

#### TuttiFrutti (TF) Method: A theoretical journey towards the 6PN two-body Hamiltonian

[Classical GR interaction of binary systems, combining information via PN, PM, MPM, SF, EOB, EFT, Delaunay averaging formalisms]

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[Based on works done in collaboration with T. Damour, A. Geralico, S. Laporta, P. Mastrolia]

*EFT Methods from Bound States to Binary Systems* University of Padova, 28-30 October, 2020



two-body Hamiltonian

two-body Hamiltonian

# Main steps of the TF method in a schematic representation





 $\rm MPM$   $\rightarrow$  Separate off the nonlocal part of the Hamiltonian and perform Delaunay averaging (leading to averaged  $\rm H_{nonloc\,)}$ 

SF  $\rightarrow$  Compute  $z_1 @ O(eccentricity^n)$ 

EOB  $\rightarrow$  Use the first law of the two-body dynamics to translate the redshift information into a O(p<sub>r</sub><sup>n</sup>) in the EOB Hamiltonian in DJS gauge

Determine  $H_{loc,1SF}$  by «subtracting» the averaged  $H_{nonloc}$ 

 $PM \rightarrow Use EOB-PM$  theory to determine most of the nonlinear-in- v dependence of the coefficients in the EOB Hamiltonian.

Evaluate flux split integrals @ 2PN to have the nonlocal Hamiltonian @6PN since nonlocal effects start @4PN

### Nonlocal effects @6PN

Strategy: Computed in h coords and translated in EOB coords

Integrated flux split equals averaged Hamiltonian

#### **First accomplishment**

Compare averaged Hamiltonian in h and EOB coordinates Fix the nonlocal part of EOB potentials up to 6PN

## Passing from nonloc to loc...

Second accomplishment

Evaluate z<sub>1</sub> from 1SF, expanded in eccentricity, which incorporates both local and nonlocal effects

Fix the nonlocal part of EOB potentials up to 6PN

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Evaluate z<sub>1</sub> in EOB using the local part of the Hamiltonian in a parametrized form with unknown coefficients

Third accomplishment

## Key tool: 1SF 7

 $\nu = \frac{m_1 m_2}{(m_1 + m_2)^2}$ 

sym mass ratio

Compare the two and fix the local EOB potentials at O(v) since we can use 1SF only Compute the scattering angle in EOB with a parametrized v dependence of unknown coefficients

9

10

8

Use a special «rule» to fix the allowed dependence on v of the local EOB potentials

Determine all missing coefficients except for 2 @5PN and 4 @6PN

#### GI quantities involved in the TF approach

- ightarrow Delaunay averaged NL Hamiltonian
- $\rightarrow$  Redshift
- $\rightarrow$  Scattering angle

#### Nonlocal action at 6PN order

As a starting point we need an explicit expression for the nonlocal part of the action Computed in:

$$S_{\text{nonloc},f}^{\leq 6\text{PN}} = S_{\text{nonloc},f}^{4+5+6\text{PN}} + S_{\text{nonloc}}^{5.5\text{PN}}$$

$$S_{\text{nonloc},f}^{4+5+6\text{PN}}[x_{1}(s_{1}), x_{2}(s_{2})] = -\int dt \, H_{\text{nonloc},f}^{4+5+6\text{PN}}(t)$$

$$H_{\text{nonloc},f}^{4+5+6\text{PN}}(t) = \frac{G^{2}\mathcal{M}}{c^{3}} Pf_{2r_{12}^{f}(t)/c} \int \frac{dt'}{|t-t'|} \mathcal{F}_{2\text{PN}}^{\text{split}}(t,t')$$

D. Bini, T. Damour, A. Geralico, *Binary dynamics at the fifth and fifth-and-a-half post-Newtonian orders*, Phys. Rev. D 102, 024062 (2020)

> Flexed version of the distance among the two bodies

Total ADM conserved mass-energy of the system At the 6PN accuracy, the nonlocal action can be linearly decomposed into its 4 + 5 + 6PN piece (depending on the choice of a flexibility function f), and a 5.5PN piece (not depending on f).

#### It is enough to work in h coordinates...

$$r^{f}(t) = f(t)r^{h}(t)$$

"Flexed" version of the harmonic coordinate radial distance between the two bodies. f(t) denotes a function of the instantaneous state of the system. Its introduction is convenient for reasons which will be clear later.

To show the essential steps of the TF method it is enough to work in h coordinates. The flexibility function simply generates another piece to be added the h-coordinate nonlocal Hamiltonian which can be cosidered later.

The h-coordinates nonlocal part of the 2-body Hamiltonian is

$$H_{\text{nonloc,h}}^{4+5+6\text{PN}}(t) = \frac{G\mathcal{M}}{c^3} \operatorname{Pf}_{2r_{12}^h(t)/c} \int \frac{dt'}{|t-t'|} \mathcal{F}_{2\text{PN}}^{\text{split}}(t,t') \,.$$

#### What should I do?

Evaluate flux split integrals @ 2PN to have the nonlocal Hamiltonian @6PN since nonlocal effects start @4PN

Compute the flux split integral

$$H_{\text{nonloc,h}}^{4+5+6\text{PN}}(t) = \frac{G\mathcal{M}}{c^3} \operatorname{Pf}_{2r_{12}^h(t)/c} \int \frac{dt'}{|t-t'|} \mathcal{F}_{2\text{PN}}^{\text{split}}(t,t') \,.$$

H nonloc,h is the integral over t' of the flux split F(t,t')/|t-t'| To have H nonloc,h at 4PN one needs the flux split at the N level To have H nonloc,h at 5PN one needs the flux split at the 1PN level To have H nonloc,h at 6PN one needs the flux split at the 2PN level

T. Damour, P. Jaranowski and G. Schafer, Phys. Rev. D 89, 064058 (2014)

Up to 2 PN the flux split follows trivially from the GW flux (function), while beyond 2PN tail contributions appear, mix, and make things not so simple.

#### I need the flux function 2PN

# What is the difference between flux and flux split at 2PN?

t,t

$$\mathcal{F}_{2PN}(t) = \frac{G}{c^5} \frac{1}{5} I_{ab}^{(3)}(t) I_{ab}^{(3)}(t) + \dots \text{ GW flux}$$

Various multipolar contributions as the building block of the flux function: electric quadrupole, electric octupole, magnetic quadrupole, electric hexadecapole, magnetic octupole.

