

# Neutrino physics with ENUBET and T2K



*Prof. Andrea Longhin*

Per gli studenti di LT e LM in Fisica

DFA UniPD, 23/02/2024



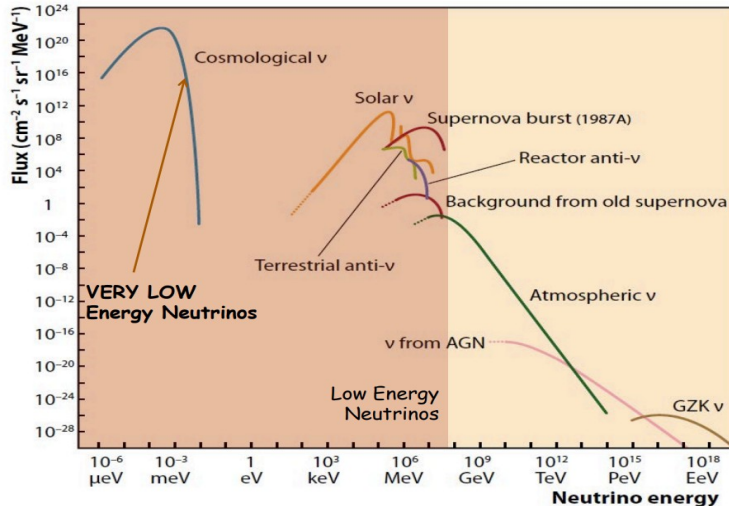
Hypothesis 1932 (Pauli), discovery 1957 (reactors, Reines-Cowan)

Extremely abundant:




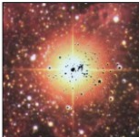
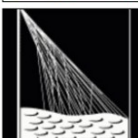

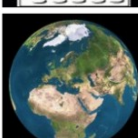
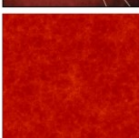
the Sun  $\rightarrow 6 \times 10^{10}/\text{s}/\text{cm}^2$  !!!

the Early Universe  $\rightarrow 336/\text{cm}^3$  (indirect)

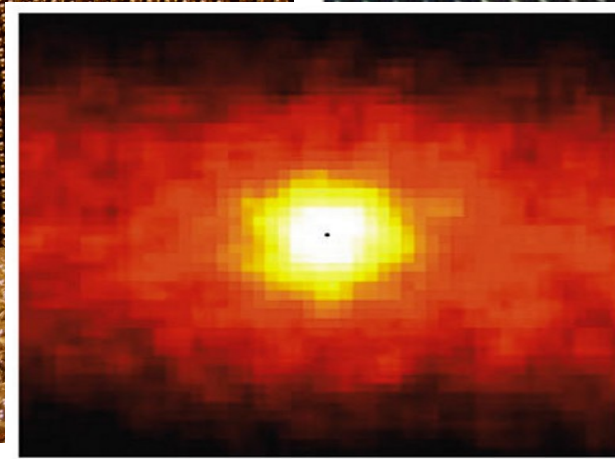
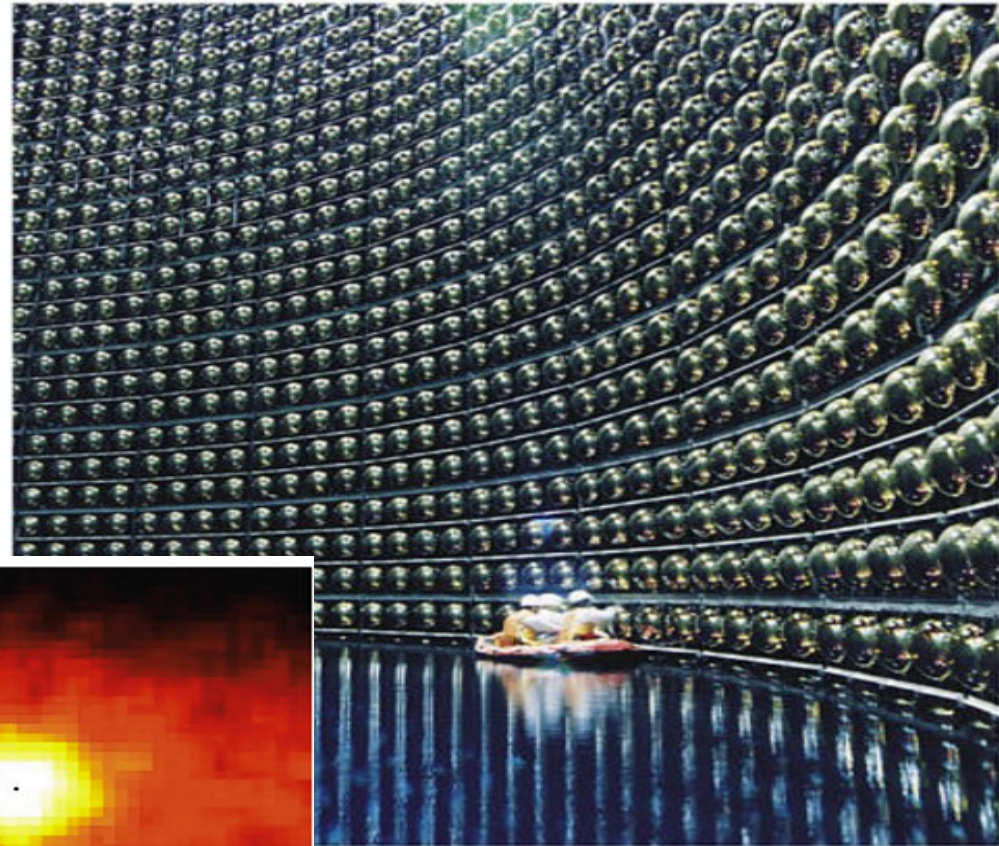
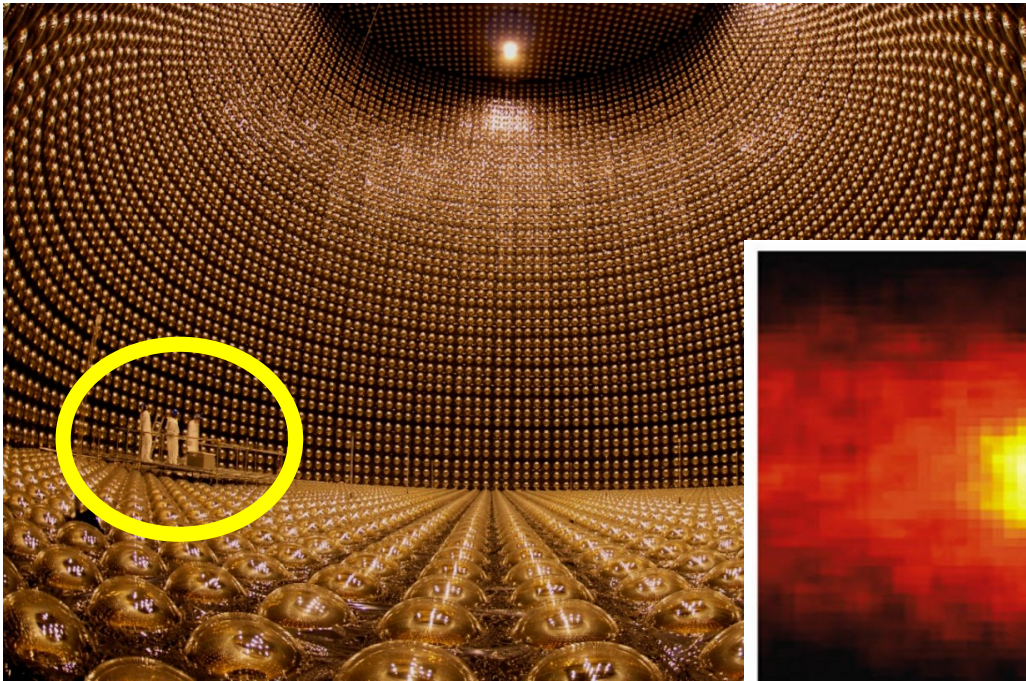
Many sources over  $\sim 20$  orders of magnitude in energy  $\rightarrow$



