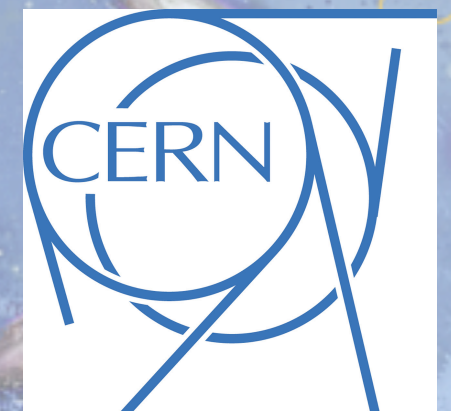


A Bumpy Ride: Distances Beyond Minkowski space

Irene Valenzuela

CERN



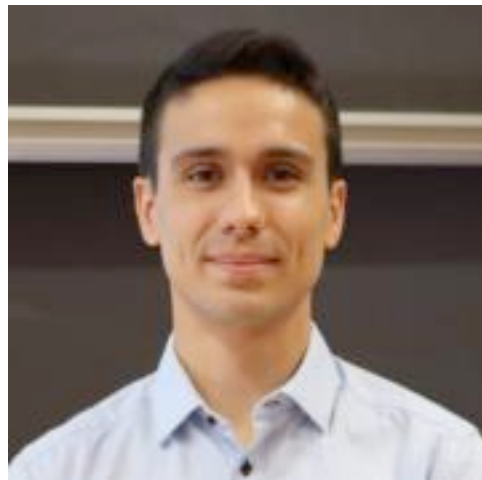
StringPheno 2024, Padova

24-28 JUNE 2024

Based on work in collaboration with members of my group:



Ignacio Ruiz
(Nacho)



José Calderón-
Infante



Alessandra
Grieco



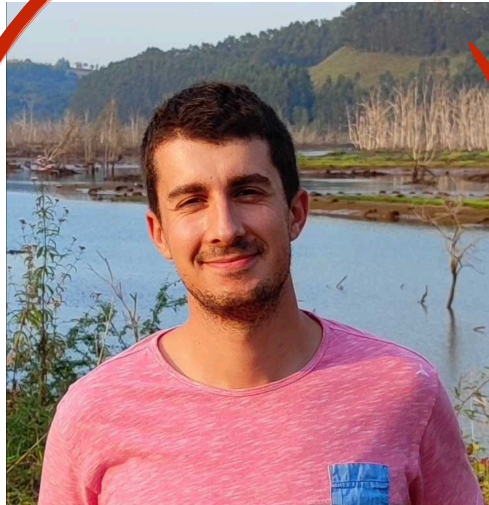
Amineh
Mohseni

and other amazing collaborators:

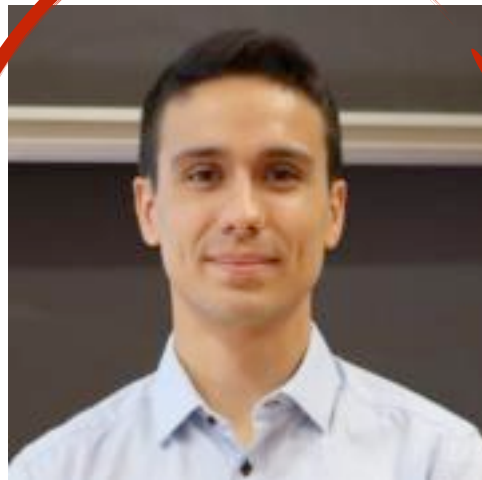
Alberto Castellano
Muldrow Etheredge
Ben Heidenreich

Miguel Montero
Tom Rudelius
Cumrun Vafa

Based on work in collaboration with members of my group:



Ignacio Ruiz
(Nacho)



José Calderón-
Infante



Alessandra
Grieco



Amineh
Mohseni

and other amazing collaborators:

Alberto Castellano
Muldrow Etheredge
Ben Heidenreich

Miguel Montero
Tom Rudelius
Cumrun Vafa

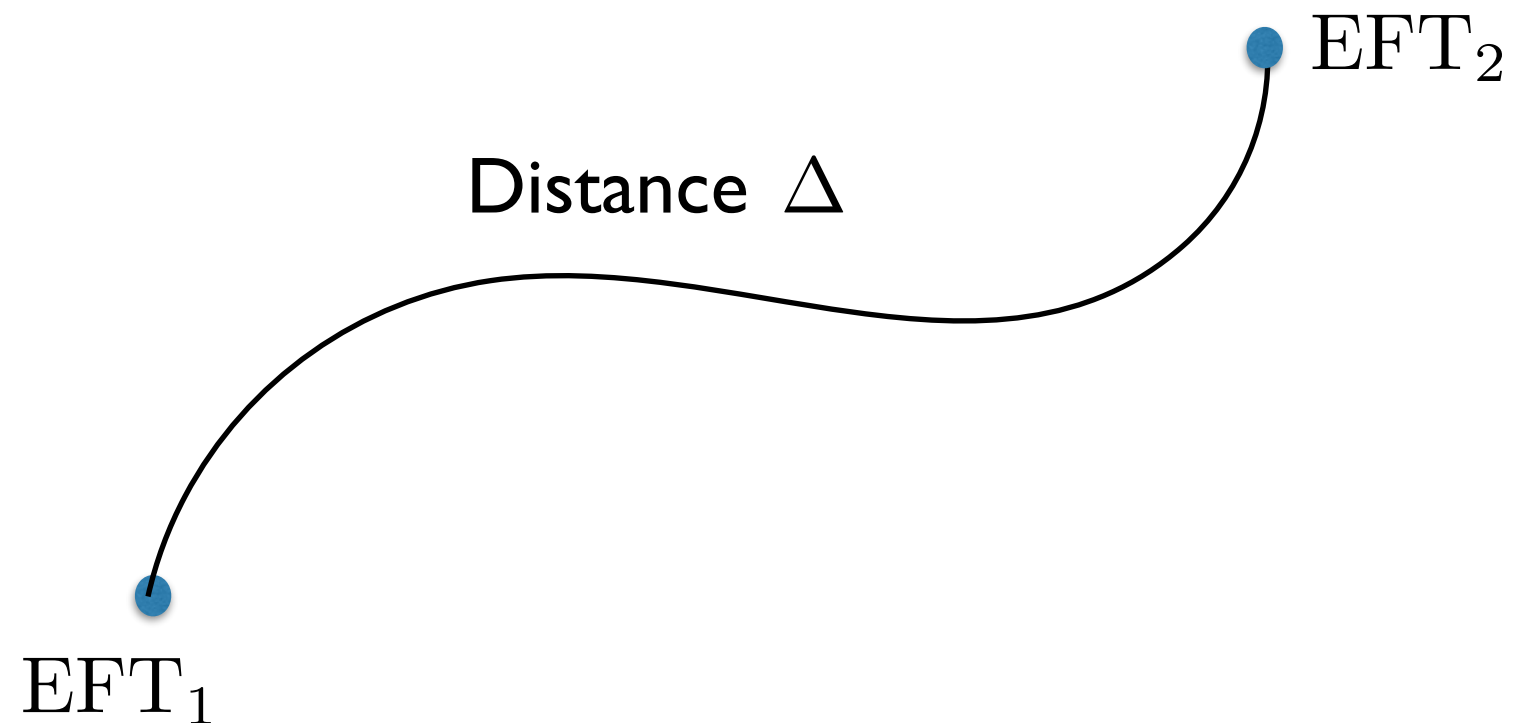
→ applying this year!!

