

Neutrino Physics with JUNO

Speaker: Alberto Garfagnini (PI)

Group members: Riccardo Brugnera, Marco Grassi

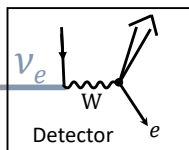
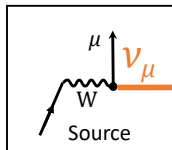
Andrea Serafini, Andrea Triossi



Istituto Nazionale di Fisica Nucleare

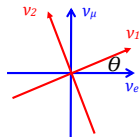
Neutrino oscillations

- neutrinos are created in one flavour, but can be detected in another



- each flavour state is a superposition of mass states

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$



Three flavour oscillation

- The mixing matrix (PMNS matrix) can be written in terms of 3 angles and 1 phase. Usually factorized into components directly related to the experiments:

- the (12) sector: Solar and Reactor

$L/E \sim 15,000 \text{ km/GeV}$

- the (23) sector: Atmospheric and Accelerator

$L/E \sim 500 \text{ km/GeV}$

- the (13) sector: Reactor and Accelerator

$L/E \sim 500 \text{ km/GeV}$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{matrix} c_{ij} = \cos \theta_{ij} \\ s_{ij} = \sin \theta_{ij} \end{matrix}$$

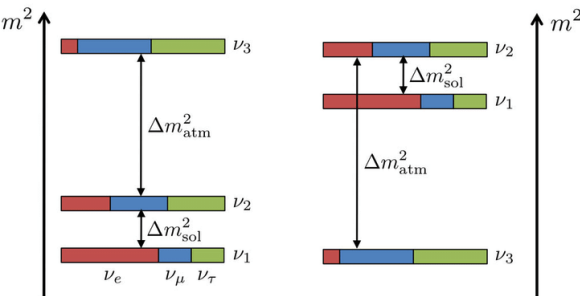
Neutrino Δmass^2 and ordering

- neutrino oscillation experiments can access the mass square differences:

$$\Delta m_{21}^2 = (7.53 \pm 0.18) \times 10^{-5} \text{ eV}^2 \quad |\Delta m_{32}^2| = (2.453 \pm 0.033) \times 10^{-3} \text{ eV}^2$$

R. L. Workman et al. (Particle Data Group), Prog. Theor. Exp. Phys. 2022, 083C01 (2022)

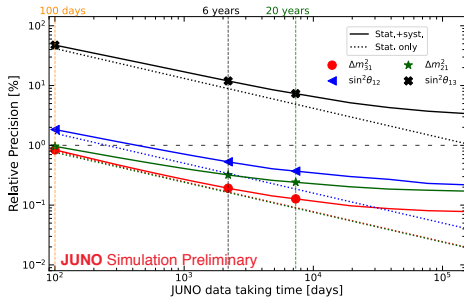
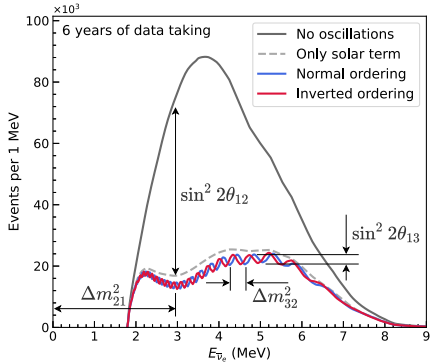
normal hierarchy (NH) inverted hierarchy (IH)



- ν_1 is (by convention) the mass eigenstate with the largest ν_e component
- we do not know which mass eigenstate is the lightest \rightarrow mass hierarchy (ordering) problem:
 - normal ordering: ν_1 lightest
 - inverted ordering: ν_3 lightest