## Where do neutrinos come from?

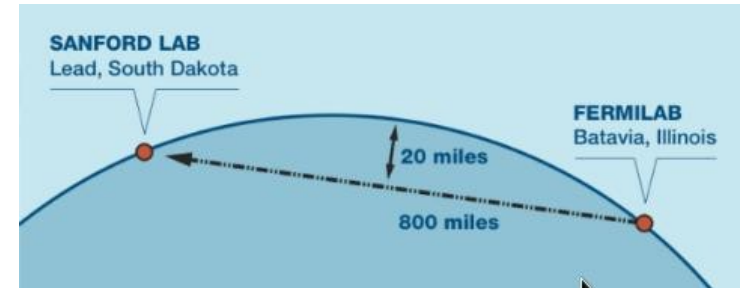
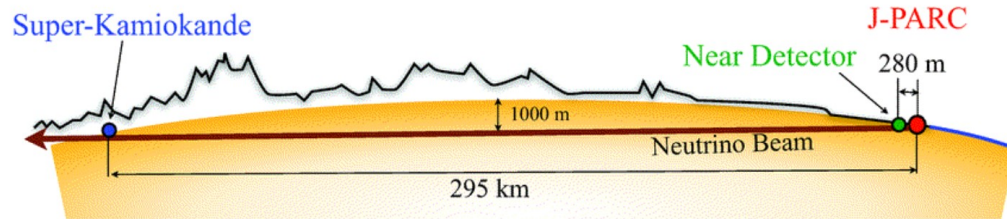
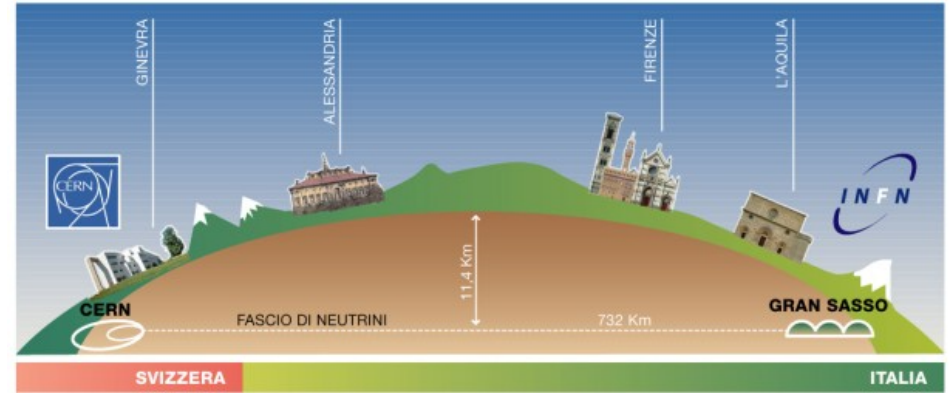
✓ Nuclear reactors			Sun ✓
✓ Particle accelerators			Supernovae SN 1987A ✓
✓ Earth Atmosphere (Cosmic rays)			Accelerators in astrophysical sources ?✓
✓ Earth interior (Natural Radioactivity)			Early Universe (today $336 \nu/\text{cm}^3$ ) Indirect evidence

The only elementary particle with weak interaction only → gigantic and refined detectors



← The Sun neutrino-graphed by Super-Kamiokande

Beams of O(100 km) or more  
feasible underground  
→ oscillation experiments  
(flavor transitions)



1998: discovery of neutrino oscillations → at least 2 out of the 3 neutrinos have a non-zero mass ( $< 1$  eV) → [Nobel 2015](#) (Kajita, McDonald)

Neutrinos are  $\sim O(1 \text{ M})$  times lighter than the electron! **why so unnaturally lighter than leptons (to which they are intimately related)?**

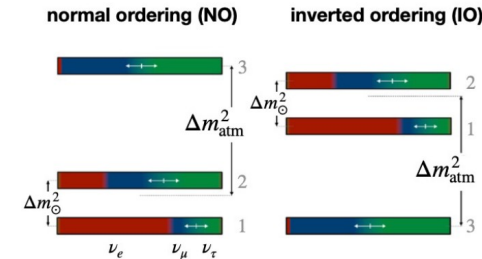
Is it due to other undiscovered neutrinos at high-mass?

→ **a “portal” towards physics beyond the standard model**

**Does neutrino masses arise from the Higgs mechanism?**

**Or are they intrinsically different particles? (“Majorana fermions”)**

- Do neutrinos violate the CP symmetry as quarks?  
Where does the **matter-over antimatter asymmetry in the universe** come from?
- Which is the **mass of the lightest neutrino**?
- Which is the **mass ordering** of the 3 mass states ?
  - We currently only know differences of mass<sup>2</sup> from oscillations.
- Are there “new” neutrinos beyond those with a standard lepton counterpart (**sterile neutrinos**)?



Neutrinos are experimentally “**difficult**”

- Need to detect them far away + weak interactions → **large statistical uncertainties**
- production+interaction only described by models → **large systematic uncertainties**

Systematics are becoming the driving limitation

→ **Desperate need for high quality experimental inputs**

I will show 2 “directions” explored by our group:

- **1)** fully exploit the information from “near detectors”
  - Bayesian analysis of the T2K experiment data (“analysis”)
- **2)** new ideas to improve the source:
  - ENUBET demonstrator → a new concept neutrino source with 1% precision (“hardware”)

Neutrinos are experimentally “**difficult**”

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Systematics are becoming the driving limitation

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... and more

- **Spoiler of the talk of Prof. Collazuol here !!!!!**
- **build brand new instrumentation for the near detector (Large angle TPC)**
- **build a huge far detector (Hyper-Kamiokande)**

→ **new source** (analysis)

→ **demonstrator** → a new concept neutrino source with 1% precision

(“hardware”)



# 1) Oscillation analysis with T2K

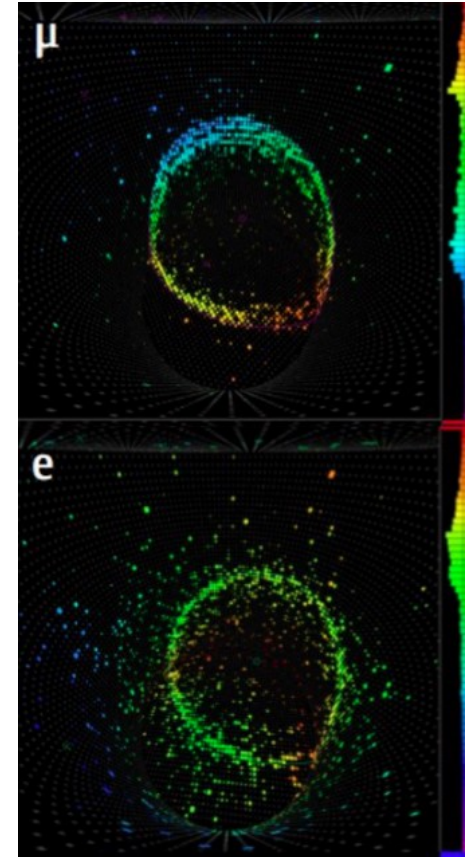
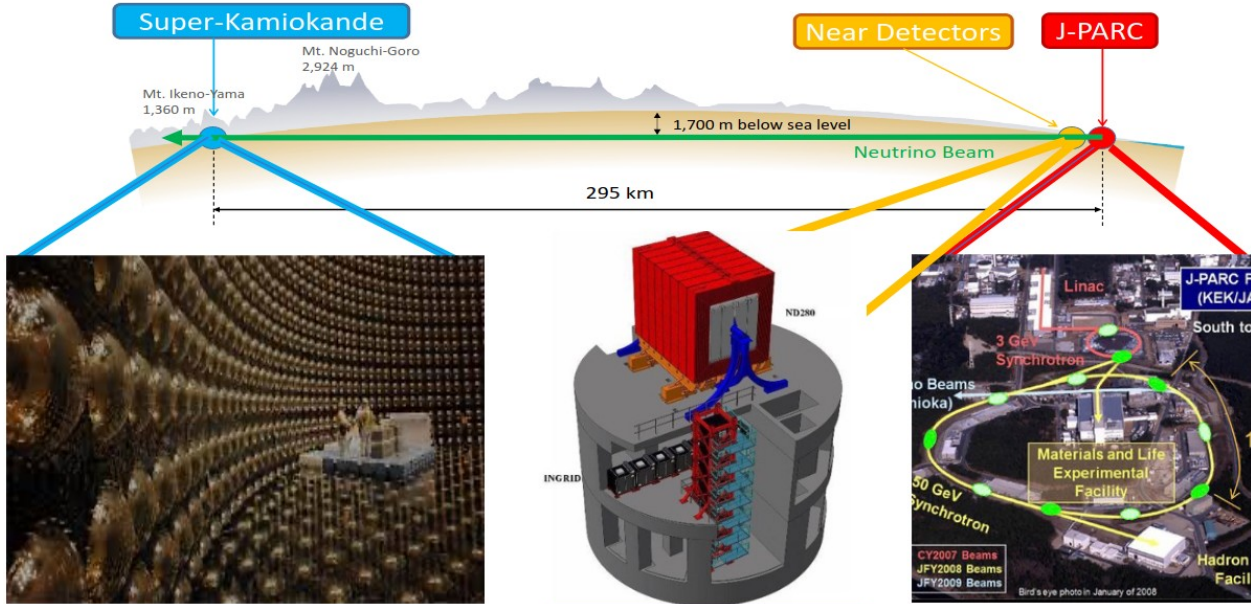
T2K uses difference in the process  $\nu_\mu \rightarrow \nu_e$  and  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  to study the matter-antimatter asymmetry over a 295 km “travel”.

Super-Kamiokande can separate  $\nu_\mu$  from  $\nu_e$

For  $\delta_{CP}$  we look at  $\nu_e/\bar{\nu}_e$  appearance.