 $\mathcal{F}_{2\rm PN}^{\rm split}(t,t') = \frac{G}{c^5} \frac{1}{5} I_{ab}^{(3)}(t) I_{ab}^{(3)}(t') + \dots$ 

t,ť

It is the flux split version of the GW flux which enters the nonlocal part of the two-body Hamiltonian!

#### I need the h multipole moments @ 2PN

Building blocks of the nonlocal Hamiltonian



See e.g. K.G. Arun, L. Blanchet, B.R. Iyer and M.S.S. Qusailah, Phys. Rev. D 77, 064034 (2008)

$$H^{4+5+6\rm PN}_{\rm nonloc,h}(t) = -\frac{G^2 H}{c^5} \mathrm{Pf}_{2s/c} \int \frac{d\tau}{|\tau|} \mathcal{F}^{\rm split}_{2\rm PN}(t,t+\tau) + 2\frac{G^2 H}{c^5} \mathcal{F}^{\rm split}_{2\rm PN}(t,t) \ln\left(\frac{r_{12}^h(t)}{s}\right)$$



Evaluate the multipole moments along a 2PN ellipticlike orbit, obtaining the integrand as a function of t and t' so to proceed with the t' and t integrations.

## I need the parametric equations of the orbit

[The integration is conveniently performed by expanding the integrands in powers of a small eccentricity!]

#### **2PN quasi-Keplerian parametrization** of ellipticlike orbits

 $\begin{aligned} r &= a_r (1 - e_r \cos u) \,, \\ \ell &= n(t - t_P) = u - e_t \sin u + f_t \sin V + g_t (V - u) \,, \\ \bar{\phi} &= \frac{\phi - \phi_P}{K} = V + f_\phi \sin 2V + g_\phi \sin 3V \,, \qquad V(u) = 2 \arctan\left[\sqrt{\frac{1 + e_\phi}{1 - e_\phi}} \tan \frac{u}{2}\right] \end{aligned}$ 

Different types of eccentricity: temporal, radial, azimuthal. All orbital elements are coordinate-dependent. They can however be expressed in terms of GI variables like E and J.

Source of computational difficulties when evaluating functions of the two-body dynamics along the orbit. Simplifications arise when taking a series expansion in eccentricity

The parametrization depends on the (coordinate dependent) orbital elements! → We need the h coordinate orbital elements at 2PN

#### And now?



I can finally take the (Delaunay) average of the nonlocal Hamiltonian in h coordinates, leading to a GI invariant quantity.

$$\langle H_{\rm nonloc,h}^h \rangle \equiv \langle H_{\rm nonloc,h}^{4+5+6\rm PN} \rangle = F^h(a_r^h, e_t^h) ,$$

$$F^{h}(a_{r}^{h}, e_{t}^{h}) = \frac{\nu^{-}}{(a_{r}^{h})^{5}} \left[ \mathcal{A}^{4\mathrm{PN}}(e_{t}^{h}) + \mathcal{B}^{4\mathrm{PN}}(e_{t}^{h}) \ln a_{r}^{h} \right]$$
$$+ \frac{\nu^{2}}{(a_{r}^{h})^{6}} \left[ \mathcal{A}^{5\mathrm{PN}}(e_{t}^{h}) + \mathcal{B}^{5\mathrm{PN}}(e_{t}^{h}) \ln a_{r}^{h} \right]$$
$$+ \frac{\nu^{2}}{(a_{r}^{h})^{7}} \left[ \mathcal{A}^{6\mathrm{PN}}(e_{t}^{h}) + \mathcal{B}^{6\mathrm{PN}}(e_{t}^{h}) \ln a_{r}^{h} \right]$$

Explicit expressions in successive papers at 4,5 and 6PN. Polynomial in eccentricity, computed up to a fixed maximum power.

#### The same average but within EOB framework... We have computed



GI

$$\langle H_{\rm nonloc,h}^{4+5+6\rm PN} \rangle = F^h(a_r^h, e_t^h)$$

in harmonic coordinates. We can do the same in EOB coordinates and the end result is the same due to its GI content

$$\langle H_{\rm nonloc,h}^{\rm eob} \rangle = \langle H_{\rm nonloc,h}^h \rangle$$

We need the EOB orbital elements and their relations with the h orbital elements.

If the EOB part (lhs) contains undetermined coefficients these will be fixed by the above equality.

#### The EOB Hamiltonian (EOB coords)

$$H = M\sqrt{1 + 2\nu(\hat{H}_{\text{eff}} - 1)}$$

- →Both the local and nonlocal information concerning the Hamiltonian are encoded into the effective Hamiltonian.
- →The effective Hamiltonian is written in terms of several potentials characterized by both a local and nonlocal part.
- →We are interested in the nonlocal part to compare with our previous computation h-coordinate based.

#### **EOB total effective Hamiltonian**

 $\widehat{H}_{\text{eff}}^2 = A(u;\nu) \Big( 1 + p_{\phi}^2 u^2 + A(u;\nu) \bar{D}(u;\nu) p_r^2 + \widehat{Q}(u,p_r;\nu) \Big)$ 

 $\widehat{Q}(u, p_r; \nu) = p_r^4 q_4(u; \nu) + p_r^6 q_6(u; \nu) + p_r^8 q_8(u; \nu) + p_r^{10} q_{10}(u; \nu) + \dots$ 

DJS gauge: expansion in powers of p<sub>r</sub>

$$\begin{aligned} A(u;\nu) &= 1 - 2u + \nu a^{\nu^{1}}(u) + \nu^{2} a^{\nu^{2}}(u) + \nu^{3} a^{\nu^{3}}(u) + \lambda \\ \bar{D}(u;\nu) &= 1 + \nu \bar{d}^{\nu^{1}}(u) + \nu^{2} \bar{d}^{\nu^{2}}(u) + \nu^{3} \bar{d}^{\nu^{2}}(u) + \dots \\ q_{4}(u;\nu) &= \nu q_{4}^{\nu^{1}}(u) + \nu^{2} q_{4}^{\nu^{2}}(u) + \nu^{3} q_{4}^{\nu^{3}}(u) + \dots \\ q_{6}(u;\nu) &= \nu q_{6}^{\nu^{1}}(u) + \nu^{2} q_{6}^{\nu^{2}}(u) + \nu^{3} q_{6}^{\nu^{3}}(u) + \dots \\ q_{8}(u;\nu) &= \nu q_{8}^{\nu^{1}}(u) + \nu^{2} q_{8}^{\nu^{2}}(u) + \nu^{3} q_{8}^{\nu^{3}}(u) + \dots \\ \delta I \\ \delta I \\ \delta I \end{aligned}$$

$$\begin{split} A &= A^{\text{loc},\text{h}} + A^{\text{nonloc},\text{h}} = A^{\text{loc},\text{f}} + A^{\text{nonloc},\text{f}},\\ \bar{D} &= \bar{D}^{\text{loc},\text{h}} + \bar{D}^{\text{nonloc},\text{h}} = \bar{D}^{\text{loc},\text{f}} + \bar{D}^{\text{nonloc},\text{f}},\\ \hat{Q} &= \hat{Q}^{\text{loc},\text{h}} + \hat{Q}^{\text{nonloc},\text{h}} = \hat{Q}^{\text{loc},\text{f}} + \hat{Q}^{\text{nonloc},\text{f}}. \end{split}$$

