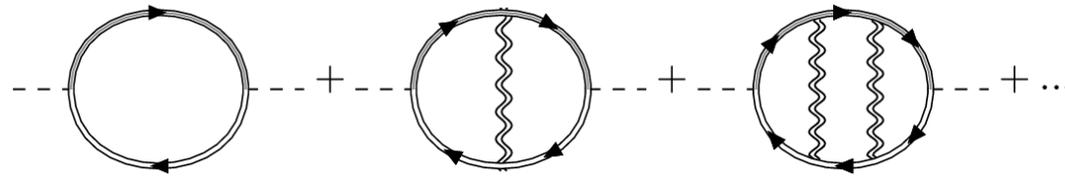
Calculation with 1PI-resummed propagators provides a more accurate rate for freeze-in



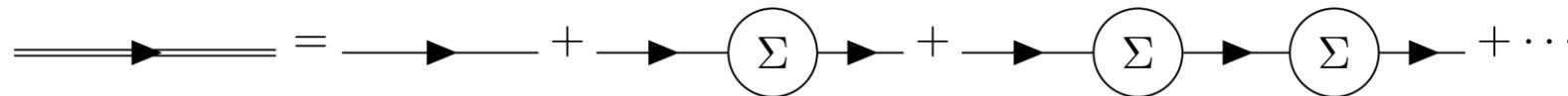
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Two key approximations:

- Self-energy truncation:  
whether we calculate  $\Pi_s^{\mathcal{A}}$  at 1-loop, 2-loop...

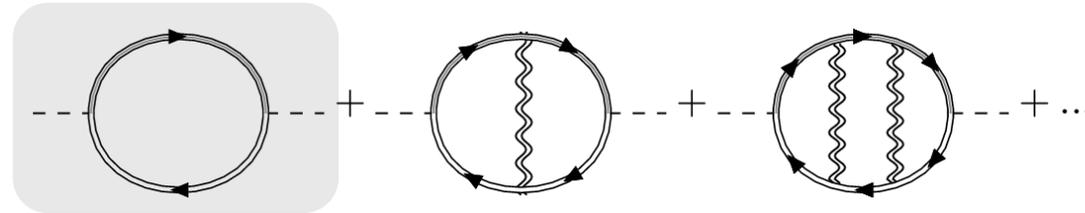


- Propagator structure:  
whether tree-level, HTL, or 1PI propagators are used inside  $\Pi_s^{\mathcal{A}}$



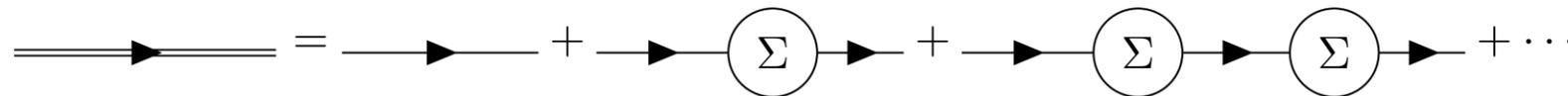
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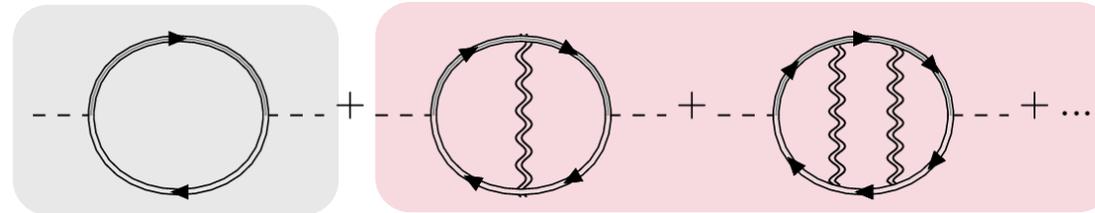
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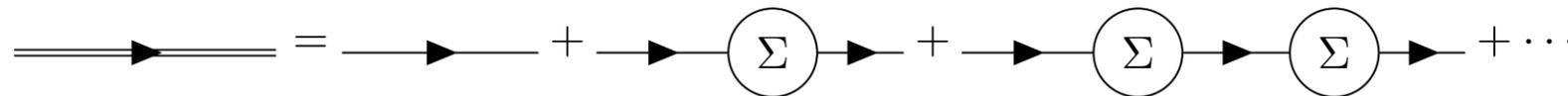


**This work**

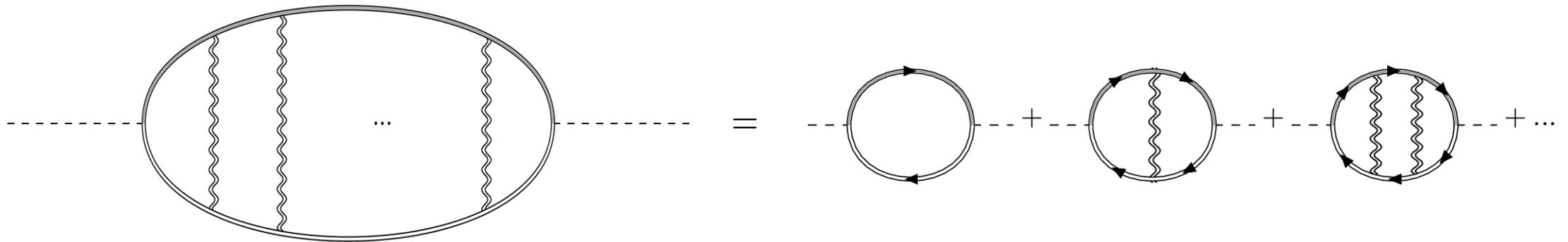
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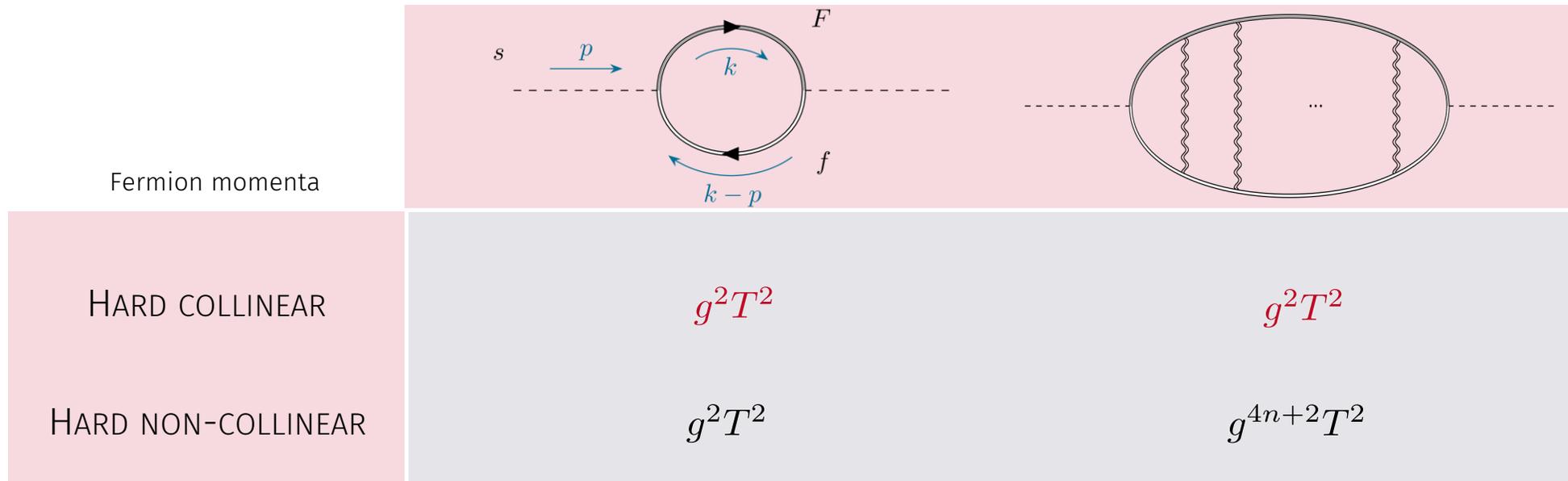


# THE LADDER DIAGRAM



The ladder diagram is a LO process in the coupling constant in the kinematical regime

$$\underbrace{p \sim T, \quad k \sim T, \quad p^2 \sim g^2 T^2, \quad k^2 \sim g^2 T^2}_{\text{Lightlike}}, \quad \underbrace{k \cdot p \sim g^2 T^2}_{\text{Collinear}}$$



# THE LPM EFFECT

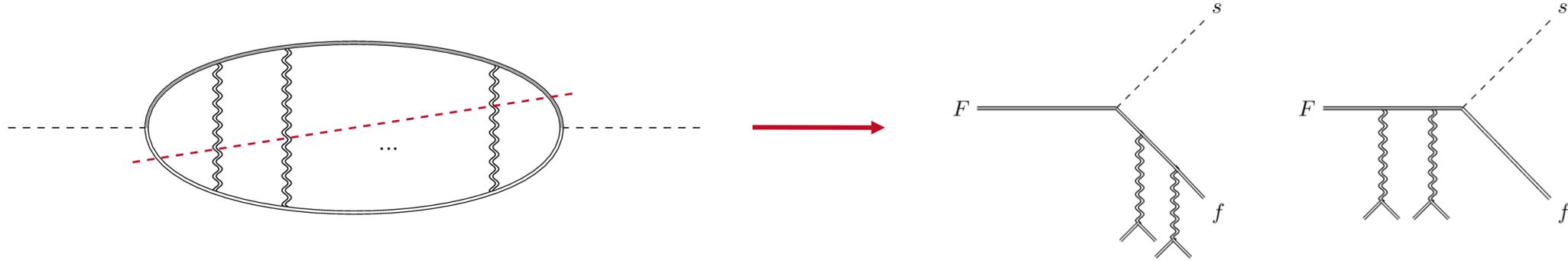
Discovered by **L**andau, **P**omeranchuk and **M**igdal (LPM) in the context of bremsstrahlung radiation.