Plan of the talk

A Bumpy Ride: Distances Beyond Minkowski space

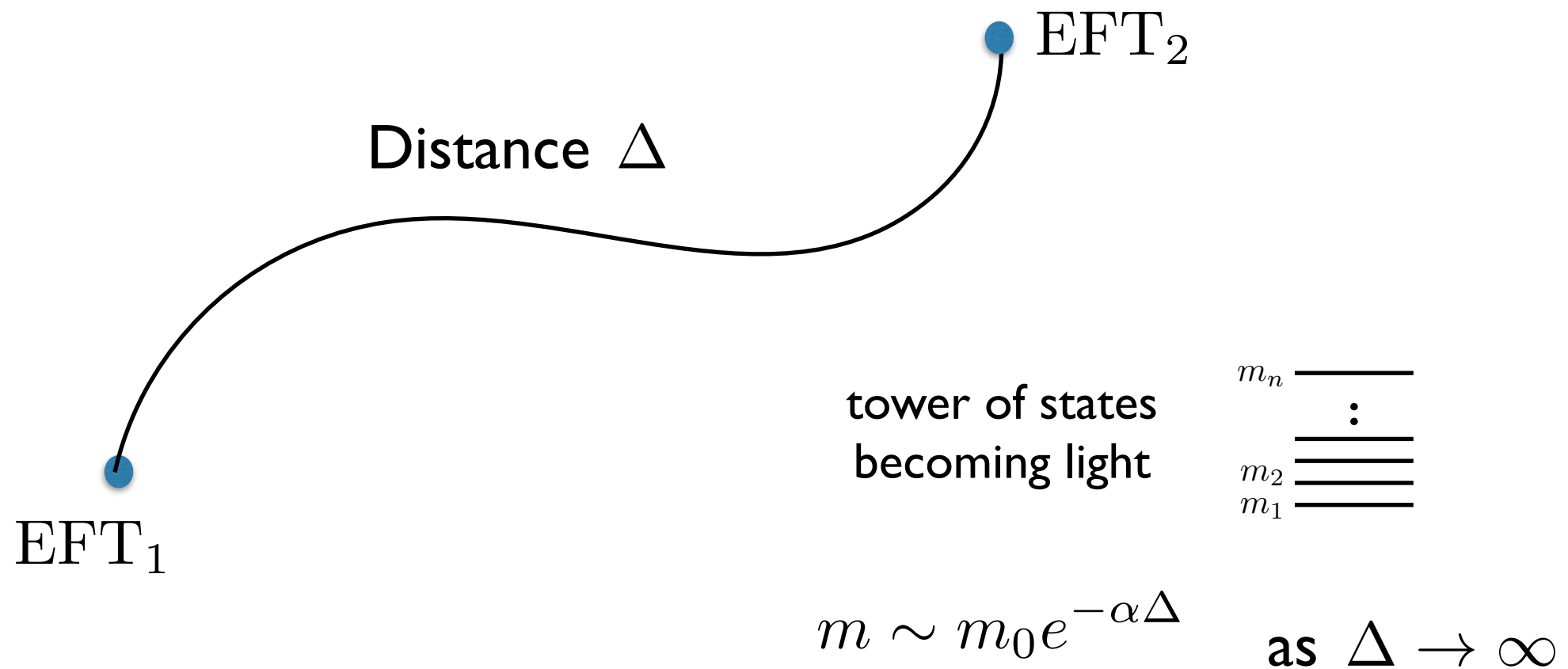
Plan of the talk

A Bumpy Ride: Distances Beyond Minkowski space



Plan of the talk

A Bumpy Ride: Distances Beyond Minkowski space



Plan of the talk

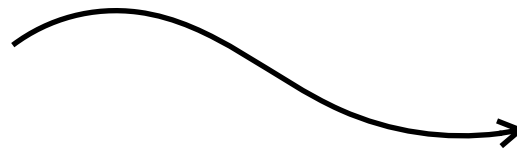
A Bumpy Ride: Distances Beyond Minkowski space