Use the GI invariance property to impose

 $\langle H_{\rm nonloc,h}^{\rm eob} \rangle = \langle H_{\rm nonloc,h}^h \rangle$ 

and fix all EOB nonloc coefficients (including logs).

Separate EOB potentials into local and nonlocal parts

$$\delta A = a_5^{\text{nonloc}} u^5 + a_6^{\text{nonloc}} u^6 + a_7^{\text{nonloc}} u^7$$

$$\begin{split} \delta \bar{D} &= \bar{d}_4^{\text{nonloc}} u^4 + \bar{d}_5^{\text{nonloc}} u^5 + \bar{d}_6^{\text{nonloc}} u^6 \,, \\ \delta \widehat{Q} &= p_r^4 (q_{43}^{\text{nonloc}} u^3 + q_{44}^{\text{nonloc}} u^4 + q_{45}^{\text{nonloc}} u^5) \\ &+ p_r^6 (q_{62}^{\text{nonloc}} u^2 + q_{63}^{\text{nonloc}} u^3 + q_{64}^{\text{nonloc}} u^4) \\ &+ p_r^8 (q_{81}^{\text{nonloc}} u + q_{82}^{\text{nonloc}} u^2 + q_{83}^{\text{nonloc}} u^3) \\ &+ p_r^{10} (q_{10,0}^{\text{nonloc}} + q_{10,1}^{\text{nonloc}} u + q_{10,2}^{\text{nonloc}} u^2) \end{split}$$

#### **Intermediate result**

 $\langle H_{\rm nonloc,h}^{\rm eob} \rangle = \langle H_{\rm nonloc,h}^h \rangle$ 



The nonlocal part of all the EOB potentials (a, d, q<sub>n</sub>) needed @ 6PN is fully known.



$$\widehat{H}_{\text{eff}}^2 = A(u;\nu) \Big( 1 + p_{\phi}^2 u^2 + A(u;\nu) \overline{D}(u;\nu) p_r^2 + \widehat{Q}(u,p_r;\nu) \Big)$$
  
$$\widehat{Q}(u,p_r;\nu) = p_r^4 q_4(u;\nu) + p_r^6 q_6(u;\nu) + p_r^8 q_8(u;\nu) + p_r^{10} q_{10}(u;\nu) + \dots$$



TABLE VI: h-route nonlocal EOB coefficients.

$a_7^{ m nl,c}$	$\left(\frac{206740}{567}\ln(2) + \frac{12664}{105} - \frac{4617}{14}\ln(3) - \frac{5044}{405}\gamma\right)\nu$	
	$+\left(-\frac{1139672}{945}\ln(2)+\frac{10132}{105}+\frac{10449}{7}\ln(3)+\frac{101272}{315}\gamma\right)\nu^2$	list
	$+\left(-\frac{112}{5}+32\gamma+\frac{1214624}{945}\ln(2)-\frac{4860}{7}\ln(3)\right)\nu^{3}$	
$a_7^{\mathrm{nl,log}}$	$-\frac{2522}{405}\nu+\frac{50636}{315}\nu^2+16\nu^3$ ,	nor
$\bar{d}_6^{\mathrm{nl,c}}$	$\left(-\frac{6381680}{189}\ln(2) + \frac{2043541}{2835} + \frac{1765881}{140}\ln(3) - \frac{64096}{45}\gamma + \frac{9765625}{2268}\ln(5)\right)\nu$	
	$+\left(\frac{28429312}{189}\ln(2)-\frac{3576231}{70}\ln(3)+\frac{167906}{105}+\frac{302752}{105}\gamma-\frac{9765625}{378}\ln(5)\right)\nu^2$	coe
	$+\left(-\frac{9908480}{63}\ln(2)-\frac{744704}{945}+\frac{9765625}{252}\ln(5)+\frac{2944}{3}\gamma+\frac{1275021}{28}\ln(3)\right)\nu^{3},$	
$\bar{d}_6^{\mathrm{nl,log}}$	$-\frac{32048}{45}\nu + \frac{151376}{105}\nu^2 + \frac{1472}{3}\nu^3$	det
$q_{45}^{\mathrm{nl,c}}$	$\left(\frac{70925884}{63}\ln(2) + \frac{13212013}{5670} - \frac{3873663}{16}\ln(3) - \frac{8787109375}{27216}\ln(5) - \frac{617716}{315}\gamma\right)\nu$	
	$+\left(\frac{92560887}{280}\ln(3)-\frac{12619052648}{2835}\ln(2)-\frac{1437979}{63}+\frac{632344}{315}\gamma+\frac{7755859375}{4536}\ln(5)\right)\nu^2$	con
	$+\left(-\frac{177316}{35}+\frac{11263031264}{2835}\ln(2)+\frac{16544}{9}\gamma-\frac{4091796875}{2268}\ln(5)+\frac{2908467}{20}\ln(3)\right)\nu^{3}$	
$q_{45}^{\mathrm{nl,log}}$	$-\frac{308858}{315}\nu + \frac{316172}{315}\nu^2 + \frac{8272}{9}\nu^3$	Del
$q_{64}^{\mathrm{nl,c}}$	$\left(-\frac{211076833264}{14175}\ln(2) - \frac{137711989}{28350} - \frac{9678652821}{5600}\ln(3) + \frac{447248}{1575}\gamma + \frac{153776136875}{23328}\ln(5)\right)$	
	$+\frac{96889010407}{116640}\ln(7)$ $\nu$	ave
	$+\left(\frac{44592947739}{2800}\ln(3) + \frac{2411178384736}{42525}\ln(2) - \frac{126070663}{4725} - \frac{26848}{175}\gamma - \frac{796015515625}{27216}\ln(5)\right)$	
	$-\frac{96889010407}{19440}\ln(7)$ $\nu^2$	
	$+\left(-\frac{40513708}{4725}-\frac{109566260523}{5600}\ln(3)+\frac{1424826953125}{54432}\ln(5)+\frac{96889010407}{12960}\ln(7)+\frac{2368}{5}\gamma\right)$	
	$-rac{431564554688}{8505} \ln(2)  ight)  u^3$	
$q_{64}^{\mathrm{nl,log}}$	$\frac{223624}{1575}\nu - \frac{13424}{175}\nu^2 + \frac{1184}{5}\nu^3$	
$q_{83}^{\mathrm{nl,c}}$	$\left(\frac{5196312336176}{35721}\ln(2) + \frac{17515638027261}{313600}\ln(3) - \frac{63886617280625}{1016064}\ln(5) - \frac{29247366220639}{933120}\ln(6)\right)$	7)
	$-rac{709195549}{132300} ight) u$	
	$+\left(-\frac{177055674739808}{297675}\ln(2)-\frac{43719724468071}{156800}\ln(3)+\frac{366449151015625}{1524096}\ln(5)+\frac{26506549233}{155520}\right)$	$\frac{199}{\ln(7)}$
	$-rac{1746293}{70} ight) u^2$	
	$+\left(\frac{57604236136064}{99225}\ln(2) + \frac{10467583300341}{39200}\ln(3) - \frac{73366198046875}{381024}\ln(5) - \frac{7709596970957}{38880}\ln(5)\right)$	n(7)
	$-\frac{154862}{21})  \nu^3$	
$q_{83}^{\mathrm{nl,log}}$	0	

