Schwinger current in de Sitter Space

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António Torres Manso

Ongoing work with M. Bastero Gil, P. B. Ferraz, L. Ubaldi, R. Vega Morales 2503.01981 and 250X.XXXXX



Spontaneous particle creation

- The Schwinger effect has been known for a long time *Fritz, Sauter (1931); W. Heisenberg, H. Euler (1936); J. Schwinger (1951)*
 - Pairs of charged particles and anti-particle created by background \vec{E}
 - Strong Electric fields are required
 - Effects are exponentially suppressed by the mass



• It can happen for constant \vec{E} , but requires time dependent vector potential \vec{A}

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Spontaneous particle creation

• A similar effect can happen in Curved backgrounds

L. Parker (1966); S. W. Hawking (1975)

- Particle production from vacuum under time dependent gravitational field
- Effects are already important in cosmology
 - LSS might be seeded by accelerated expansion during inflation

Cosmological Schwinger effect, J. Martin 0704.3540



Sloan Digital Sky Survey, in Saraswati supercluster. Credit: IUCAA

Spontaneous particle creation by time-varying backgrounds

Combining the two examples: Schwinger effect in de Sitter

Inflationary Magnetogenesis

• Generate the **observed magnetic fields** present in voids our universe

T. Kobayashi, N. Afshordi 2014

Generation of Dark Sectors

• Candidates for non-thermal dark matter

M. Bastero-Gil, P. Ferraz, L. Ubaldi, R. Vega-Morales 2023

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During inflation (Φ), in practice this could be realized with

$$S = -\int d^4x \sqrt{-g} \left[rac{1}{2} \partial_\mu \Phi \partial^\mu \Phi + V(\Phi) + rac{1}{4} F_{\mu
u} F^{\mu
u} + rac{lpha}{4f} \Phi F_{\mu
u} ilde{F}^{\mu
u} + \mathcal{L}_{ch}(\phi, A_
u)
ight],$$

• But no analytical solutions, difficult to test if (renormalization) results make sense

Scalar QED in de-Sitter

Forget about inflation

- Fix a de-Sitter background
- Set a constant electric field \vec{E} (along z direction)

$$m{A}_{\mu}=rac{m{E}}{m{H}^{2} au}\delta_{\mu}^{z}, \qquad m{F}_{\mu
u}m{F}^{\mu
u}=-2m{E}^{2}$$

$$S = \int d^4x \sqrt{-g} \left\{ -g^{\mu\nu} \left(\partial_\mu - i e A_\mu \right) \phi^* \left(\partial_\nu + i e A_\nu \right) \phi - \left(m_\phi^2 + \xi R \right) \phi^* \phi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right\}$$

Equations of motion

$$\phi_{k}^{\prime\prime} + 2aH\phi_{k} + \omega_{k}^{2}\phi_{k} = 0 \quad \longrightarrow \quad \phi_{k} = \frac{e^{-\pi\lambda r/2}}{a\sqrt{2k}}W_{i\lambda r,\mu}(2ik\tau)$$
$$\nabla^{\nu}F_{\mu\nu} = J_{\mu}^{\phi} \quad \text{with} \quad J_{\mu}^{\phi} = \frac{ie}{2}\left[\phi^{\dagger}\left(\partial_{\mu} + ieA_{\mu}\right)\phi - \phi\left(\partial_{\mu} - ieA_{\mu}\right)\phi^{\dagger}\right]$$

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Renormalization (Scalar) QED in de-Sitter

• Divergent expectation value.

$$\left\langle 0\left|J_{z}^{\phi}\right|0
ight
angle =rac{2e}{a^{4}}\intrac{d^{3}k}{(2\pi)^{3}}\left(k_{z}+eA_{z}
ight)\left|\phi_{k}
ight|^{2}.$$

Renormalized current

$$\langle J_z \rangle_{\rm ren} = aH rac{e^2 E}{4\pi^2} \left[rac{1}{6} \ln rac{m_{\xi}^2}{H^2} - rac{2\lambda^2}{15} + F_{\phi}(\lambda,\mu)
ight].$$

 $\lambda = rac{eE}{H^2}, \qquad \mu^2 = rac{9}{4} - rac{m_{\xi}^2}{H^2} - \lambda^2 \quad \text{and} \quad m_{\xi}^2 = m_{\phi}^2 + 12\,\xi H^2.$

- Adiabatic Subtraction
- Point Splitting
- Pauli Villars

T. Kobayashi, N. Afshordi 2014

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T. Hayashinaka, J. Yokoyama 2016

