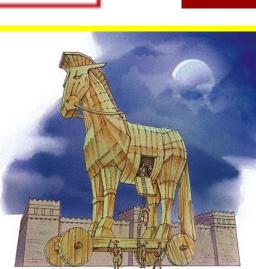


### Nuclear Astrophysics with Indirect Methods

Marco Mazzocco







# **Nuclear Astrophysics**

Nuclear Astrophysics investigates how nuclear processes generate the energy of stars over their lifetime and, in doing so, synthesize heavier elements from the primordial hydrogen and helium produced in the Big Bang.

We **humans** are mostly **(83%) oxygen** and **carbon**. We **understand** in a general way the **chemistry** and **biology** involved, but we certainly **do not understand** the **nuclear astrophysics** which produced the oxygen and carbon in our bodies.

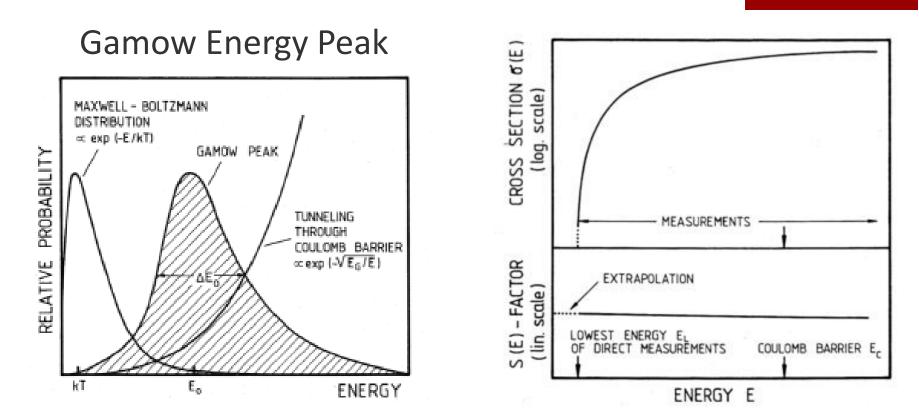
Each of **us**, in a sense, has been **inside a star** and truly and literally **consists of stardust**. Every molecule in our bodies contains matter that once was subjected to the **tremendous temperatures** and **pressures** at the **center** of a **star**.

This is where the **iron** in our **blood cells** originated, the **oxygen** we **breathe**, the **carbon** and **nitrogen** in our **tissues** and the **calcium** in our **bones**.

(W.A. Fowler – Nobel Price in Physics 1983)

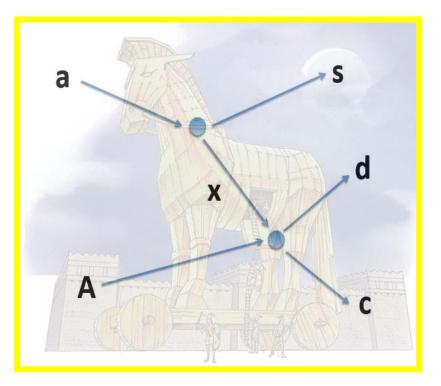


## **ThermoNuclear Reactions**



Nuclear reactions of astrophysical interest occur at energies **deeply below the Coulomb barrier**, their **cross sections** (i.e. probabilities) might be extremely **small** and **background** might be very **high**.

## **Trojan Horse Method**



#### **Basic Concept**:

It it possible to extract the astrophysically relevant **two-body cross section**:

 $A + x \rightarrow c + d$ 

from the quasi-free contribution of a suitable **three-body reaction**:

 $A + a \rightarrow c + d + s$ 

**Method**: to hide the **projectile x** inside a **Trojan Horse Nucleus a**, with a prominent cluster structure **x** + **s**, and to make sure that **particle s** acts as a **spectator** in the nuclear reaction (quasi-free mechanism)