List of the EOB nonlocal coefficients determined by comparison of Delaunay averaging

# How can we determine the local part of the EOB potentials?



→Compute a GI invariant quantity (the redshift)
 which incorporates both local and nonlocal effects
 in two different ways:
 1) 1SF (computations only linear in v)

2) Using the full EOB Hamiltonian

 $z_1 = \langle \frac{\partial H}{\partial m_1} \rangle$ 

Compare the two («subtract» nonlocal from the full local plus nonlocal objects), and identify the local part of the EOB potentials! But only linear in v!

### **Computing z<sub>1</sub> 1SF**

$$\begin{split} \delta z_1^{e^8}(u) &= C_3 u^3 + C_4 u^4 + (C_5^c + C_5^{\ln} \ln u) u^5 \\ &+ (C_6^c + C_6^{\ln} \ln u) u^6 + C_{13/2} u^{13/2} \\ &+ (C_7^c + C_7^{\ln} \ln u) u^7 + C_{15/2} u^{15/2} \\ &+ (C_8^c + C_8^{\ln} \ln u + C_8^{\ln^2} \ln^2 u) u^8 + C_{17/2} u^{17/2} \\ &+ (C_9^c + C_9^{\ln} \ln u + C_9^{\ln^2} \ln^2 u) u^9 \\ &+ (C_{19/2}^c + C_{19/2}^{\ln} \ln u) u^{19/2} \\ &+ O_{\ln(u)}(u^{10}) \,, \end{split}$$

 $C_3$ 

 $C_4$ 

 $\begin{array}{c} C_5^{\rm c} \\ C_5^{\rm ln} \end{array}$ 

 $C_6^c$  $C_6^{\ln}$  SF people could have computed this before. Unfortunately, this was not the case and we had to do ourselves.

## External computation, performed by using standard 1SF technology.

 $\frac{\frac{15}{64}}{\frac{3001}{384} - \frac{287}{4096}\pi^2} \frac{\pi^2}{\frac{4597}{96} - \frac{162109375}{2304}\ln(5) - \frac{11332791}{1280}\ln(3) + \frac{55}{6}\gamma + \frac{15967961}{90}\ln(2) - \frac{474715}{196608}\pi^2}{\frac{55}{12}} \pi^2 - \frac{\frac{9863051}{40320}}{\frac{40320}{442368}} + \frac{96889010407}{442368}\ln(7) - \frac{64481546637}{114688}\ln(3) - \frac{5977}{240}\gamma - \frac{16605499789}{5040}\ln(2) + \frac{1466047}{196608}\pi^2 + \frac{4761539921875}{3096576}\ln(5) - \frac{\frac{5977}{480}}{\frac{5977}{480}} + \frac{1000}{1000} + \frac{1000}{10$ 

#### Converting z<sub>1</sub> 1SF information in EOB potentials (analytic 1SF encodes both local+nonlocal effects)

$$q_8^{\nu^1}(u) = B_1 u + B_2 u^2 + B_{5/2} u^{5/2} + B_3 u^3 + B_{7/2} u^{7/2} + (B_4^c + B_4^{\ln} \ln u) u^4 + B_{9/2} u^{9/2} + (B_5^c + B_5^{\ln} \ln u) u^5 + (B_{11/2}^c + B_{11/2}^{\ln} \ln u) u^{11/2} + O_{\ln(u)}(u^6),$$

Subtract the (sum of local and nonlocal) EOB potential obtained from 1SF from their nonlocal parts obtained by using Delaunay averaging to obtain the local part of the EOB potentials.

1SF allows to compute only terms linear in v: all other coefficients should be included in a parametrized form.

 $B_1$ 

 $B_2$ 

 $B_{5/2}$ 

$$\begin{aligned} & -\frac{27734375}{126}\ln(5) + \frac{6591861}{350}\ln(3) + \frac{21668992}{45}\ln(2) - \frac{35772}{175} \\ & \frac{13841287201}{17280}\ln(7) - \frac{393786545409}{156800}\ln(3) - \frac{16175693888}{1575}\ln(2) + \frac{875090984375}{169344}\ln(5) + \frac{5790381}{2450} \\ & +\frac{5994461}{12700800}\pi \end{aligned}$$

#### Intermediate result

All the EOB potentials have been determined for both their **nonlocal part (completely)** and their **local part (at linear order in v).** The high-order in v of the local part of the Hamiltonian is unkown, and written by using parameters!



