

# Dynamical Cobordism and Intersecting End-of-the-World Branes

Based on 2312.16286 with R. Angius, A. Uranga

Andriana Makridou String Phenomenology 2024 Padova, June 25th 2024

## Background

Cobordism Conjecture: [McNamara, Vafa '19]

All Cobordism Classes should be trivial:  $\Omega_k^{QG} = 0$ 



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How can we study them when we don't strictly have a moduli space?

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Infinite distance limits

How can we study them when we don't strictly have a moduli space?

Idea:

[Buratti, Calderon-Infante, Delgado, Uranga '21]

An infinite field distance limit can be realized as running into a cobordism wall of nothing.

When is this true?





. . .



[Sugimoto '99] [Antoniadis, Dudas, Sagnotti '99] [Angelantonj '99] ...

> Dynamical tadpoles (vs RR tadpoles)

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 $V(\phi)$ 

[Mininno, Uranga '20] [Basile, Raucci, Thomée '22] [Mourad, Sagnotti '23]

Spacetime-dependent solutions (instead of maximally-symmetric vacuum) extending over finite spacetime distance  $\Delta$ featuring a Ricci curvature singularity at infinite field distance  $D \to \infty$ 

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Interpretation: [Buratti, Delgado, Uranga '21]

Spacetime is cut-off due to running onto a cobordism defect of the initial theory

**Dynamical Cobordism** 

[Sugimoto '99] [Antoniadis, Dudas, Sagnotti '99] [Angelantonj '99]





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**Dynamical Cobordism** 

Dynamical cobordisms obey the scaling relations

$$\Delta \sim e^{-\frac{1}{2}\delta D}, \quad |\mathscr{R}| \sim e^{\delta D}, \quad \delta > 0.$$

[Buratti, Calderon-Infante, Delgado, Uranga '21]

## Universal description

[Angius, Calderon-Infante, Delgado, Huertas, Uranga '21]

critical exponent  $\delta \leftrightarrow$  universal local description

$$S = \int d^d x \sqrt{-g} \left(\frac{1}{2}R - \frac{1}{2}(\partial\phi)^2 - V(\phi)\right)$$

$$ds_d^2 = e^{-2\sigma(y)} ds_{d-1}^2 + dy^2$$

Here  $ds_{d-1}^2$  flat, see talk by Jesús for AdS case!



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Local description near ETW brane:

$$\phi(y) \sim -\frac{2}{\delta}\log y, \qquad \sigma(y) \sim -\frac{4}{(d-2)\delta^2}\log y + \frac{1}{2}\log c$$

Leading behaviour of potential:

$$V(\phi) \sim -ace^{\delta\phi}, \qquad \delta = 2\sqrt{\frac{d-1}{d-2}(1-a)}$$



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Many examples fall within this universal description [Angius, Calderon-Infante, Delgado, Huertas, Uranga '21]

[Witten '82] Bubble of Nothing $\delta = \sqrt{2/7}, d = 4$ a = 0D2-brane $\delta = \sqrt{6}, d = 4$ a = 20/21[Sugimoto '99] USp(32) string $\delta = \sqrt{6}, d = 10$ a = 0



### Why do we care?

[Angius, Calderon-Infante, Delgado, Huertas, Uranga '21]



Dynamical Cobordism solutions are quite constrained

$$\delta$$
 is bound from above for  $V \le 0$ :  $\delta \le 2\sqrt{\frac{d-1}{d-2}}$ 

and bound from below for  $V \ge 0$ :  $\delta \ge 2\sqrt{\frac{d-1}{d-2}}$ 

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Can Dynamical Cobordism be used as guiding principle for new solutions? Yes! [e.g. Blumenhagen, Cribiori, Kneissl, AM '22, Blumenhagen, Kneissl, Wang '23]

see [Blumenhagen, Kneissl, Wang '23] for such a discussion  
Parallels with (Sharpened) Distance Conjecture: 
$$\lambda \ge \frac{1}{\sqrt{d-2}}$$
. Anything special at  $\delta_{crit}$ ?  
[Etheredge, Heidenreich, Kaya, Qiu, Rudelius '22]

Can this be viewed as a criterion for "good" singularities? See e.g. [Gubser '00]

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See talks by Jakob Moritz, Björn Friedrich

$$S = \int d^{n+2}x \sqrt{-g} \left( \frac{1}{2}R - \frac{1}{2}(\partial\phi_1)^2 - \frac{1}{2}(\partial\phi_2)^2 - \frac{\alpha}{2}\partial_\mu\phi_1\partial^\mu\phi_2 - V(\phi_1, \phi_2) \right)$$

$$ds_{n+2}^2 = e^{2A(y_1, y_2)} ds_n^2 + e^{2B(y_1, y_2)} dy_1^2 + e^{2C(y_1, y_2)} dy_2^2$$

 $\phi_1 = \phi_1(y_1), \ \phi_2 = \phi_2(y_2)$ 



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We want: ETW-1 solutions for constant  $y_i$ Both scalars exploding around the origin

$$A(y_1, y_2) = -\sigma_1(y_1) - \sigma_2(y_2)$$
$$B(y_1, y_2) = -\sigma_2(y_2)$$
$$C(y_1, y_2) = -\sigma_1(y_1)$$

Intersecting solution conformally flat!

