

28.05.2025

ALP production from Abelian Gauge interactions: Beyond the HTL.



JOHANNES GUTENBERG
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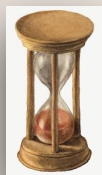
*Cristina Puchades
Ibáñez*

Rates in the early universe

◆ DM , Baryon asymmetry, thermal GW...

◆ ! Massless particles: IR divergences

$p \sim gT$: perturbative approach not reliable.



✗ Unphysical rates for $p < gT$ (soft ALP)

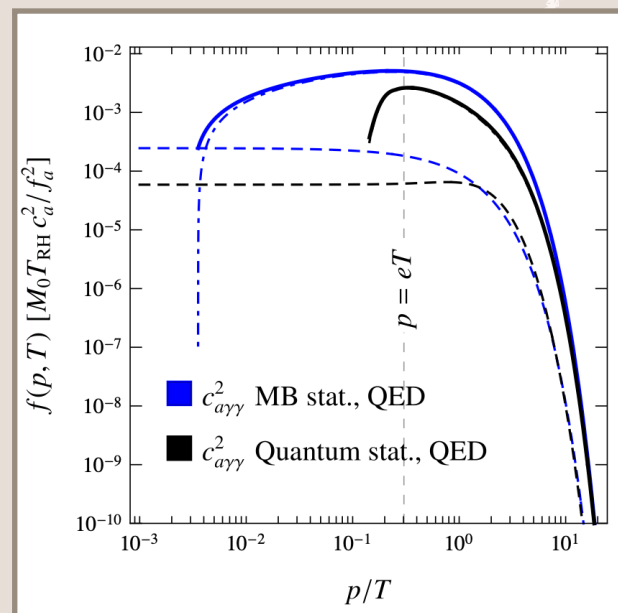


S.Baumholzer, V.Brdarand
E. Morgante

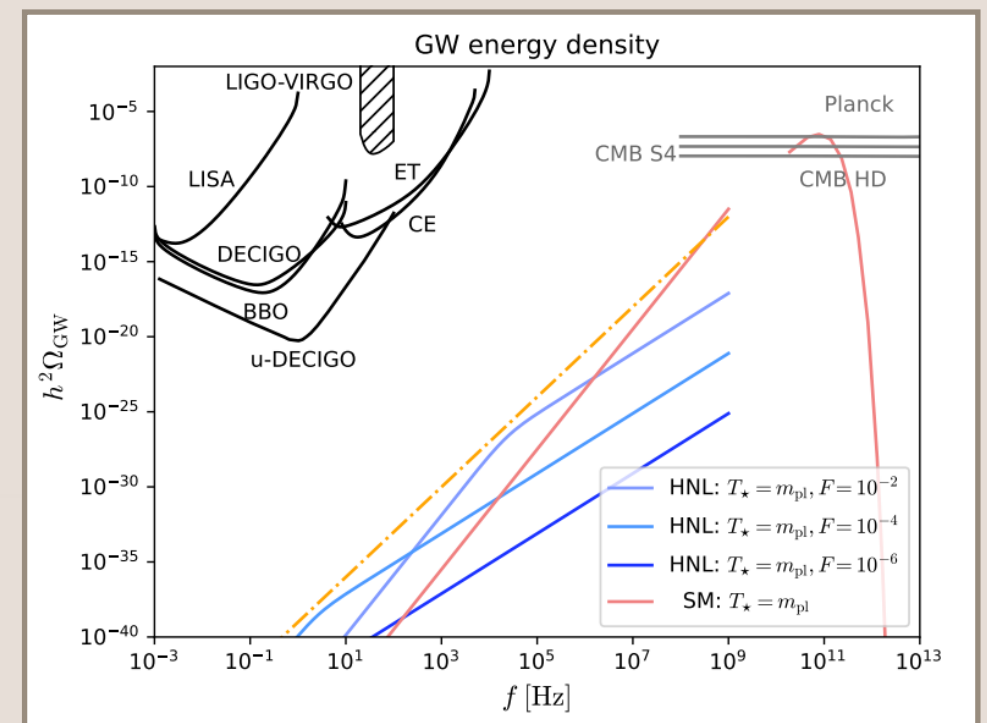
See also



E.Braaten and M.H. Thoma



◆ Thermal GW spectrum



M.Drewes, Y.Georis, J. Klarick and P.Klose

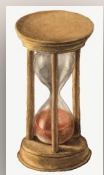
Rates in the early universe

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+

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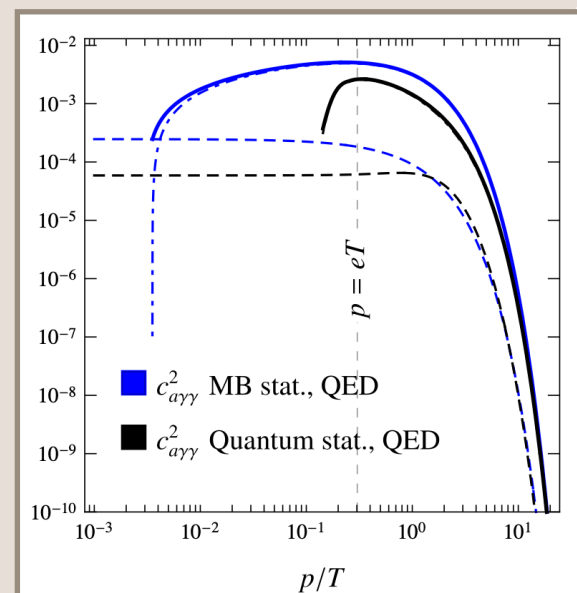


✗ Unphysical rates for $p < gT$

📖 S.Baumholzer, V.Brdarand
E. Morgante

See also

📖 E.Braaten and M.H. Thoma



Today:

Full form of 1PI-resummed TFT propagator.

📖 arXiv:**2502.01729**: Becker, Harz,
Morgante, CPI, Schwaller.

Different approach and non-abelian

📖 K.Bouzoud and J.Ghiglieri

Model: “photophilic” axion

$$\mathcal{L} = \frac{1}{2} \partial_\mu a \partial^\mu a + \frac{1}{2} m_a^2 a^2 - \frac{1}{4} B_{\mu\nu} B^{\mu\nu} + \frac{c_{aBB}}{4f_a} a B_{\mu\nu} \tilde{B}^{\mu\nu}$$

◆ $B_{\mu\nu}$: field strength of the abelian gauge field.

➡ “photon”: $U(1)_Y$ ➡ DM relic abundance

Applied as well to SM photon or any other $U(1)$ dark field

We need:

$$\left[\frac{\partial}{\partial t} - H p \frac{\partial}{\partial p} \right] f(p, t) = \mathcal{C}(p)$$

$$\mathcal{C}(p) = \frac{\Pi^<(P)}{2p_0} = \frac{\epsilon(p_0)}{p_0(1 - e^{p_0/T})} \text{Im } \Pi(P)$$

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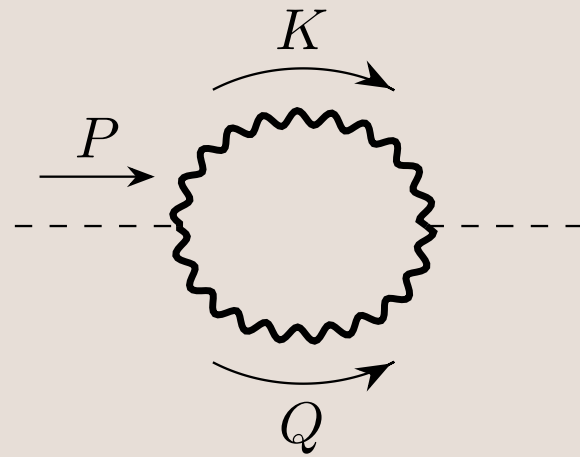
Integrate over
 $z = m_a/T$

