

Flavor, Compositeness & Supersymmetry

"La Grande Bouffe"

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- K. Agashe, G. Giudice, RR, R. Sundrum in progress

Hierarchy Paradox

$$L_{SM} = \underbrace{L^{(d \leq 4)}}_{\text{seen}} + \underbrace{\frac{1}{M_*} L^{(5)}}_{\text{unseen or } p\text{-decay} \dots} + \underbrace{\frac{1}{M_*^2} L^{(6)}}_{\text{tiny } \nu\text{-mass}} + \dots$$

■ $m_* \gg m_{\text{weak}}$

• B, L, GIM



Simple
yet

• M_H^2



unnatural

■ $m_* \sim m_{\text{weak}}$

• M_H^2



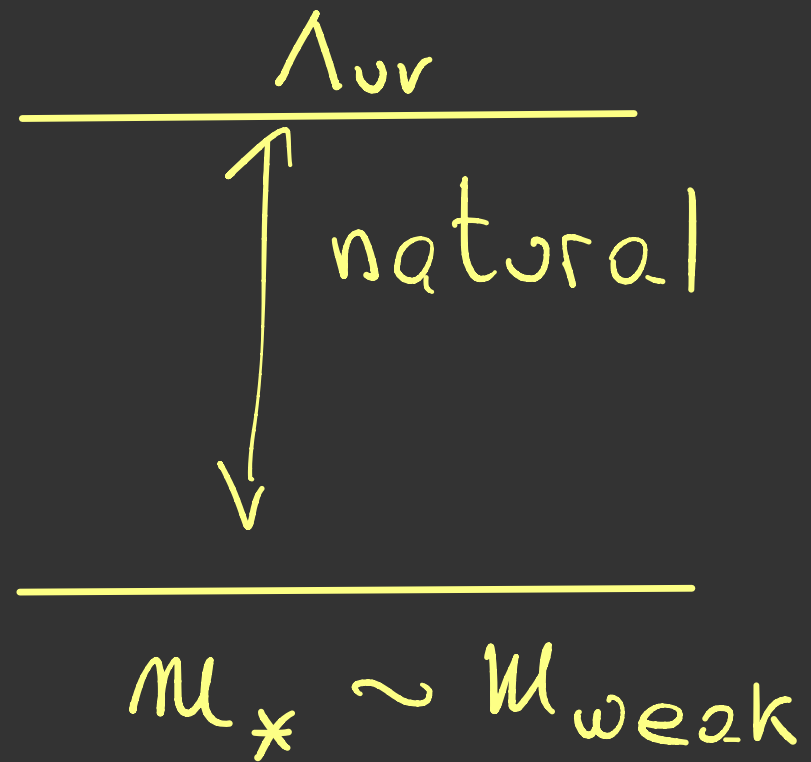
Natural
yet

• B, L, GIM



"not so simple"

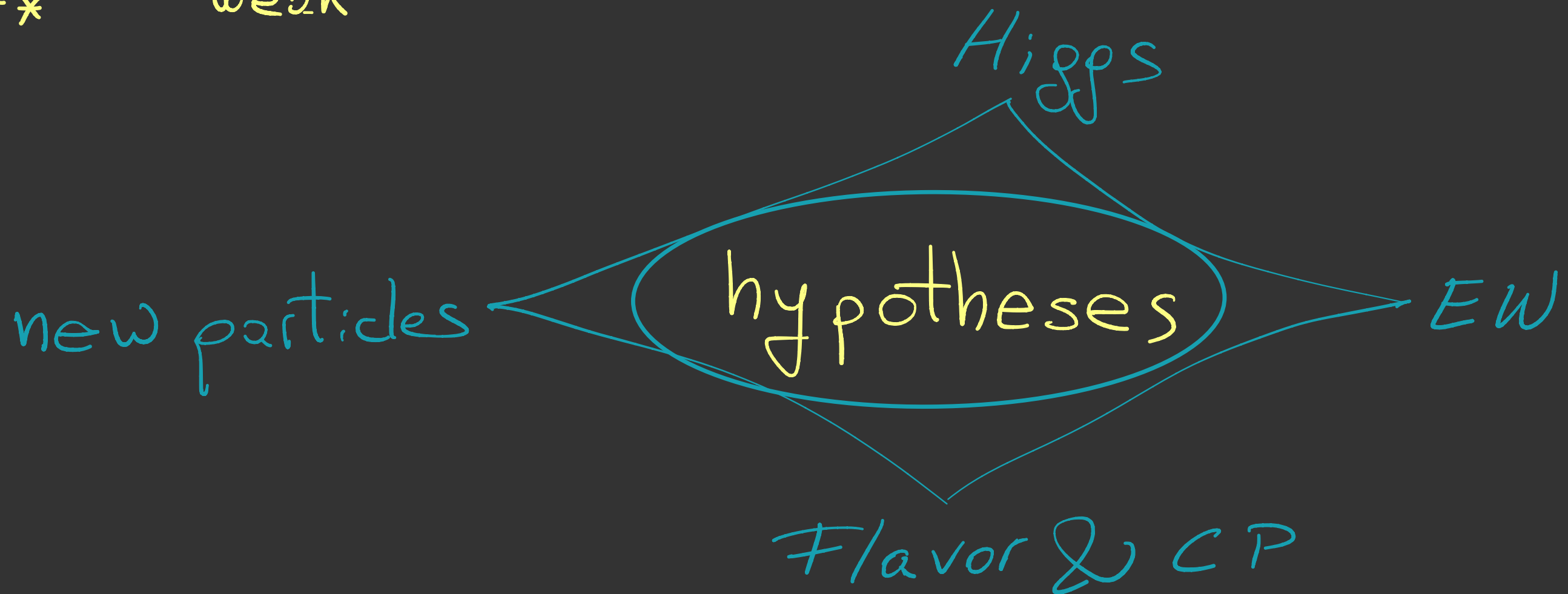
"Not so simple": but how much?



$\Rightarrow$

- explain $V(H)$
- deal with Flavor

} Model Building:
Exploration of
Structured hypotheses



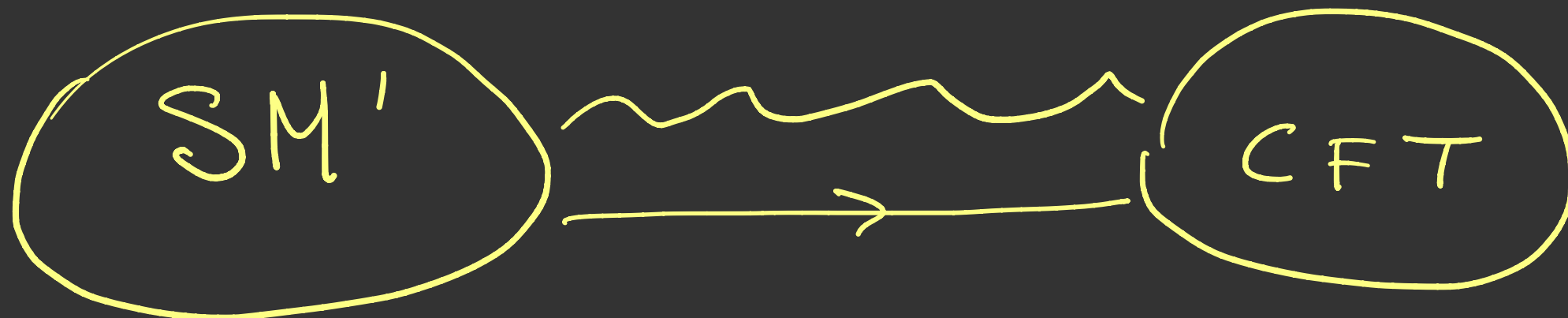
Modern Composite Higgs

Agashe,
Contino
Powarol
'04

$$\mathcal{L} = \mathcal{L}_{\text{SM}'} + \mathcal{L}_{\text{CFT}} + g A_\mu J_{\text{CFT}}^\mu + y_{i\alpha} \psi_i \mathcal{O}_\alpha$$