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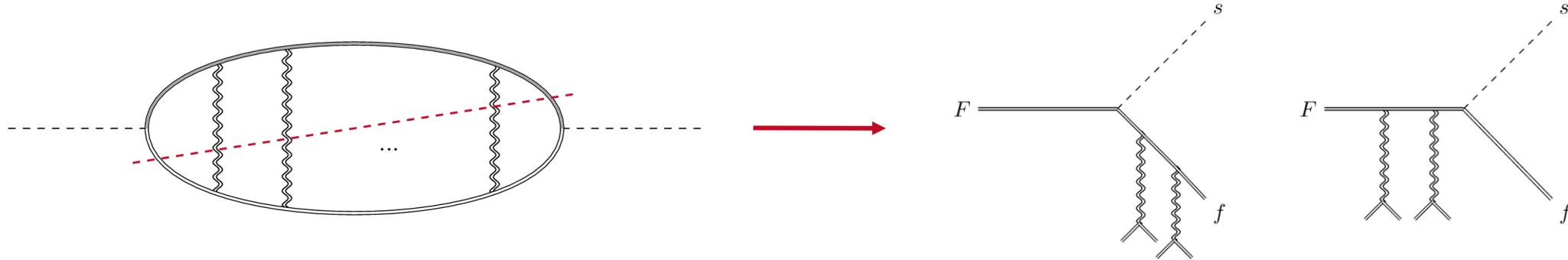
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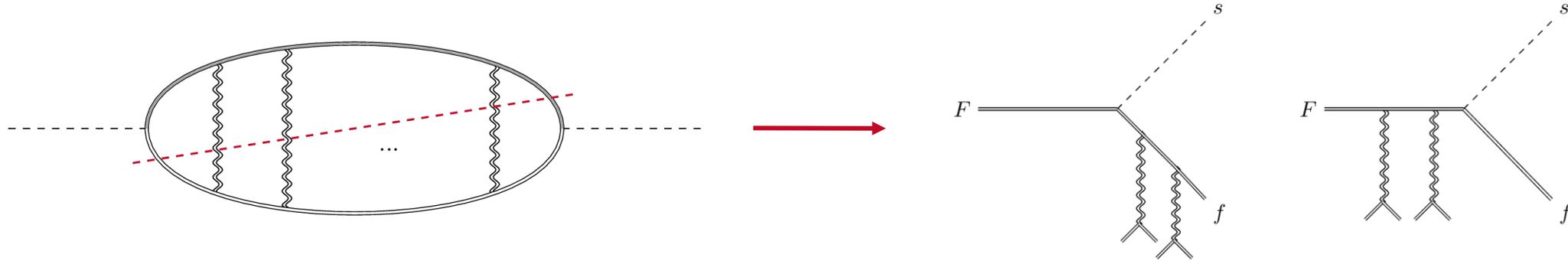


**Multiple soft scattering** of high-energy gauge bosons with the particles in a hot plasma.

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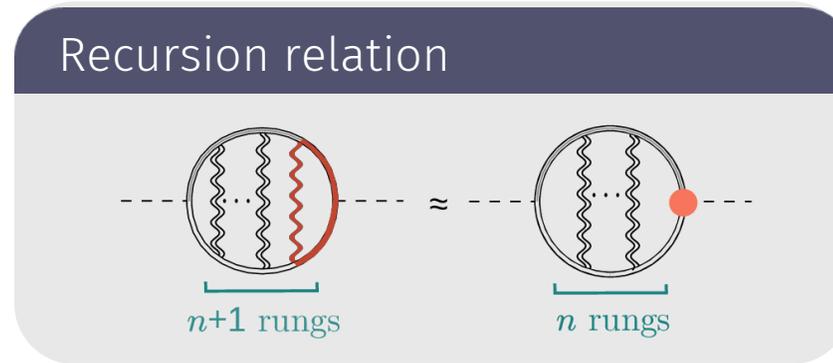
fermion self energies

- Leptogenesis [Besak, Bodeker, 1202.1288]
- Fermionic FIMPs [Biondini, Ghiglieri, 2012.09083]

photon self energies

- QGP [Arnold, Moore, Yaffe, hep-ph/011107]

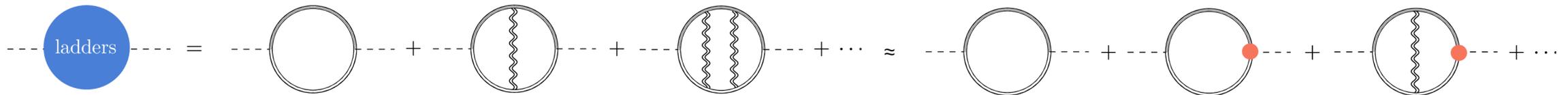
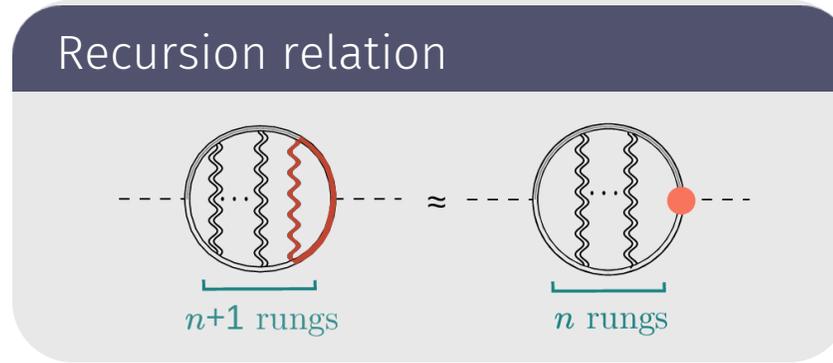
# CALCULATION OF THE LPM RATE



[Anisimov, Besak,  
Bodeker, 1202.1288]

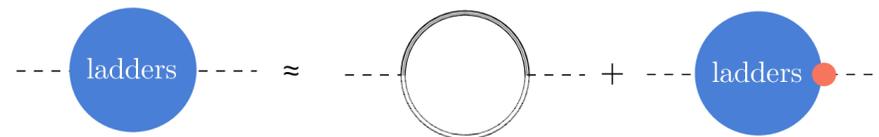
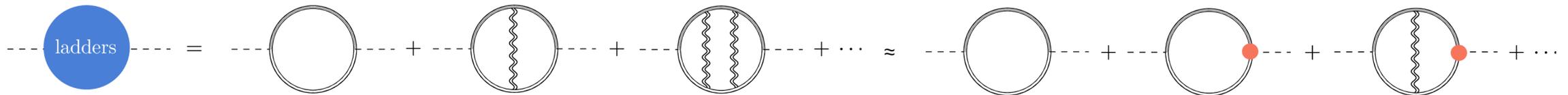
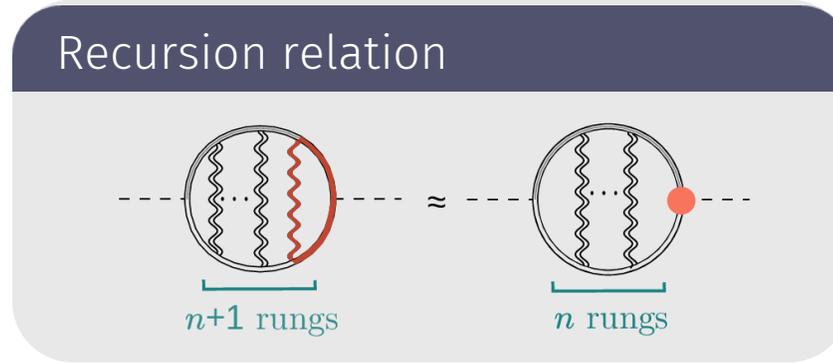
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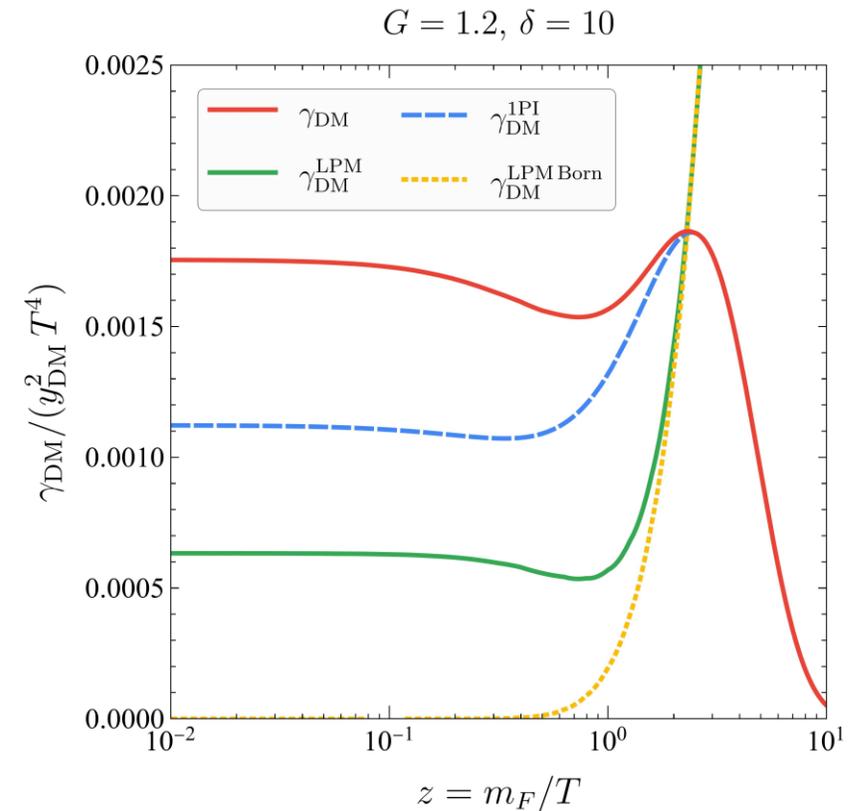