We need:

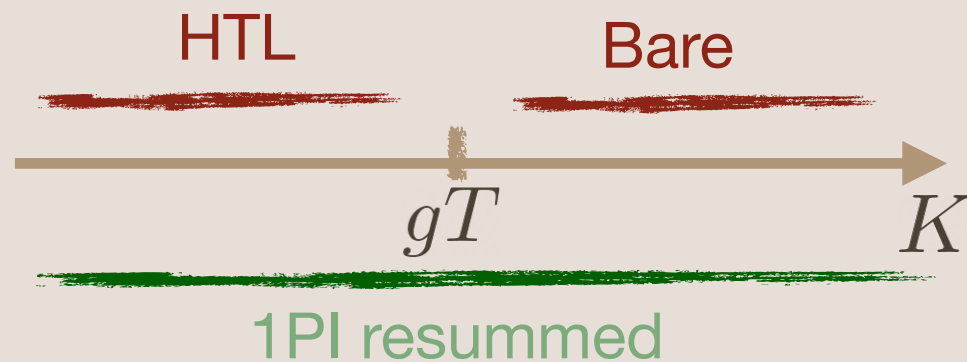
$$\left[\frac{\partial}{\partial t} - H p \frac{\partial}{\partial p} \right] f(p, t) = \mathcal{C}(p) \quad \longrightarrow \quad \frac{\partial f(p/T, z)}{\partial z} = \frac{1}{2p_0 z H} \Pi^<(P, z)$$

$$\mathcal{C}(p) = \frac{\Pi^<(P)}{2p_0} = \frac{\epsilon(p_0)}{p_0(1 - e^{p_0/T})} \text{Im } \Pi(P)$$

Thermal axion production.



◆ Our implementation:



$$\text{wavy line} = \text{bare wavy line} + \text{one-loop bubble} + \text{two-loop bubble} + \dots$$

$$\diamond {}^*D_{\mu\nu}^<(K) = f_B(k_0) \left[\mathcal{P}_{\mu\nu}^t \rho_t(K) + \mathcal{P}_{\mu\nu}^l \frac{k^2}{K^2} \rho_l(K) + \xi \frac{k_\mu k_\nu}{K^2} \right]$$

Spectral densities

! We resummed **both** gauge bosons

- ◆ The gauge dependent parallel polarization cancels
- ◆ Our spectral densities are gauge independent

➡ gauge independent

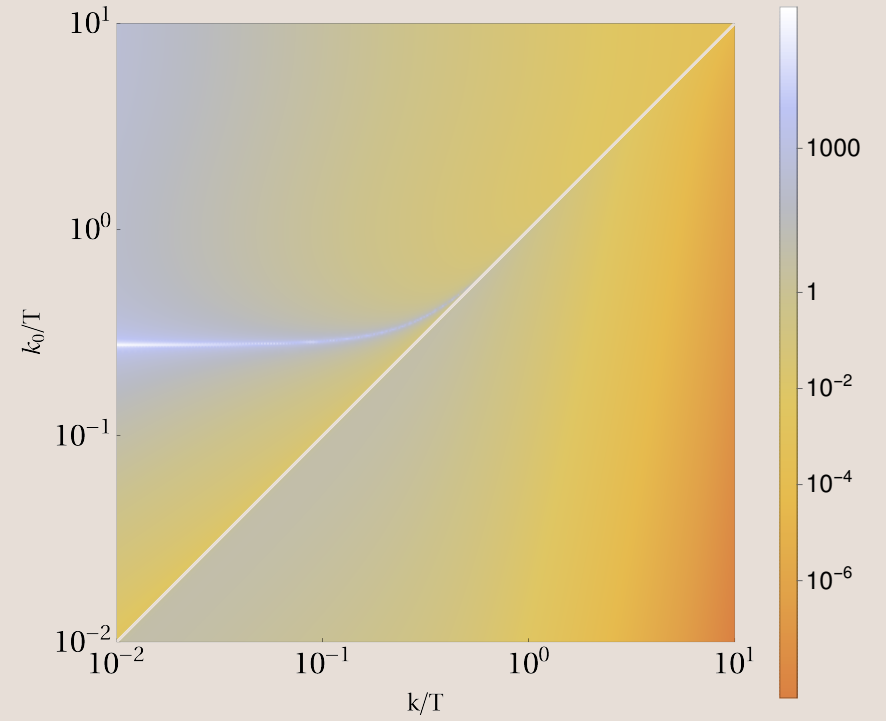
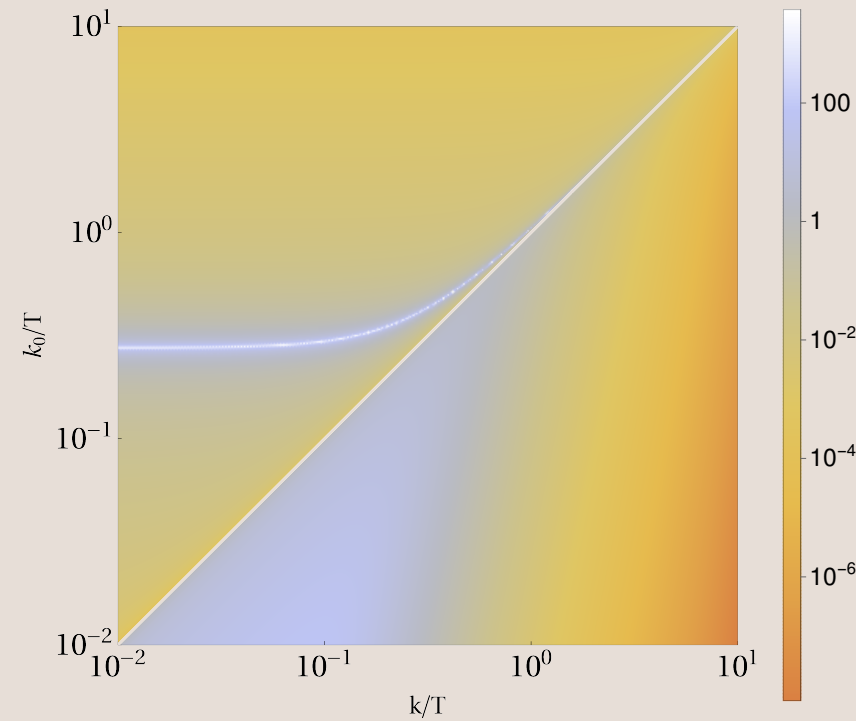
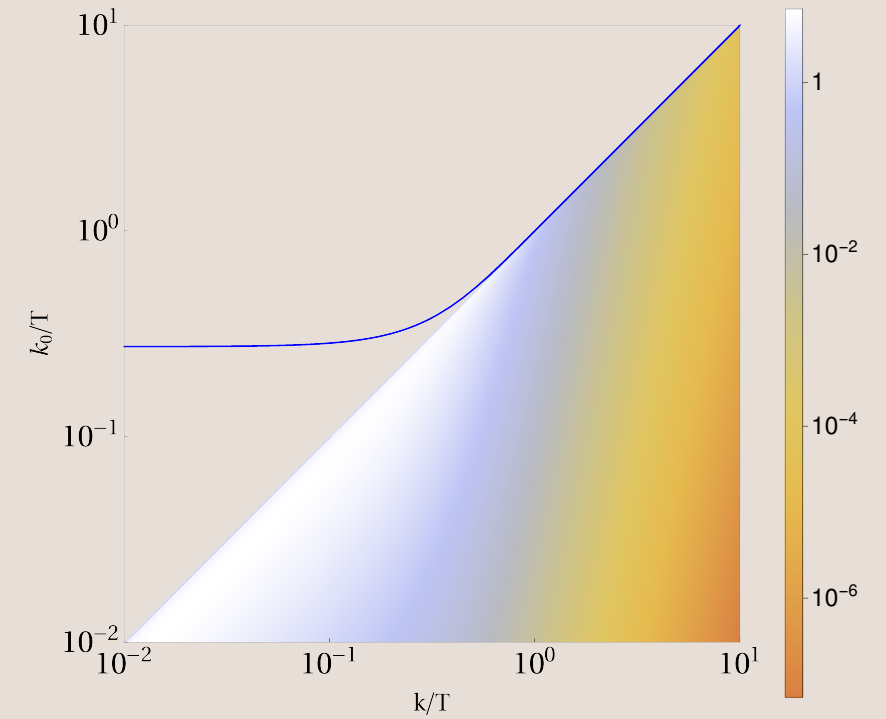
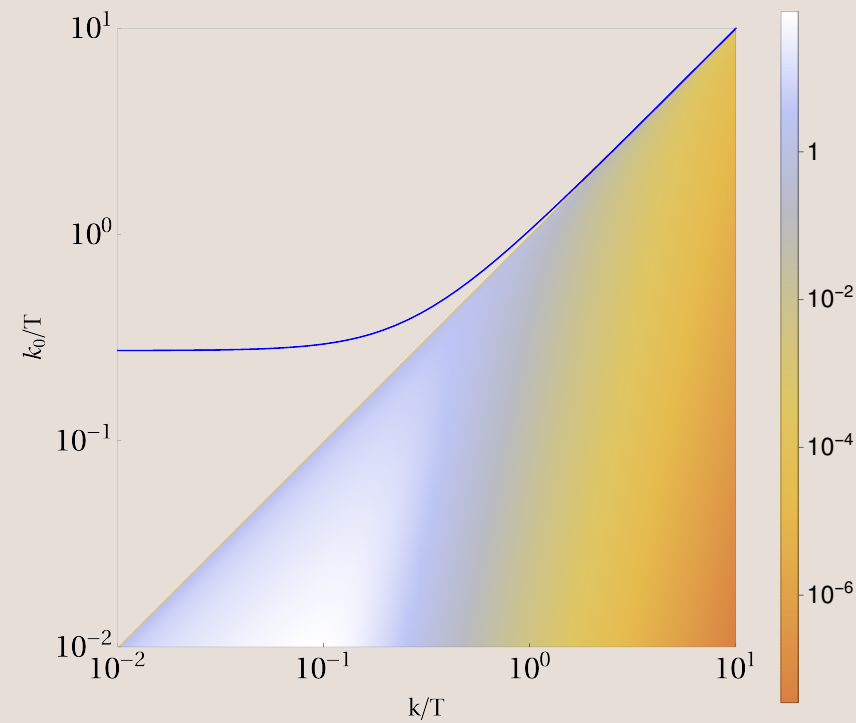
Different approach and non-abelian