Minkowski space
(EFTs connected by a moduli space)

Plan of the talk

A Bumpy Ride: Distances Beyond Minkowski space



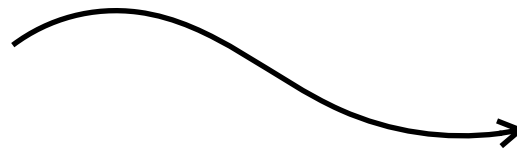
Minkowski space
(EFTs connected by a moduli space)

Distances in AdS/CFT

- New corners: Non-critical strings
- Absence of scale separation in AdS

Plan of the talk

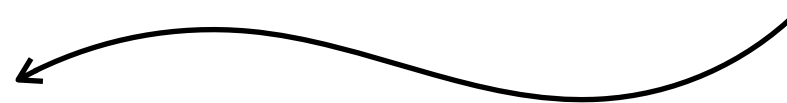
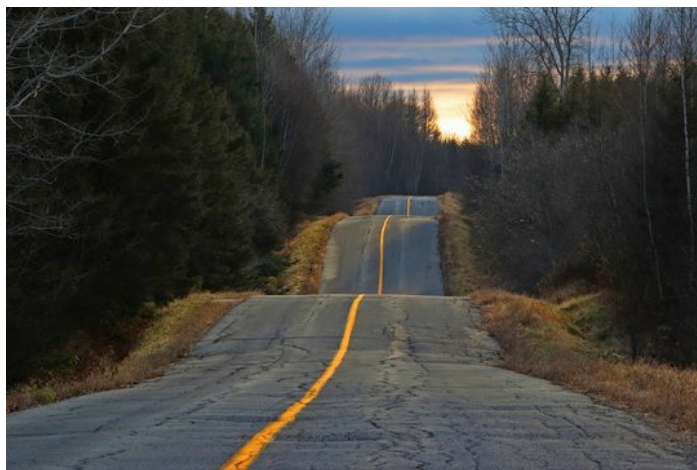
A Bumpy Ride: Distances Beyond Minkowski space



Minkowski space
(EFTs connected by a moduli space)

Distances in AdS/CFT

- New corners: Non-critical strings
- Absence of scale separation in AdS



Generalised distance beyond moduli spaces

I) Minkowski space

$$\mathcal{L} = \frac{M_{\text{pl},d}^{d-2}}{2} \left(R - \frac{1}{2} g_{ij}(\phi) (\partial\phi)^2 + \dots \right)$$

↪ field metric



[Castellano,Ruiz,IV'23]

Moduli space distance: $\Delta\phi = \int_Q^P \sqrt{g_{ij} \frac{d\phi^i}{ds} \frac{d\phi^j}{ds}} ds$