# SWITCHING OFF THE LPM RATE

We have calculated the LPM rate in the ultrarelativistic regime. To move to non-relativistic, we use a simple prescription:

[Biondini, Ghiglieri, 2012.09083]

$$\gamma_{\text{DM}} = \left( \gamma_{\text{DM}}^{\text{LPM}} - \gamma_{\text{DM}}^{\text{LPM Born}} \right) f(m_F) + \gamma_{\text{DM}}^{\text{1PI}}$$

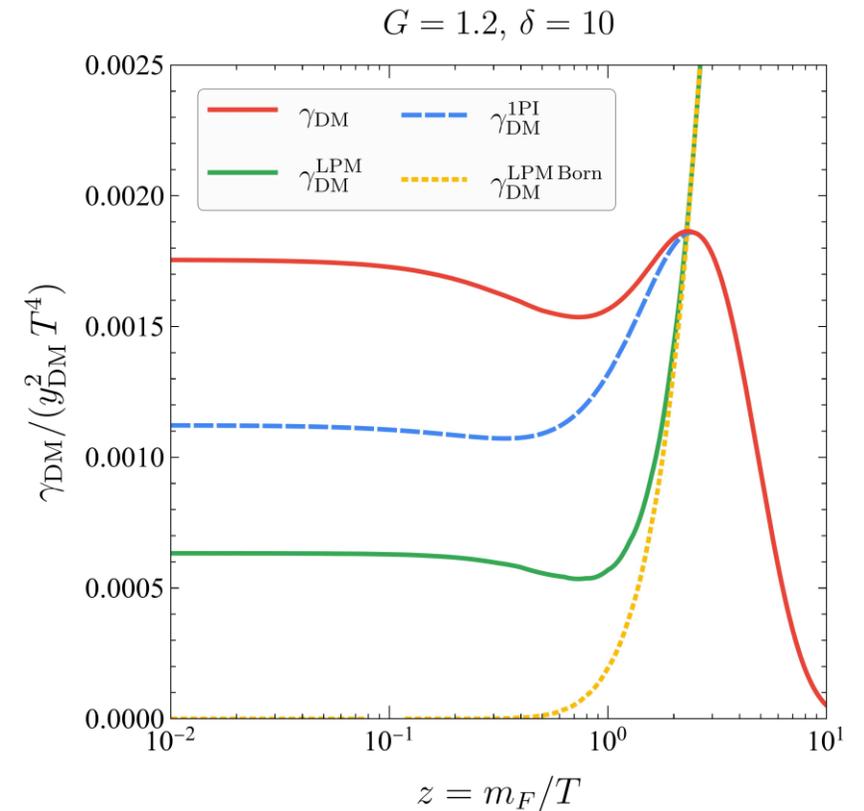
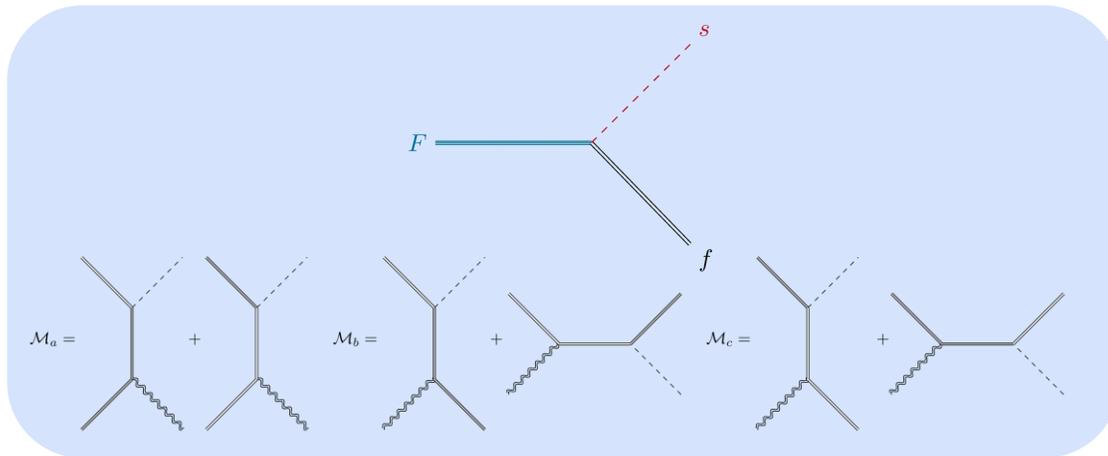
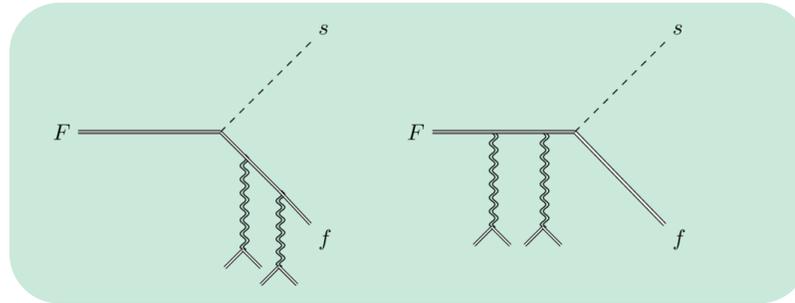