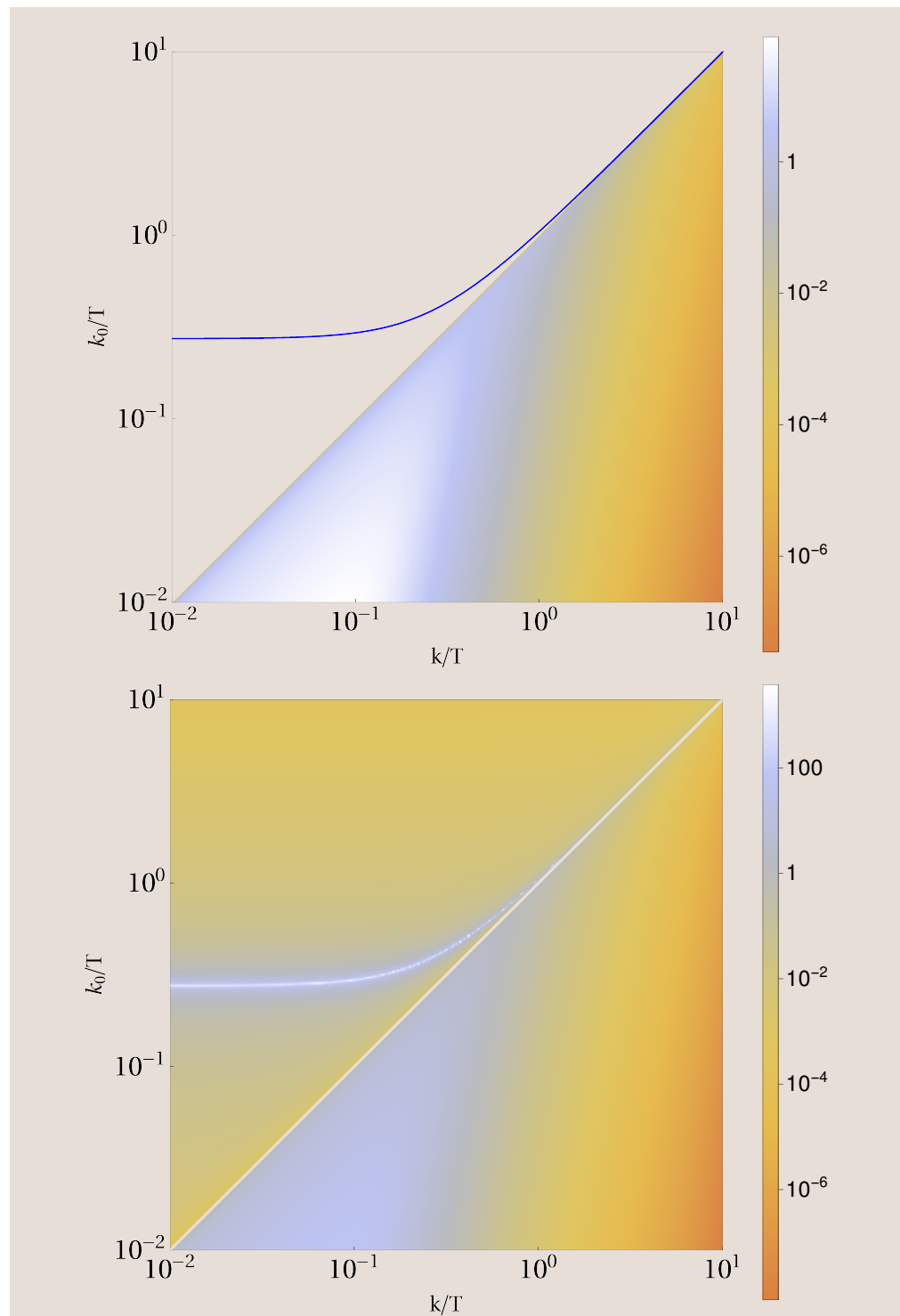
K.Bouzoud and J.Ghiglieri

$$\rho_t = -2 \frac{\text{Im } \pi_0 + \text{Im } \pi_t}{(K^2 - \text{Re } \pi_0 - \text{Re } \pi_t)^2 + (\text{Im } \pi_0 + \text{Im } \pi_t)^2}$$

$$\rho_l = -2 \frac{K^2}{k^2} \frac{\text{Im } \pi_0 + \text{Im } \pi_l}{(K^2 - \text{Re } \pi_0 - \text{Re } \pi_l)^2 + (\text{Im } \pi_0 + \text{Im } \pi_l)^2}$$



$$\rho_t = -2 \frac{\text{Im } \pi_0 + \text{Im } \pi_t}{(K^2 - \text{Re } \pi_0 - \text{Re } \pi_t)^2 + (\text{Im } \pi_0 + \text{Im } \pi_t)^2}$$



■ $\Pi_{TT}^{\leq}(p)$:

(Inverse) Decays for **HTL**

includes $2 \leftrightarrow 3$ scatterings for **1PI**

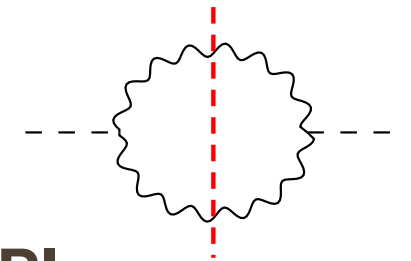
■ $\Pi_{TS}^{\leq}(p)$:

$2 \leftrightarrow 2$ scatterings

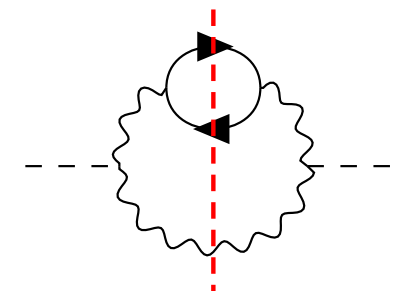
■ $\Pi_{SS}^{\leq}(p)$:

$2 \leftrightarrow 3$ scatterings

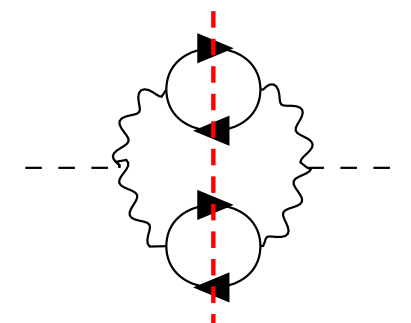
1 loop, included



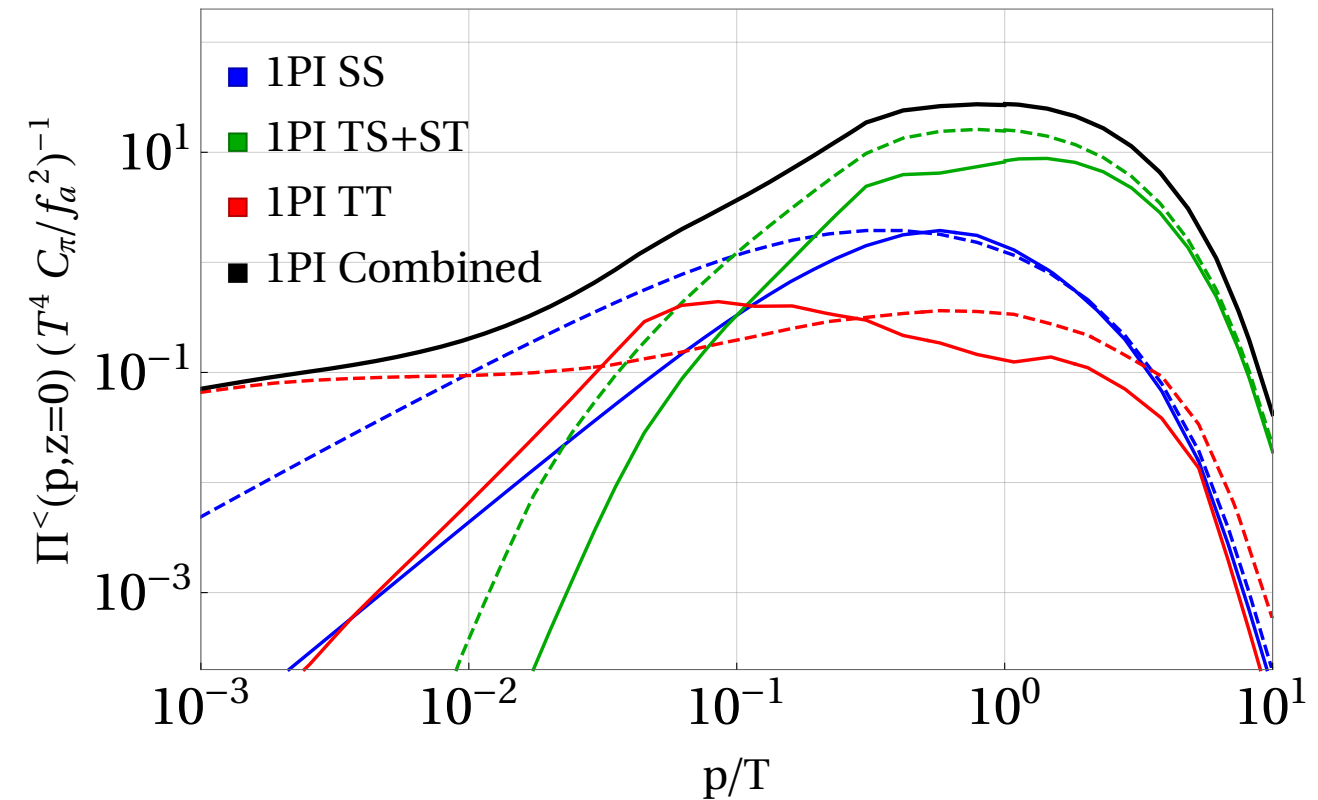
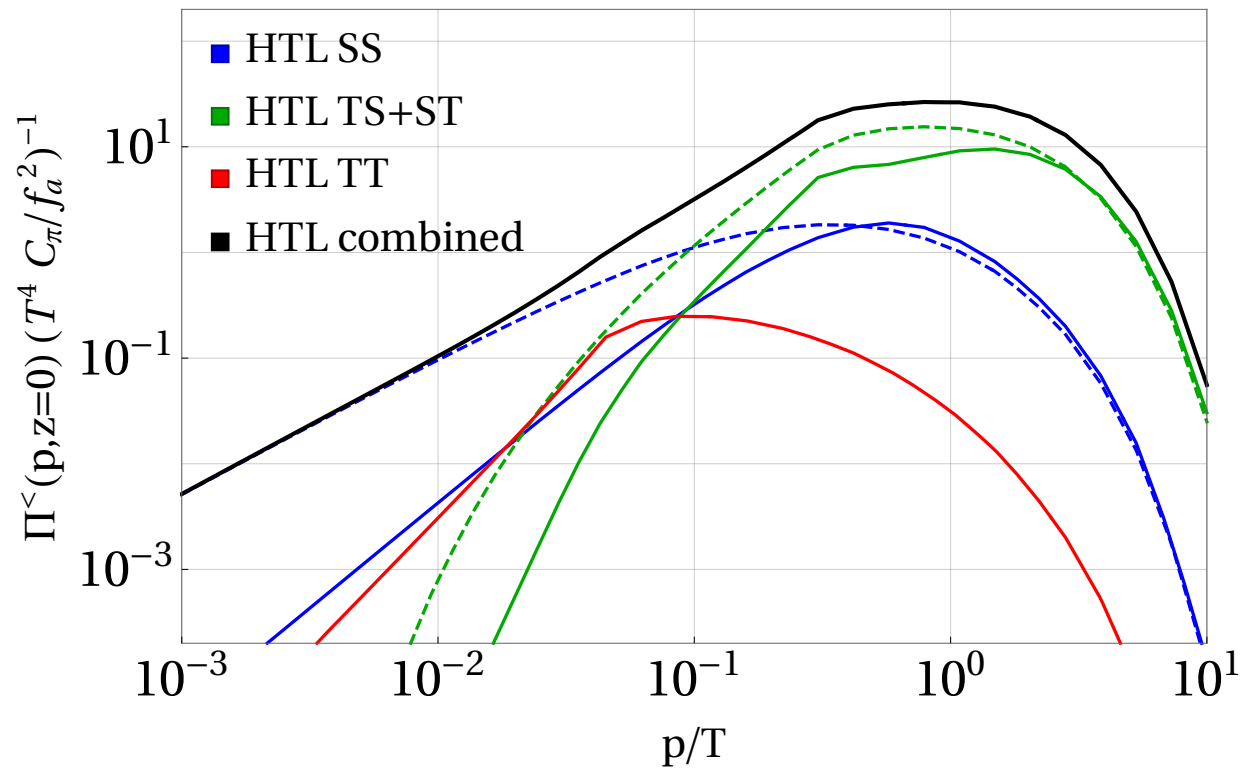
2 loops, included



3 loops, included



Results: Production rate.