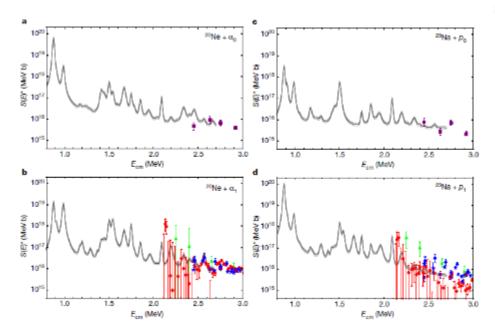
## Recent Measurements Massive Star Evolution (LNS)

### LETTER

https://doi.org/10.1038/s41586-018-0149-4

### An increase in the ${}^{12}C + {}^{12}C$ fusion rate from resonances at astrophysical energies

A. Tumino<sup>1,2</sup>\*, C. Spitaleri<sup>2,3</sup>, M. La Cognata<sup>2</sup>, S. Cherubini<sup>2,3</sup>, G. L. Guardo<sup>2,4</sup>, M. Gulino<sup>1,2</sup>, S. Hayakawa<sup>2,5</sup>, I. Indelicato<sup>2</sup>, L. Lamia<sup>2,3</sup>, H. Petrascu<sup>4</sup>, R. G. Pizzone<sup>2</sup>, S. M. R. Puglia<sup>2</sup>, G. G. Rapisarda<sup>2</sup>, S. Romano<sup>2,3</sup>, M. L. Sergi<sup>2</sup>, R. Spartá<sup>2</sup> & L. Trache<sup>4</sup>

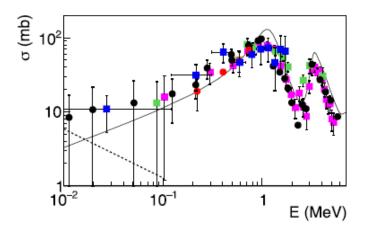


### Big-Bang Nucleosynthesis (LNL)

THE ASTROPHYSICAL JOURNAL, 879-23 (Spp), 2019 July 1 ID 2003 The American Automatical Society, All rights reserved. https://doi.org/10.3847/1538-4357/ab223

Cross-section Measurement of the Cosmologically Relevant  ${}^{7}Be(n, \alpha){}^{4}He$  Reaction over a Broad Energy Range in a Single Experiment

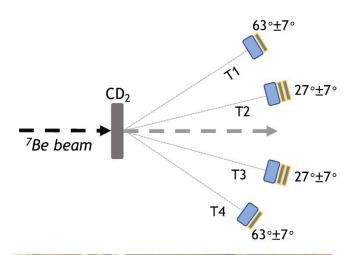
L. Lamia<sup>1,2</sup><sup>(2)</sup>, M. Mazzocco<sup>3,4</sup>, R. G. Pizzone<sup>2</sup>, S. Hayakawa<sup>5</sup>, M. La Cognata<sup>2</sup>, C. Spitaleri<sup>1,2</sup>, C. A. Bertulani<sup>6</sup><sup>(2)</sup>, A. Boiano<sup>7</sup>, C. Boiano<sup>8</sup>, C. Broggini<sup>4</sup>, A. Caciolli<sup>3,4</sup>, S. Cherubini<sup>1,2</sup>, G. D'Agata<sup>1,2,13</sup><sup>(2)</sup>, H. da Silva<sup>9</sup>, R. Depalo<sup>3,4</sup>, F. Gatarossa<sup>10</sup>, G. L. Guardo<sup>1,2</sup>, M. Gulino<sup>2,11</sup>, I. Indelicato<sup>1,2</sup>, M. La Commura<sup>7,12</sup>, G. La Runa<sup>7,12</sup>, R. Menegazzo<sup>4</sup>, J. Mrazek<sup>13</sup>, A. Pakou<sup>14</sup>, C. Parascandolo<sup>7</sup>, D. Patti<sup>1,4</sup>, D. Pierroutsakou<sup>7</sup>, S. M. R. Puglia<sup>7</sup>, S. Romano<sup>1,7</sup>, G. G. Rapisarda<sup>7</sup>, A. M. Sánchez-Benítez<sup>15</sup>, M. L. Sergi<sup>2</sup>, O. Sparta<sup>1,4</sup>, F. Sonamel<sup>3,4</sup>, V. Soukers<sup>2,14</sup>, R. Spartá<sup>1,4</sup>, E. Stano<sup>5,4</sup>, D. Torresi<sup>7</sup>, A. Tumino<sup>2,11</sup>, H. Yamaguchi<sup>6</sup>, and G. L. Zhang<sup>16</sup>