$\nu_e, \nu_e, \nu_\mu, \nu_\mu, \nu_\mu$   
 $\bar{\nu}_e, \bar{\nu}_\mu, \bar{\nu}_\mu, \bar{\nu}_\mu$

$\nu_\mu, \nu_\mu, \nu_\mu, \nu_\mu, \nu_\mu, \nu_\mu, \nu_\mu$   
 $\bar{\nu}_\mu, \bar{\nu}_\mu, \bar{\nu}_\mu, \bar{\nu}_\mu, \bar{\nu}_\mu, \bar{\nu}_\mu, \bar{\nu}_\mu$

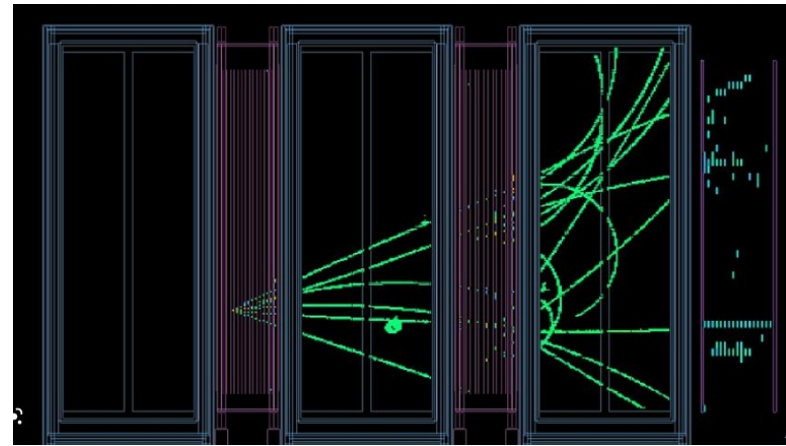
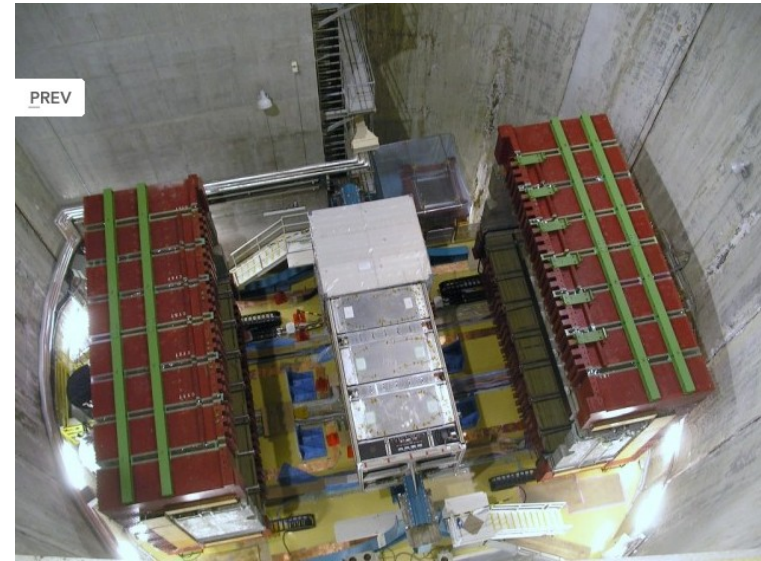


## Many unknowns:

- the neutrino flux (different between near and far)
- the neutrino interaction parameters (that depend on nuclear models, from the target composition)
- There are several different final states each with a different efficiency in the detectors.

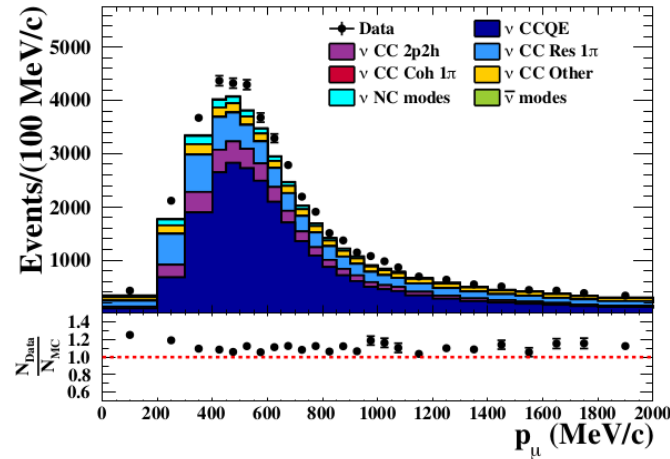
The goal is to **predict as precisely as possible the number of neutrinos that we expect at the far detector** → the difference is due to oscillations (appearance of electron neutrinos, disappearance of muon neutrinos) that we can then study.

We use the data collected at the **near detector** (at only 280 m → high statistics) to get a constraint on unknown quantities (“nuisance parameters”).

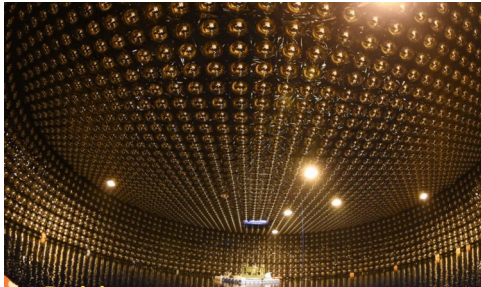
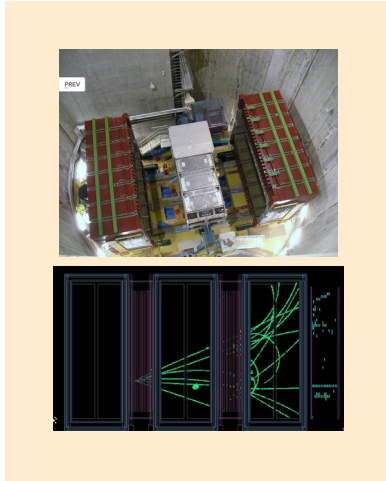
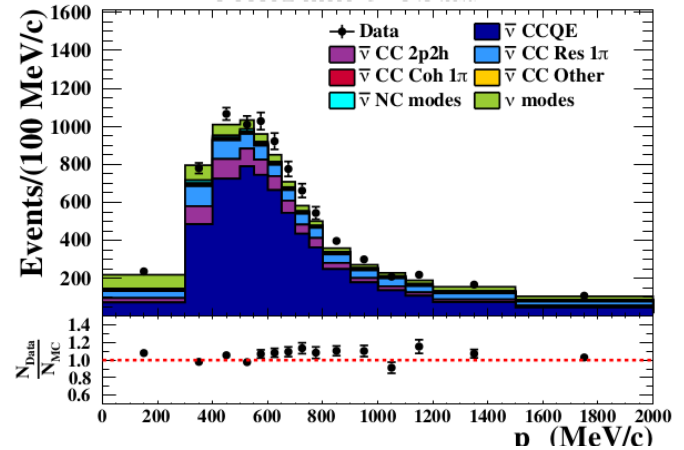


## The "Near detector fit"

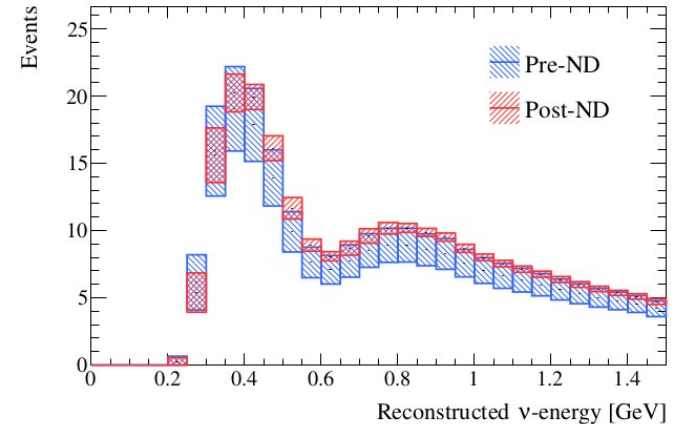
### Near detector data BEFORE fit



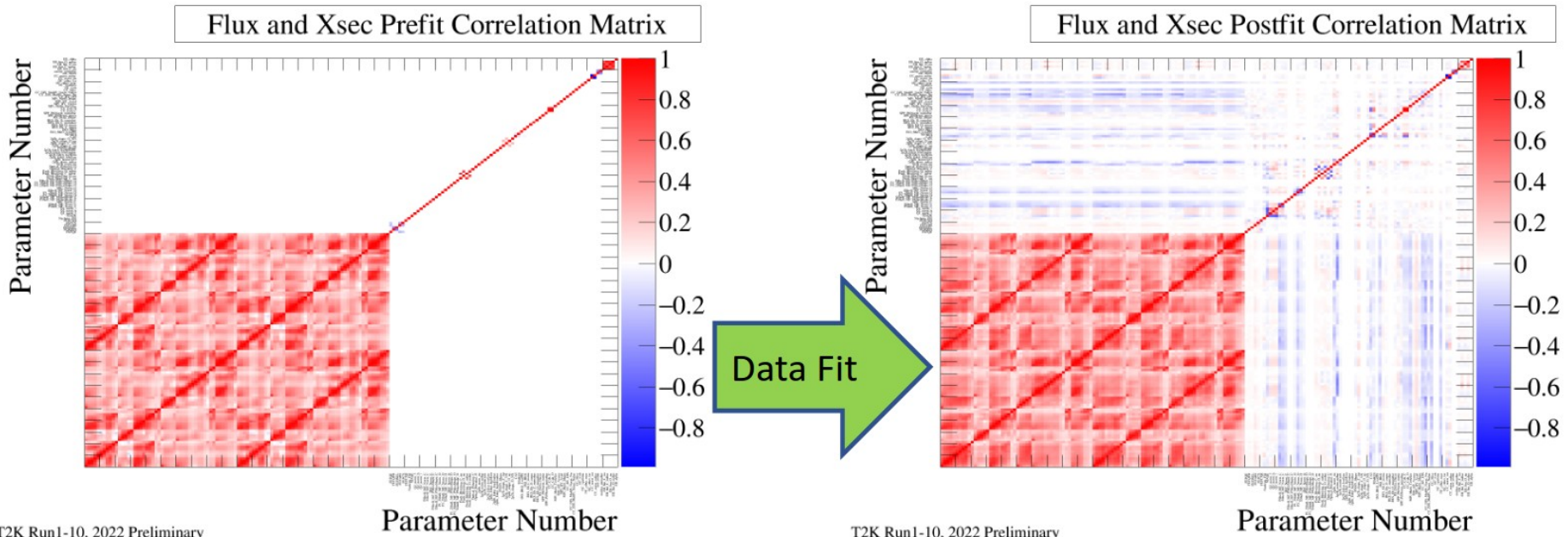
### Near detector data AFTER fit



Uncertainty on the prediction at far  
**BEFORE** and  
**AFTER** the near detector fit:



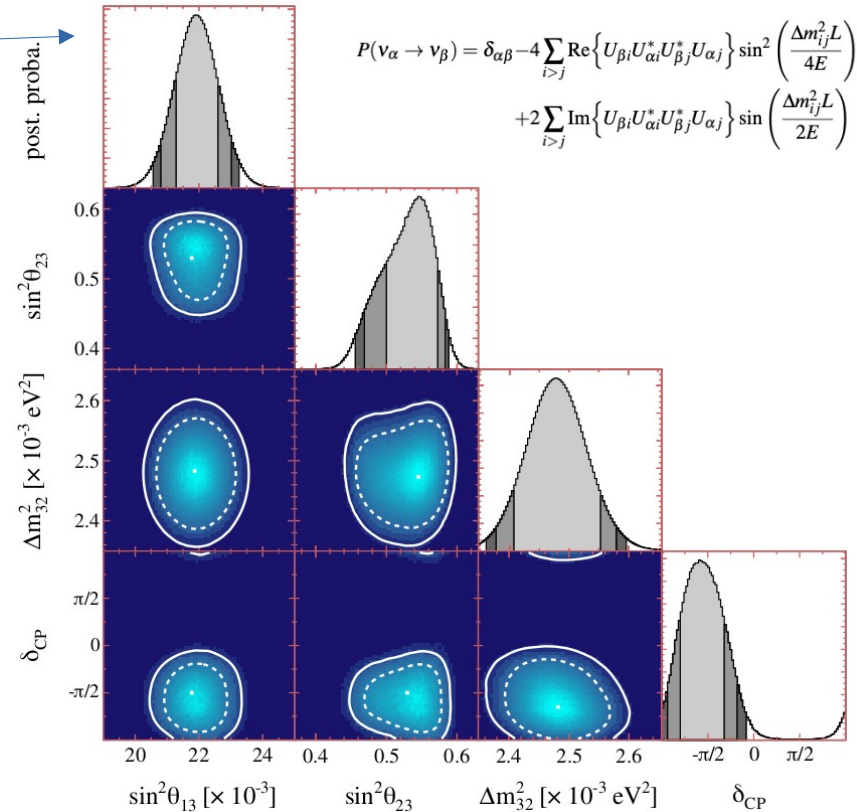
The fit is very complex. Besides the oscillation parameters it involves **hundreds of unknown parameters** (“nuisance”) from uncertainties in nuclear models, final state interactions, flux, detector parameters. These are “fitted out” to finally access the interesting oscillation parameters (3).



**Bayesian statics** → posterior probabilities on physical quantities.

The search of maxima in the complex multi-dim space of the likelihood function and marginalization of nuisance parameters is achieved with a **Metropolis-Hastings Markov-chain Monte Carlo**.

$$\begin{aligned}
 -\ln(L) = & \sum_i^{ND280bins} N_i^p(\vec{b}, \vec{x}, \vec{f}, \vec{d}) - N_i^d + N_i^d \ln[N_i^d / N_i^p(\vec{b}, \vec{x}, \vec{f}, \vec{d})] \\
 & + \sum_i^{SK1R_ebins} N_i^p(\vec{b}, \vec{x}, s\vec{k}d) - N_i^d + N_i^d \ln[N_i^d / N_i^p(\vec{b}, \vec{x}, s\vec{k}d)] \\
 & + \sum_i^{SK1R_\nu bins} N_i^p(\vec{b}, \vec{x}, s\vec{k}d) - N_i^d + N_i^d \ln[N_i^d / N_i^p(\vec{b}, \vec{x}, s\vec{k}d)] \\
 & + \frac{1}{2} \sum_i^{E_\nu bins} \sum_j^{E_\nu bins} \Delta b_i (V_b^{-1})_{i,j} \Delta b_j \\
 & + \frac{1}{2} \sum_i^{xsecpars} \sum_j^{xsecpars} \Delta x_i (V_x^{-1})_{i,j} \Delta x_j \\
 & + \frac{1}{2} \sum_i^{fsipars} \sum_j^{fsipars} \Delta f_i (V_f^{-1})_{i,j} \Delta f_j \\
 & + \frac{1}{2} \sum_i^{nd280det} \sum_j^{nd280det} \Delta d_i (V_d^{-1})_{i,j} \Delta d_j \\
 & + \frac{1}{2} \sum_i^{skdet} \sum_j^{skdet} \Delta skd_i (V_{skd}^{-1})_{i,j} \Delta skd_j
 \end{aligned}$$



Plenty of room for students with a passion in **advanced analysis tools and statistics**.

Work would be in a **highly qualified international group** (in Padova A.L. convener of the oscillation analysis group)

# 2) monitored $\nu$ beams: ENUBET

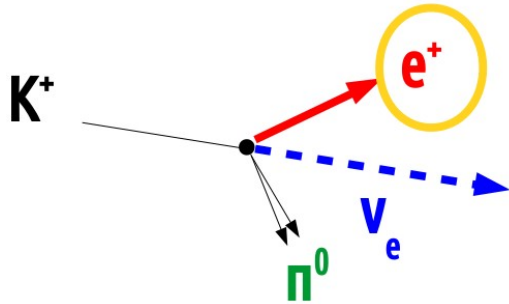
Reach a O(1 %) control on the flux  $\rightarrow$  precise physics

How ? “counting” positrons from kaon decays with a large/scalable detector

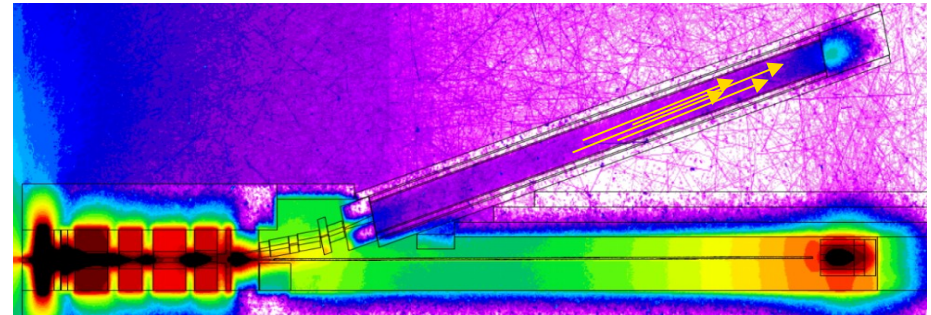
ENUBET  $\rightarrow$  Enhanced Neutrino Beams from kaon Tagging



European Research Council  
Established by the European Commission



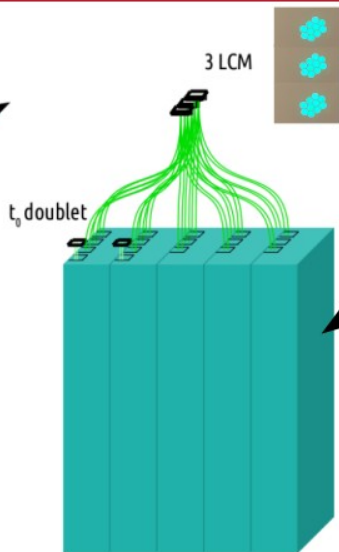
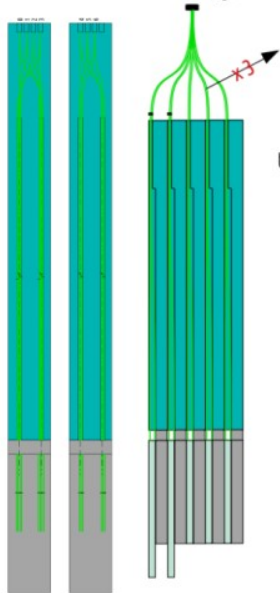
(1  $e^+$  = 1 neutrino)



Challenges for the detector:

fast electronics / radiation tolerance / good particle identification with a cost-effective solution