■ $\Pi_{SS}^<(p) : 2 \leftrightarrow 3$ scatterings

■ $\Pi_{TS}^<(p) : 2 \leftrightarrow 2$ scatterings

■ $\Pi_{TT}^<(p) : (\text{Inverse}) \text{ Decays}$
includes $2 \leftrightarrow 3$ scatterings for **1PI**

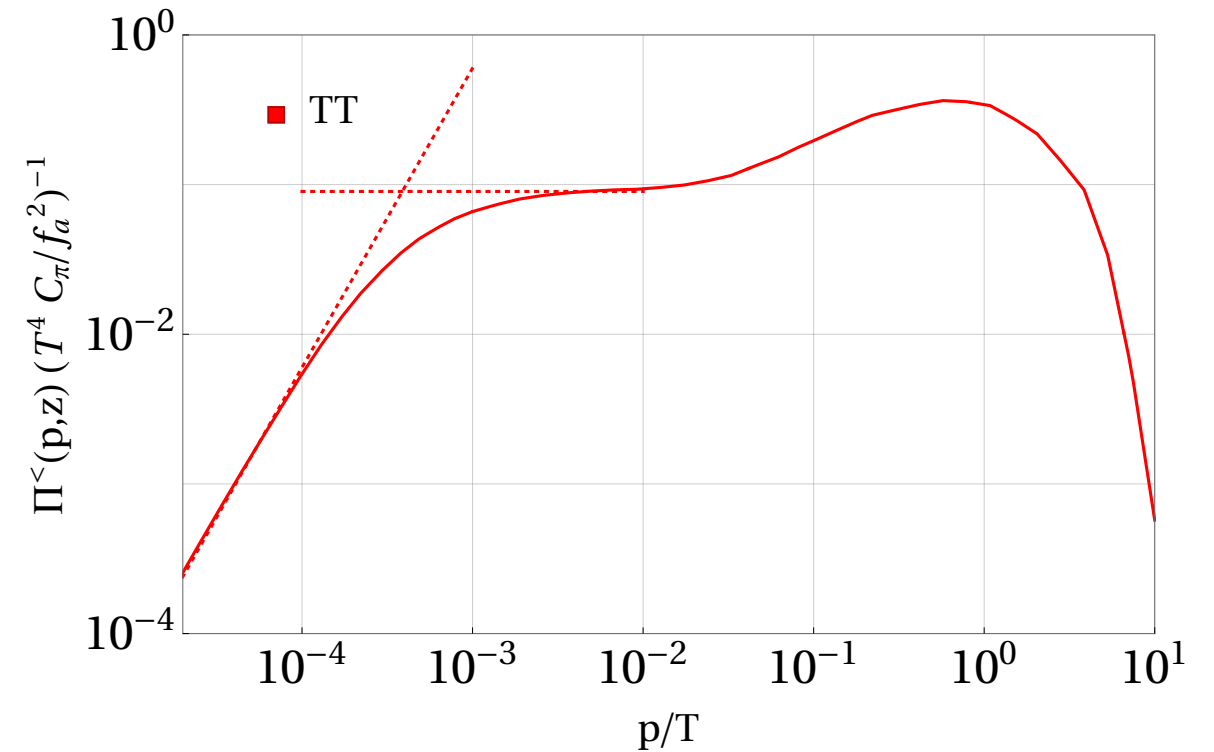
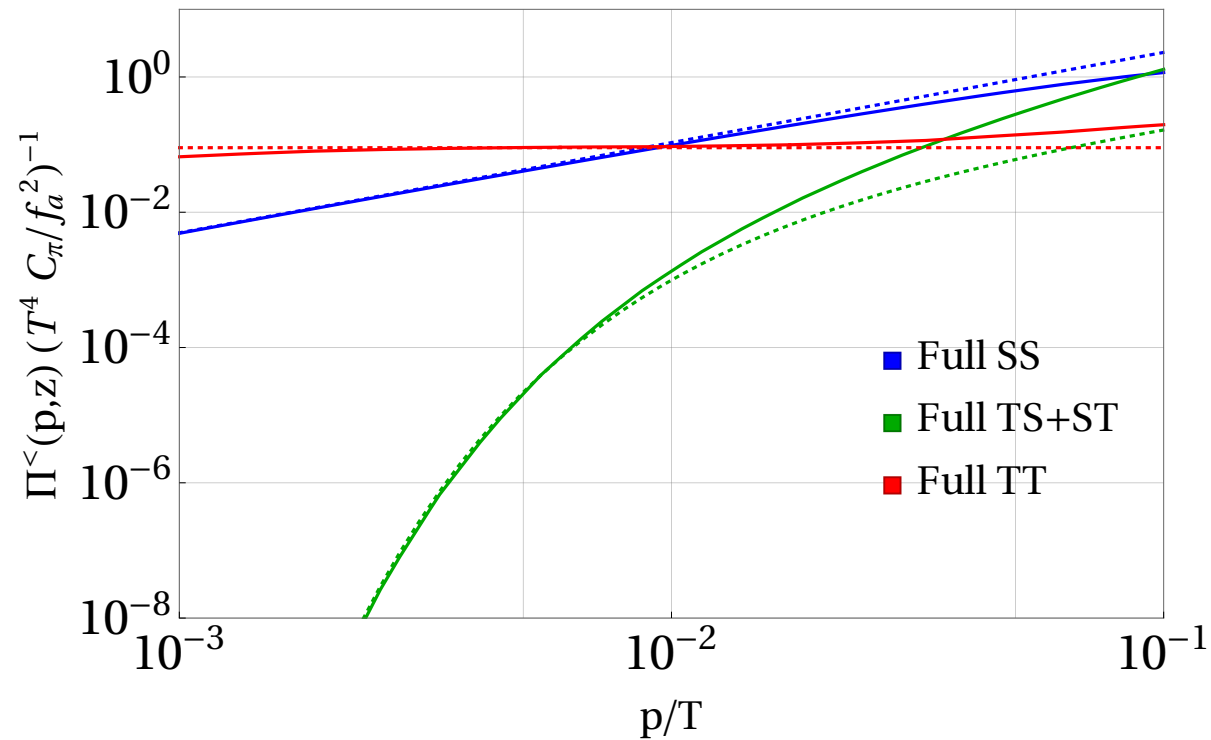
◆ Production dominated by high temperatures around T_{rh}

➡ $z = m_a/T = 0$

————— Longitudinal-Transversal

..... Transversal-Transversal

Results: Production rate.