# Determining the local part of the EOB potential @ O(v<sup>1</sup>)

$$\begin{aligned} a_{4+5+6} PN, loc, f &= \left[ \left( \frac{2275}{512} \pi^2 - \frac{4237}{60} \right) \nu + \left( \frac{41}{32} \pi^2 - \frac{221}{6} \right) \nu^2 \right] u^5 + \left[ \left( -\frac{1026301}{1575} + \frac{246367}{3072} \pi^2 \right) \nu \cdot \left( a_{6,f}^{(\nu)} \right) u^6 \right. \\ &+ \left[ \left( -\frac{2800873}{262144} \pi^4 + \frac{608698367}{1769472} \pi^2 - \frac{1469618167}{907200} \right) \nu + \left( a_{7,f}^{(\nu)} \right) u^6 \right] \\ d_{4+5+6} PN, loc, f &= \left[ \left( \frac{1679}{9} - \frac{23761}{1536} \pi^2 \right) \nu + \left( -260 + \frac{123}{16} \pi^2 \right) \nu^2 \right] u^4 + \left( \frac{331054}{175} \nu - \frac{63707}{512} \nu \pi^2 + \left( a_{5,f}^{(\nu)} \right) u^5 \right. \\ &+ \left[ \left( \frac{229504763}{98304} \pi^2 + \frac{135909}{262144} \pi^4 - \frac{99741733409}{6350400} \right) \nu \cdot \left( a_{6,f}^{(\nu)} \right) u^6 \right] u^6 \right] \\ &+ \left[ \left( \frac{81030481}{65536} \pi^2 - \frac{3492647551}{423360} \right) \nu \cdot \left( a_{4,f}^{(\nu)} \right) u^5 \right] u^5 \right] u^6 \\ &+ \left[ \left( \frac{81030481}{65536} \pi^2 - \frac{3492647551}{423360} \right) \nu \cdot \left( a_{4,f}^{(\nu)} \right) u^5 \right] u^5 \right] u^5 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{294000} \right) \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{294000} \right) \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{294000} \right) \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{294000} \right) \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{294000} \right) \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{294000} \right) \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{294000} \right) \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{3294000} \right) u^4 \\ &+ \left[ \left( -\frac{9733841}{327680} \pi^2 - \frac{112218283}{3294000} \right) u^4 \\ &+ \left[ \left( -\frac{97447}{560} \nu + \left( a_{6,f}^{(\nu)} \right) u^3 \right] u^4 \\ &+ \left[ \left( -\frac{97447}{560} \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{97447}{560} \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{976}{560} \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{976}{560} \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{976}{560} \nu + \left( a_{6,f}^{(\nu)} \right) u^4 \right] u^4 \\ &+ \left[ \left( -\frac{976}{560} \mu + \left( a_{6,f}^{(\nu)} \right) u$$

# Using the v dependence of the scattering angle to determine most of the v structure of the local Hamiltonian

 $\chi_5^{\rm loc,f}$ 

The scattering angle depend on unknown EOB parameters (at either 5PN and 6PN)

$$\frac{1}{2}(\chi(\gamma,j)+\pi) = -\int_0^{u_{\max}} \frac{\partial}{\partial j} p_r(u;\gamma,j) \frac{du}{u^2}$$

$$\gamma^{2} = \widehat{\mathcal{E}}_{\text{eff}}^{2} = \widehat{H}_{\text{eff,loc,f}}^{2 \text{ EG}}(u, p_{r}, j; \nu)$$
$$= H_{S}^{2} + (1 - 2u)\widehat{Q}_{H \text{ loc,f}}^{\text{EG}}(u, H_{S}; \nu)$$

$$p_r^2(\gamma, j, u) = \frac{\gamma^2 - (1 - 2u) \left(1 + j^2 u^2 + \widehat{Q}_{E \,\text{loc}, f}^{\text{EG}}(u, \gamma; \nu)\right)}{(1 - 2u)^2}$$

$$p_r = p_r^{(0)} + G p_r^{(1)} + G^2 p_r^{(2)} + \dots,$$

$$\begin{split} &= \frac{1}{5p_{\infty}^{5}} - \frac{2}{p_{\infty}^{3}} \eta^{2} + \frac{32 - 8\nu}{p_{\infty}} \eta^{4} \\ &+ \left[ 320 + \left( -\frac{1168}{3} + \frac{41}{8} \pi^{2} \right) \nu + 24\nu^{2} \right] p_{\infty} \eta^{6} \\ &+ \left[ 640 + \left( \frac{5069}{144} \pi^{2} - \frac{227059}{135} \right) \nu \right. \\ &+ \left( -\frac{287}{24} \pi^{2} + \frac{7342}{9} \right) \nu^{2} - 40\nu^{3} \right] p_{\infty}^{3} \eta^{8} \\ &+ \left[ \frac{1792}{5} + \left( -\frac{1460479}{525} + \frac{111049}{960} \pi^{2} \right) \nu \right. \\ &+ \left( \frac{41026}{15} - \frac{40817}{640} \pi^{2} - \frac{4}{15} \bar{d}_{5}^{\nu^{2}} \right) \nu^{2} \\ &+ \left( -\frac{11108}{9} + \frac{451}{24} \pi^{2} \right) \nu^{3} + 56\nu^{4} \right] p_{\infty}^{5} \eta^{10} \\ &+ \left[ \left( \frac{93031}{2304} \pi^{2} - \frac{498343703}{604800} \right) \nu \right. \\ &+ \left( \frac{2827607}{1152} - \frac{31633}{768} \pi^{2} \right) \nu^{2} \\ &+ \left( \frac{205}{16} \pi^{2} - \frac{253361}{96} \right) \nu^{3} + \frac{212879}{384} \nu^{4} + \frac{63}{64} \nu^{5} \\ &- 2q_{3\rm EG}^{4} - 4q_{4\rm EG}^{3} - \frac{4}{3} q_{5\rm EG}^{2} \right] p_{\infty}^{7} \eta^{12} \,. \end{split}$$





$$\widetilde{\chi}_n^{\mathrm{loc,f}}(\gamma;\nu) \equiv \left[h(\gamma;\nu)\right]^{n-1} \chi_n^{\mathrm{loc,f}}(\gamma;\nu)$$

$$h(\gamma; \nu) \equiv \sqrt{1 + 2\nu(\gamma - 1)} = \frac{H}{Mc^2}.$$

The total (local plus nonlocal) scattering angle satisfies the property shown by Damour

$$C_n^{\text{tot}}: \widetilde{\chi}_n^{\text{tot}}(\gamma; \nu) = P_{d_n}^{\gamma}(\nu); \text{ with } d_n \equiv \left[\frac{n-1}{2}\right]$$

The introduction of the flexibility function allows both the local and the nonlocal parts to satisfy this rule!

T. Damour, *Classical and quantum scattering in post-Minkowskian gravity*, Phys. Rev. D **102** (2), 024060 (2020). [arXiv:1912.02139 [gr-qc]].