Λ_{UV}

$\Downarrow$



m_*

$\downarrow$

$$\mathcal{L}_{\text{SM}'} + \mathcal{L}(H) + Y H \bar{\Psi} \Psi$$

$\nwarrow$

Fermion Mass Spectrum from RG flow

$$y_{ia}^q q_i \mathcal{O}_a^q + y_{ia}^d d_i \mathcal{O}_a^d + y_{ia}^u u_i \mathcal{O}_a^u + y_{ia}^l l_i \mathcal{O}_a^l + y_{ia}^e e_i \mathcal{O}_a^e$$

$$\downarrow \quad \downarrow$$

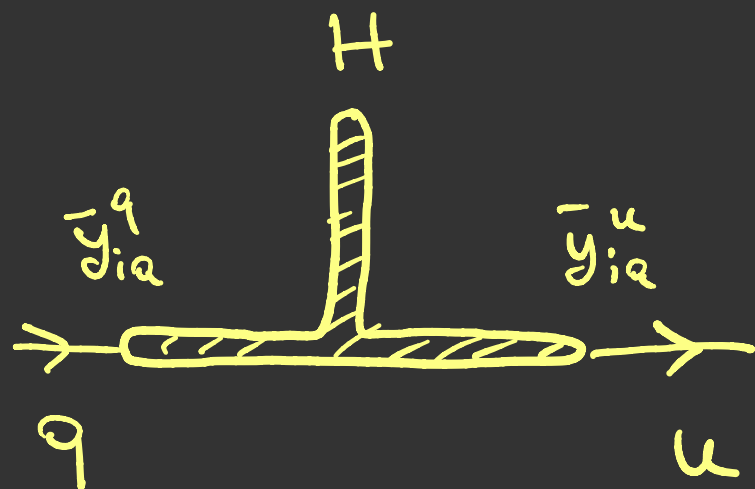
$$\frac{3}{2} \quad \Delta_a^q = \frac{5}{2} + \gamma_a^q \quad \text{etc}$$

$$\Rightarrow y_{ia}^F(\mu) = y_{ia}^F(\Lambda_w) \left(\frac{\mu}{\Lambda_w} \right)^{\gamma_a^F} \xrightarrow{u_*} \bar{y}_{ia}^F = y_{ia}^F(\Lambda_w) \left(\frac{u_*}{\Lambda_w} \right)^{\gamma_a^q}$$

• Imagine $\gamma_1^F \gtrsim \gamma_2^F \gtrsim \gamma_3^F$ [Ex $\gamma_1 = 0.6, \gamma_2 = 0.4, \gamma_3 = 0.2$]

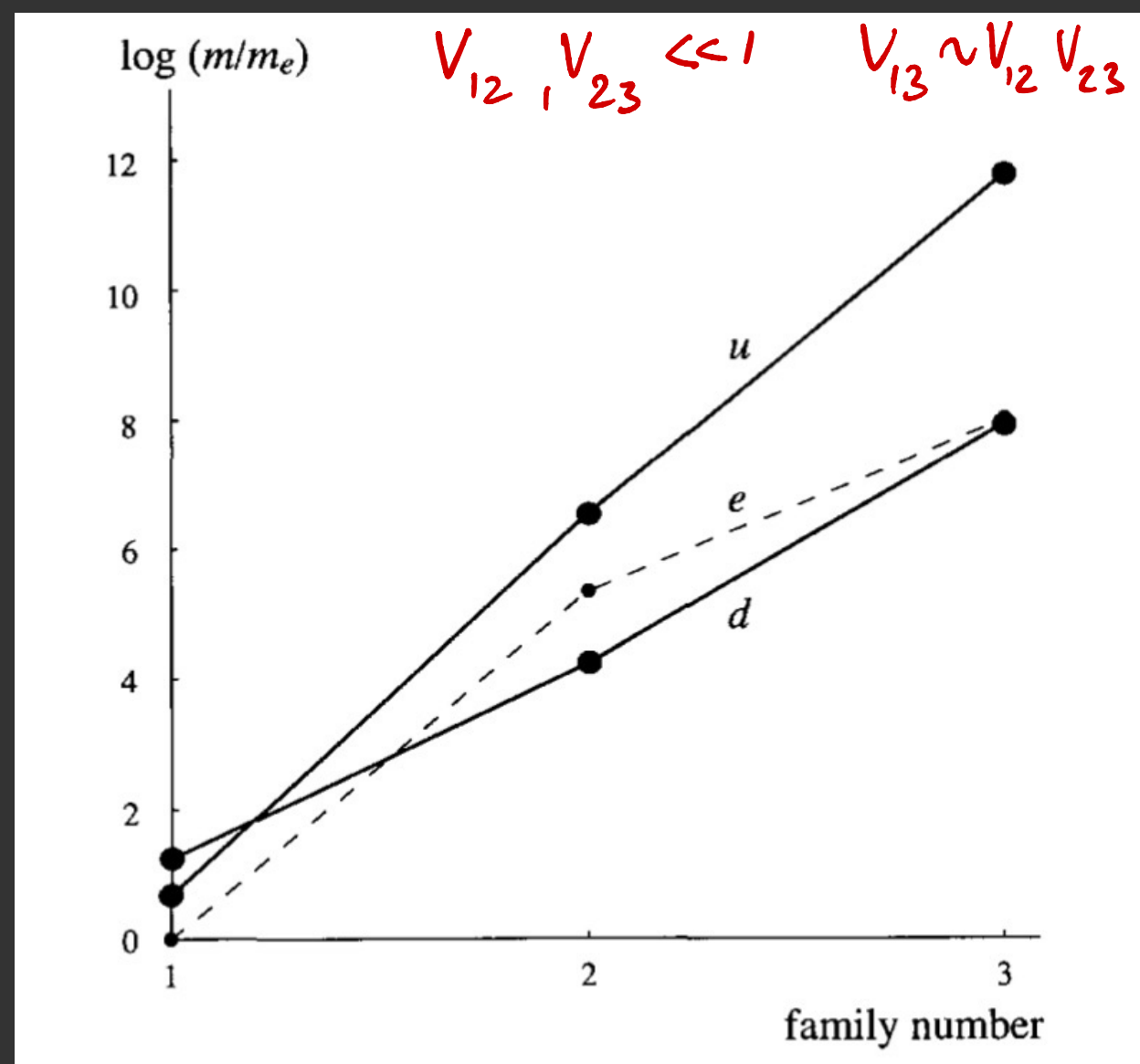
$$|\bar{y}_{i1}| \sim \left(\frac{u_*}{M} \right)^{0.6} \sim (10^{-6})^{0.6} \sim 10^{-3}, \quad |\bar{y}_{i2}| \sim 10^{-2}, \quad |\bar{y}_{i3}| \sim 10^{-1}$$