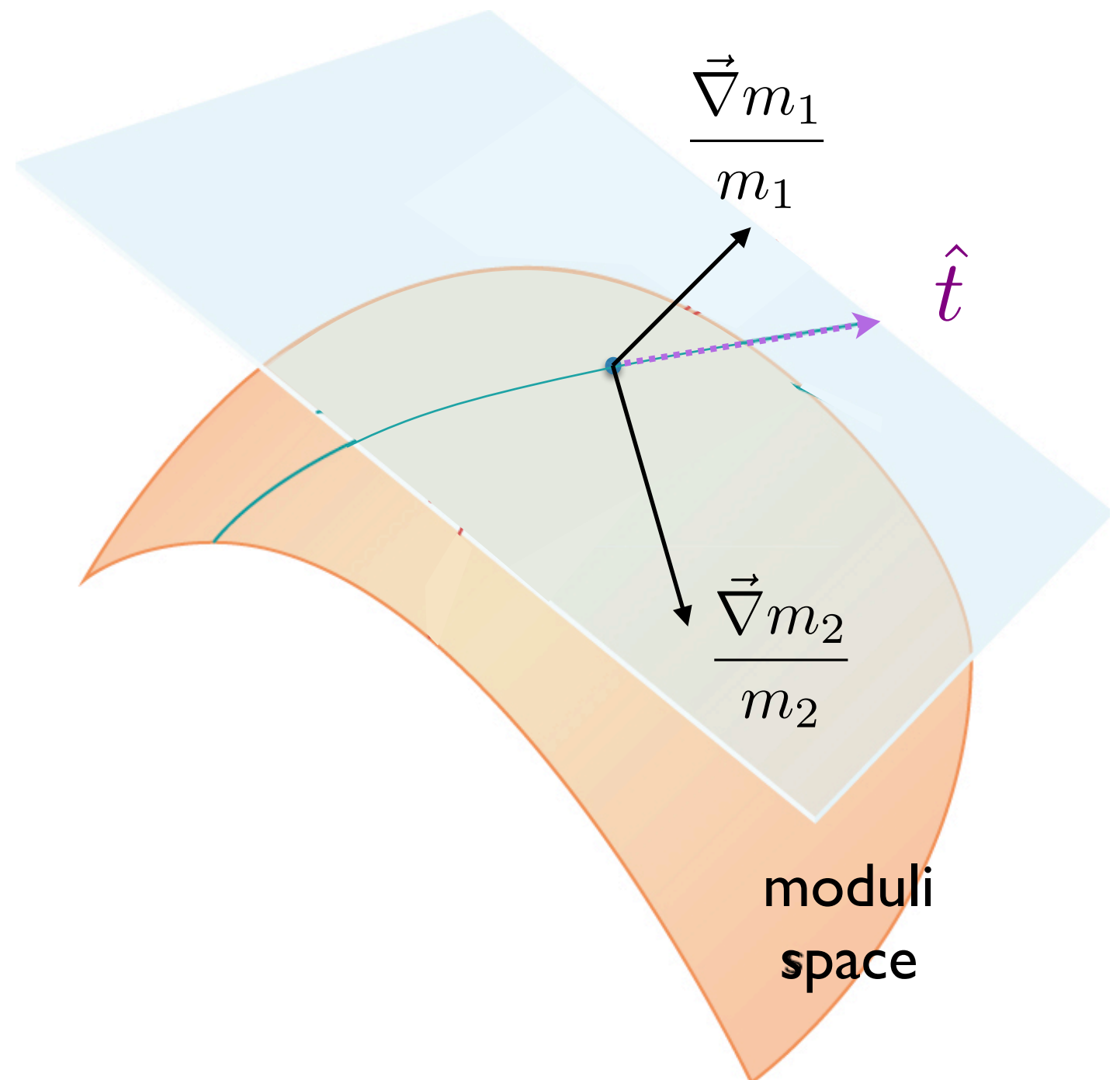
[Etheredge,Heidenreich,Rudelius,Ruiz,IV'24]

Tower and species vectors

Story of success for the Distance conjecture: $m \sim m_0 e^{-\alpha \Delta}$ [Ooguri-Vafa'06]