M. Banyeres, G. Domenèch, J. Garriga 2018

• Similar for fermions w. Adiabatic Subtraction

T. Hayashinaka, T. Fujita, J. Yokoyama 2016

Renormalization QED in de-Sitter



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- Schwinger effect with **classical** \vec{E}
 - A_{μ} not quantized (charged particles do not accelerate in dS)
 - Only charged particles (ϕ/ψ) are quantized
 - No photon loops $\rightarrow \phi/\psi$ propagator not corrected at loop level
 - Running of charge $e \iff A_{\mu}$ (from Ward Identity)

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 - Running of charge $e \iff A_{\mu}$ (from Ward Identity)
 - Only one counter-term in Lagrangian

$${\cal L} = - {1 \over 4} ({\it F}_{\mu
u})^2 - {1 \over 4} \delta_3 ({\it F}_{\mu
u})^2 - e {\it A}_\mu {\it J}^\mu + ... ~,$$

And the corrected equations of motion will be

$$(\delta_3 + 1) \nabla^{
u} F_{\mu
u} = \langle J_{\mu} \rangle.$$

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• For a constant electric field in de-Sitter ,

$$(\delta_3+1)\nabla^{\nu}F_{\mu\nu}=(\delta_3+1)(-2aHE\delta_{\nu}^z).$$

• Define the **renormalized current**

$$abla^{
u} F_{\mu
u} = \langle J_{\mu}
angle_{
m ren}
\langle J_{\mu}
angle_{
m ren} = \langle J_{\mu}
angle_{
m reg} - (-2aHE\delta_{
u}^z)\delta_{3\,
m reg} \,.$$

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m reg} - (-2aHE\delta_{
u}^z)\delta_{
m 3 reg}.$

• To get physical renormalized current, on-shell counter-term!

$$\Pi(p^2 = m_A^2) = 0 \ o \ \delta_3 = -e^2 \Pi_2(m_A^2)$$

• With classical $A_{\mu\nu}$, Π_2 is fully defined (at one loop) by



• Standard literature result is obtained with by treating Π_2 as in Minkowski

$$\delta_3 = -e^2 \Pi_2 (p^2 = 0) \ o \ \delta_3^{PV} = - rac{e^2}{48 \pi^2} \ln rac{\Lambda^2}{m^2}$$

- This results in $\ln m/H$ term that creates **negative conductivies** when $m \ll H$
- But does this condition actually hold for our setting?

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- This results in $\ln m/H$ term that creates **negative conductivies** when $m \ll H$
- But does this condition actually hold for our setting?

$$S = -\int d^4x \sqrt{-g} \, rac{1}{4} F^{\mu
u} F_{\mu
u} \ o \ g^{lpha
u} g^{eta\sigma}
abla_{lpha} F_{
u\sigma} = 0 \, .$$

Taking $A_{\mu} = \frac{E}{H^2 \tau} \delta^z_{\mu}$, in e.o.m. we find

$$g^{lpha
u}g^{eta\sigma}
abla_{lpha}F_{
u\sigma}=-2a^{-4}rac{E}{ au^{3}H^{2}}\delta_{i}^{z}
eq 0$$

• Just a kinetic term is not compatible with constant \vec{E} in de-Sitter

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5 × 3 × 5 × 5 × 0 0

Introduce an effective mass in Lagrangian

$$\mathcal{S}=-\int d^4x\sqrt{-g}\,\left(rac{1}{4}\mathcal{F}^{\mu
u}\mathcal{F}_{\mu
u}+rac{1}{2}m_A^2\mathcal{A}_\mu\mathcal{A}^\mu
ight)\,.$$

e.o.m. gives

$$-a^{-4}2\frac{E}{\tau^{3}H^{2}}\delta_{i}^{z}-m_{A}^{2}a^{-2}\frac{E}{\tau H^{2}}\delta_{i}^{z}=0 \quad \rightarrow \quad m_{A}^{2}=-2H^{2}.$$

- Get effective tachyonic mass
- Interpreted as effective source that ensures that \vec{E} is not diluted with expansion
- Consistency with constant electric field background implies

$$\Pi(p^2 = m_A^2) = 0 \ \rightarrow \ \delta_3 = -e^2 \Pi_2(p^2 = -2H^2)$$

Computing δ_3 as in Minkowski with the external momentum fixed by $p^2 = m_A^2 = -2H^2$