Yukawa Couplings



$$Y_{ij}^u \sim \frac{\bar{y}_{iq}^q \bar{y}_{jb}^u m_{ab}}{g_*}$$

$$\approx \varepsilon_{iq}^q \varepsilon_{jb}^u m_{ab} \cdot g_*$$

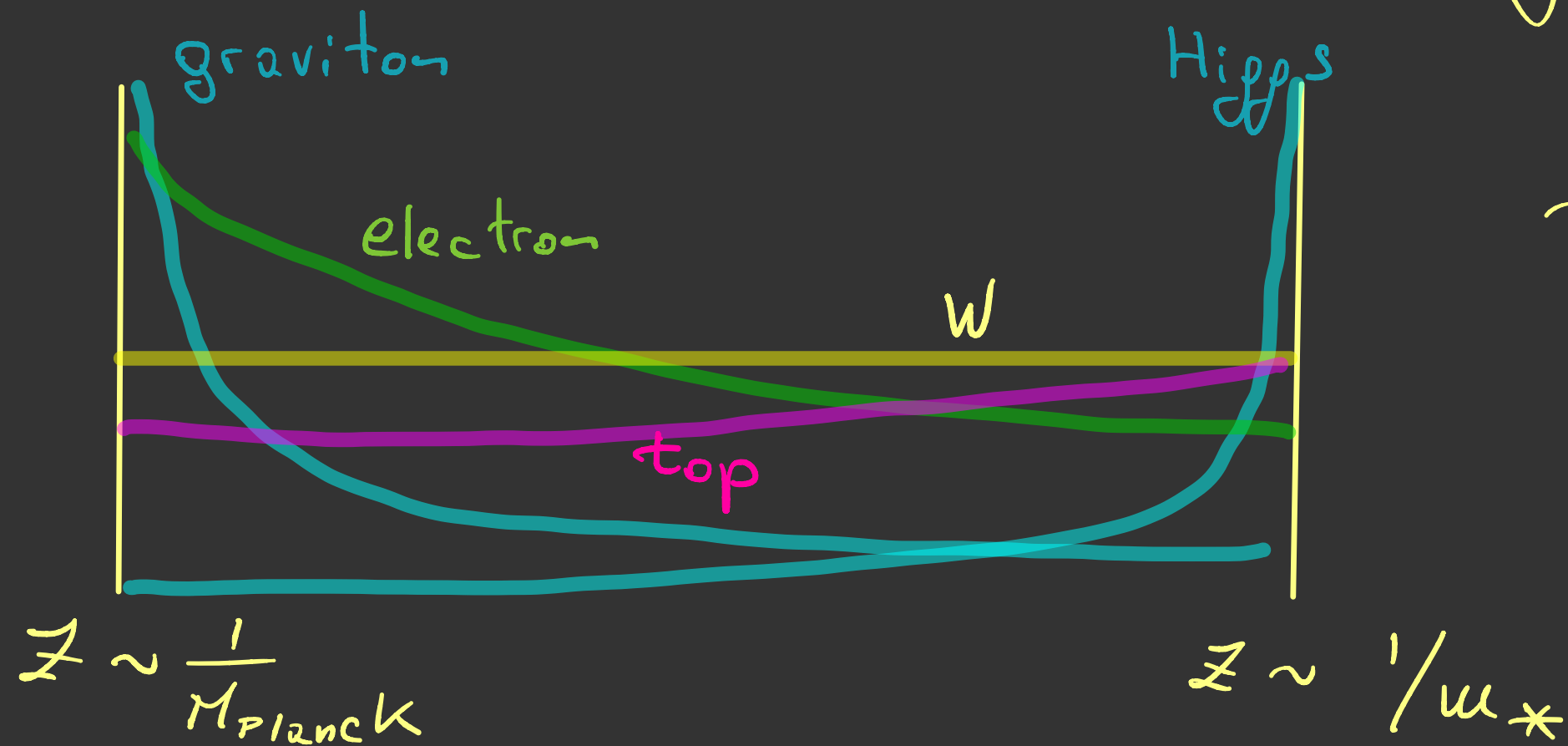


$$\left\{ \begin{array}{l} y_{q3} \rightarrow \text{rank} = 1 \\ \oplus y_{q2} \rightarrow \text{rank} = 2 \\ \oplus y_{q1} \rightarrow \text{rank} = 3 \end{array} \right.$$

A full fledged CFT realizing all that?

... of course not

- but we have at least "holographic realizations"



Randall-Sundrum's
slice of AdS5

$$m_* \sim m_{KK} \quad , \quad g_* \sim g_{KK} \sim \frac{4\pi}{\sqrt{N}} \quad , \quad \epsilon's \equiv \text{wave function overlap}$$

Pseudo NG-Higgs

$$\text{---} m_*$$

$$\text{---} m_h$$

$$\blacksquare m_h^2 \sim \frac{1}{8\pi^2} (\# y_t^2 + \# g^2) m_*^2$$

$$\blacksquare v^2 \sim \frac{m_*^2}{g_*^2} \equiv f^2$$

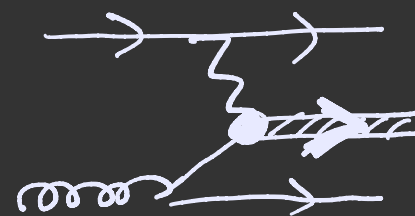
Naturalness

$$\left(\frac{v}{f}\right)^2 \sim 1$$

$$m_* \sim 0.5 \text{ TeV}$$

LHC

$$\frac{\delta g_h}{g_h|_{SM}} \sim \left(\frac{v}{f}\right)^2 \leq 0.1 \div 0.2$$

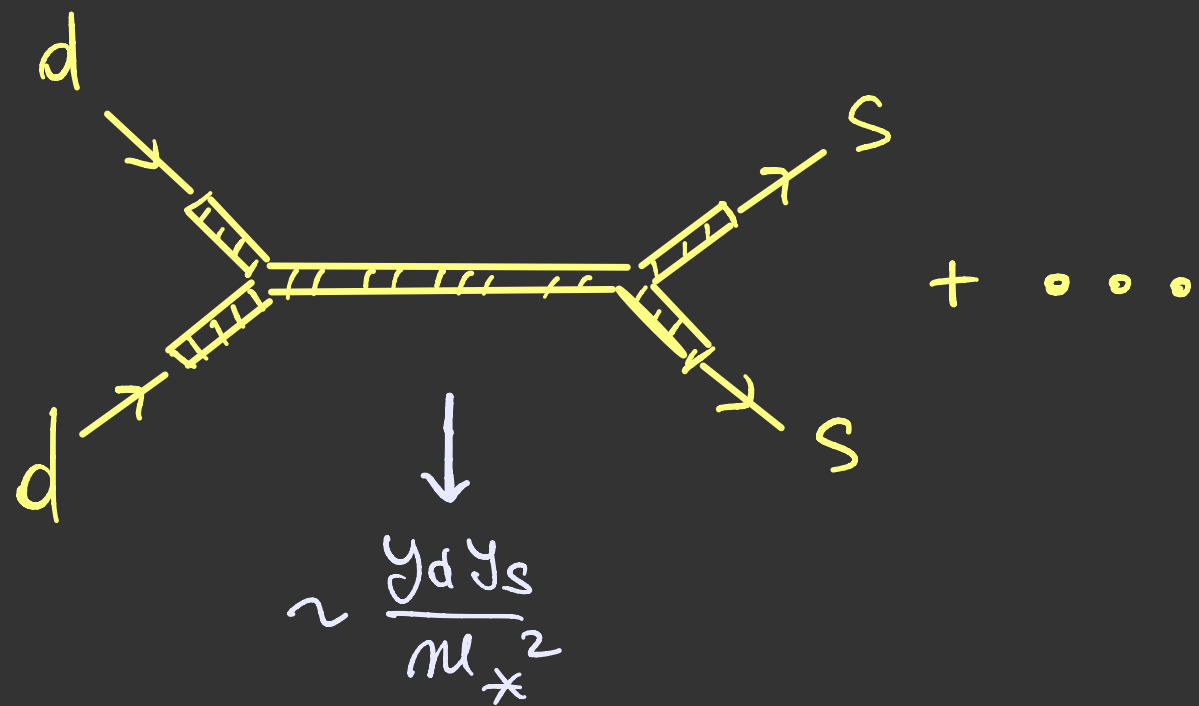


$$m_* \gtrsim 1.5 \text{ TeV}$$

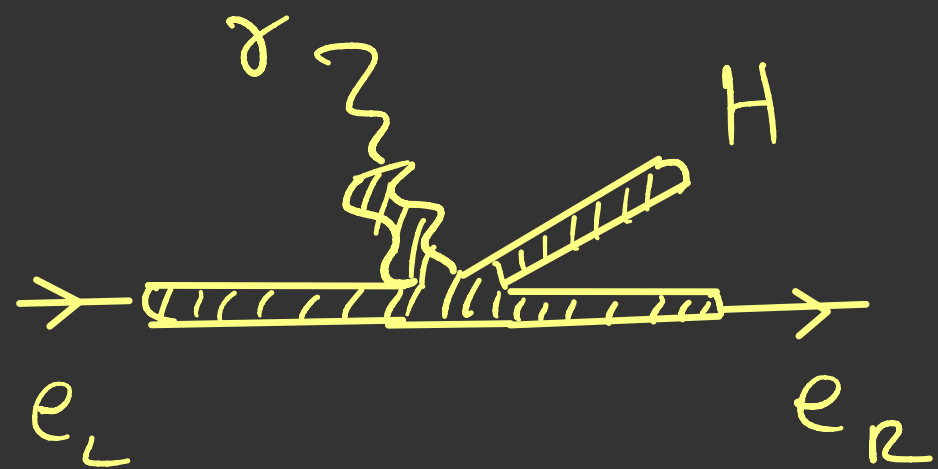


$$m_* \gtrsim 4.5 \text{ TeV}$$

▲ The Flavor Catastrophe



$$\left\{ \begin{array}{l} \varepsilon_K \\ b \rightarrow s \gamma \\ c \rightarrow u \gamma \\ e d u_{u,d} \end{array} \right. \Rightarrow m_* \gtrsim 20 \div 50 \text{ TeV}$$