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Local description ansatz:

$$\sigma_1(y_1) = -\frac{4}{n\delta_1^2} \log y_1 \qquad \sigma_2(y_2) = -\frac{4}{n\delta_2^2} \log y_2$$
  
$$\phi_1(y_1) = -\frac{2}{\delta_1} \log y_1 \qquad \phi_2(y_2) = -\frac{2}{\delta_2} \log y_2$$

$$V = -c_1 v_1 y_1^{-2} y_2^{-2a_2} - c_2 v_2 y_1^{-2a_1} y_2^{-2} \equiv V_1 + V_2$$
$$= -c_1 v_1 e^{\delta_1 \phi_1} e^{a_2 \delta_2 \phi_2} + c_2 v_2 e^{a_1 \delta_1 \phi_1} e^{\delta_2 \phi_2}$$

## Example: $S^{p_1} \times S^{p_2}$ compactifications

Setup: Einstein gravity for  $(n + p_1 + p_2 + 2)$ -dimensional space, reduced over  $S^{p_1} \times S^{p_2}$ 

$$S_{n+2} = \frac{1}{2} \int d^{n+2}x \sqrt{-g_{n+2}} \Big( R_{n+2} - |\partial\rho_1|^2 - |\partial\rho_2|^2 - \frac{2}{n+1} \partial_\mu \rho_1 \partial^\mu \rho_2 + \frac{p_1(p_1-1)}{2} (\frac{n}{n+p_1})^2 e^{(\alpha_1+\beta_1)\rho_1 + \alpha_2\rho_2} + \frac{p_2(p_2-1)}{2} (\frac{n}{n+p_2})^2 e^{(\alpha_2+\beta_2)\rho_2 + \alpha_1\rho_1} \Big)$$

ETW-2 solution: 
$$ds_{n+2}^2 = y_1^{\frac{2p_1}{n+p_1}} y_2^{\frac{2p_2}{n+p+2}} ds_n^2 + y_2^{\frac{2p_2}{n+p_2}} dy_1^2 + y_1^{\frac{2p_1}{p_1+1}} dy_2^2$$



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Spacetime distance:

$$\Delta = \int [e^{-2\sigma_1} dy_1^2 + e^{-2\sigma_2} dy_2^2]^{1/2}$$

Field space distance:

$$\mathcal{D} = \int [d\phi_1^2 + d\phi_2^2 + \alpha d\phi_1 d\phi_2]^{1/2}$$

Pick a path 
$$y_i(t)$$
  
 $\Delta = \int [\gamma_1^2 t^{2r_1} + \gamma_2^2 t^{2r_2}]^{1/2} dt$   
 $y_1(t) = t^{\gamma_1}$   
 $y_2(t) = t^{\gamma_2}$   
 $r_1 = \frac{4\gamma_2}{n\delta_2^2} + \gamma_1 - 1$   
 $\Im = -2\left(\frac{\gamma_1^2}{\delta_1^2} + \frac{\gamma_2^2}{\delta_2^2} + \frac{\alpha\gamma_1\gamma_2}{\delta_1\delta_2}\right)$ 



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Scaling relations still hold But are the critical exponent is **path-dependent** 

$$\delta_{int} = \left(\frac{\gamma_1^2}{\delta_1^2} + \frac{\gamma_2^2}{\delta_2^2} + \frac{\alpha \gamma_1 \gamma_2}{\delta_1 \delta_2}\right)^{-1/2} (r_i + 1)$$





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Consider the case of a single scalar:  $S = \left[ d^{n+2}x \sqrt{-g} \left( \frac{1}{2}R - \frac{1}{2}(\partial \phi)^2 - V(\phi) \right) \right],$  $ds_{n+2}^2 = e^{2A(y_1, y_2)} ds_n^2 + e^{2B(y_1, y_2)} dy_1^2 + e^{2C(y_1, y_2)} dy_2^2$ 

with the logarithmic dependence:  $A = a_1 \log y_1 + a_2 \log y_2$ ,  $B = b_2 \log y_2$ ,  $C = c_1 \log y_1$ 

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Eoms satisfied for:  $d_1 = d_2 = -\sqrt{n}, a_1 = a_2 = b_2 = c_1 = 1$ 

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#### Explanation:

After a change of coordinates the scalar depends non-trivially on a single (new) coordinate

We have a codimension-1 ETW brane - recombined brane instead of intersection!

See [Angius, Delgado, Uranga '22] for an explicit such realization!

## Summary and Outlook

Can construct intersecting Dynamical Cobordism solutions

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A universal local description exists in terms of the critical exponents \delta_1, \delta_2
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What is the interplay with relevant Swampland Conjectures?

e.g. Convex Hull Conjecture, Sharpened Distance Conjecture [Calderón-Infante, Uranga, Valenzuela '20] [Etheredge, Heidenreich, Kaya, Kiu, Rudelius '22]

#### What is the interplay with other ETW/BON approaches?

e.g. Blanco-Pillado, Friedrich, Garcia-Etxebarria, Hebecker, Huertas, Montero, Moritz, Sousa, Sugimoto, Suzuki, Valenzuela, Walcher ....