## WLS routing

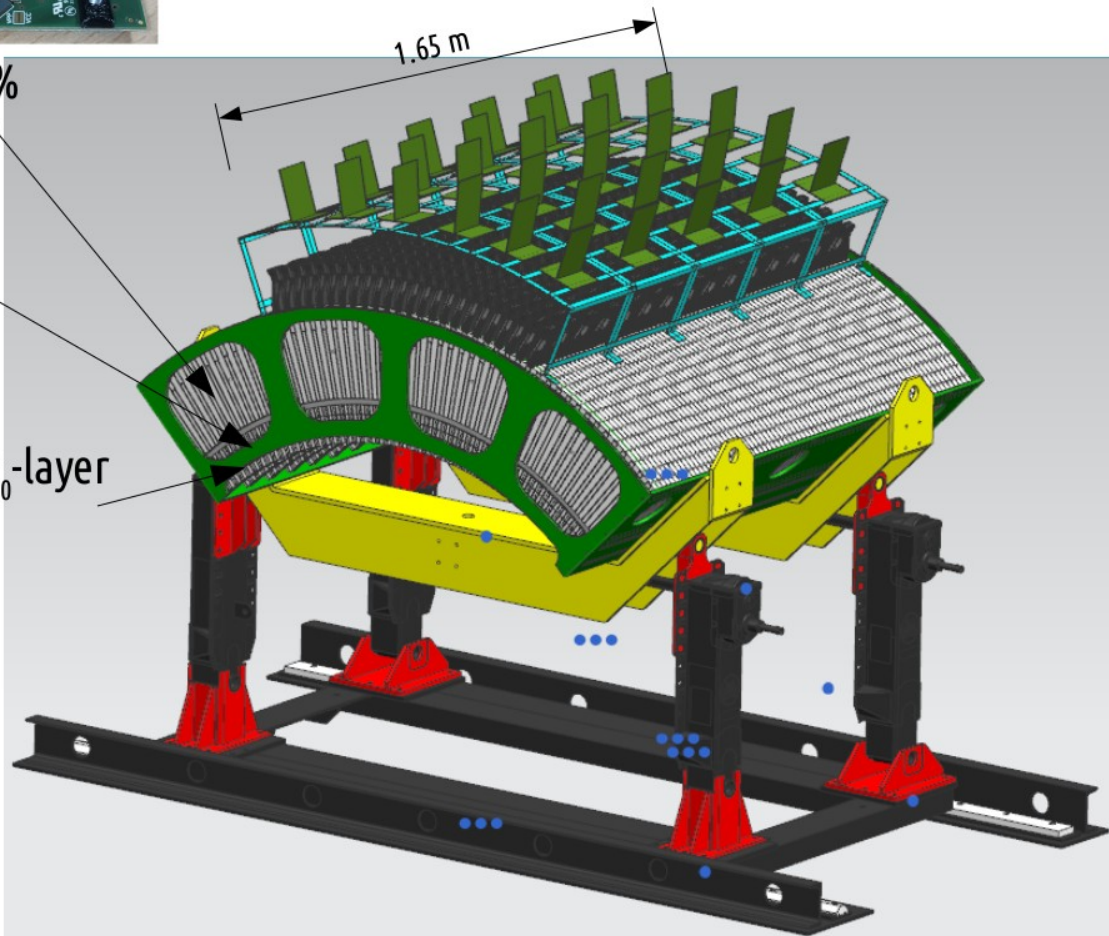
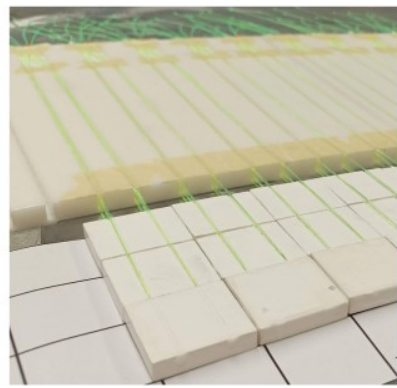
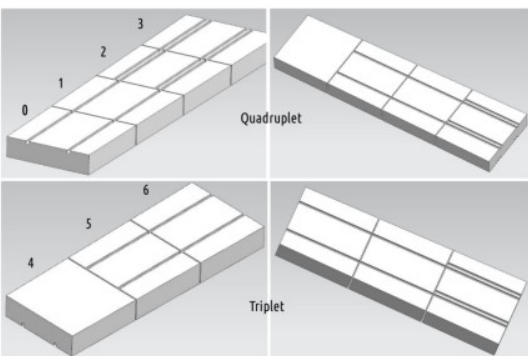


BPE 5%

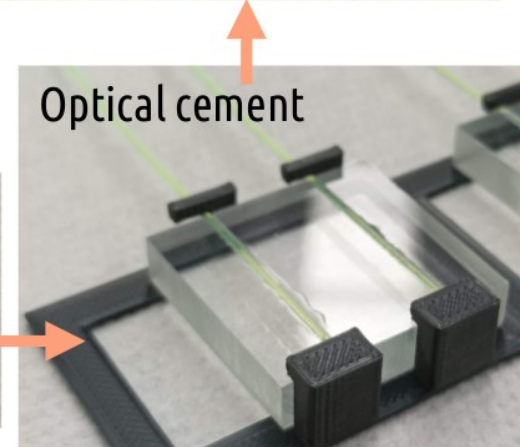
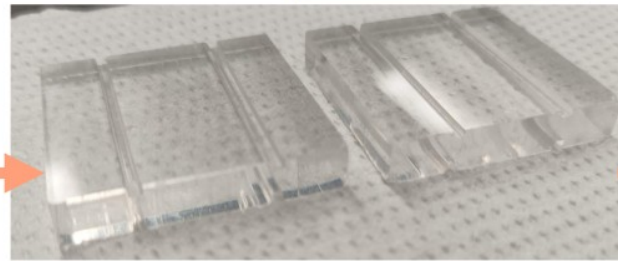
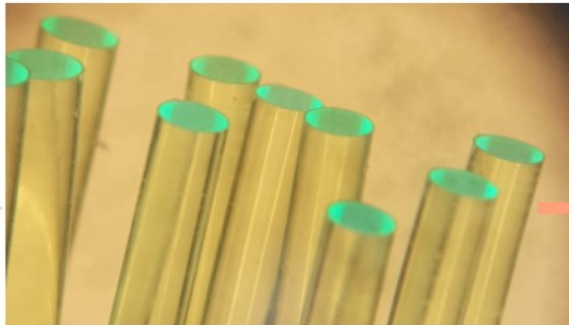
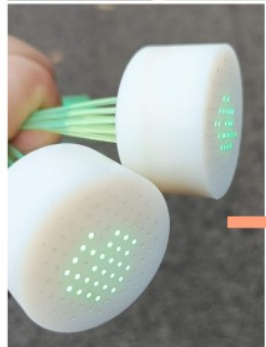
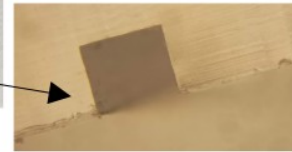
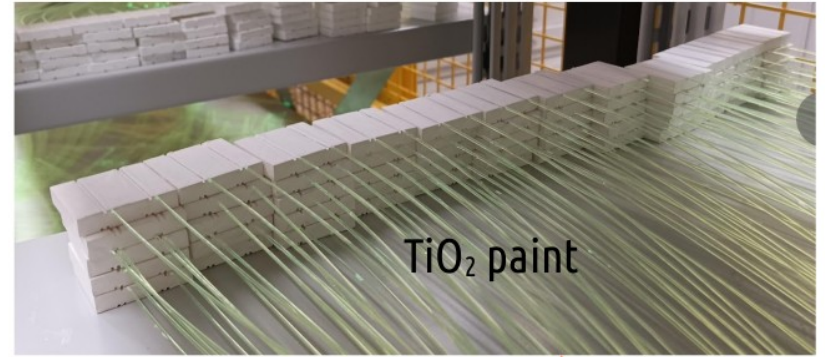
Sampling  
iron/scint calo



t<sub>0</sub>-layer



Summer 2022 @ INFN-LNL

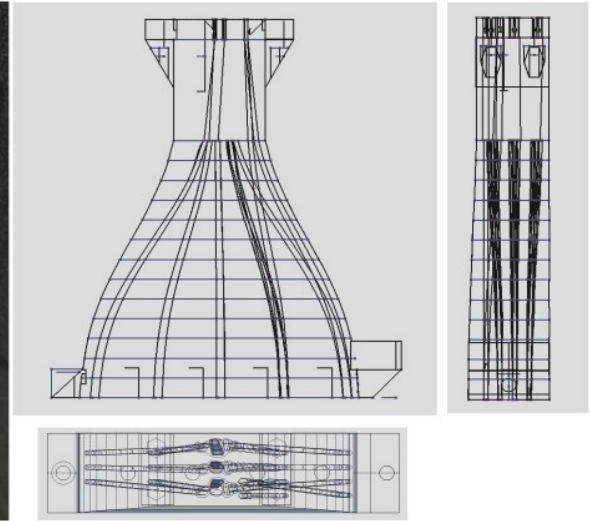




# Construction of the ENUBET demonstrator

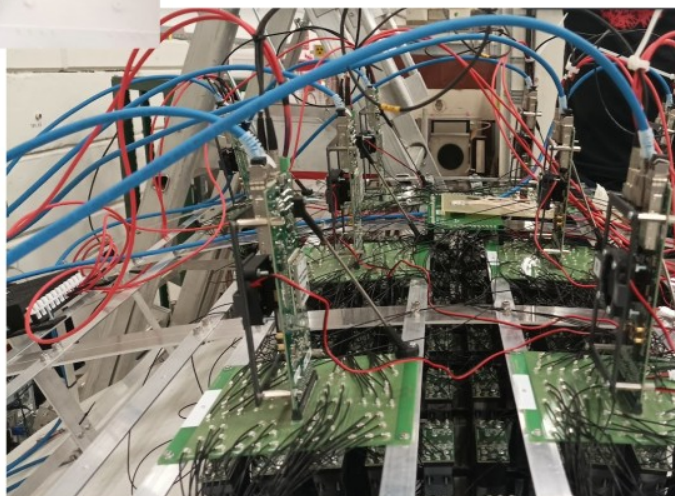
Summer 2022 @ INFN-LNL

bundling of the WLS fibers with 3D printed  
“fiber concentrators”+ in situ polishing