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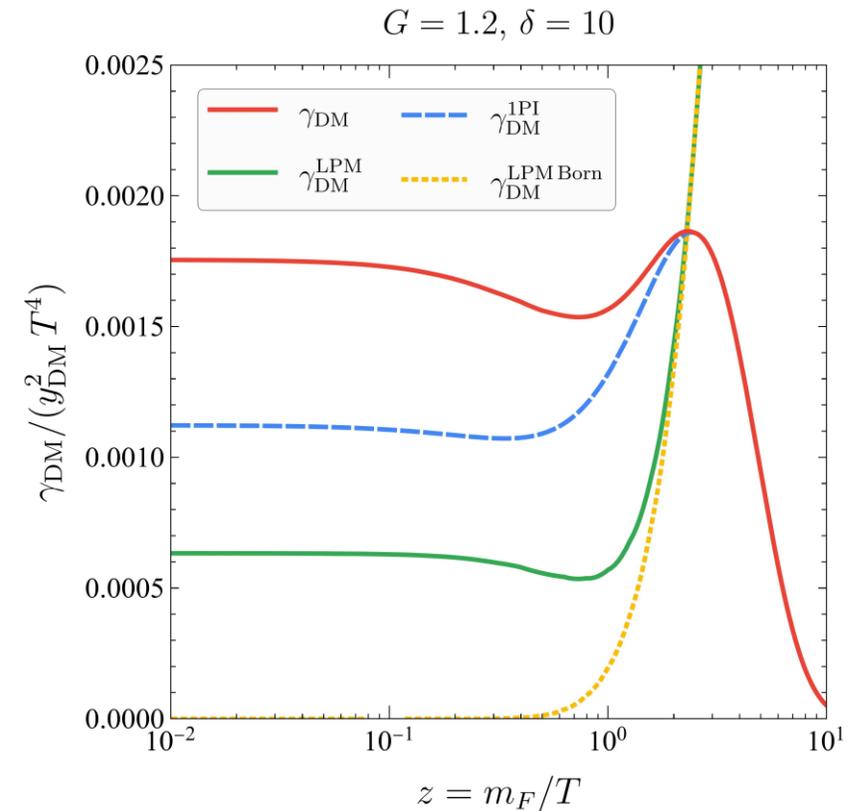
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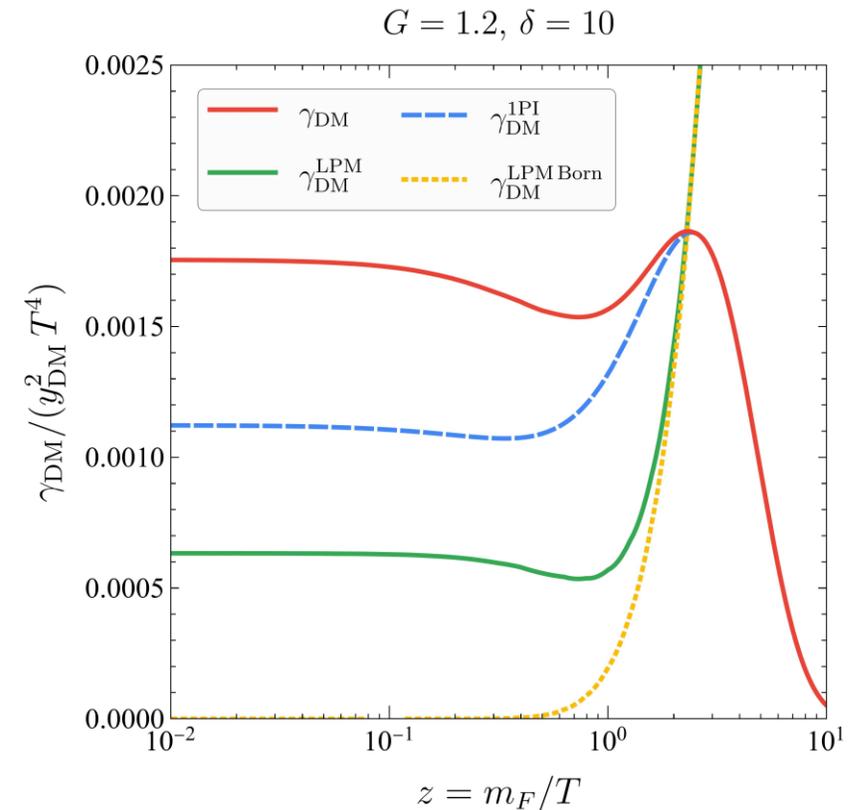
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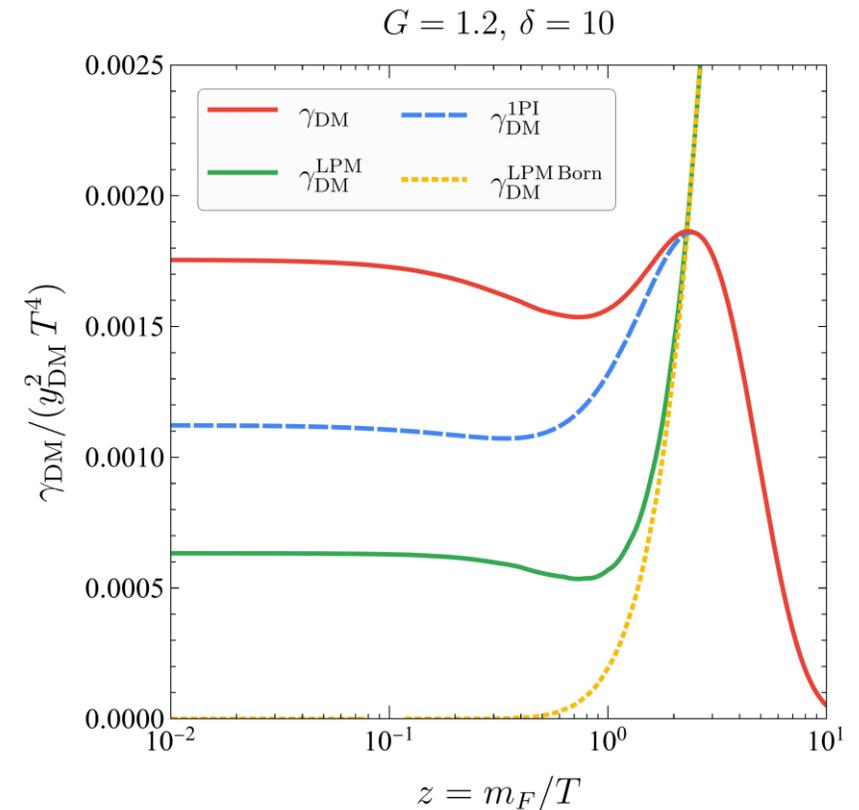
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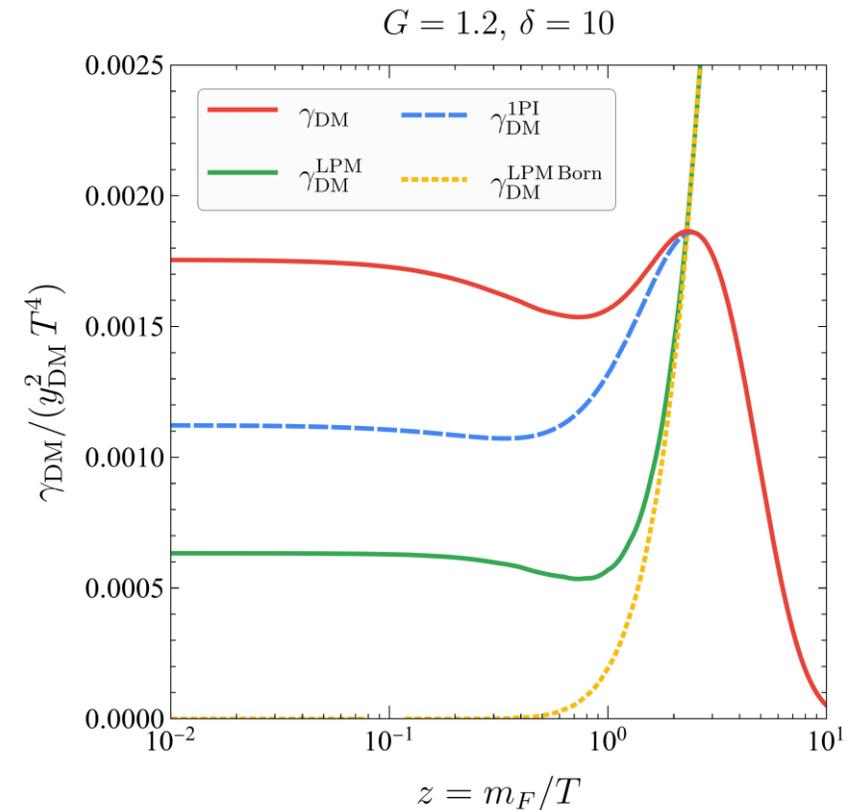
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## Switch-off function

Switches off the unphysical rate. Must satisfy

$$f\left(\frac{m_F}{T} \rightarrow 0\right) = 1, f\left(\frac{m_F}{T} \rightarrow \infty\right) = 0$$



# DIFFERENT SWITCH-OFF PROCEDURES

[Ghiglieri, Laine,  
1605.07720]

## Susceptibility function

- Slower switch-off might lead to DM overestimation
- Doesn't switch-off scatterings fast enough