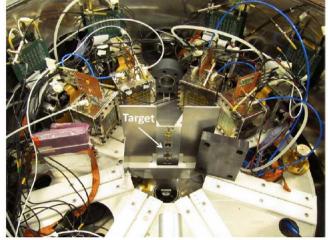


# Thesis Activity (I)

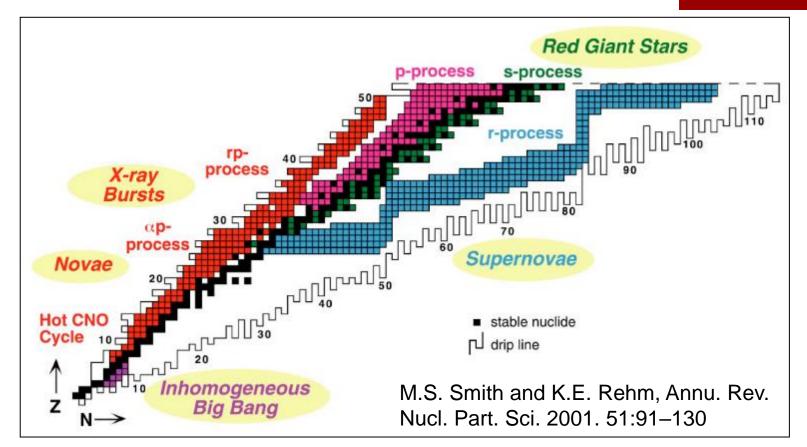
- Setup design and construction,
- Setup characterization,
- Data taking,
- Data analysis,
- Study the impact of the cross section results in astrophysical models.

Join experimental runs performed at the RIB facility EXOTIC located at the Legnaro National Laboratories, Laboratori Nazionali del Sud (Catania) or abroad (France, Czech Republic, Canada, Romania, Japan, ...)





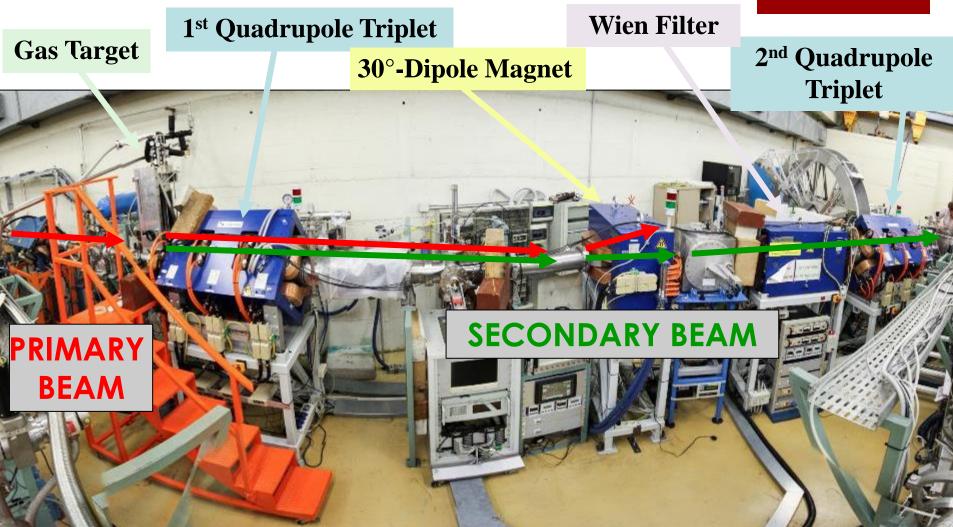
## Nuclear Astrophysics and Radioactive Ion Beams



Radioactive Ion Beams (RIBs) are deeply involved in several scenarios of interest for Nuclear Astrophysics.