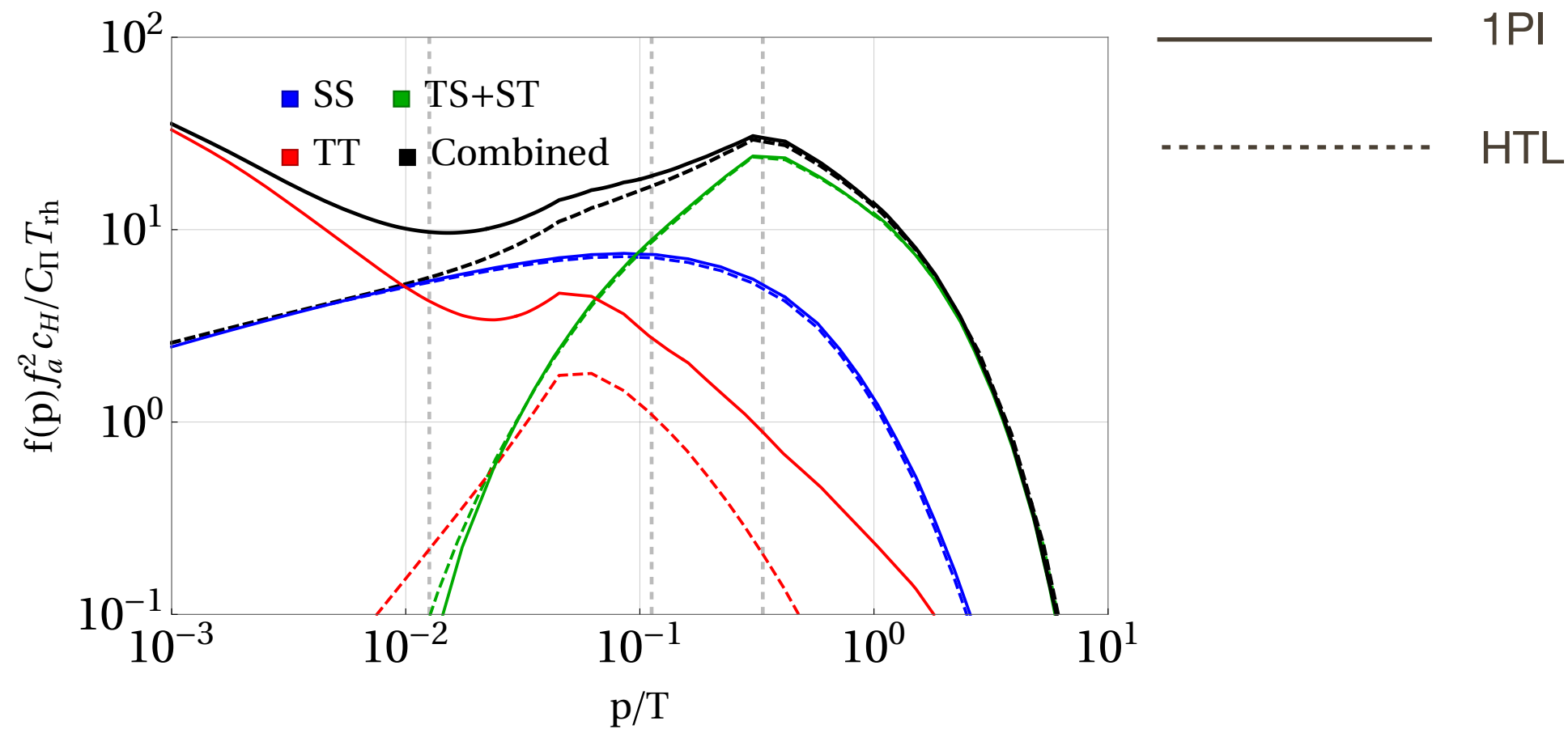


Small p scaling

■ $\Pi_{SS}^{\leq}(p) \propto p^{\frac{4}{3}} \quad 2 \leftrightarrow 3$
■ $\Pi_{TS}^{\leq}(p) \propto p \exp\left(-\frac{m_V^2}{4p}\right) \quad 2 \leftrightarrow 2$
■ $\Pi_{TT}^{\leq}(p) \propto \begin{cases} p^0 & p \gtrsim p_c \\ p^2 & p \lesssim p_c \end{cases} \quad 2 \leftrightarrow 3$
 $p_c \approx 5 \times 10^{-4}$

- ◆ At soft momentum ($p < gT$) $2 \leftrightarrow 2$ scatterings suffer a kinematic suppression.
- ◆ 1PI captures time-like gauge boson effects

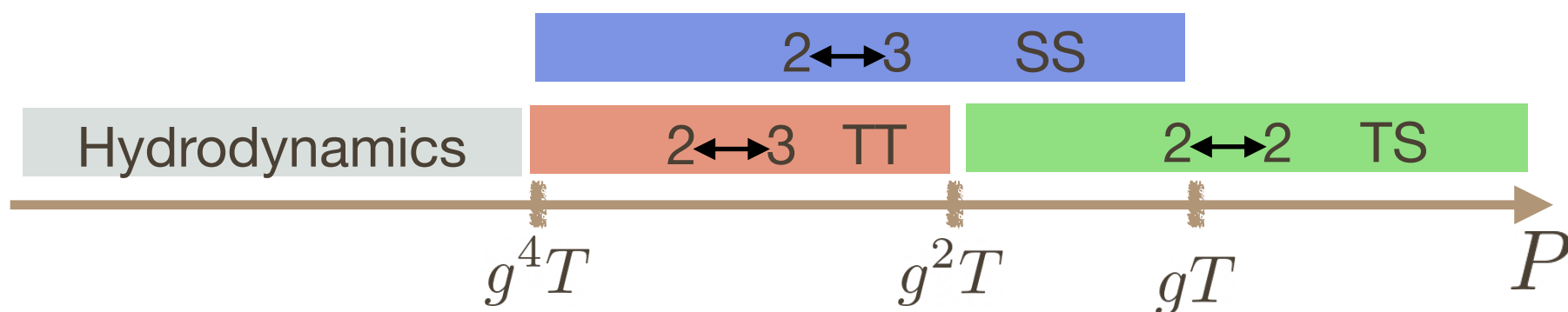
Results: Distribution function.