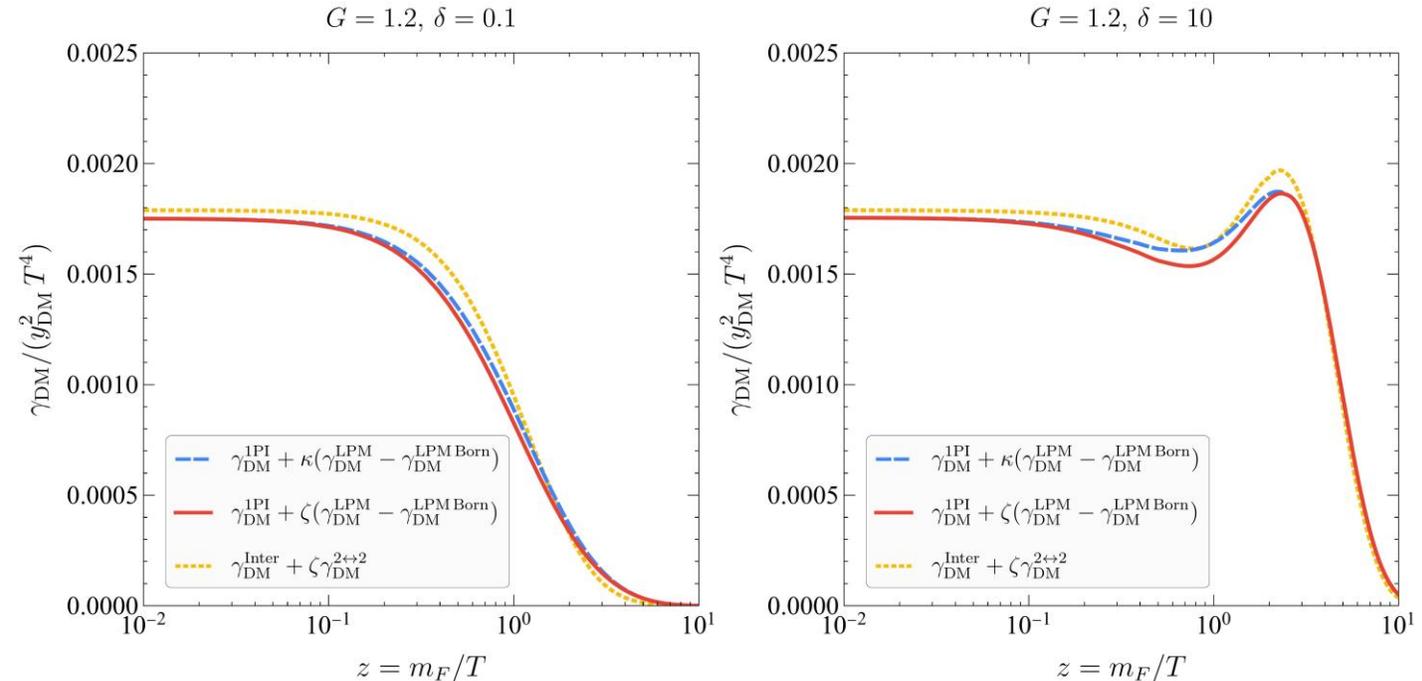
## Thermal function

- + Captures 1PI behaviour
- + Most conservative approach

## Smooth interpolation

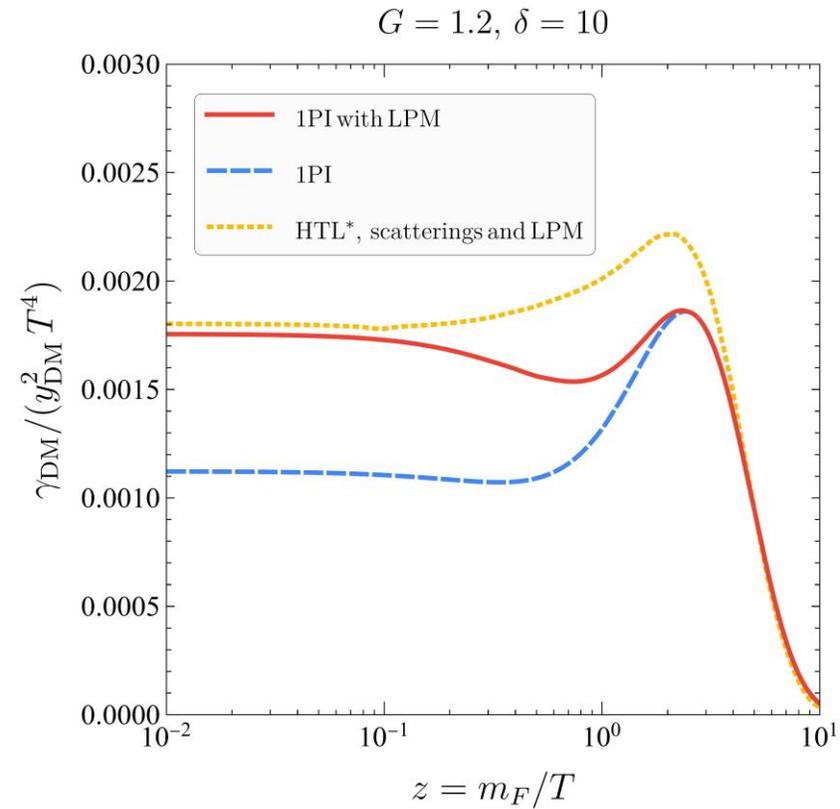
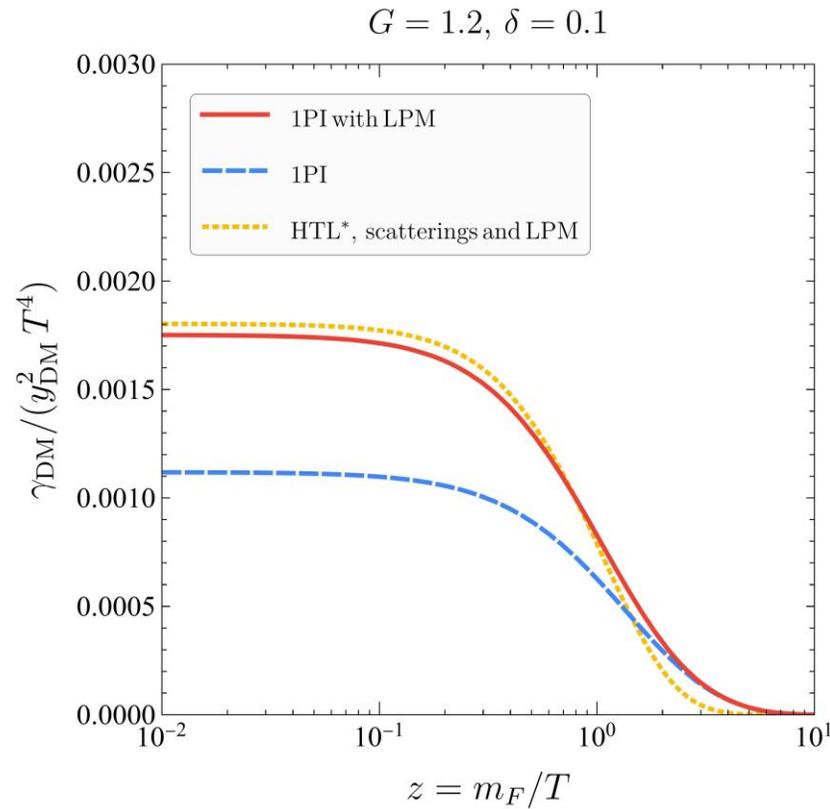
[Ghiglieri, Laine  
2110.07149]

- Scatterings must be added manually
- Decay peak is overestimated

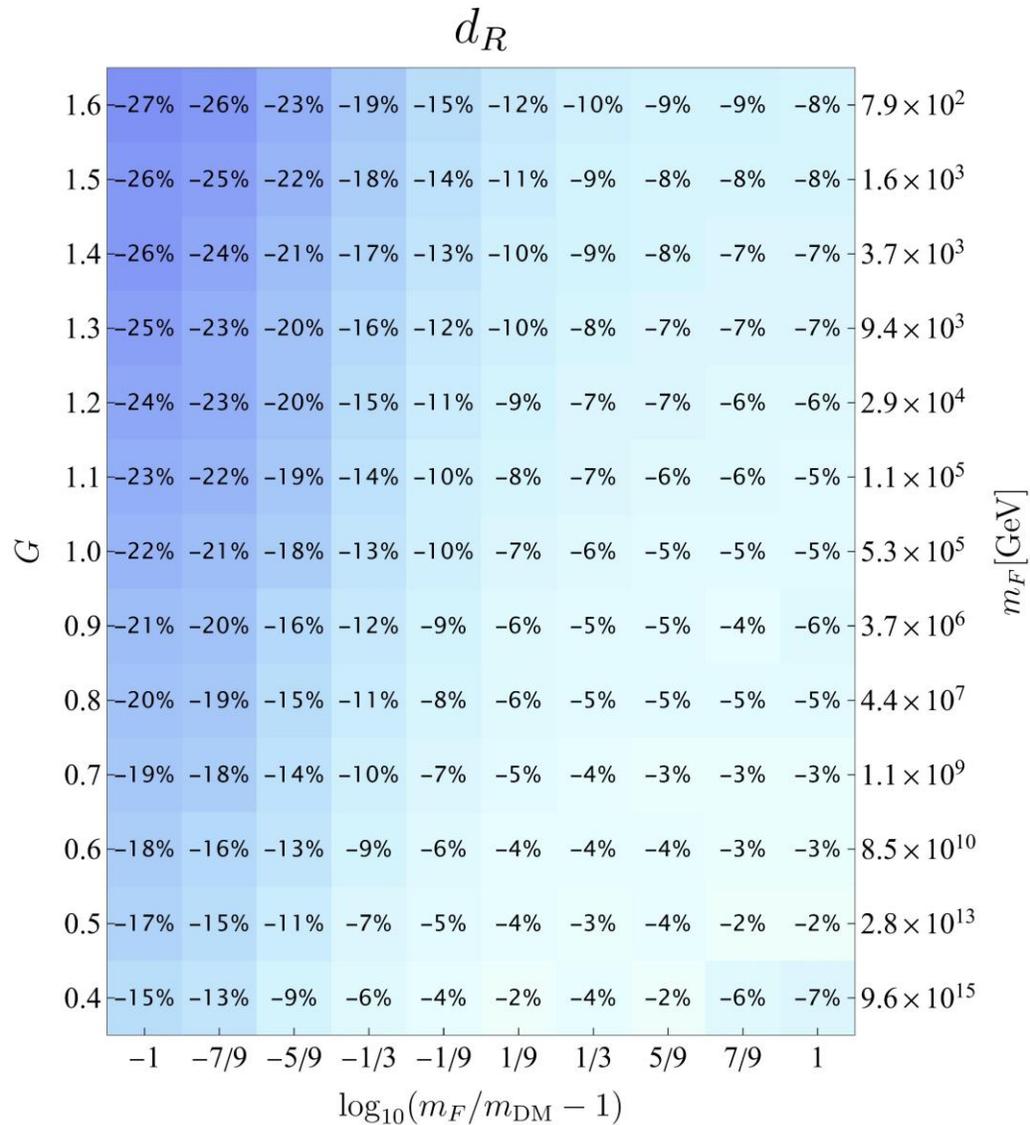


Discrepancies of  $\sim 30\%$  and  $\sim 5\%$  at the  $\Omega h^2$  level

# THE FINAL DM RATE



We provide a fitting of the LPM in the ultrarelativistic limit  $\left. \gamma_{\text{DM}}^{\text{LPM}} / (y_{\text{DM}}^2 T^4) \right|_{z=0.01} = a + bG$