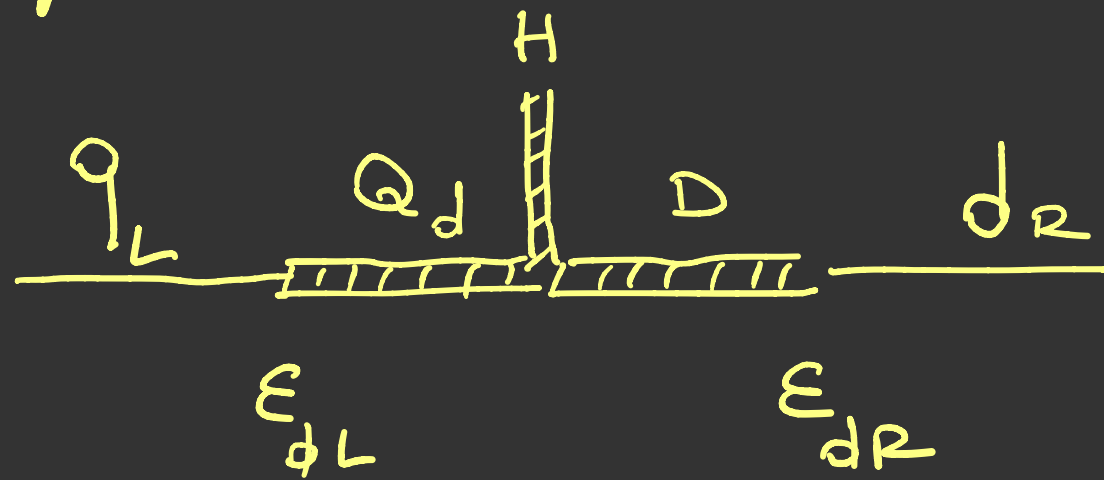
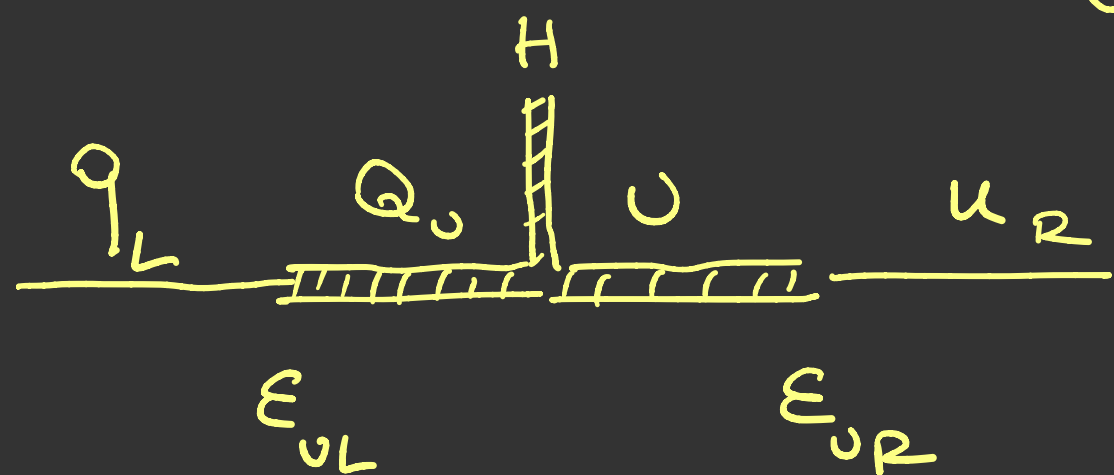


$$\Rightarrow d_e \sim \frac{m_e e}{m_*^2} \Rightarrow m_* \gtrsim 2 \cdot 10^3 \text{ TeV}$$

▲ Not only unNatural, but out of FCC hh reach!!!

▲ To allow low m_* must assume Flavor+CP Symms
 ⇒ explanation of fermion spectrum is lost

▲ Largest Symmetry Group at play



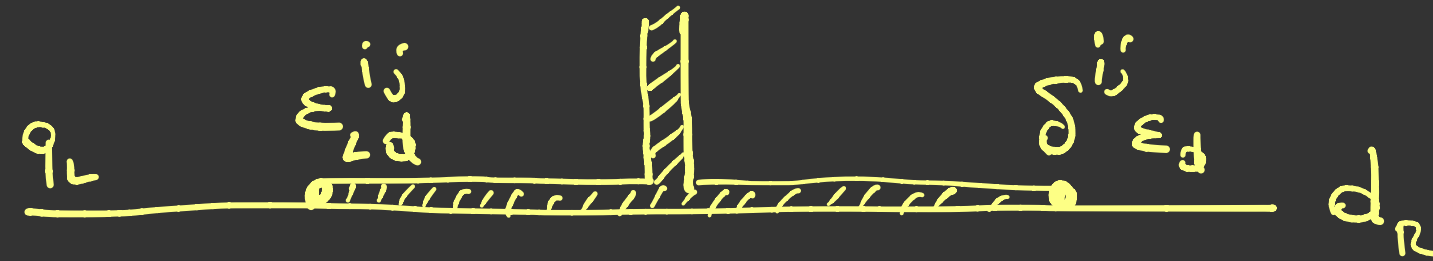
$$U(3)_q \times U(3)_u \times U(3)_q \times U(3)_d \times U(3)_D$$

● Scenarios $\Leftrightarrow$ E -induced breaking patterns

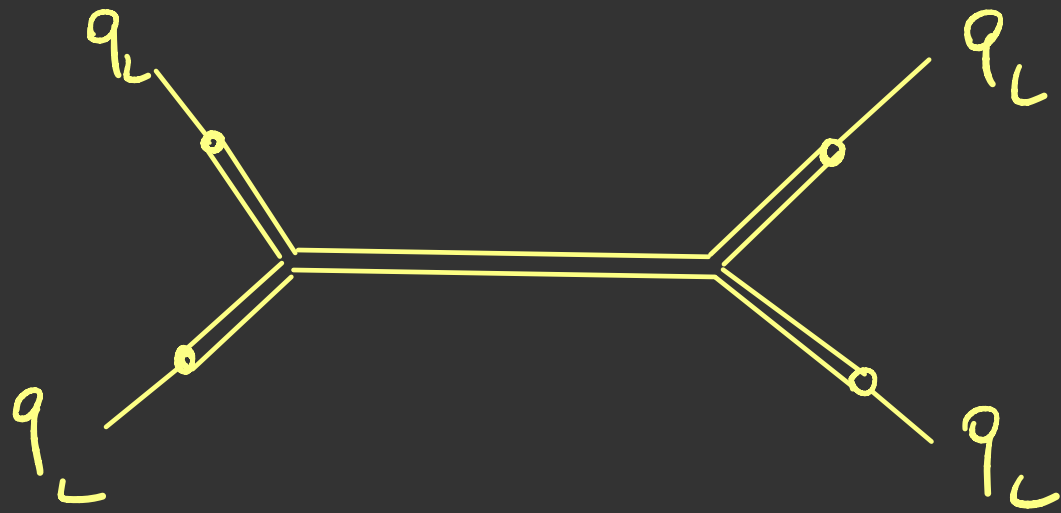
Ex "Right
Universality"