JUNO contribution to neutrino oscillations

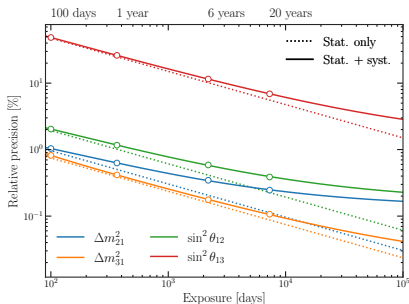
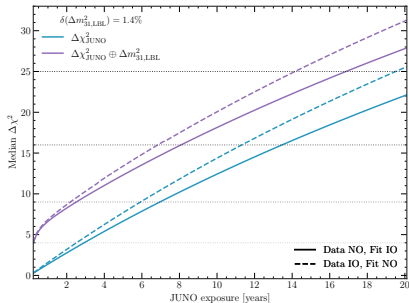
- JUNO will detect reactor $\bar{\nu}_e$ and
 - determine neutrino mass ordering, at $\sim 3 - 4\sigma$ with 6 years of reactor data
 - the measurement is independent of δ_{CP} and $\sin^2 \theta_{23}$
 - measure at sub-percent level the neutrino oscillation parameters
 - $\sin^2 2\theta_{12}$, Δm_{21}^2 and $|\Delta m_{32}^2|$



	Δm_{31}^2	Δm_{21}^2	$\sin^2 \theta_{12}$	$\sin^2 \theta_{13}$
JUNO 6 years	$\sim 0.2\%$	$\sim 0.3\%$	$\sim 0.5\%$	$\sim 12\%$
PDG2020	1.4%	2.4%	4.2%	3.2%

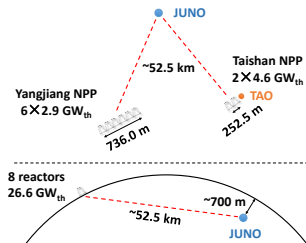
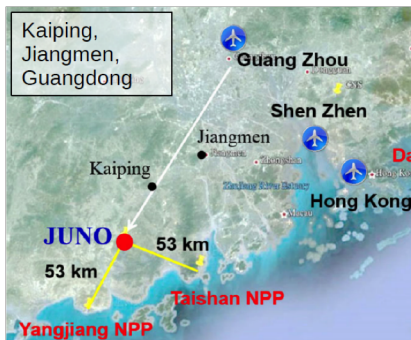
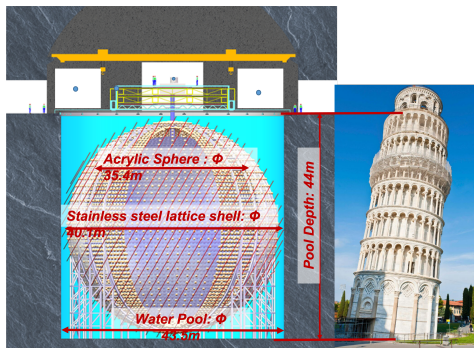
A. Abusleme et al., Chin. Phys. C 46 (2022) 123001

- First experiment to be simultaneously sensitive to two oscillation frequencies
- Neutrino Mass Ordering at 3σ with 6 years of data-taking
- No dependence on δ_{CP} and θ_{23}
- Complementary to long baseline experiments
- Sub-percent precision in less than 2 years on three parameters: Δm_{21}^2 , $\sin^2 \theta_{12}$, and Δm_{31}^2
- Precision measurement of oscillation parameters as a powerful tool to test the standard 3-flavor neutrino framework
- Unitarity of the PMNS matrix (Electron Row Unitarity test)

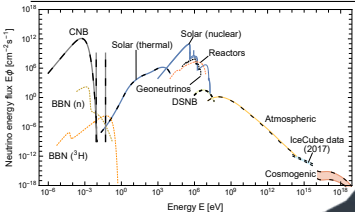


JUNO in a nutshell

- it's going to be **the largest** ever built **liquid scintillator (LS) detector** for neutrino and rare events physics (including dark matter)
- the **main target** is the determination of the **neutrino mass hierarchy**, one of the still unanswered questions in neutrino physics
- thanks to the large mass (20 kt) and overburden (1800 m.w.e.), JUNO will be **able to exploit several neutrino physics channels**



JUNO rich physics programme



SuperNova ν
5k in 10 s (10 kpc)



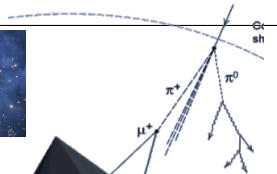
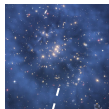
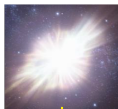
Solar ν
 $O(10 - 1000)/\text{day}$

Reactor ν
 $\sim 60 - 80/\text{day}$

Wimp
(dark matter)
?