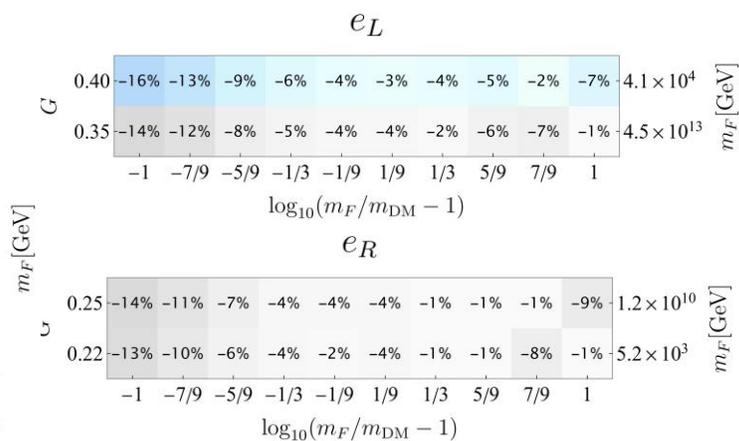
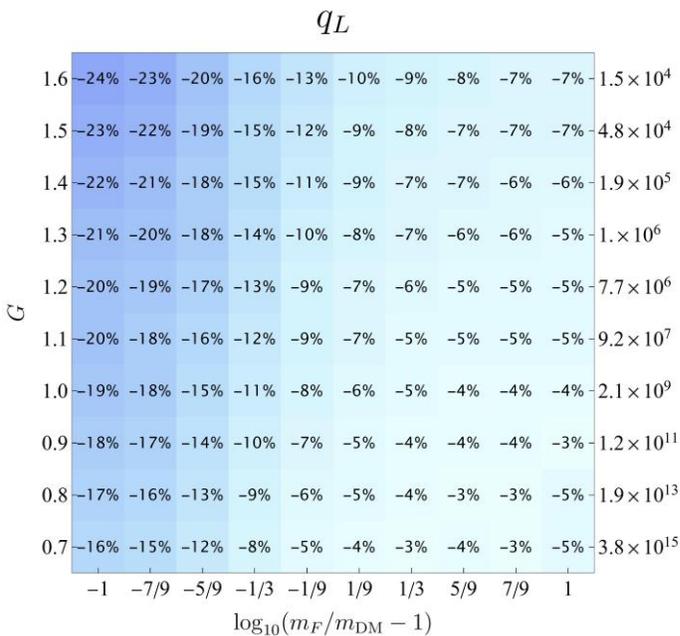
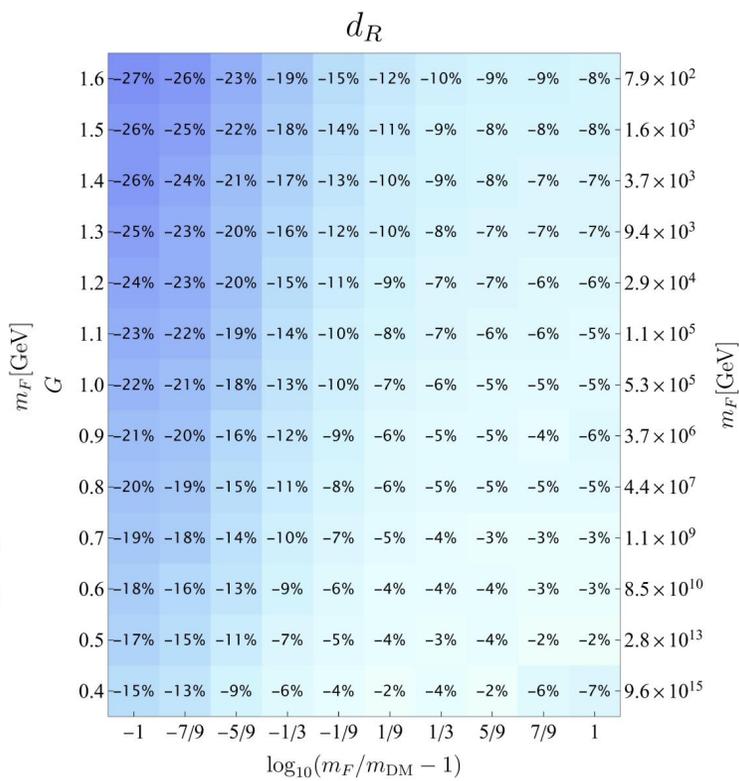
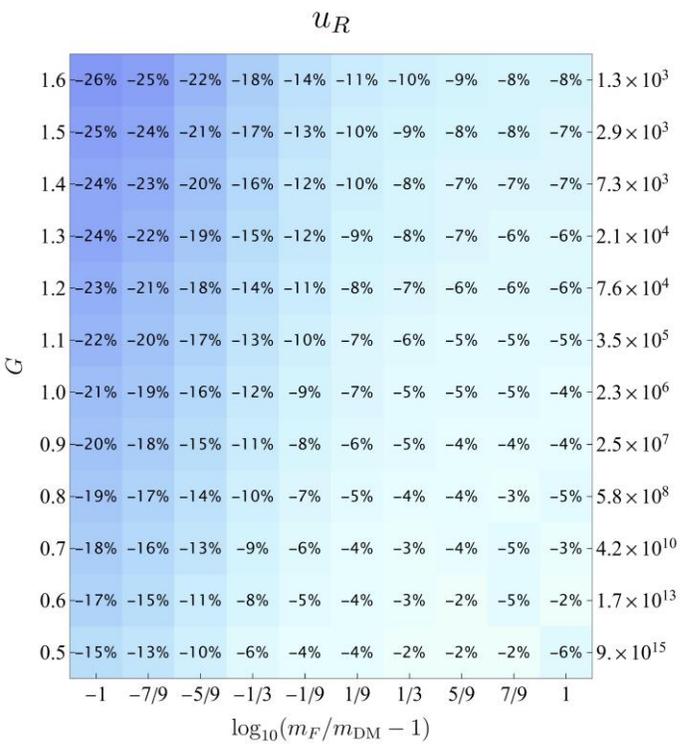
## DM RELIC ABUNDANCE $\Omega h^2$

The amount of DM that is underestimated if one does not take the LPM into account

- Greater  $G$  leads to bigger LPM effect
- Smaller  $\delta$  leads to bigger LPM effect (no decay contribution)
- Five possible realizations of DM mediator

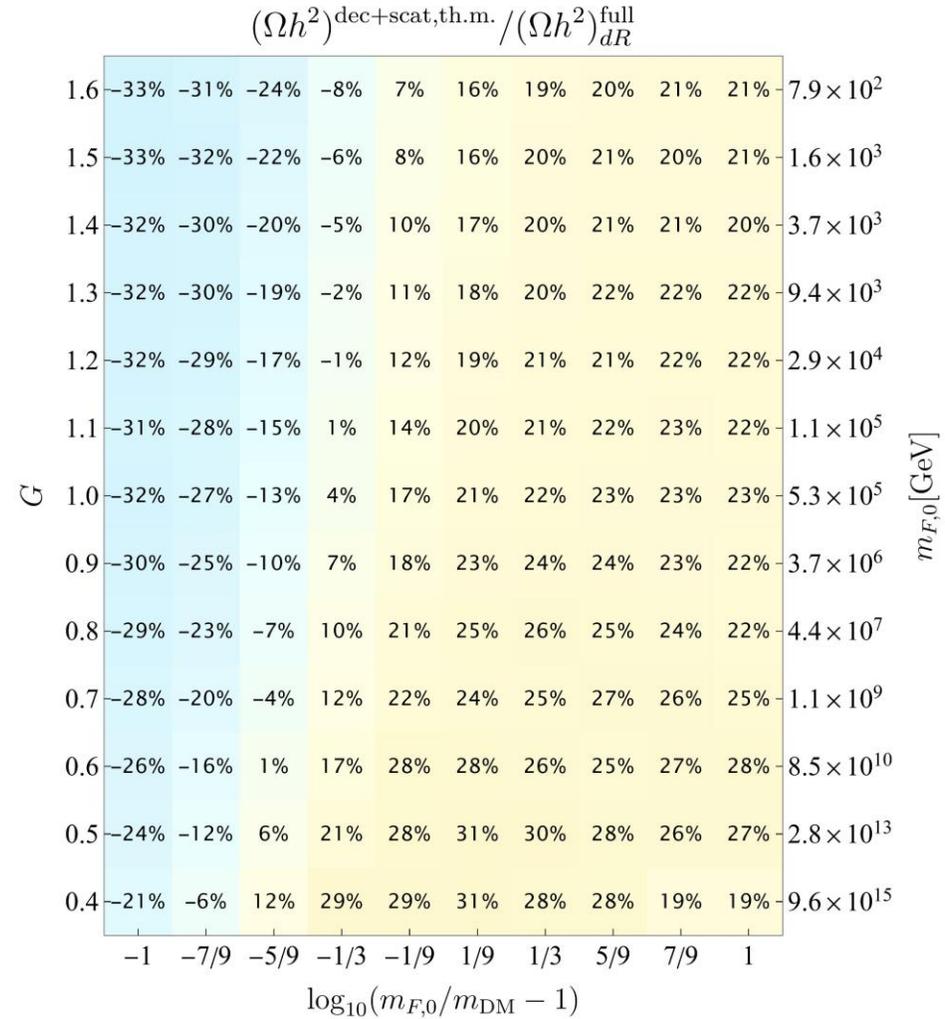
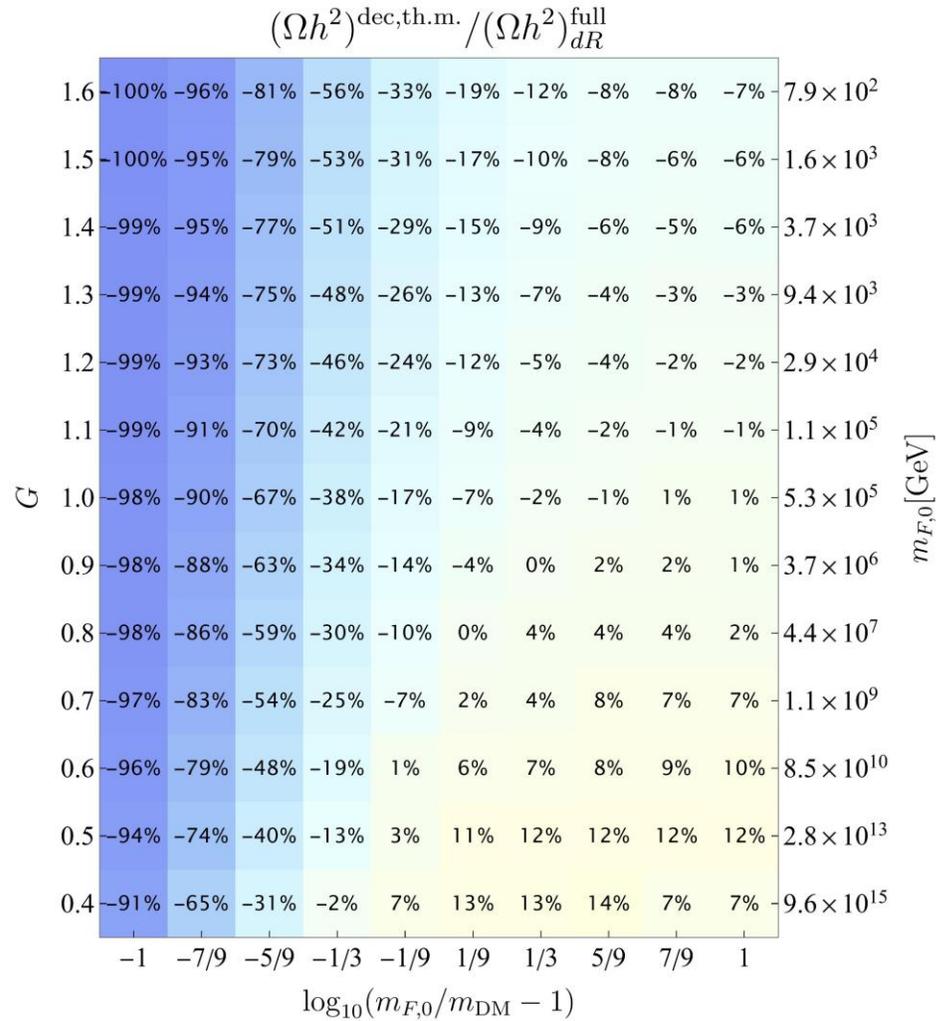
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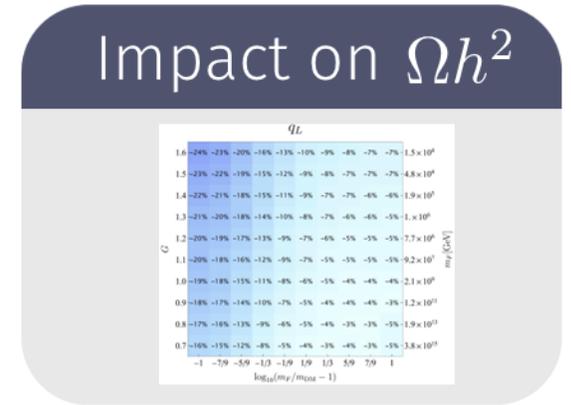
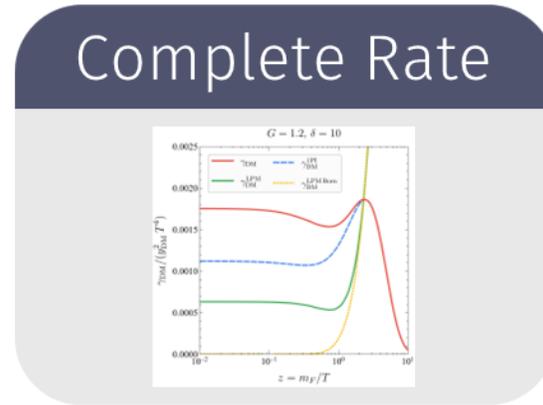
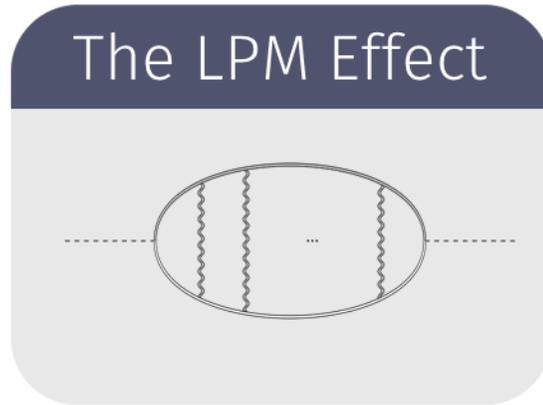
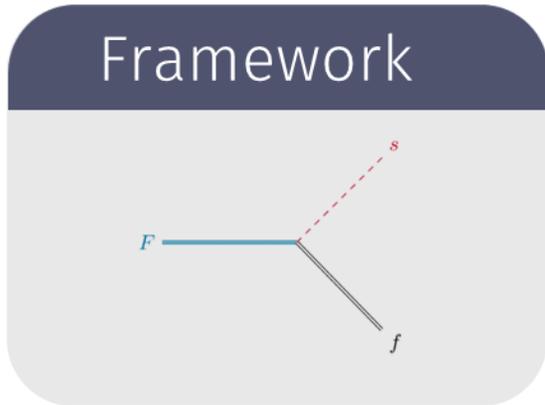
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# COMPARISON TO SIMPLIFIED METHODS