#### Result

We have determined the local 5PN (first) and 6PN (later) EOB Hamiltonian modulo

2 unknown at 5PN and 4 more unknown at 6PN.

Note that thinking of all possible terms entering the 6PN real Hamiltonian, we have determined 147 terms of the 151 total number!





#### What remains....

$A(u; \nu)$ $\bar{D}(u; \nu)$	$= 1 - 2u + \nu a^{\nu^{1}}(u) + \nu^{2} a^{\nu^{2}}(u) + \nu^{3} a^{\nu^{3}}(u) + \cdots$ $= 1 + \nu d^{\nu^{1}}(u) + \nu^{2} d^{\nu^{2}}(u) + \nu^{3} d^{\nu^{2}}(u) + \cdots$ List of the <i>j</i> -route EOB potentials in <i>p<sub>r</sub></i> -gauge. Reminder of EOB
$\begin{array}{c} a_5^{\rm loc,f} \\ a_6^{\rm loc,f} \\ a_7^{\rm loc,f} \\ \bar{d}_4^{\rm loc,f} \\ \bar{d}_5^{\rm loc,f} \\ \bar{d}_6^{\rm loc,f} \end{array}$	$ \begin{array}{l} \left(-\frac{4237}{60}+\frac{2275}{512}\pi^2\right)\nu+\left(\frac{41}{32}\pi^2-\frac{221}{6}\right)\nu^2 \\ \left(-\frac{1026301}{1575}+\frac{246367}{3072}\pi^2\right)\nu+a_6^{\nu}\nu^2+4\nu^3 \\ \left(-\frac{2800873}{262144}\pi^4+\frac{608698367}{1769472}\pi^2-\frac{1469618167}{907200}\right)\nu+a_7^{\nu}\nu^2+a_7^{\nu}\nu^3 \\ \left(\frac{1679}{9}-\frac{23761}{1536}\pi^2\right)\nu+\left(\frac{123}{16}\pi^2-260\right)\nu^2 \\ \left(\frac{331054}{175}-\frac{63707}{512}\pi^2\right)\nu+d_5^{\nu}\nu^2+\left(-\frac{205}{16}\pi^2+\frac{1069}{3}\right)\nu^3 \\ \left(\frac{229504763}{98304}\pi^2+\frac{135909}{262144}\pi^4-\frac{99741733409}{6350400}\right)\nu+d_6^{\nu}\nu^2 + \left(\frac{45089}{72}-\frac{44489}{1536}\pi^2+d_5^{\nu}\nu^2-15a_6^{\nu}\nu^2\right)\nu^3-48\nu^4 \end{array}$
$\begin{array}{c} q_{43}^{\rm loc,f} \\ q_{44}^{\rm loc,f} \\ q_{45}^{\rm loc,f} \\ q_{62}^{\rm loc,f} \\ q_{63}^{\rm loc,f} \\ q_{63}^{\rm loc,f} \\ q_{64}^{\rm loc,f} \\ q_{82}^{\rm loc,f} \end{array}$	$ \begin{array}{c} 20\nu - 83\nu^2 + 10\nu^3 \\ \left(\frac{1580641}{3150} - \frac{93031}{1536}\pi^2\right)\nu + \left(-\frac{2075}{3} + \frac{31633}{512}\pi^2\right)\nu^2 + \left(640 - \frac{615}{32}\pi^2\right)\nu^3 \\ \left(\frac{81030481}{65536}\pi^2 - \frac{3492647551}{423360}\right)\nu + q_{45}^{\nu^2}\nu^2 + \left(-\frac{14}{3}\vec{d_5}^{\nu^2} + \frac{36677}{1152}\pi^2 - \frac{474899}{216}\right)\nu^3 + \left(\frac{1435}{32}\pi^2 - \frac{7375}{6}\right)\nu^4 \\ - \frac{9}{5}\nu - \frac{27}{5}\nu^2 + 6\nu^3 \\ \frac{123}{10}\nu - \frac{69}{5}\nu^2 + 116\nu^3 - 14\nu^4 \\ \left(-\frac{9733841}{327680}\pi^2 - \frac{112218283}{294000}\right)\nu + \left(\frac{156397}{1280}\pi^2 - \frac{21996581}{21000}\right)\nu^2 + \left(\frac{6977}{6} - \frac{29665}{256}\pi^2\right)\nu^3 + \left(\frac{287}{8}\pi^2 - \frac{3640}{3}\right)\nu^4 \\ \frac{6}{7}\nu + \frac{18}{7}\nu^2 + \frac{24}{7}\nu^3 - 6\nu^4 \end{array} $

$$\begin{pmatrix} -\frac{9733841}{327680}\pi^2 - \frac{112218283}{294000} \end{pmatrix} \nu + \left( \frac{156397}{1280}\pi^2 - \frac{21996581}{21000} \right) \nu^2 + \left( \frac{6977}{6} - \frac{29665}{256}\pi^2 \right) \nu^3 + \left( \frac{287}{8}\pi^2 - \frac{3640}{3} \right) \nu^4 \\ \frac{6}{7}\nu + \frac{18}{7}\nu^2 + \frac{24}{7}\nu^3 - 6\nu^4 \\ -\frac{7447}{560}\nu - \frac{963}{56}\nu^2 - \frac{117}{10}\nu^3 - 147\nu^4 + 18\nu^5 \\ -\frac{11}{21}\nu - \frac{11}{7}\nu^2 - \frac{20}{7}\nu^3 - \frac{5}{3}\nu^4 + 6\nu^5$$

#### H EOB boxed

 $q_{83}^{\mathrm{loc},\mathrm{f}}$  $q_{10,2}^{\mathrm{loc},\mathrm{f}}$ 

## Schematic representation of the results obtained at 5PN

The real Hamiltonian @5PN contains 97 coefficients of which we have determined 95!



#### What is needed to complete the 2 body dynamics at a fixed PM level



Each vertical column of dots describes the PN expansion (keyed by powers of p<sup>2</sup>) of an energy-dependent function parametrizing the scattering angle.

The various columns at a given PM level correspond to increasing powers of v.

D. Bini, T. Damour and A.Geralico, ``Sixth post-Newtonian local-in-time dynamics of binary systems," Phys. Rev. D 102, no.2, 024061 (2020) [arXiv:2004.05407 [gr-qc]].

Schematic representation of the irreducible information contained, at each PM level (keyed by a power of u = GM/r), in the local dynamics.



The next challenge...