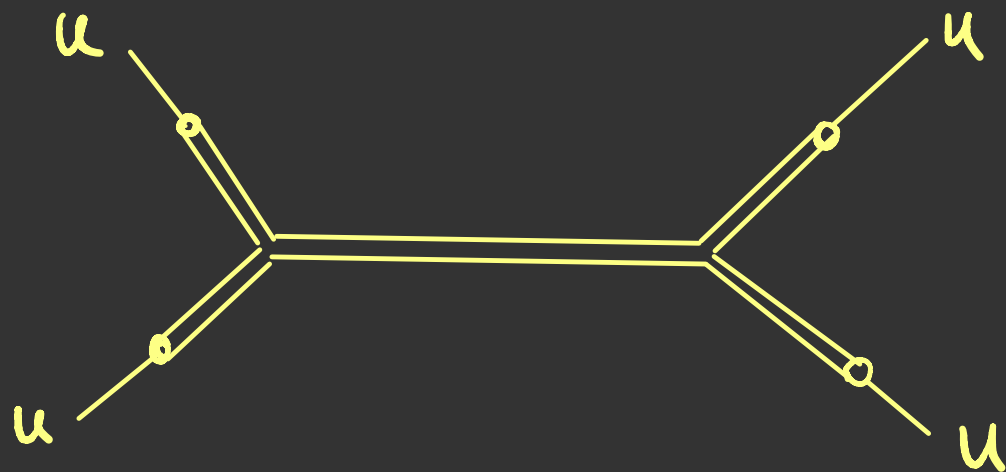
$$Y_u^{ij} = \epsilon_L^{ij} \epsilon_u g_*$$



$$Y_d^{ij} = \epsilon_L^{ij} \epsilon_d g_*$$



$$\sim \frac{(Y_u^\dagger Y_u)^2}{g_*^2 \epsilon_u^4} \frac{1}{u_*^2} \Rightarrow u_* \gtrsim \frac{6.6 \text{ TeV}}{g_* \epsilon_u^2}$$



$$\sim \frac{g_*^2 \epsilon_u^4}{u_*^2} \Rightarrow u_* \gtrsim (5-8) \text{ TeV} \cdot g_* \epsilon_u^2$$

▲ Broad Scenarios

|Redi-Weiler '11| Barbieri et al. '13
|Matsedonskyi '13| Barbieri '23

- R.U.

$$U(3)_q \times U(3)_{u+u} \times U(3)_{D+d}$$

- partial up R.U.

$$U(3)_q \times [U(2) \times U(1)]_{u+u} \times U(3)_{D+d}$$

- partial R.U.