Does  $\delta$  have a physical meaning?

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Thank you!

For more Dynamical Cobordism see talks by Jesus, Roberta!



## Possible generalizations

Up to now: Einstein gravity, 2 DC scalars  $\phi_1, \phi_2$  with mixed kinetic term, potential

Solutions: 
$$ds_{n+2}^2 = e^{2A(y_1, y_2)} ds_n^2 + e^{2B(y_1, y_2)} dy_1^2 + e^{2C(y_1, y_2)} dy_2^2$$
  
 $\phi_1 = \phi_1(y_1), \ \phi_2 = \phi_2(y_2)$   
 $A(y_1, y_2) = -\sigma_1(y_1) - \sigma_2(y_2), \ B(y_1, y_2) = -\sigma_2(y_2), \ C(y_1, y_2) = -\sigma_1(y_1)$ 

Generalizations have a certain "cost":

1) 
$$\phi_1 \rightarrow \psi_1 = \psi_1(y_1, y_2) = b_{11} \log y_1 + b_{12} \log y_2$$
,  $\rightarrow \alpha = 0$ : mixed kinetic term killed  
 $\phi_2 \rightarrow \psi_2 = \psi_2(y_2) = b_2 \log y_2$  but now  $\psi_1$  explodes at both  $y_1 = y_2 = 0$ 

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[Angius '24] [Angius, AM, Uranga '23]

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 $A(y_1, y_2) = -\sigma_1(y_1) - \sigma_2(y_2), \ B(y_1, y_2) = -\sigma_2(y_2), \ C(y_1, y_2) = -\sigma_1(y_1)$ 

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$$||) \quad A(y_1, y_2) = -a_1 \log y_1 + a_2 \log y_2,$$

$$B(y_2) = a_2 \log y_2 - \frac{1}{2} \log c_2,$$
  
$$C(y_1) = (1 - a_1 n) \log y_1 - \frac{1}{2} \log c_1$$

 $\rightarrow \alpha = 0$ : mixed kinetic term killed but now  $\psi_1$  explodes at both  $y_1 = y_2 = 0$ 

 $\rightarrow V = -V_1 \cdot V_2 \sim -e^{\lambda_1 \phi_1} e^{\lambda_2 \phi_2}$ 

single contribution to the potential, but solution no longer conformally flat

#### Intersection at an angle

Metric Ansatz: 
$$ds_{n+2}^2 = e^{-2\sigma_1 - 2\sigma_2}(ds_n^2 + dx_1^2 + dx_2^2 + fdx_1dx_2)$$
  
=  $e^{-2\sigma_1 - 2\sigma_2}ds_n^2 + e^{-2\sigma_2}dy_1^2 + e^{-2\sigma_1}dy_2^2 + fe^{-\sigma_1 - \sigma_2}dy_1dy_2$ 

Using the logarithmic ansatz for the scalars and warp factors, the potential is now:

$$V = V_1 + V_2 + V_{12} = -c_1 v_1 y_1^{-2} y_2^{-2a_2} - c_2 v_2 y_1^{-2a_1} y_2^{-2} - c_1 c_2 \frac{2n^2 f a_1 a_2}{f^2 c_1 c_2 - 4} y_1^{-1-a_1} y_2^{-1-a_2}$$

$$V_1 = \frac{2c_1 n a_1}{f^2 c_1 c_2 - 4} [(na_1 + a_1 - 1)] y_1^{-2} y_2^{-2a_2}$$
 Subleading for  $a_1, a_2 < 1$ 

$$V_2 = \frac{2c_2na_2}{f^2c_1c_2 - 4}[(na_2 + a_2 - 1)]y_1^{-2a_1}y_2^{-2}$$

with 
$$a_i = \frac{n \pm \sqrt{n + 2(n+1)v_1(-c_1c_2 - 4)}}{2n(n+1)}$$
 and  $\delta_i^2 = \frac{4}{a_i n} = \frac{8(n+1)^2}{n \pm \sqrt{n + 2(n+1)v_i(-c_1c_2 - 4)}}$ 

$$S = \int d^{n+2}x \sqrt{-g} \left( \frac{1}{2}R - \frac{1}{2}(\partial\phi_1)^2 - \frac{1}{2}(\partial\phi_2)^2 - \frac{\alpha}{2}\partial_\mu\phi_1\partial^\mu\phi_2 - V(\phi_1, \phi_2) \right)$$

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Local description ansatz:

$$\sigma(y_1) = -a_1 \log y_1 + \frac{1}{2} \log c_1$$

$$\sigma(y_2) = -a_2 \log y_2 + \frac{1}{2} \log c_2$$
(assuming  $a_1, a_2 < 1$ )

$$\phi_1(y_1) = -b_1 \log y_1$$
  $\phi_2(y_2) = -b_2 \log y_2$ 

$$V = -c_1 v_1 y_1^{-2} y_2^{-2a_2} - c_2 v_2 y_1^{-2a_1} y_2^{-2} \equiv V_1 + V_2$$