# Readout electronics

Oct 2022 @ CERN



16:20 0,2KB/s

Post

francesco.terrano.tel



Place a vatee\_terra e altri 18

francesco.terrano.tel An hairy detector for neutrino physics 🤖 #enubet #cern

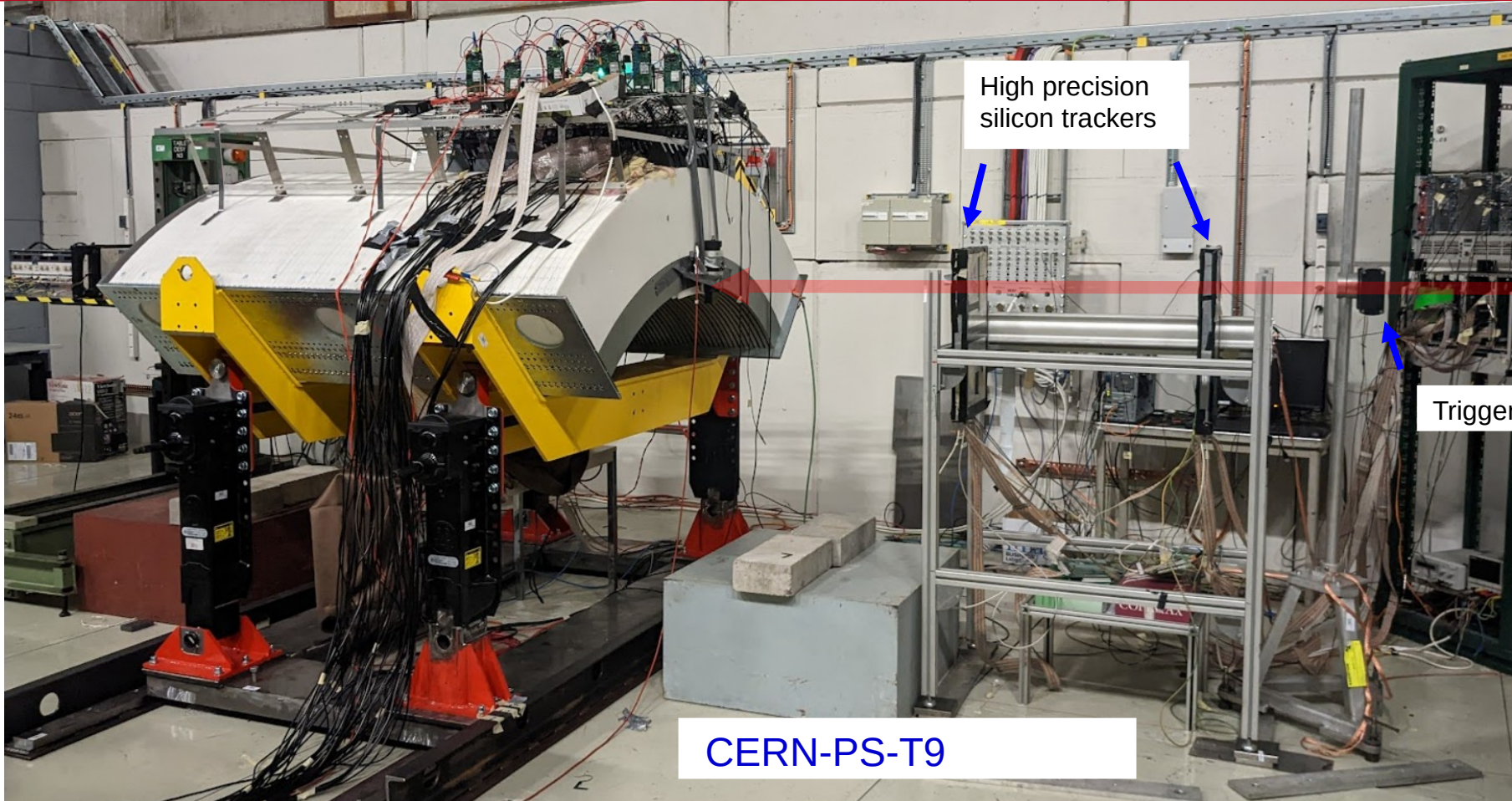


building 157, CERN  
Meyrin PS East Hall  
T9 area



Movable platform  
“landing site” @ T9  
test beam area.





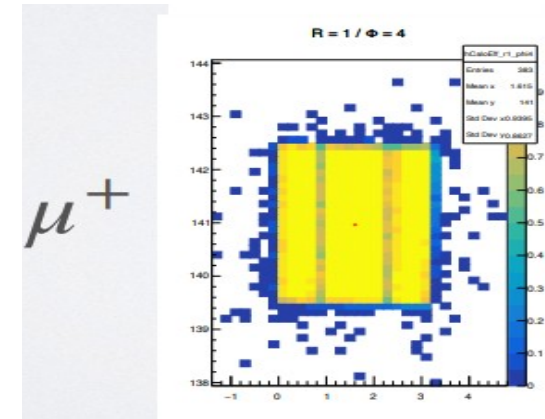
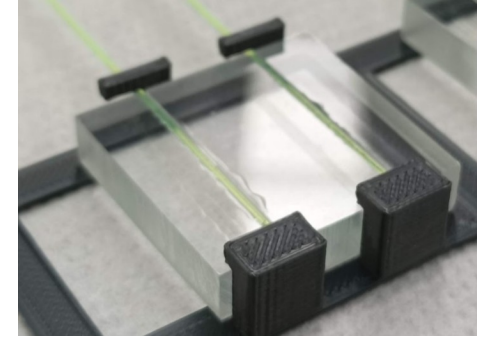
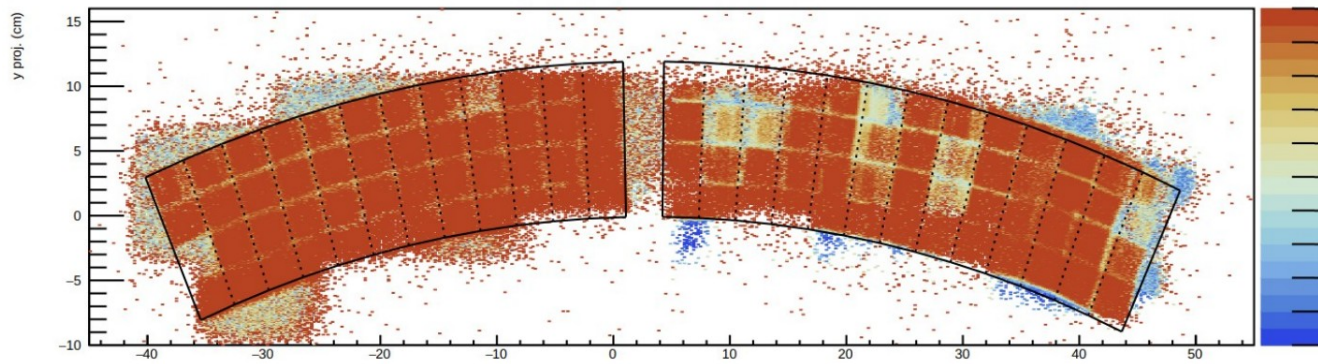
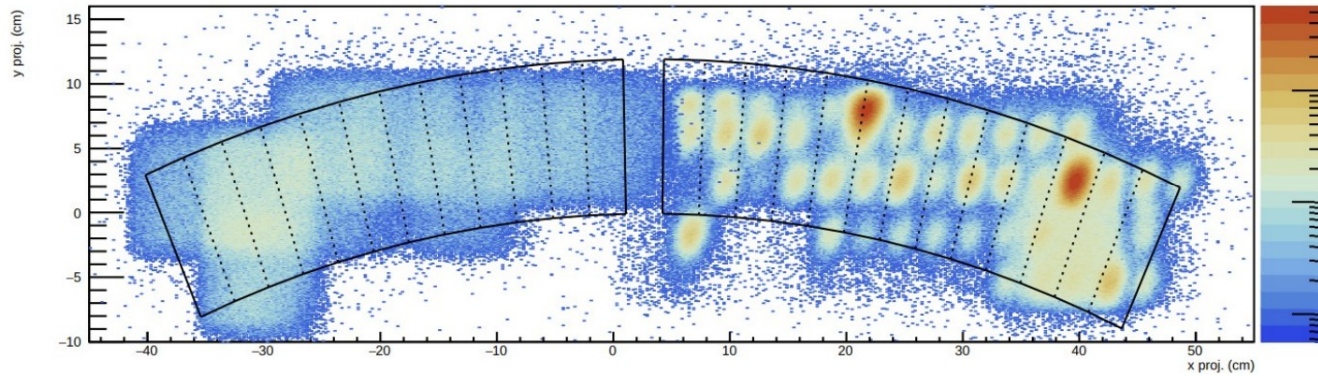
High precision  
silicon trackers

$e, \pi, \mu$  (0.5-15 GeV)

Trigger scint.

CERN-PS-T9

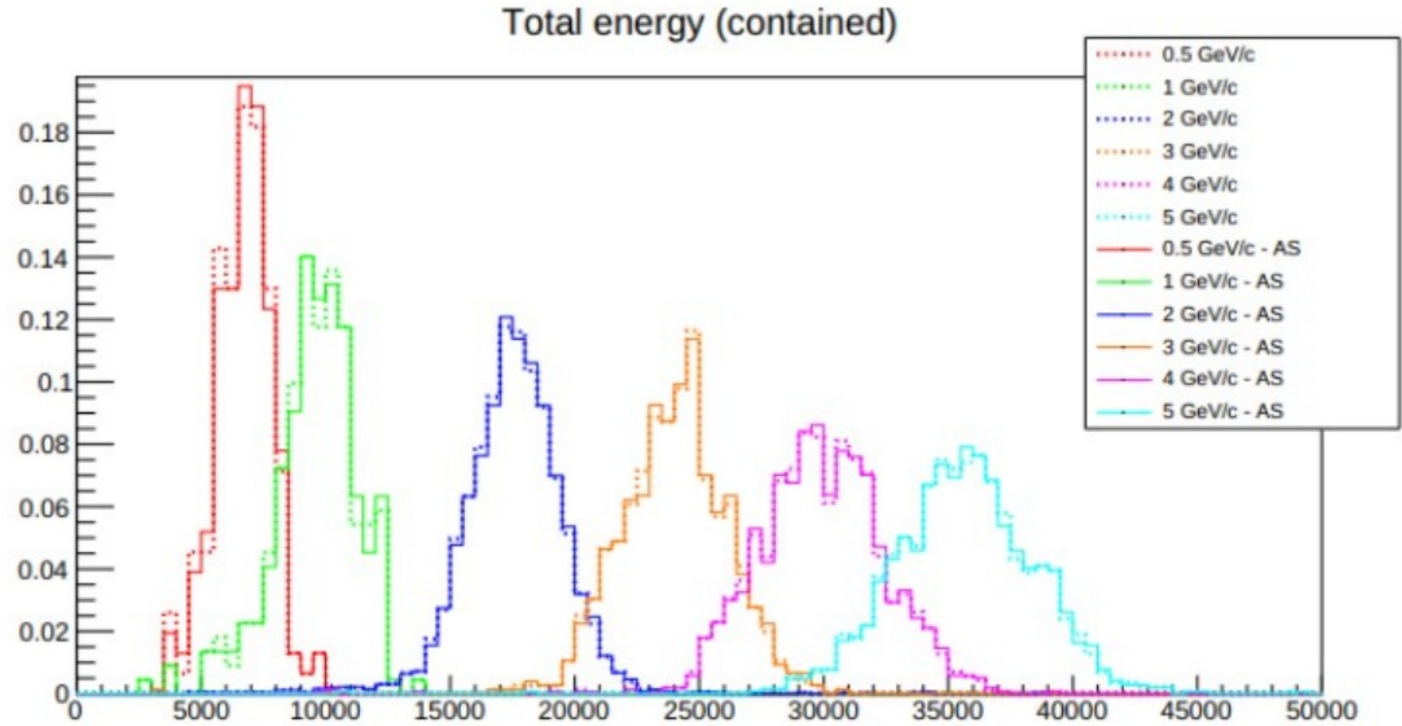
Efficiency maps → “radiography” using muons