Atmospheric ν
10-20/day

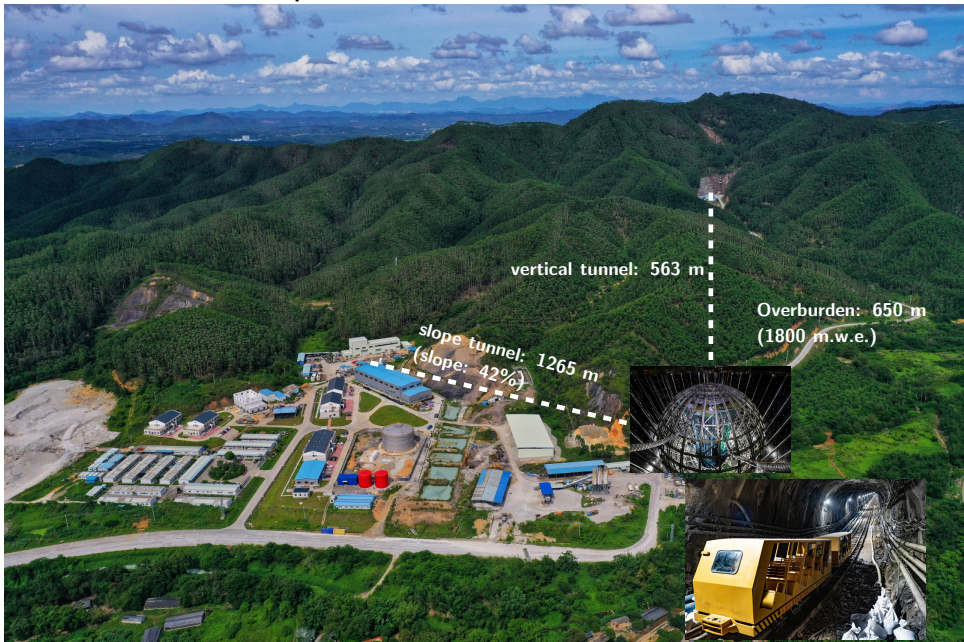
Geo- ν
 $\sim 1 - 2/\text{day}$



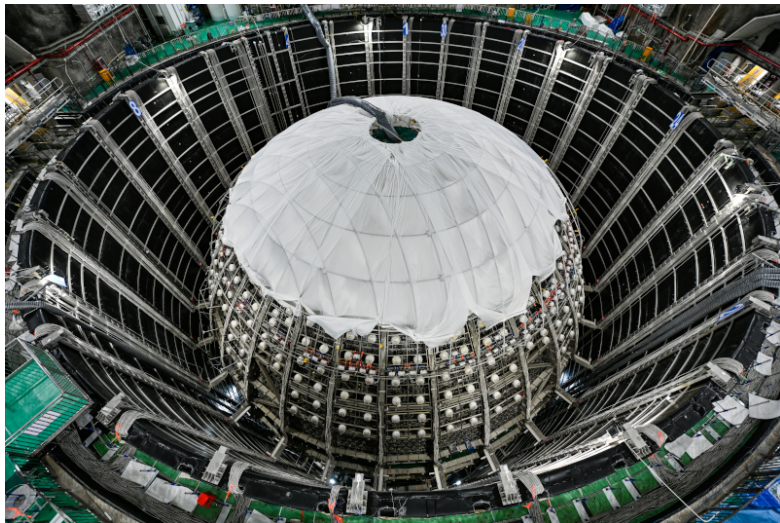
F. An et al., *Neutrino Physics with JUNO*, J. Phys. G 43 (2016) 030401, [arXiv:1507.05613](https://arxiv.org/abs/1507.05613)