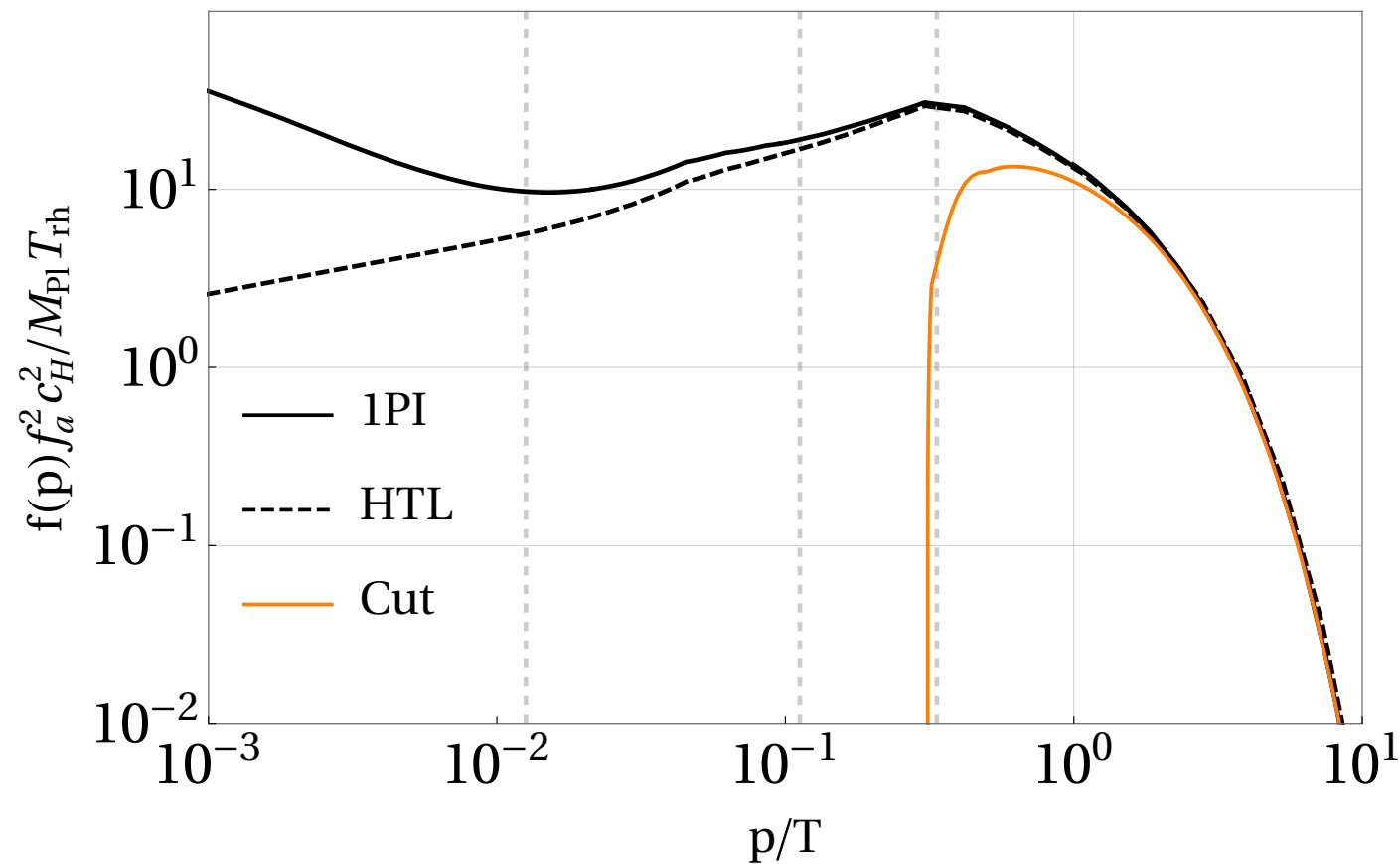
Scales

$P \sim T$ “Hard”

$P \sim gT$ “Soft”



Results: Comparision.



◆ Mild impact on Lyman- α constraints.

$$\lambda_{\text{fs}} \propto \frac{\langle p \rangle / T}{m_a}$$

$$\langle p \rangle = \frac{\int dp p^3 f_a(p)}{\int dp p^2 f_a(p)}$$



S.Baumholzer, V.Brdar, and E. Morgante

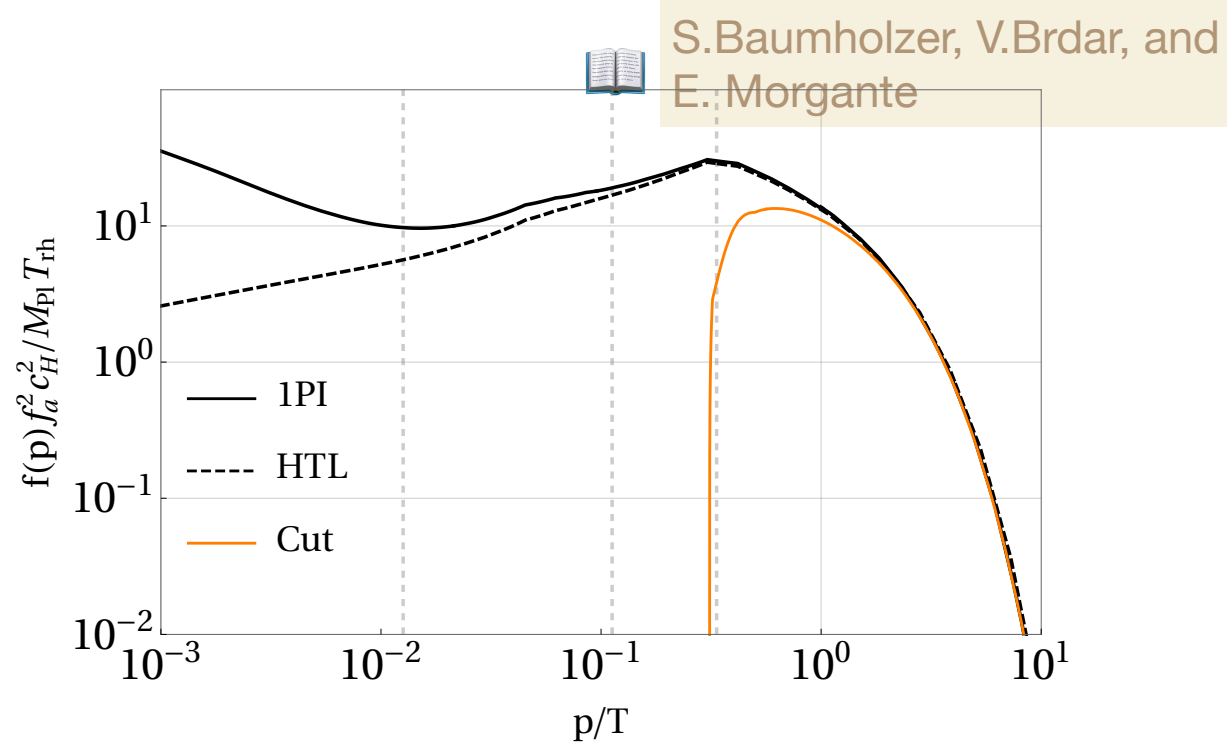
$$\langle p \rangle_{1PI} = 3.06$$

$$\langle p \rangle_{HTL} = 3.17$$

$$\langle p \rangle_{Cut} = 3.14$$

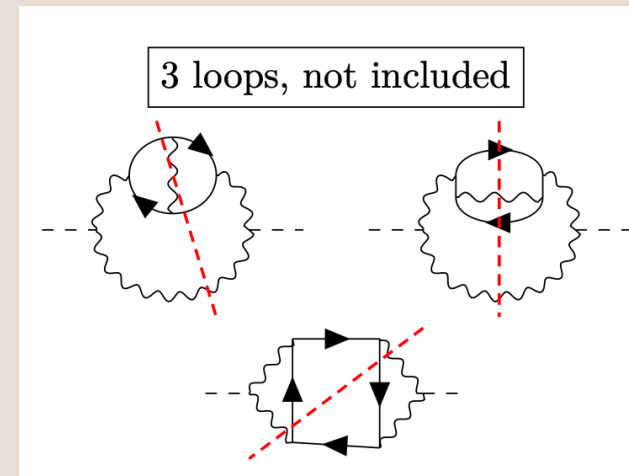
$$\langle p \rangle_{BE} = 2.70$$

However.....

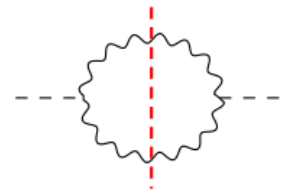


However we still don't have the LO calculation....

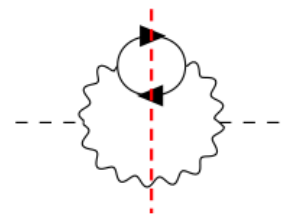
- ◆ We include soft ALP radiation from virtual photons
- ◆ We do not include soft ALP radiation from real photons yet → Goes beyond 1-loop 1PI resummation