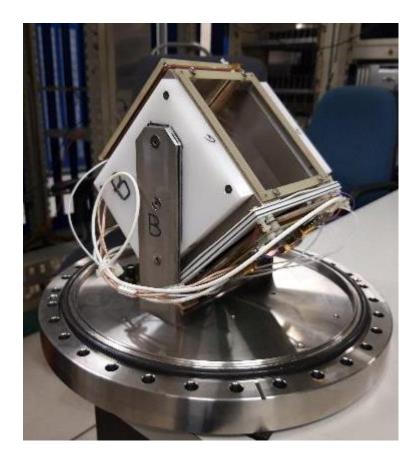
# Thesis Activity (II)

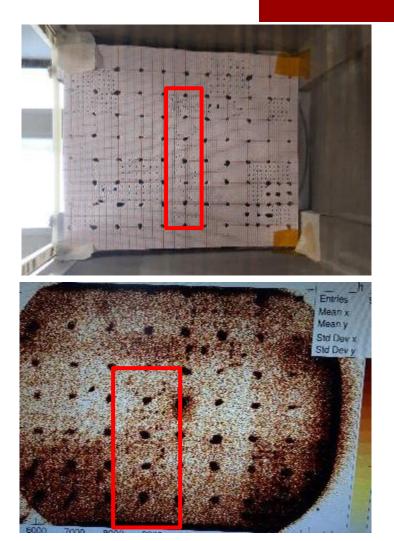
• Study of the ion-optics for the connection of EXOTIC and AGATA



# Thesis Activity (III)

 Development and characterization of a MCP tracking detector.





### ASFIN partnerships and collaborations







@ <u>Catania/LNS</u>: A. Bonasera, S. Cherubini, G. D'Agata, A. Di Pietro, P. Figuera, G.L. Guardo, M. Gulino, M. La Cognata, L. Lamia, D. Lattuada, A.A. Oliva, G.G. Rapisarda, R.G. Pizzone, S. Romano, D. Santonocito, M.L. Sergi, R. Spartà, A. Tumino
@ <u>Padova</u> M. Mazzocco, S. Pigliapoco, F. Soramel
@ <u>Perugia</u> M. Busso, S. Palmerini, M. Limongi, A. Chieffi, M.C. Nucci. N. Vukman
@ Napoli M. La Commara

#### International Collaborations

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- Cyclotron Institute, Texas A&M, USA: R. Tribble, V. Goldberg
- Texas A&M Commerce USA: C. Bertulani
- Florida State University USA: I. Wiedenhofer
- C.N.S. Riken, Wako, Japan: S. Kubono, H. Yamaguchi, S. Hayakawa
- University of Taskent: B. Irgaziev, R. Yarmukhanmedov
- CIAE, Beijing, China: S. Zhou, C. Li, Q. Wen
- Nuclear Physics Institute, ASCR, Rez, Czech Rep.: V. Kroha, V. Burjan, J. Mrazek
- Nipne IFIN Bucharest: L. Trache
- ELI-NP Bucharest: C. Matei, D. Balabanski
- Atomki, Debrecen, Hungary: G. Kiss
- CSNSM, Orsay, France : A. Coc , F. Hammache, N. De Sereville
- University of Catalunya: J. Jose
- Rudjer Boskovic Institute Zagreb Croatia: N. Soic, M. Milin
- INFN Sez. Napoli e UniNA: D. Pierroutsakou, C. Parascandolo
- University of Pisa: S. Degl'Innoccenti, P. Prada Moroni