A. Abusleme et al., *Prog. Part. Nucl. Phys.* 123 (2022) 103927 [arXiv:2104.02565](https://arxiv.org/abs/2104.02565)

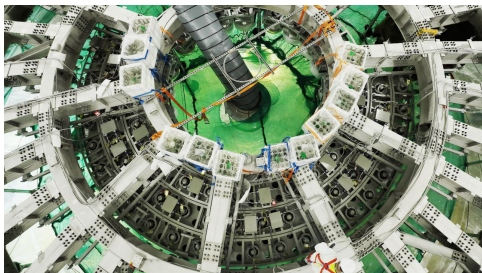
The JUNO experimental site



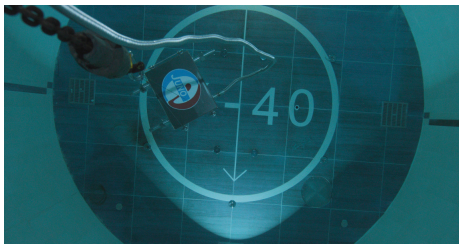
The JUNO detector construction



JUNO large PMT Electronics



Under Water Electronics

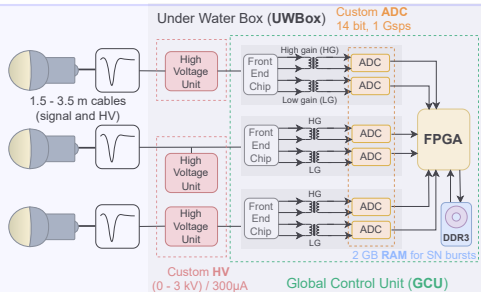


UWBox under water (-40 m)

Wet electronics

Dry electronics

UWBox with electronics



Up to 100 m CAT6 + low Z power cables

LV

Trigger Electronics

Sync link

Back End Card

CLK

Async link

Gbit Enterprise Switch

DAQ

Up to 100 m CAT5 cables

Dry Electronics

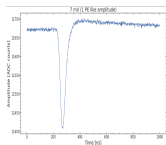


Real-Time charge reconstruction in JUNO

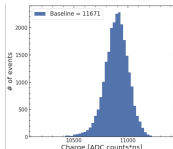
- real-time monitoring of transient phenomena acquired by the JUNO detector (PMT tubes)
- based on a Baseline Tracking algorithm and a Continuous Over-Threshold Integration (COTi)
- firmware developing (VHDL) for the FPGAs of the GCUs



PMT



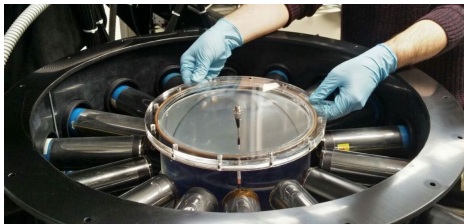
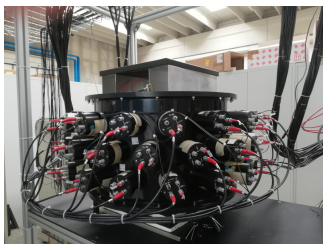
GCU acquisition



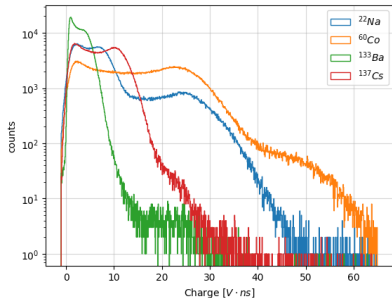
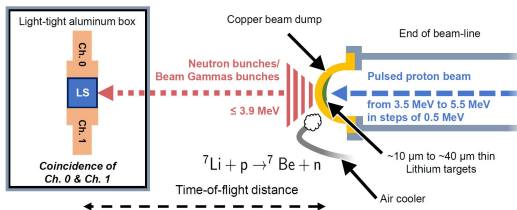
Reconstruction



JUNO Electronics setup (LNL)