- Multiple soft scatterings give up to a 27% (8%) contribution for mass splittings of  $\delta = 0.1$  ( $\delta = 10$ )
- We calculate for the first time the LPM effect for scalar DM
- We provide the most accurate state-of-art calculation

# BACKUP SLIDES



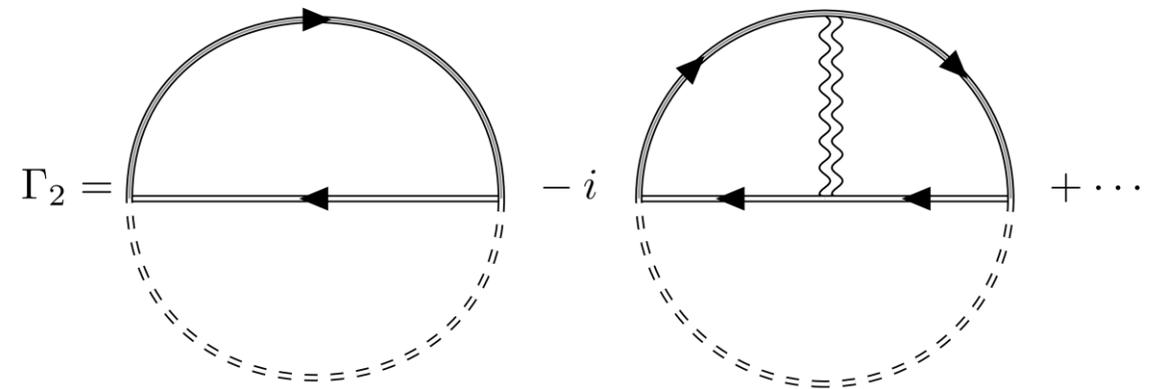
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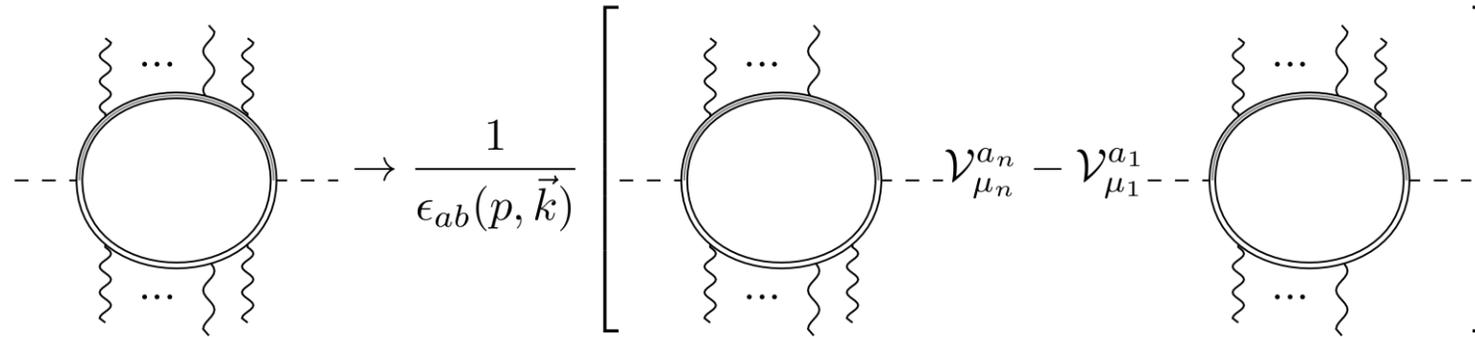
[M. Becker, E. Copello, J. Harz, C. Tamarit (2023)]

The self-energy is given by the functional derivative of the 2PI effective action

$$\Pi^{ab}(x, y) = iab \frac{\delta\Gamma_2[\Delta, S]}{i\delta\Delta^{ba}(y, x)}$$



# STRATEGY OF THE CALCULATION



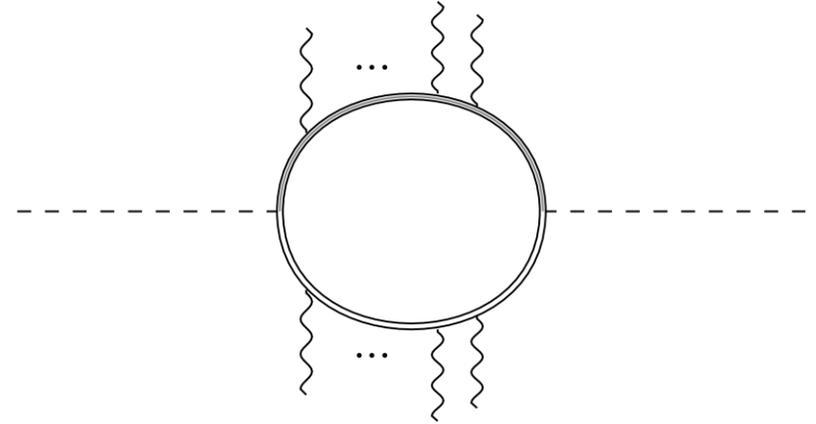
1. In the CTL regime, relate the n-point function to the (n-1)-point function where only one gauge boson was removed
2. Define a current by integrating over all external momenta and contracting with external fields
3. Integrate out the soft gauge boson background, so that external gauge bosons appear as rungs or in the self-energy insertions

*D.Besak (2010)*

$$\hat{\Pi}_\phi(P, k_{||}, \vec{k}_\perp) = \frac{1}{\epsilon_{ab}(P, \vec{k})} \left\{ d(r) \mathcal{C}_\phi(K, P) + ig^2 C_2(r) T \int \frac{d^2 q_\perp}{(2\pi)^2} \mathcal{K}(\vec{q}_\perp) \left[ \hat{\Pi}_\phi(P, k_{||}, \vec{k}_\perp) - \hat{\Pi}_\phi(P, k_{||}, \vec{k}_\perp - \vec{q}_\perp) \right] \right\}$$

External and internal momenta are collinear and in the near-lightcone scale.