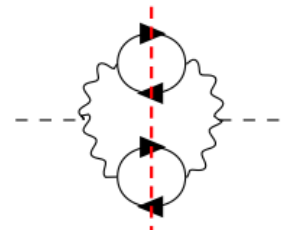
1 loop, included



2 loops, included

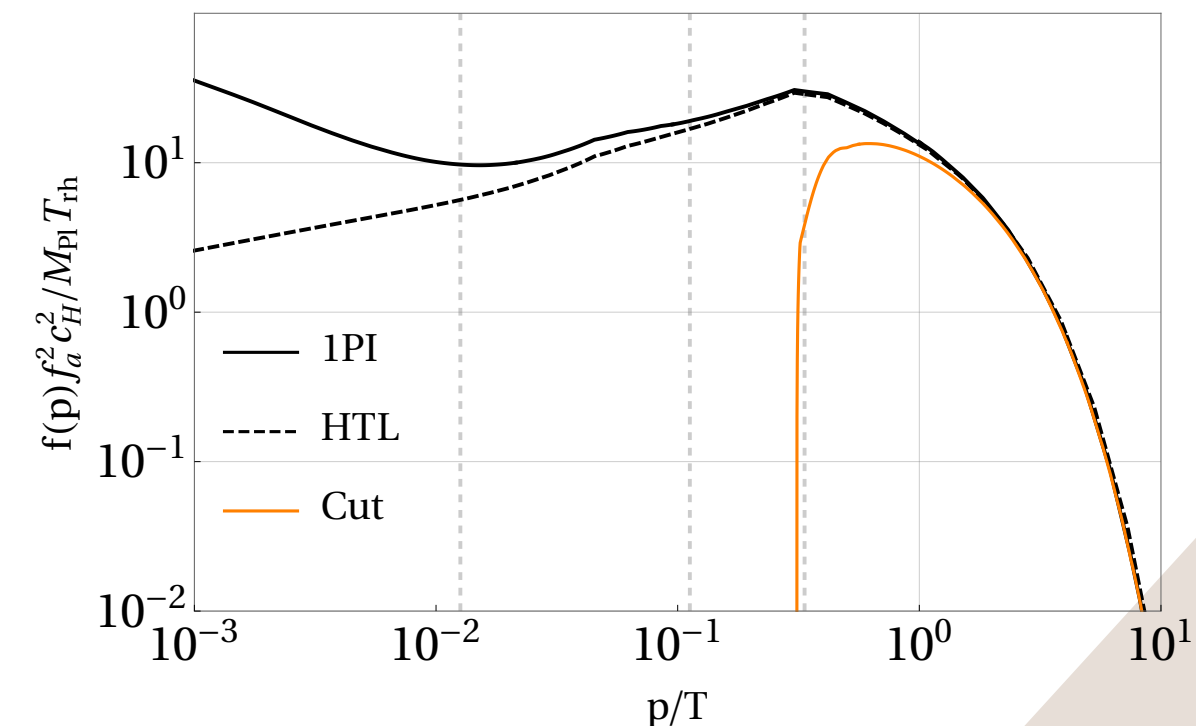


3 loops, included



Conclusions.

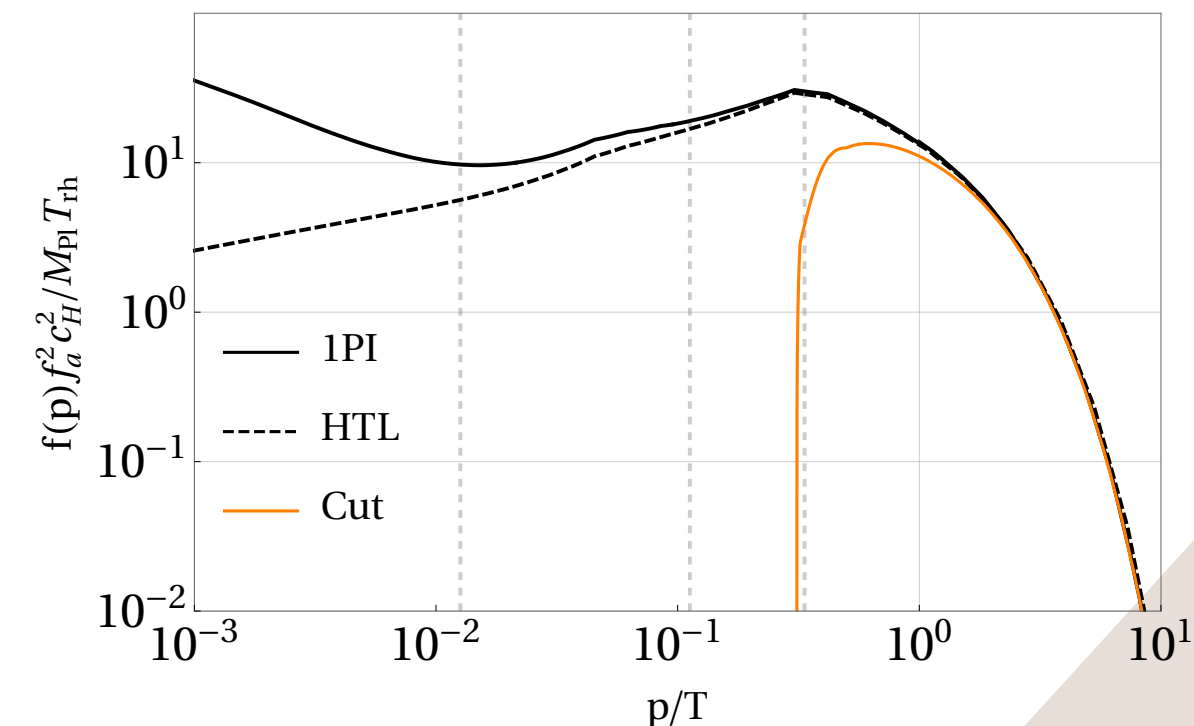
- A gap exist for $g^4 T \leq P \leq gT$
- Our HTL and 1PI calculations resolve negative rates $\longrightarrow$ Below gT
- 1PI captures contributions from 2 time-like gauge bosons.
- $2 \leftrightarrow 3$ scatterings dominate for soft ALPs but we need to keep working for full accuracy at LO for $P < g^2 T$
- Mild impact on Lyman- α constraints.



Conclusions.

- A gap exist for $g^4 T \leq P \leq gT$
- Our HTL and Full calculations resolve negative rates $\longrightarrow$ Below gT
- 1PI captures contributions from 2 time-like gauge bosons.
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Thank
you!



Thermal axion production.

$$\Pi^<(P) = \frac{C_\Pi}{f_a^2 p} \int \frac{d^4 K}{2\pi^4} \epsilon^{\mu\nu\alpha\beta} \epsilon^{\mu'\nu'\alpha'\beta'} K_\alpha Q_\beta K_{\alpha'} Q_{\beta'} {}^* D_{\mu\mu'}^<(K) {}^* D_{\nu\nu'}^<(Q)$$



$$\diamond {}^* D_{\mu\nu}^<(K) = f_B(k_0) \left[\mathcal{P}_{\mu\nu}^t \rho_t(K) + \mathcal{P}_{\mu\nu}^l \frac{k^2}{K^2} \rho_l(K) + \xi \frac{k_\mu k_\nu}{K^2} \right]$$

$$\diamond C_\Pi = \frac{c_{aBB} d_i \alpha_i^2}{8(2\pi)^5} = \frac{c_{aBB} \alpha_Y^2}{8(2\pi)^5}$$

$$\diamond f_B(k_0) = \frac{1}{e^{k_0/T} - 1}$$

$$\Pi^<(P) = \frac{C_\Pi}{f_a^2 p} \int_{-\infty}^{\infty} dk^0 \int_0^{\infty} dk \int_{|p-k|}^{|p+k|} dq k q f_B(k^0) f_B(p^0 - k^0) [\mathcal{I}_{lt} + \mathcal{I}_{tt}]$$

$$\diamond \mathcal{I}_{lt} = (\rho_t(K) \rho_l(Q) + \rho_l(K) \rho_t(Q)) [(k+q)^2 - p^2] [p^2 - (k-q)^2]$$

$$\diamond \mathcal{I}_{tt} = \rho_t(K) \rho_t(Q) \left[\left(\frac{k_0}{k} \right)^2 + \left(\frac{q_0}{q} \right)^2 ((k^2 - p^2 + q^2)^2 + 4k^2 q^2) + 8k^0 q^0 (k^2 + q^2 - p^2) \right]$$