#### Can we improve the present knowledge of the nonlocal part of the scattering angle?

#### Reminder

To be evaluated along hyperboliclike orbits, expanding the flux split (@2 PN) in 1/eccentricity, i.e., computing LO, NLO, NNLO,....

$$W^{\text{nonloc,h}}(E, J; \nu) = \frac{GE}{c^5} \times \\ \operatorname{Pf}_{2r_{12}^h(t)/c} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \frac{dtdt'}{|t-t'|} \mathcal{F}_{\mathrm{GW}}^{\mathrm{split}}(t, t')$$

$$W^{\rm nonloc,h}(E,J;\nu) = \int_{-\infty}^{+\infty} dt \, H^{\rm nonloc,h}(t)$$

$$\chi^{\text{nonloc,h}}(E,J,\nu) = \frac{\partial W^{\text{nonloc,h}}(E,J,\nu)}{\partial J}$$

$$\frac{1}{2}\chi^{\text{tail,h}} = \frac{1}{2M^{2}\nu}\frac{\partial}{\partial j}W^{\text{tail,h}} 
= \frac{\chi_{4}^{\text{tail,h}}(\bar{E},\nu)}{j^{4}} + \frac{\chi_{5}^{\text{tail,h}}(\bar{E},\nu)}{j^{5}} + \frac{\chi_{6}^{\text{tail,h}}(\bar{E},\nu)}{j^{6}} 
+ \frac{\chi_{7}^{\text{tail,h}}(\bar{E},\nu)}{j^{7}} + O\left(\frac{1}{j^{8}}\right), 
\frac{\chi_{4}^{\text{tail,h}}}{j^{4}} = \frac{1}{2M^{2}\nu}\frac{\partial}{\partial j}W^{\text{tail,h LO}_{j}}, \quad O(G^{4}) 
\frac{\chi_{5}^{\text{tail,h}}}{j^{5}} = \frac{1}{2M^{2}\nu}\frac{\partial}{\partial j}W^{\text{tail,h NLO}_{j}}, 
\frac{\chi_{6}^{\text{tail,h}}}{j^{6}} = \frac{1}{2M^{2}\nu}\frac{\partial}{\partial j}W^{\text{tail,h NNLO}_{j}} 
\frac{\chi_{6}^{\text{tail,h}}}{j^{7}} = \frac{1}{2M^{2}\nu}\frac{\partial}{\partial j}W^{\text{tail,h NNLO}_{j}}. \quad O(G^{7})$$

#### Notation used in our paper(s)

$$\frac{1}{2}\chi^{\text{tail,h}}(p_{\infty}, j; \nu) = \nu \frac{p_{\infty}^{4}}{j^{4}} \left( A_{0}^{\text{tail,h}}(p_{\infty}; \nu) + \frac{A_{1}^{\text{tail,h}}(p_{\infty}; \nu)}{p_{\infty}j} + \frac{A_{2}^{\text{tail,h}}(p_{\infty}; \nu)}{(p_{\infty}j)^{2}} + \frac{A_{3}^{\text{tail,h}}(p_{\infty}; \nu)}{(p_{\infty}j)^{3}} + \cdots \right) + \frac{A_{2}^{\text{tail,h}}(p_{\infty}; \nu)}{(p_{\infty}j)^{3}} + \cdots \right)$$

$$MIO \qquad \text{NNLO}$$

$$A_{0}(p_{\infty}; \nu) = A_{0}^{\text{tail,h,N}} + A_{0}^{\text{tail,h,1PN}} + A_{0}^{\text{tail,h,2PN}} + \cdots + A_{1}(p_{\infty}; \nu) = A_{0}^{\text{tail,h,N}} + A_{1}^{\text{tail,h,1PN}} + A_{1}^{\text{tail,h,2PN}} + \cdots + A_{2}(p_{\infty}; \nu) = A_{2}^{\text{tail,h,N}} + A_{2}^{\text{tail,h,1PN}} + A_{2}^{\text{tail,h,2PN}} + \cdots + A_{3}(p_{\infty}; \nu) = A_{3}^{\text{tail,h,N}} + A_{3}^{\text{tail,h,1PN}} + A_{3}^{\text{tail,h,2PN}} + \cdots$$

#### NNLO O(G<sup>6</sup>)

At NNLO we found large integrands which made difficult even their numerical integration.

The collaboration with **S. Laporta** and **P. Mastrolia** has allowed to obtain **high-precision numerical values first**, e.g.,

$Q_{20}$	524.7672921802125843427359557031017584761419995573690119377287112384988398300977120939070371581066666666666666666666666666666666666
	9606083170623899520567705206794678374496647513473011101045588318417017082934721207112410611316558831841701708293472120711241061131656666666666666666666666666666666
	8613485679

Successively, we could fit these results with **PSLQ** algorithm

$Q_{20}$	$\frac{25883}{1800}$ -	$-\frac{22333}{140}$ K -	$\frac{625463}{3360}\pi$ -	$\frac{361911}{560}\pi\ln 2$ +	$-\frac{99837}{160}\pi\zeta(3)$

And finally we could analytically check the results by using HPLs of w=4

Noticeably, in the intermediate results the Catalan constant and  $\zeta(3)$  enter. However, the first cancels out in the scattering angle while BOTH cancel out in the periastron advance, leading to a rational coefficient only (at NNLO)

#### The NNLO accomplishment

Integral entering our computation (in its reduced, final form)

$$J(x) \equiv \int_{0}^{1} dT (1 - x^{2}) \times$$
Expression in terms of HPLs w=4  
$$\times \frac{16 \operatorname{arctanh}^{3}(T) - 3 \operatorname{Li}_{3} \left[ \left( \frac{(1 - T)(1 - x)}{(1 + T)(1 + x)} \right)^{2} \right]}{2i(T + x)(T + 1/x)}$$
Expression in terms of HPLs w=4  
sought for result  
$$J(i) = -\frac{1}{2}\pi^{2} \operatorname{K} + \frac{9}{2}\pi\zeta(3)$$

# NNNLO: Further developments and recent accomplishments

→ Nonlocal part of the scattering angle  $@N^{3}LO$  (still unpublished) → Interplay of works in the ordinary space and in the Fourier space (better understanding of these two complementary povs)

Bini, Damour, Geralico, Laporta, Mastrolia, in its final phase....