$\mu^+$

## Electron energy resolution

A. Moro,  
 tesi triennale,  
 Dec. 2023



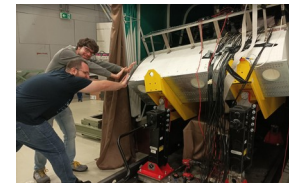
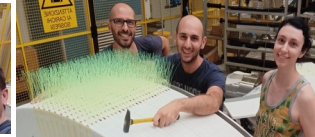
# New horizons for ENUBET

We are currently upgrading the Data Acquisition system for a **new exposure at CERN in 2024**. Precious occasion for a thesis work with the possibility of participating to the construction (Legnaro) and testing (CERN) of a complex detector and analysis of the associated data.



[andrea.longhin@unipd.it](mailto:andrea.longhin@unipd.it)

Uff. 129, via Marzolo 8





## Demonstrator tests paper

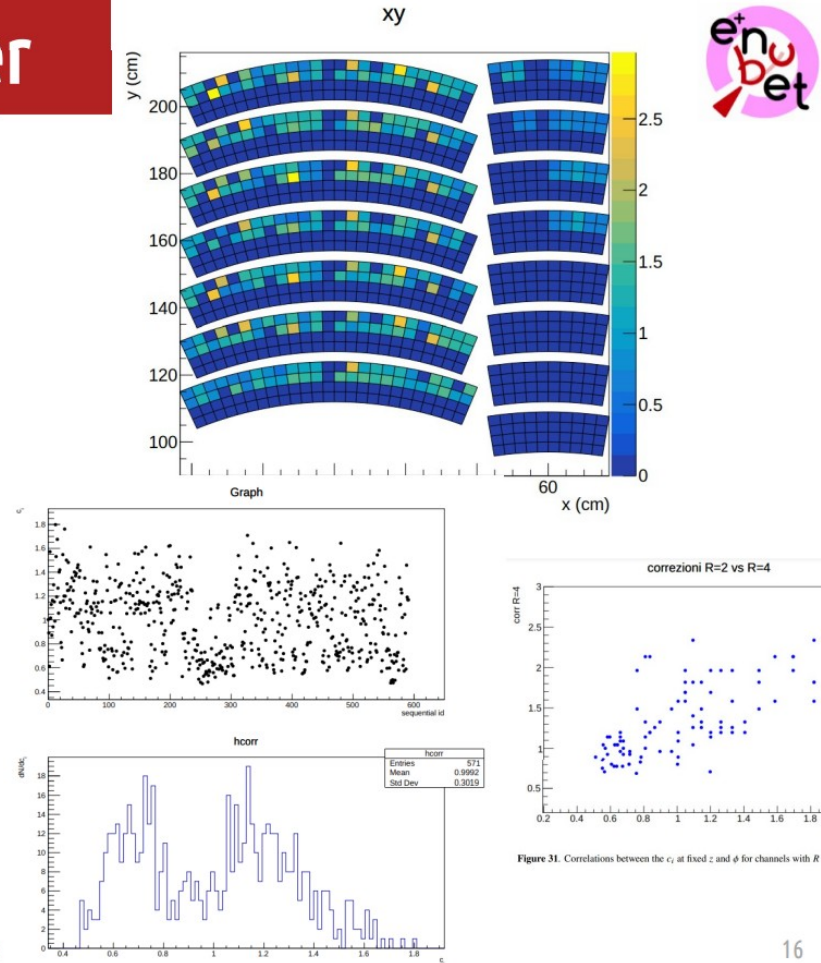
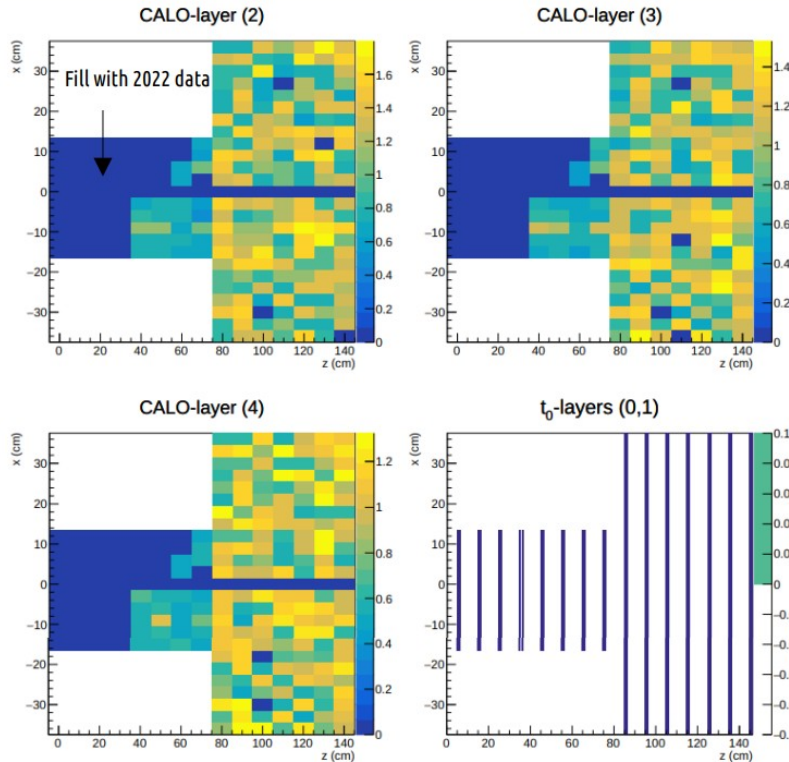
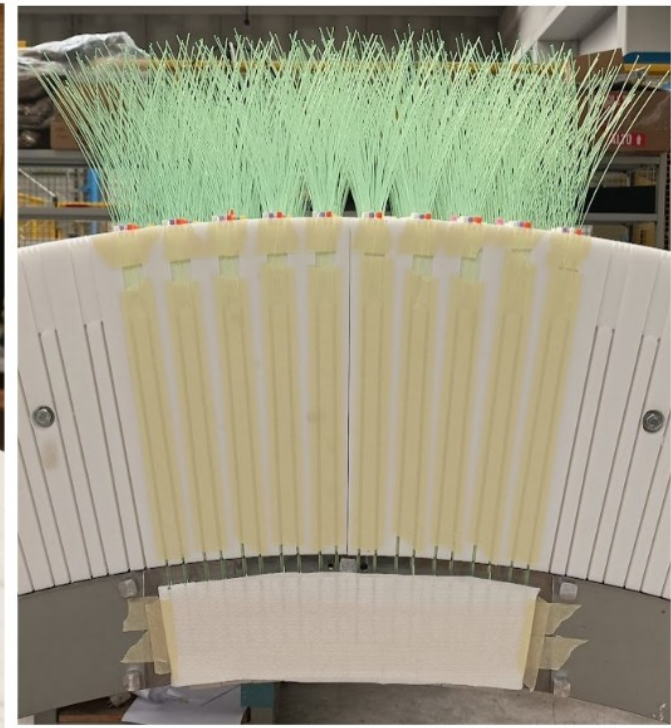


Figure 31. Correlations between the  $c_i$  at fixed  $z$  and  $\phi$  for channels with  $R = 2$  and  $R = 4$

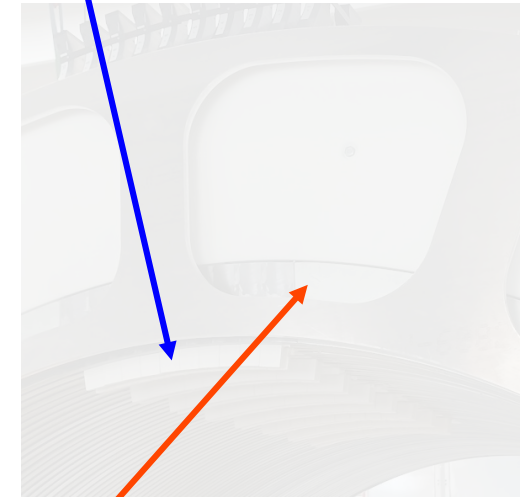
Summer 2022 @ INFN-LNL



# Event displays



Tracker layers ("t<sub>0</sub>")