- Corrected $\ln m/H$ factor \rightarrow Currents in the massless limit become **finite**
- But for fermions and conformal scalars ($\xi = 1/6$), when $eE \ll H^2$ they are negative

Computing δ_3 as in Minkowski with the external momentum fixed by $p^2 = m_A^2 = -2H^2$



- Corrected $\ln m/H$ factor \rightarrow Currents in the massless limit become **finite**
- But for fermions and conformal scalars ($\xi = 1/6$), when $eE \ll H^2$ they are negative
- Minkowski propagators in the loop are not accurate
- Do not capture correctly **IR effects**
- We try a correction as exact de-Sitter does not seem doable (to us)

Scalars
$$m^2 \rightarrow m^2 + \xi R$$
 Fermions $m^2 \rightarrow m^2 + \frac{1}{4}R$

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Finally δ_3

• Pauli-Villars to regularize both δ_3 and $\langle J_\mu
angle$

$$abla^{
u} \textit{\textit{F}}_{\mu
u} = \langle \textit{J}_{\mu}
angle_{\textit{ren}} = \langle \textit{J}_{\mu}
angle_{\textit{reg}}^{\textit{PV}} - (-2a\textit{HE}\delta_{
u}^{\ z})\delta_{3}^{\textit{PV}}$$

With

$$\delta_{3} = \left(\frac{e}{12\pi}\right)^{2} \left(3\ln\left(\frac{m^{2}}{\Lambda^{2}}\right) - 12\left(\frac{m}{H}\right)^{2} + 6\left(2\left(\frac{m}{H}\right)^{2} + 1\right)^{3/2} \operatorname{coth}^{-1}\left(\sqrt{2\left(\frac{m}{H}\right)^{2} + 1}\right) - 8\right)$$

• We find the **renormalized** current to be

$$\left\langle J_{z}^{\phi} \right\rangle_{\text{ren}}^{PV} = aH \frac{e^{2}E}{4\pi^{2}} \left[\frac{1}{3} \ln \frac{m}{H} - \frac{4}{9} - \frac{2}{3} \left(\frac{m}{H} \right)^{2} - \frac{2\lambda^{2}}{15} + F_{\phi} \right] \\ + \frac{\left(1 + 2 \left(\frac{m}{H} \right)^{2} \right)^{3/2}}{3} \operatorname{coth}^{-1} \left(\sqrt{2 \left(\frac{m}{H} \right)^{2} + 1} \right) \right]$$

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Results



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Conclusions

- We have revised Schwinger pair production for constant **E** in de-Sitter
- Obtain the renormalised Lagrangian and correspondent parameters
- Imposed physical renormalisation conditions
- We were able to address and clarify literature's negative conductivities in H > m case
 - Unphysical result comes from wrong physical conditions
 - Minkowski propagators are inadequate for IR behavior
- Obtained an UV and IR divergence free Schwinger current.

Backup

Revising PV

- An arbitrary number of additional auxiliary fields are introduced to cancel divergences
- The mass of these extra fields will then be sent to infinity, making them **non-dynamical**

Introduce **3** fields

$$\sum_{i=0}^{3} (-1)^{i} = 0 \quad \text{and} \quad \sum_{i=0}^{3} (-1)^{i} m_{i}^{2} = 0,$$

$$m_{0} = m, \quad m_{2}^{2} = 4\Lambda^{2} - m^{2} \quad \text{and} \quad m_{1}^{2} = m_{3}^{2} = 2\Lambda^{2}, \quad \Lambda \to \infty$$
The **regularized** current

$$\langle J_{z} \rangle_{\text{reg}} = \lim_{\Lambda \to \infty} \sum_{i=0}^{3} (-1)^{i} \langle J_{z} \rangle_{i}.$$

$$\langle J_{z}^{\phi} \rangle_{\text{reg}} = aH \frac{e^{2}E}{4\pi^{2}} \lim_{\Lambda \to \infty} \left[\frac{1}{6} \ln \frac{\Lambda^{2}}{H^{2}} - \frac{2\lambda^{2}}{15} + F_{\phi}(\lambda, \mu, r) \right]$$

• $\ln \Lambda / H$ divergence to be reabsorbed with renormalization of the charge

$$\begin{split} \langle J_{\mu} \rangle_{reg} &= (\delta_{3} + 1) \, \nabla^{\nu} F_{\mu\nu} \\ \nabla^{\nu} F_{\mu\nu} &= \langle J_{\mu} \rangle_{ren}^{PV} - (-2aHE\delta_{\nu}^{\ z}) \delta_{3}^{PV} \end{split}$$