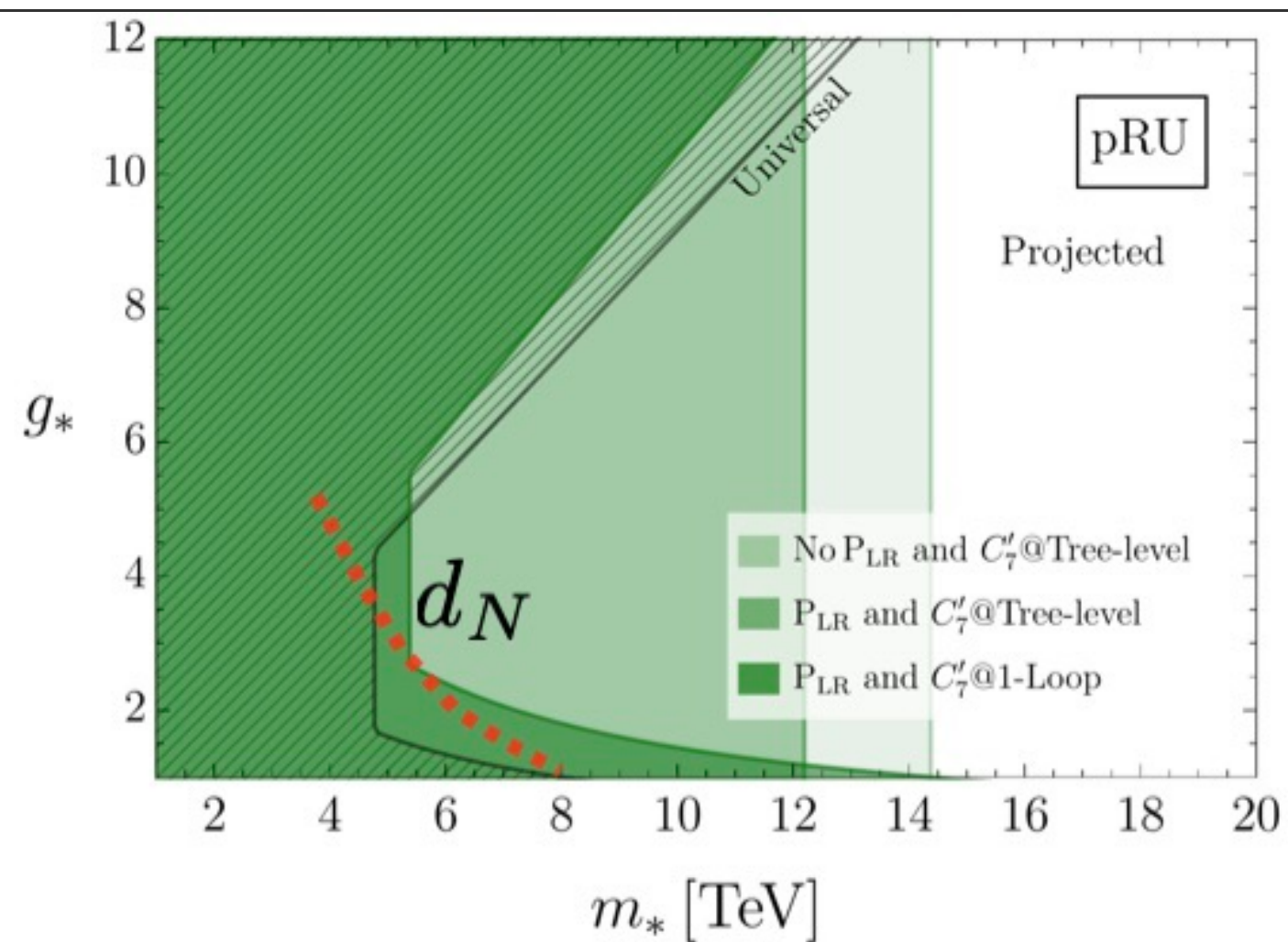
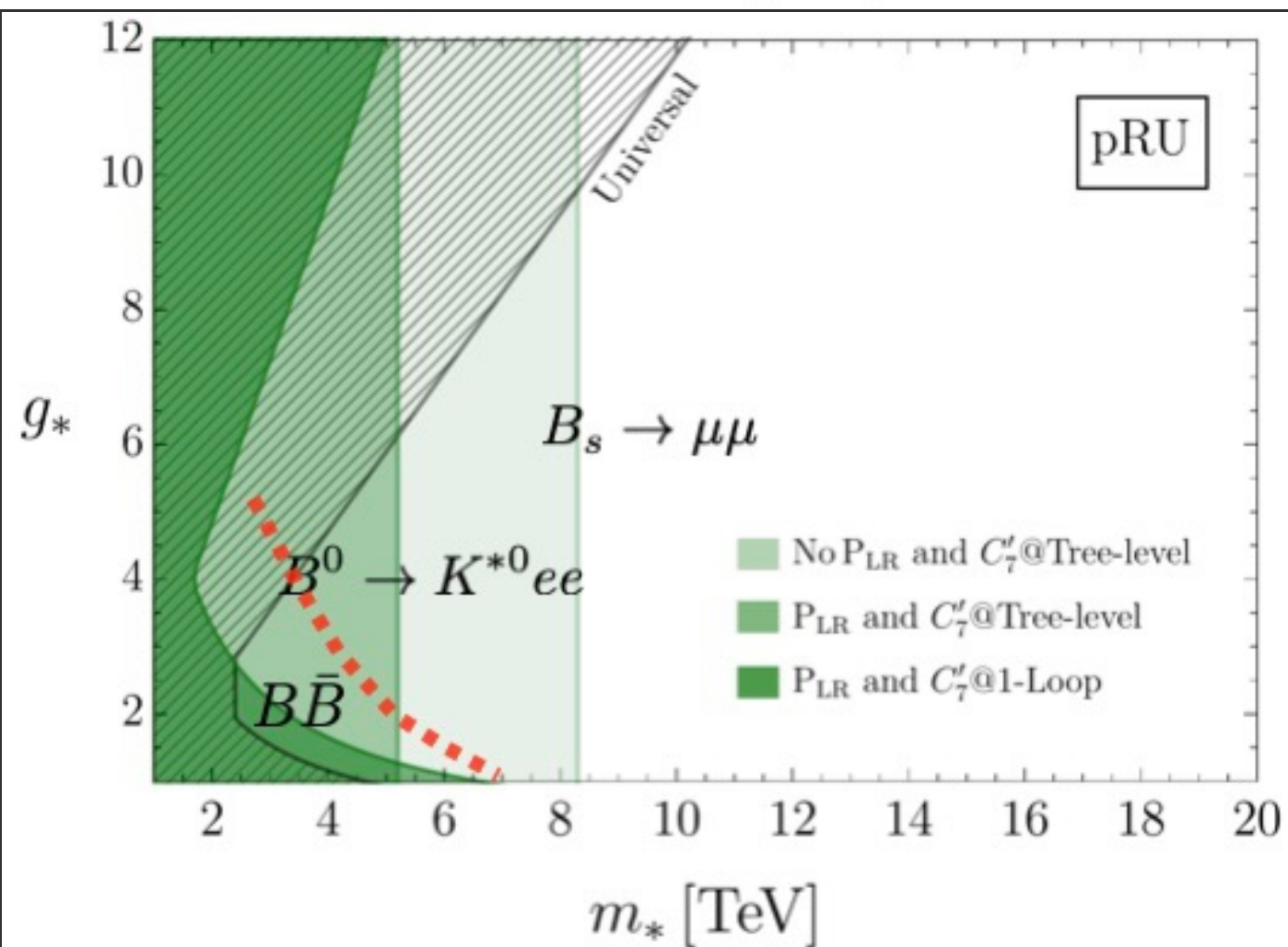
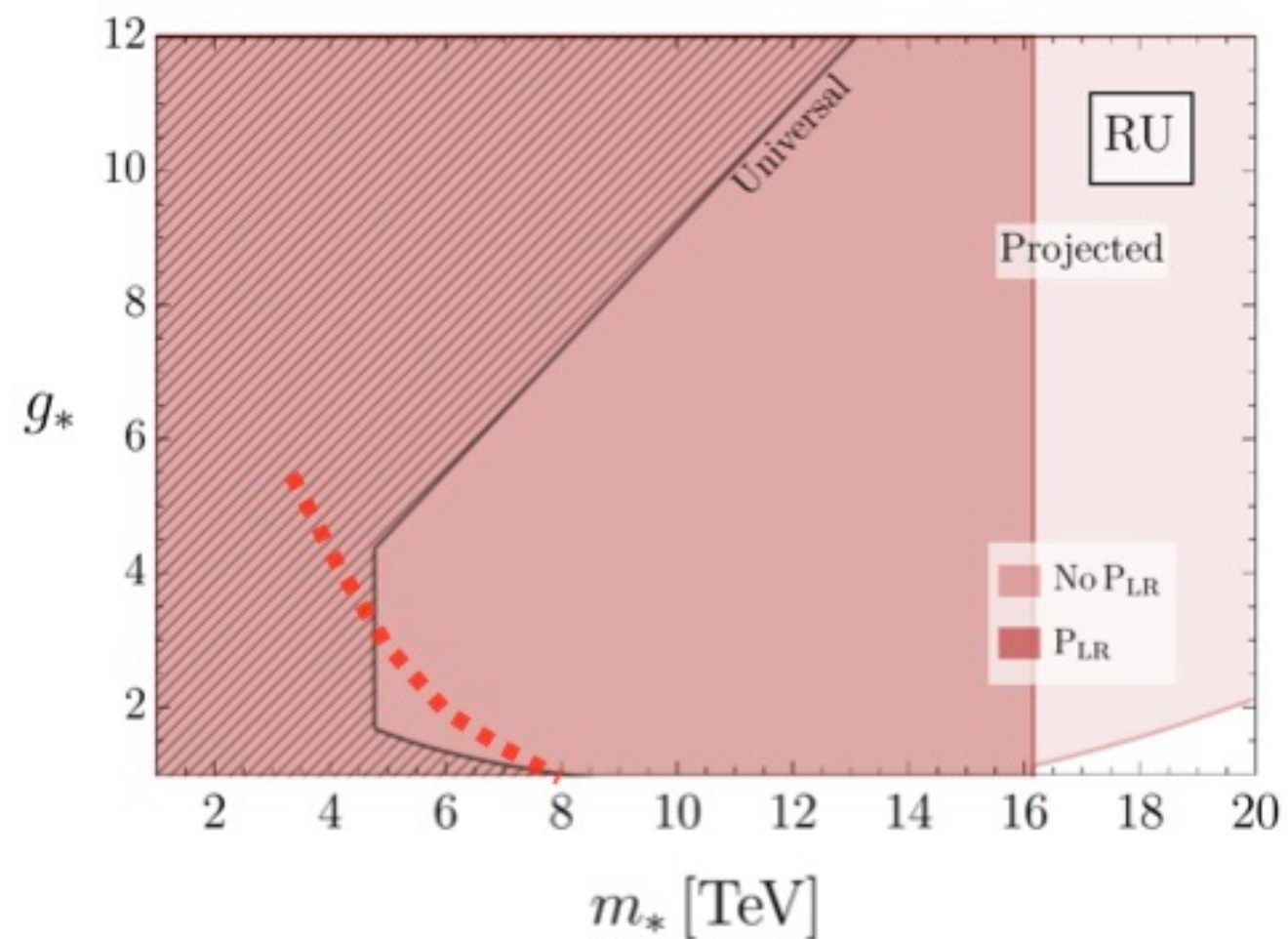
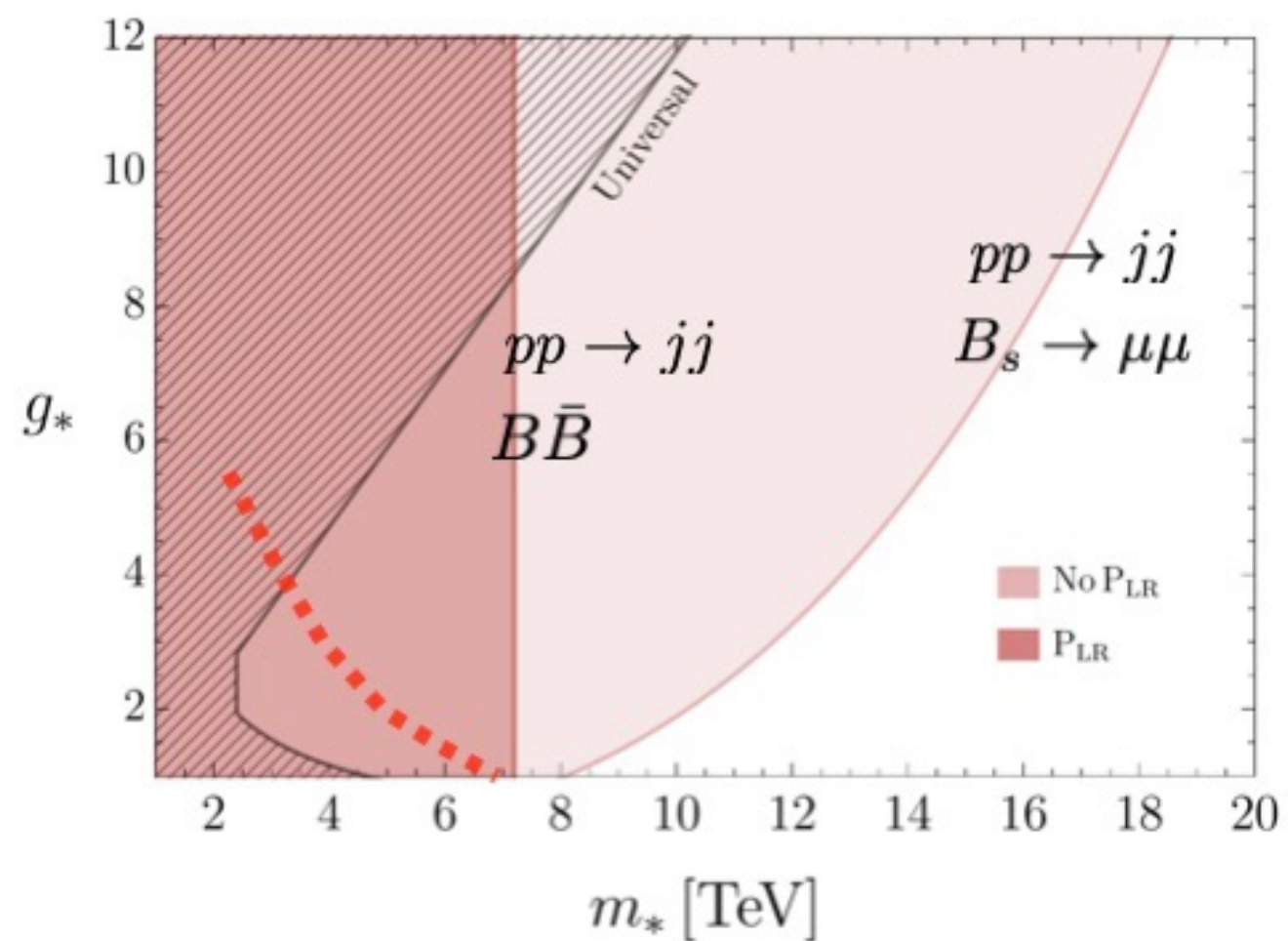
$$U(3)_q \times [U(2) \times U(1)]_{u+u} \times [U(2) \times U(1)]_{D+d}$$