$$p \sim T, p^2 \sim g^2 T^2, p \cdot k \sim g^2 T^2$$



We can define the lightlike coordinates through a lightlike vector  $V^\mu = (1, \hat{v})$

$$p_{\parallel}, k_{\parallel} \sim \mathcal{O}(T)$$

$$\vec{p}_{\perp}, \vec{k}_{\perp} \sim \mathcal{O}(gT)$$

Finite temperature analogue of Soft Collinear Effective Theory (SCET)

$$p_{\pm} = p_0 \pm p_{\parallel}$$

$$p^\mu \sim (p_+, p_-, \vec{p}_{\perp})$$

$$p_+ \sim T, \vec{p}_{\perp} \sim gT, p_- \sim g^2 T$$

	$Y$	SU(2)	SU(3)	$G$	$\mu = M_Z$	$10^4 \text{ GeV}$	$10^7 \text{ GeV}$	$10^{10} \text{ GeV}$
$e_L$	$-1/2$	2	1	$\frac{g_1^2}{4} + \frac{3g_2^2}{4}$	0.38	0.4	0.46	0.52
$q_L$	$+1/6$	2	3	$\frac{g_1^2}{36} + \frac{3g_2^2}{4} + \frac{4g_3^2}{3}$	2.3	1.6	1.2	1.0
$e_R$	$-1$	1	1	$g_1^2$	0.21	0.22	0.24	0.26
$u_R$	$+2/3$	1	3	$\frac{4g_1^2}{9} + \frac{4g_3^2}{3}$	2.1	1.3	0.9	0.7
$d_R$	$-1/3$	1	3	$\frac{g_1^2}{9} + \frac{4g_3^2}{3}$	2.0	1.2	0.8	0.6

# THE LPM EFFECT

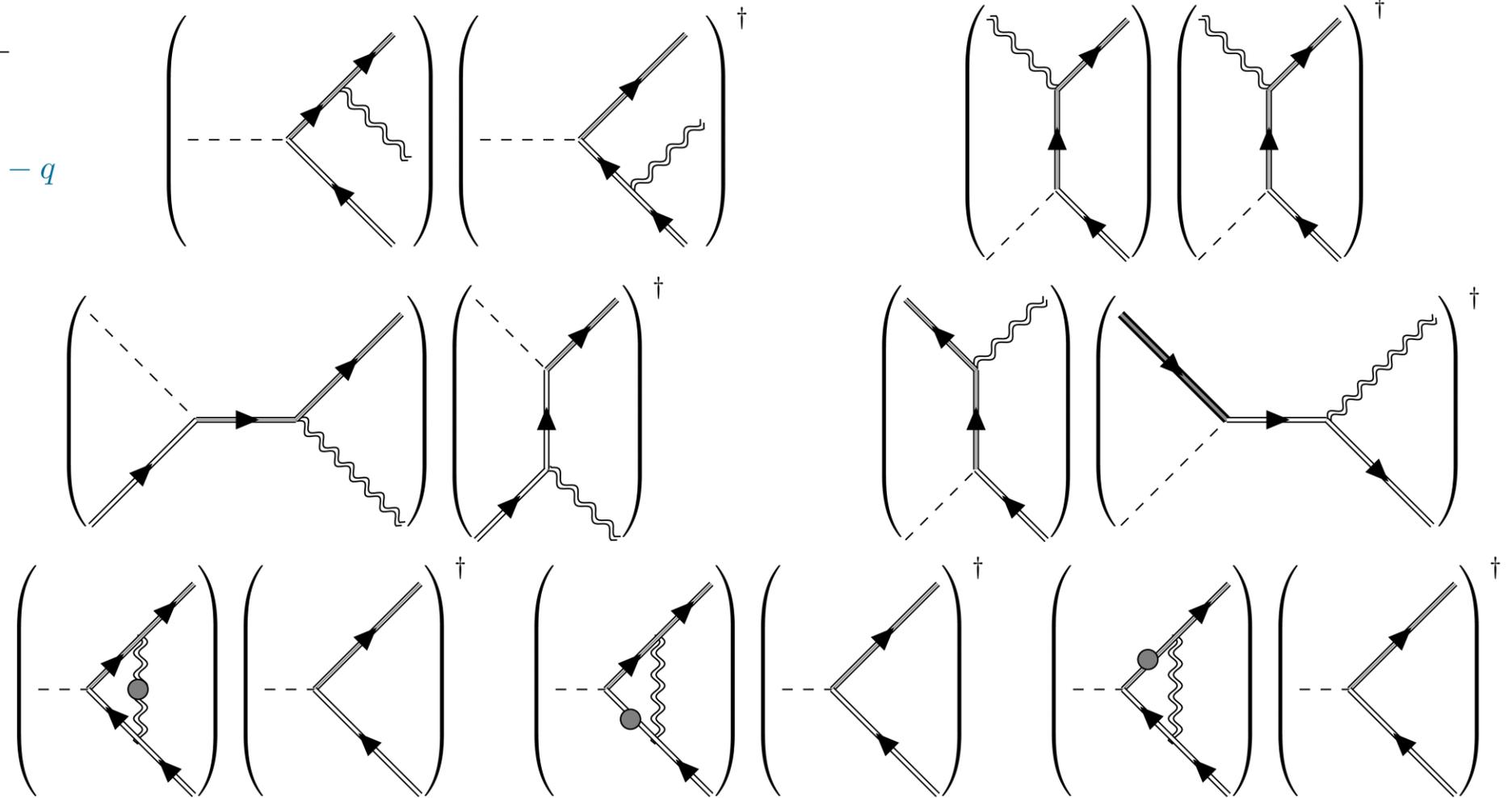
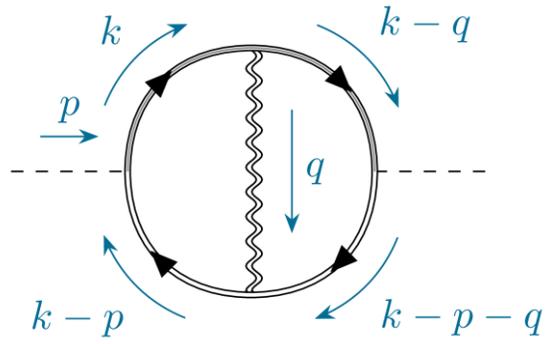
If all masses are thermal masses,

$$k^2 = m_F^2 = (p + q)^2 = m_{\text{DM}}^2 + m_f^2 + 2pq$$

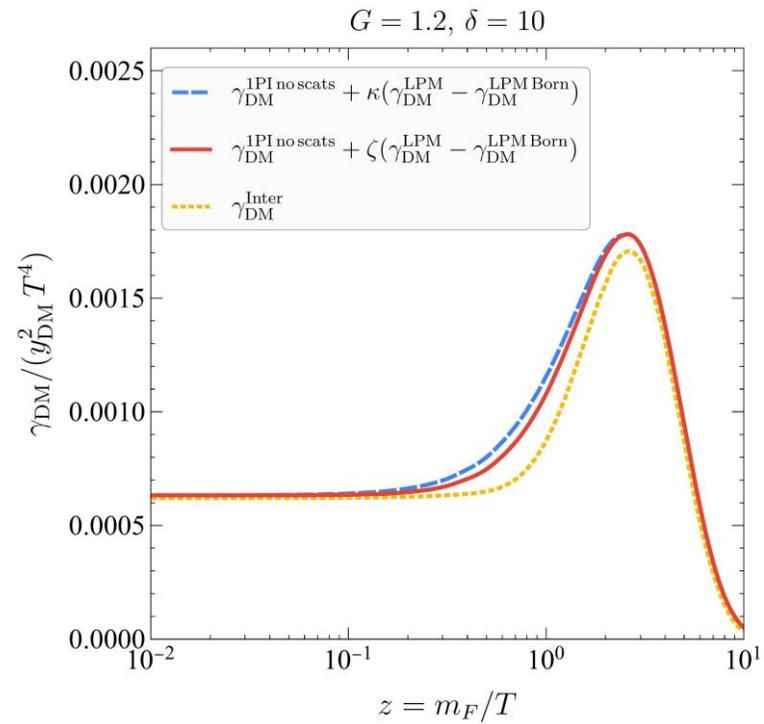
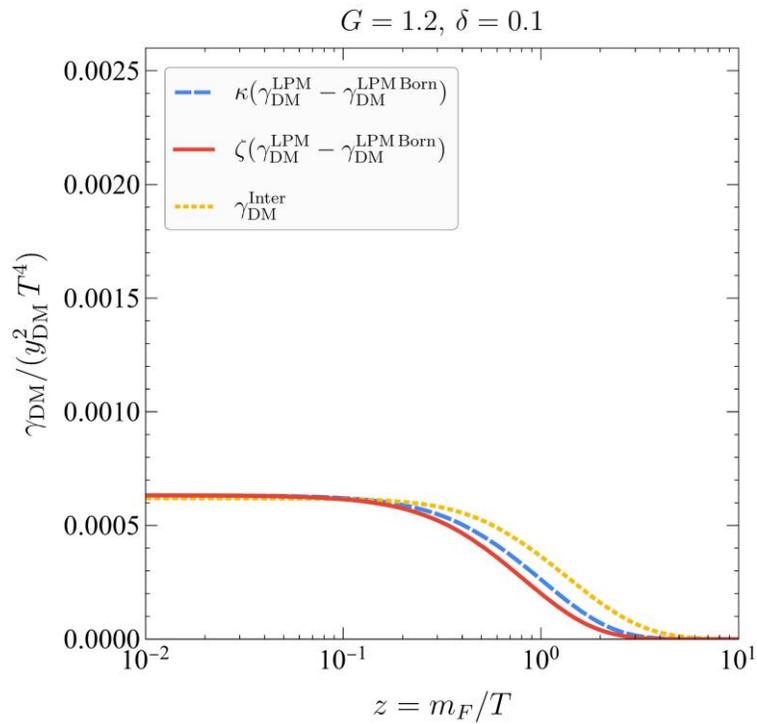
then the emission occurs collinearly  $\theta \sim g$

$$pq = |\vec{p}||\vec{q}|(1 - \cos \theta) \sim g^2 T^2$$

The mean free time between collisions is of the same order as particle formation. It is impossible to treat emission and scatterings as independent processes.



# PRESCRIPTIONS. RATE WITHOUT SCATTERINGS



# PRESCRIPTIONS. RATE WITHOUT LPM.