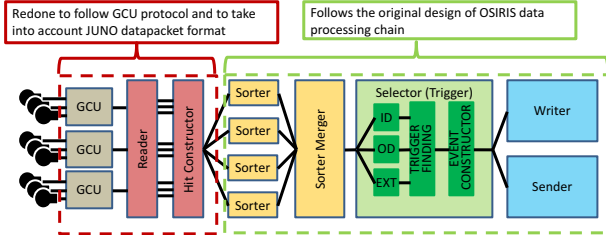
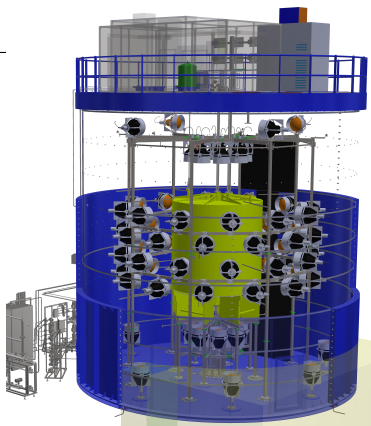


- small complete setup with LS (20 liters), 48 PMTs + JUNO electronics readout (full chain)
- allows several measurements with different sources: laser, γ s, cosmic μ s and test beam particles



JUNO Electronics for OSIRIS

- Online Scintillator Internal Radioactivity Investigation System (OSIRS) : stand-alone (20 t) LS detector to verify the quality and radiopurity of JUNO LS during filling
- OSIRIS : *small JUNO* : 20 t LS read by 75 LPMTs
- in installation phase: it's commissioning will help to verify the performances of electronics and the whole DAQ and reconstruction phase
- now detector filled with water, replacement with LS ongoing
- the JUNO LNL setup is fundamental in commissioning the radout software and analysis chain



The JUNO Padova Group



A. Garfagnini



R. Brugnera



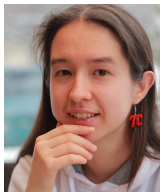
M. Grassi



A. Triossi



A. Serafini



M. Redchuk



B. Jelmini



L. Lastrucci



V. Cerrone



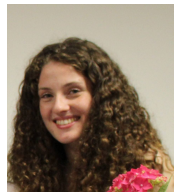
A. Gavrikov



L. V. D'Auria



R. M. Guizzetti



B. Rasera

Recent Publications

- R. Triozzi et al., *Implementation and performances of the IPbus protocol for the JUNO Large-PMT readout electronics*, submitted to Nucl. Instr. Meth. **A**
<https://arxiv.org/abs/2302.10133>,
- A. Coppi et al., *Mass testing of the JUNO experiment 20-inch PMTs readout electronics*, submitted to Nucl. Instr. Meth. **A**
<https://arxiv.org/abs/2301.04379>,
- V. Cerrone et al., *Validation and integration tests of the JUNO 20-inch PMTs readout electronics*, submitted to Nucl. Instr. Meth. **A**
<https://arxiv.org/abs/2212.08454>,
- A. Abusleme et al., JUNO Collaboration, *JUNO physics and detector*, Progr. Part. Nucl. Phys. **123**, (2022) 103927
<https://arxiv.org/abs/2104.02565>
<https://doi.org/10.1016/j.pnpnp.2021.103927>
- Z. Qian et al., *Vertex and Energy Reconstruction in JUNO with Machine Learning Methods*, Nucl. Instr. Meth. **A 1010** (2021) 165527.
<https://arxiv.org/abs/2101.04839>
[doi:10.1016/j.nima.2021.165527](https://doi.org/10.1016/j.nima.2021.165527)
- A. Bellato et al., *Embedded readout electronics R&D for the large PMTs in the JUNO experiment*, Nucl. Instr. Meth. **A 986** (2021) 164600.
<https://arxiv.org/abs/2003.08339>
[doi:10.1016/j.nima.2020.164600](https://doi.org/10.1016/j.nima.2020.164600)

JUNO thesis topics

- analysis of the **first data collected by JUNO** and determination of the **neutrino oscillation parameters**
- **supernova detection** in JUNO: development of **online triggering** algorithms and data reconstruction
- **machine learning** inspired algorithms applied to **event reconstruction, data selection** and **analysis in JUNO**
- commissioning of the **OSIRIS detector** and reconstruction of the first physics events: calibration data, cosmic muons, radioactive background (U/Th chains)
- real-time **charge reconstruction** with the large PMTs of the JUNO detector
- measurement of LS performances with test beam particle at LNL