- L.U.

$$U(3)_{q+q} \times U(3)_u \times U(3)_d$$

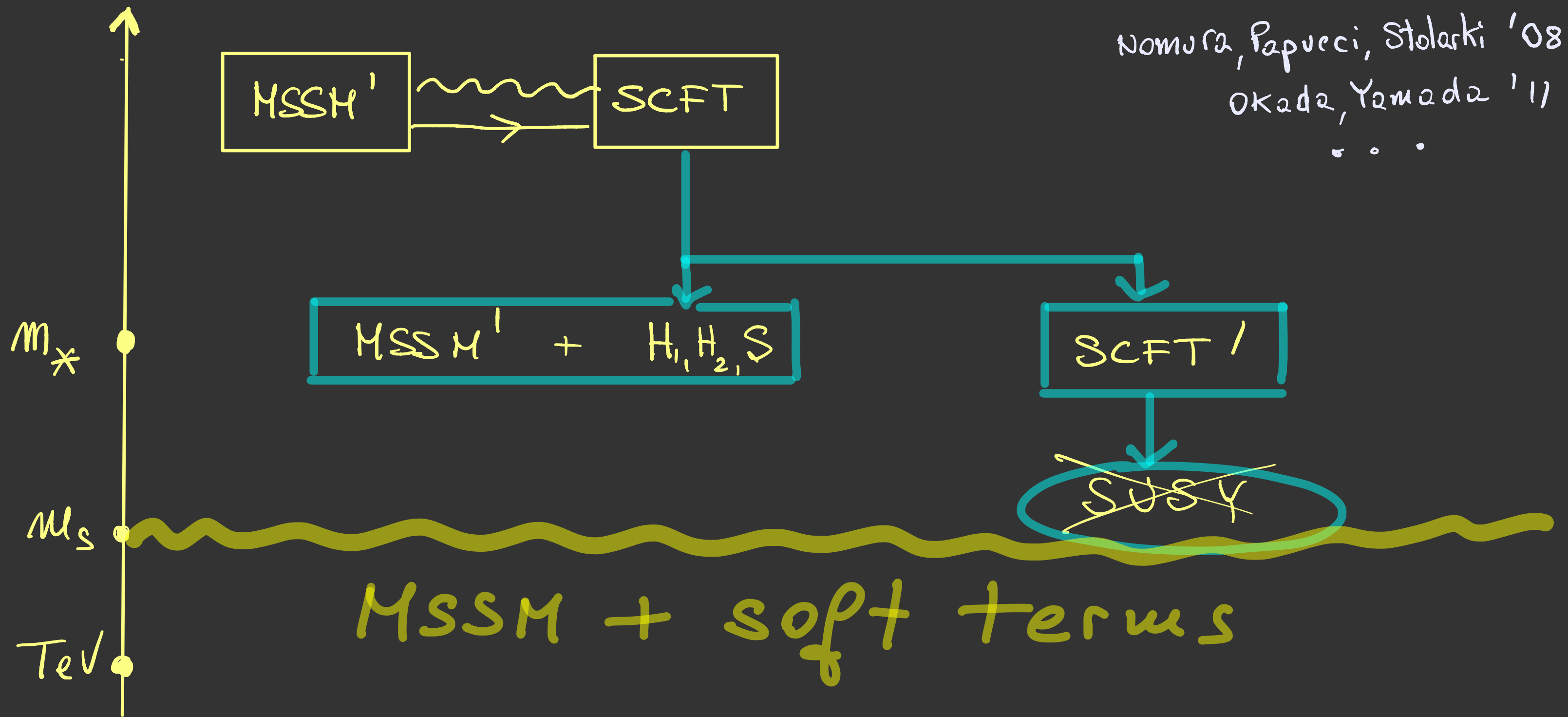
- partial L.U.

$$[U(2) \times U(1)]_{q+q} \times U(3)_u \times U(3)_d$$

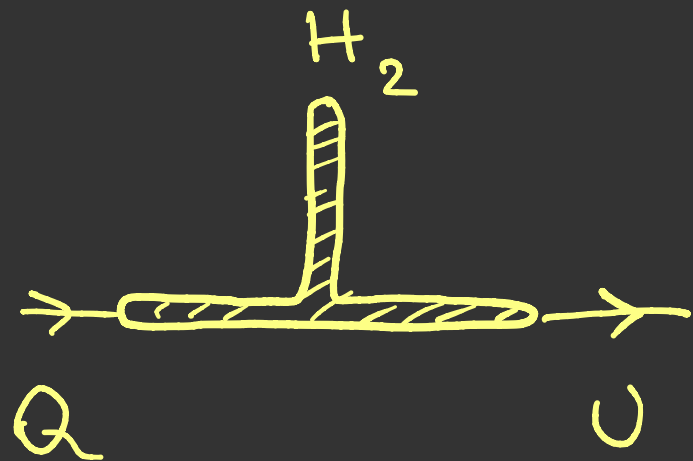


SuperSymmetric Composite Higgs

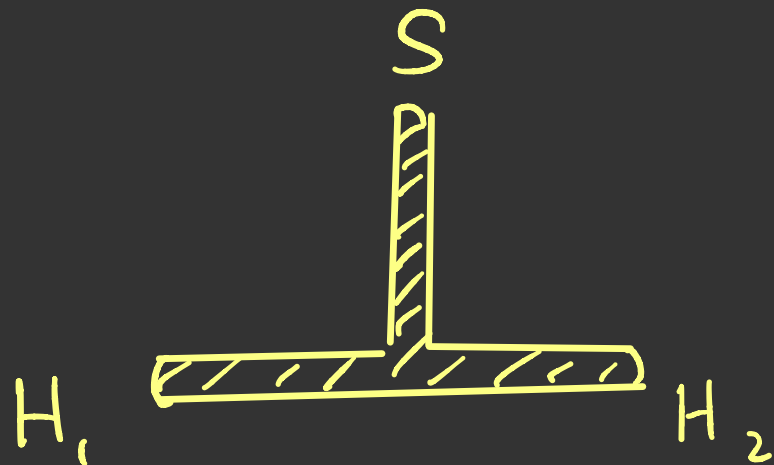
$$\mathcal{L} = \mathcal{L}_{\text{MSSM}'} + \mathcal{L}_{\text{SCFT}} + \int d^4\theta \, g V_{\text{SH}} \mathcal{Z}_{\text{CFT}} + \int d^2\theta \, y_{ia} \bar{F}_i \mathcal{O}_a$$



At m_*



$$\Rightarrow Y_{ij}^u \sim \epsilon_{ia}^Q \epsilon_{jb}^U u_{ab} \cdot \mathcal{G}_*$$