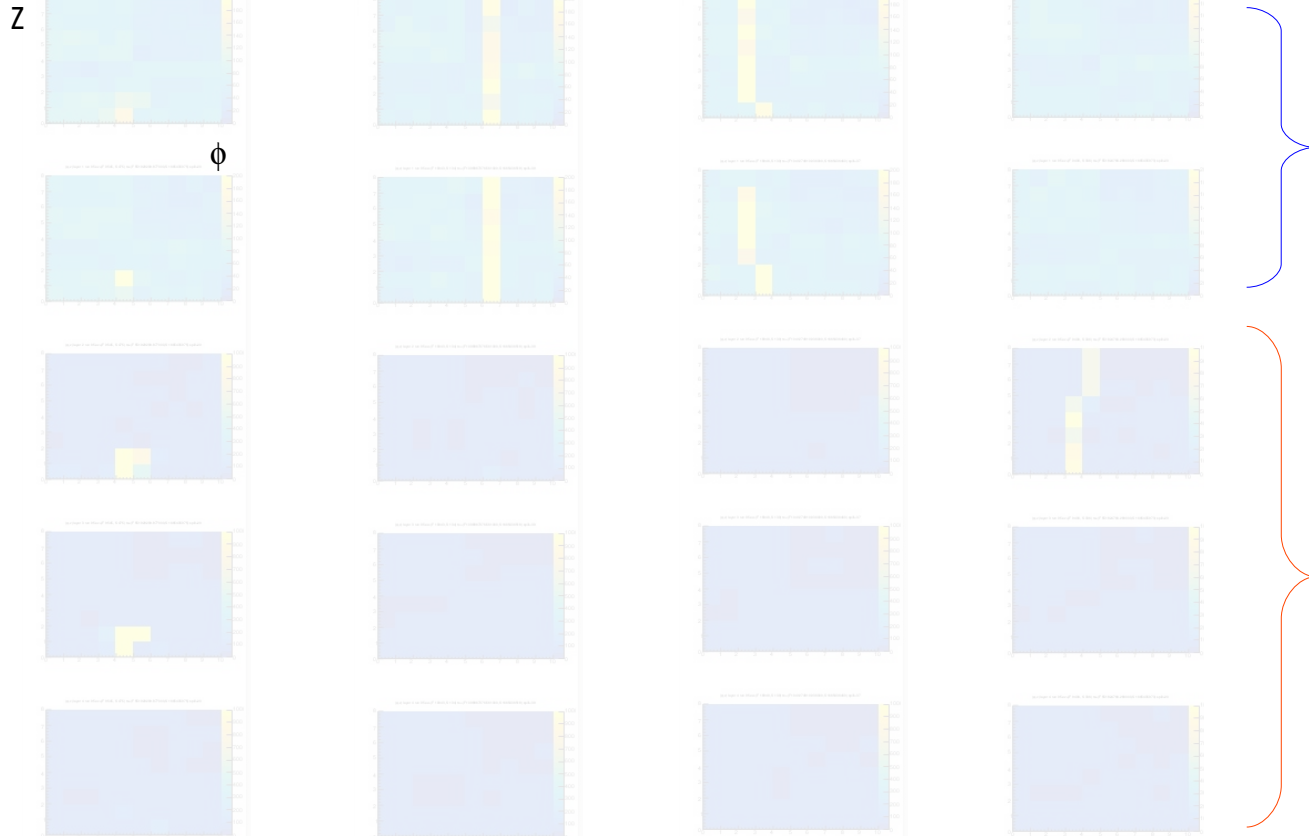
calorimeter layers

e-like

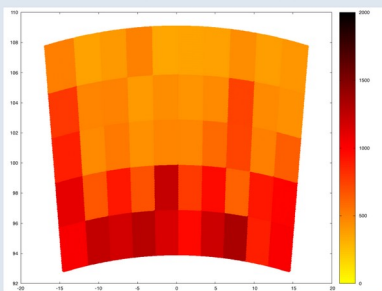
mip-like in t<sub>0</sub>-layer

mip-like in t<sub>0</sub>-layer

mip-like in 1 layer of calo



## Calibration - MIPs z = 4

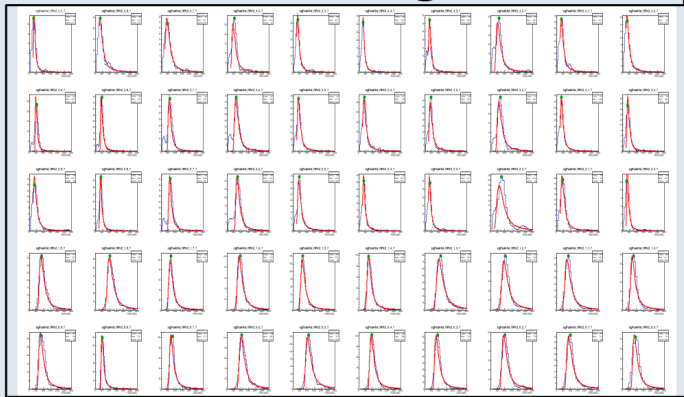


MIP peak (ADC) is color coordinate

F. Iacob, Padova Univ., 2023-02-03

20

## Calibration – Signal z = 7



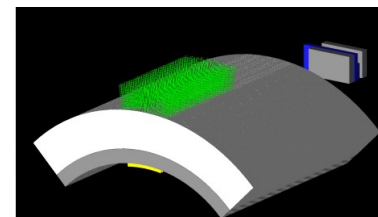
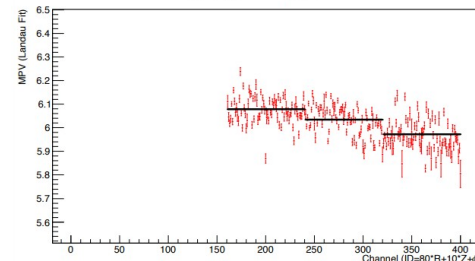
$$\varepsilon(r, \varphi, z) = \frac{\#signals(r, \varphi, z) > thr}{\#signals}$$

F. Iacob, Padova Univ., 2023-02-03

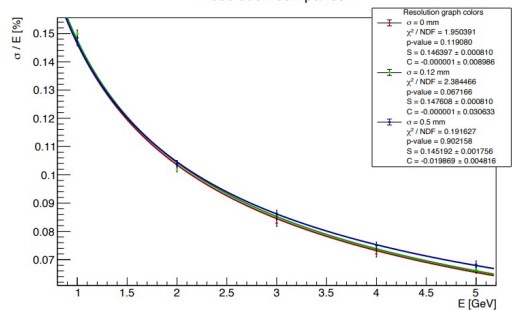
- Sqrt optimized nbin
- Moving average 3 nsamples, 2 cycle

15

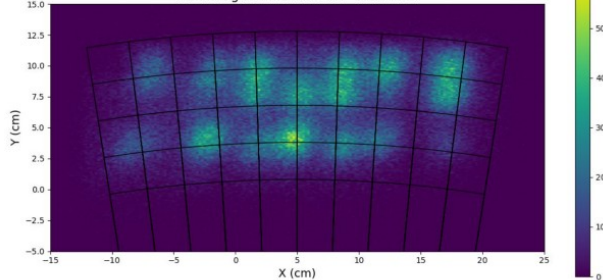
Calibration MIP peaks



Resolution comparison

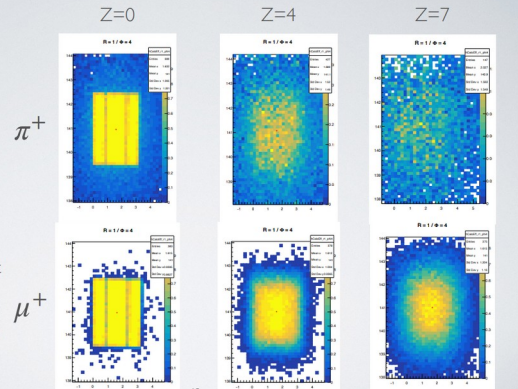


2D Histogram without cut - channel 9



## EFFICIENCY MAP

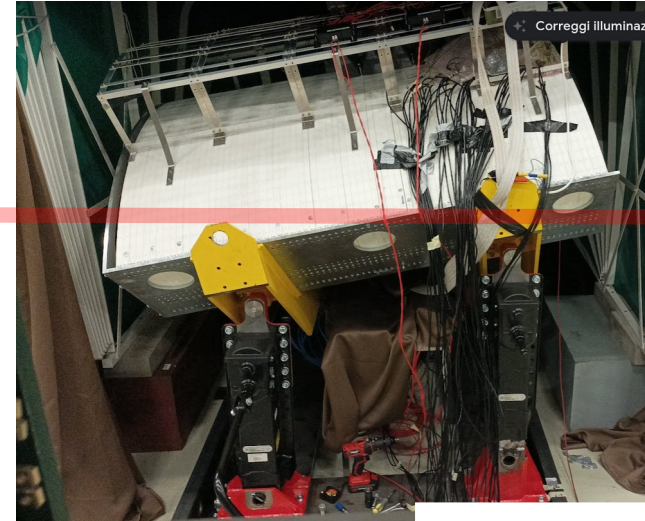
- Decreasing effect with increasing Z;
- This is way more significant for pions than for muons;
- The extrapolation from the silicon trackers is a guess that is correct for the first layer (Z=0), but it becomes less reliable as increasing depths, because of the multiple scattering ( $\mu + \pi$ ) / showering ( $\pi$ ) effects.



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# Data taking with the demonstrator

horizontal run with darkening cover



200 mrad tilt run

Beam spot at the detector  
upstream face after several runs  
illuminating different regions of the  
detector