Computations in the Fourier space are apparently simpler than the corresponding in the ordinary space Worki space but pr integra

Working in the ordinary space is more immediate but produces very large integrals to be computed

$$\begin{split} A_0^{\text{tail,h,N}} &= \pi \left[ -\frac{37}{5} \ln \left( \frac{p_{\infty}}{2} \right) - \frac{63}{4} \right], \\ A_0^{\text{tail,h,1PN}} &= \pi \left[ \left( -\frac{1357}{280} + \frac{111}{10} \nu \right) \ln \left( \frac{p_{\infty}}{2} \right) - \frac{2753}{1120} + \frac{1071}{40} \nu \right] p_{\infty}^2, \\ A_0^{\text{tail,h,2PN}} &= \pi \left[ \left( -\frac{27953}{3360} + \frac{2517}{560} \nu - \frac{111}{8} \nu^2 \right) \ln \left( \frac{p_{\infty}}{2} \right) - \frac{155473}{8960} + \frac{109559}{40320} \nu - \frac{186317}{5040} \nu^2 \right] p_{\infty}^4 \end{split}$$

$$\begin{split} A_1^{\text{tail,h,N}} &= -\frac{6656}{45} - \frac{6272}{45} \ln \left( 4\frac{p_{\infty}}{2} \right), \\ A_1^{\text{tail,h,1PN}} &= \left[ \left( -\frac{74432}{525} + \frac{13952}{45} \nu \right) \ln \left( 4\frac{p_{\infty}}{2} \right) + \frac{114368}{1125} + \frac{221504}{525} \nu \right] p_{\infty}^2, \\ A_1^{\text{tail,h,2PN}} &= \left[ \left( -\frac{881392}{11025} + \frac{288224}{1575} \nu - \frac{21632}{45} \nu^2 \right) \ln \left( 4\frac{p_{\infty}}{2} \right) + \frac{48497312}{231525} - \frac{5134816}{23625} \nu - \frac{25465952}{33075} \nu^2 \right] p_{\infty}^4 \end{split}$$

$$\begin{split} A_2^{\text{tail,h,N}} &= \pi \left[ -122 \ln \left( \frac{p_{\infty}}{2} \right) - \frac{99}{4} - \frac{2079}{8} \zeta(3) \right], \\ A_2^{\text{tail,h,1PN}} &= \pi \left[ \left( \frac{811}{2} \nu - \frac{13831}{56} \right) \ln \left( \frac{p_{\infty}}{2} \right) - \frac{41297}{112} - \frac{9216}{7} \ln(2) + \frac{49941}{64} \zeta(3) + \left( \frac{3303}{4} \zeta(3) + \frac{1937}{8} \right) \nu \right] p_{\infty}^2, \\ A_2^{\text{tail,h,2PN}} &= \pi \left[ \left( \frac{64579}{1008} - 785\nu^2 + \frac{75595}{168} \nu \right) \ln \left( \frac{p_{\infty}}{2} \right) - \frac{40711}{128} \zeta(3) + \frac{1033549}{4536} + \frac{10704}{7} \ln(2) \right. \\ &+ \left( \frac{75520}{21} \ln(2) + \frac{8008171}{8064} - \frac{660675}{256} \zeta(3) \right) \nu + \left( -\frac{100935}{64} \zeta(3) - \frac{583751}{864} \right) \nu^2 \right] p_{\infty}^4, \end{split}$$

#### **At NNNLO?**

#### **One of these integrands**

#### (-86222/315\*(-1+T)\*

«About 2-300 times this block»



Q40num = 2324.\

2945518755079244796761032384673115986096012600915287113669334105968155 9651694952785135498410096031353631295560553413266360924201075481405964 2335939609295101495919259244145297218612521402336653114469276802586585 5818941914970953912386638933986533600312540324370182179717623944381629 7914887792750942760325300191696425379993558158933940567719749455512295 0654024975716659024974807825147802357227383231

```
Q40PSLQ = -(21644756692/7640325) + (13986515191 Pi)/
19051200 - (5616 Pi^2)/25 + (122517 Pi^4)/
2240 - (11235344 Zeta[3])/51975
```

# A comment on our fruitful collaboration with HEP-EFT people

Bini, Damour, Geralico, Laporta, Mastrolia

Fitting is easy only once you know the structure of the expected result!

Producing numerical results for the sought for integrals with high numerical precision (about 500 estimated correct digits).

Fitting of results with available algorithms (PSLQ) and having an «experimental math» solution.

Obtaining the complete analytical solution by using HLP.

Complete agreement with the recent work by J. Blumlein, A. Maier, P. Marquard, G. Schafer, arXiv: 2010.13672 @5PN.

See Eqs. 63-64 there with the determination of the Pi^2 part of the 5PN missing coefficients in TF

$$\begin{split} \bar{d}_{5} &= r_{d_{5}} + \frac{306545}{512}\pi^{2} \\ a_{6} &= r_{a_{6}} + \frac{25911}{256}\pi^{2} \end{split}$$
Typo in our 6PN nonloc paper  
arXiv:2007.11239 [gr-qc]  
Eq. (8.27), first line  $-\frac{155}{12}\ln 2 \longrightarrow -\frac{155}{112}\ln 2$ 

Warning

#### **Conclusions and plans for future works**

Can TF method be applied to reach 7PN?

In principle yes...for both local and nonlocal H!

Can one push forward our knowledge of the nonlocal in time dynamics and obtain more information about the nonlocal part of the scattering angle?

 Large-eccentricity expansion of flux split integrals computed along hyperboliclike motion. [Fruitful collaboration with S. Laporta and P. Mastrolia...We went to N3LO]

Forthcoming 5PN h-coordinate-based complete Hamiltonian? [See a recent work by J. Blumlein, A. Maier, P. Marquard, G. Schafer, arXiv: 2010.13672]

Scattering angle from 1SF?

[Maybe, hopefully we will see first numerical 1SF results before their analytical counterparts]

Scattering angle @ 4PM from amplitudes? [Maybe, these are all new and promising avenues...See e.g., Z. Bern, et al. Phys. Rev. Lett. 122, no.20, 201603 (2019)]

Other recent EFT approaches/results [See S. Foffa , R. Sturani, W. J. Torres Bobadilla, arXiv: 2010.13730; G.Kalin and R.A. Porto, JHEP 01, 072 (2020); L.Blanchet, S.Foffa, F.Larrouturou and R.Sturani, Phys. Rev. D 101, no.8, 084045 (2020); ...]





GCOD IDEA

I would like to thank the organizers of this meeting especially for stimulating interactions among different communities of theoretical physicists, with the aim of making immediate progress from mutual collaborations.