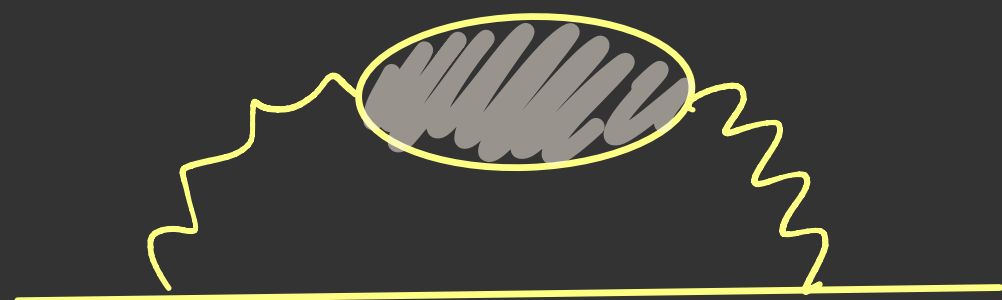
$$\Rightarrow \int d^2\theta \mathcal{G}_* H_1 H_2 S$$

$\mathcal{G}_* \sim 2 \div 4 \Rightarrow$ helps $m_h = 125 \text{ GeV}$ with "light stops"

At m_s



$$m_{1/2} \sim \frac{g^2}{g_*^2} m_s \sim \frac{\alpha N}{4\pi} m_s$$



$$m_{\text{sfermions}}^2 \sim \frac{m_{1/2}^2}{N} + \text{RG.}$$

Flavor universal masses mediated by gauge fields
 $\equiv$ Gauge Mediated Supersymmetry Breaking

The μ -B μ -A Situation | Giudice-Masiero '88

- $\mathcal{O}$ chiral, $R \sim \mathcal{O} \times \mathcal{O}^\dagger$ vector $\Rightarrow \Delta_R = 2\Delta_{\mathcal{O}} + \mathcal{O}(\frac{1}{N})$
- ~~Susy~~ $\Rightarrow \langle \mathcal{O} \rangle = u^{\Delta_{\mathcal{O}}} \theta^2_s$ $\langle R \rangle = \dots$

$$I. \int d^4\theta (H_1 H_2 + H_1^\dagger H_1 + \dots) \left(\frac{\mathcal{O} + \mathcal{O}^\dagger}{u_*^{\Delta_{\mathcal{O}}}} + \frac{R}{u_*^{\Delta_R}} \right) + \dots$$

$$\mu \sim B \sim A \sim \sqrt{5} u_H^2 \sim \frac{u_s^{\Delta_{\mathcal{O}}+1}}{u_*^{\Delta_{\mathcal{O}}}}$$

$$\mu \sim u_{1/2} \Rightarrow$$

$$100 \text{ TeV} \lesssim u_* \lesssim 2000 \text{ TeV}$$

$$N=10 \quad \Delta_{\mathcal{O}}=2$$

$$N=5 \quad \Delta_{\mathcal{O}}=1$$

II. $\mathbb{Z}_3^{PQ} \oplus \text{singlet}$

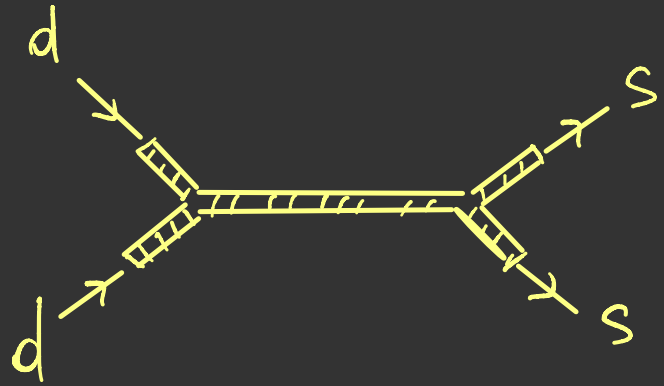
$$\int d^4\theta \left(H_1 H_2 + S^2 \right) \frac{\mathcal{O}^+}{\omega_* \Delta_0} + \left(S^\dagger S + H_1^\dagger H_1, \dots \right) \frac{\mathcal{R}}{\omega_* \Delta_R} \\ + \int d^2\theta \quad \mathcal{Q}_* \left(H_1 H_2 S + S^3 \right) + \cancel{SQ} \\ \text{holomorphy}$$

$$\mu \sim \mu_S \sim \sqrt{\delta \omega_H^2} \sim \sqrt{\delta \omega_S^2} \gg A, B$$

Study EWSB to fully assess

- $\omega_h = 125$?
- range of $\tan\beta$

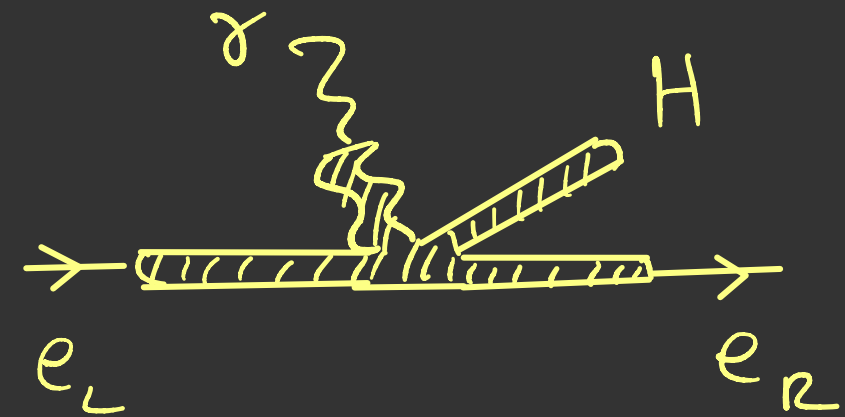
SUSY saves Flavor and edms



$$\longrightarrow \frac{y_d y_s}{u_*^2} e^{i\varphi} d d \bar{s} \bar{s}$$

$$\frac{u_*}{\tan \beta} \gtrsim 30 TeV$$

$\Rightarrow$ for $\tan \beta \lesssim 3 \div 5$ u_* allowed at FCC edge



$\longrightarrow$ "O" in SUSY limit $\Rightarrow$ can relax bound

Ferrara-Remiddi '74

d_e : plethora of possible effects

$$\int d^4\theta y_{ia}^L y_{jb}^R c_{ab} L_i E_j H_1 \frac{R}{M_*^{\Delta_R+1}} \Rightarrow \text{Im} \mathcal{A}_{ee} = \sqrt{\frac{y_\mu}{y_\tau}} \frac{(\text{TeV})^2}{M_*}$$

$\Downarrow$

$$\int d^4\theta (\dots) L_i E_j H_1 \bar{W}_\alpha \bar{D}^\alpha \frac{\mathcal{O}^+}{M_*^{\Delta_O+3}} \Rightarrow$$

$$d_e < 4 \cdot 10^{-30} \text{ e.cm}$$

$$\Rightarrow M_* \gtrsim 100 \text{ TeV}$$

$$N_0 P\theta \Rightarrow \int d^4\theta (\dots) L_i E_j H_2^+ \Rightarrow \tilde{e}_i \tilde{h}_2 \bar{\phi} e_j \frac{1}{M_*}$$

$$M_* \gtrsim 300 \cdot \tan\beta \text{ TeV}$$

Many details to work out

- Supersymmetric SILH-EFT : $\mathcal{G}_*$, $\mathcal{U}_*$, $H_{1,2}$, S
⊕ constrained $\rightarrow$
 - X , $X^2=0$ Goldstino
 - Π , $X(\Pi-\Pi^\dagger)=0$ PQ-Goldstone
- EWSB and m_h , with and without S , $\mathbb{Z}_3^{\text{PQ}}$, $U(1)_{\text{PQ}}$...
- Cosmology : light gravitino affects LSS & DM
if in thermal equilibrium $m_s \lesssim 100 \text{ TeV}$
- edms beyond standard SUSY contributions

I. Higgs mass scale

II. Fermion spectrum

no explanation
within SM

- ▲ I. forces us to deal with II.
- ▲ Easiest and dullest $\Rightarrow$ Flavor Symmetries
- ▲ SUSY-Compositeness helps explain I & II at "testable" energy
- ▲ electron edm is crucial