Tower vector: $\frac{\vec{\nabla} m_a}{m_a}$

a : labels number of light towers



Tower and species vectors

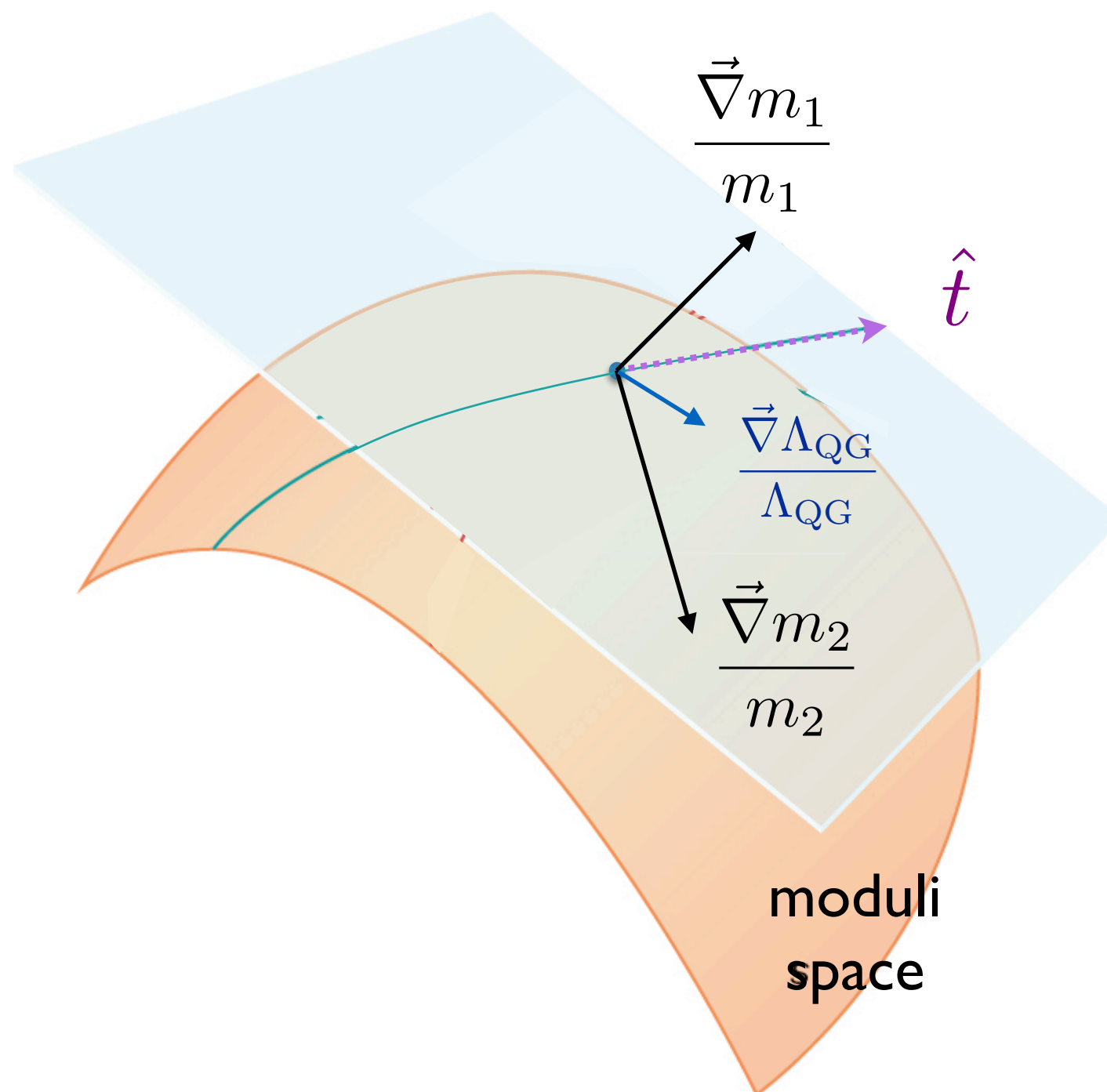
Story of success for the Distance conjecture: $m \sim m_0 e^{-\alpha \Delta}$ [Ooguri-Vafa'06]

Tower vector: $\frac{\vec{\nabla} m_a}{m_a}$

a : labels number of light towers

Species vector: $\frac{\vec{\nabla} \Lambda_{\text{QG}}}{\Lambda_{\text{QG}}}$

→ scale at which the local QFT description breaks down



Tower and species vectors

Story of success for the Distance conjecture: $m \sim m_0 e^{-\alpha \Delta}$ [Ooguri-Vafa'06]

Tower vector: $\frac{\vec{\nabla} m_a}{m_a}$

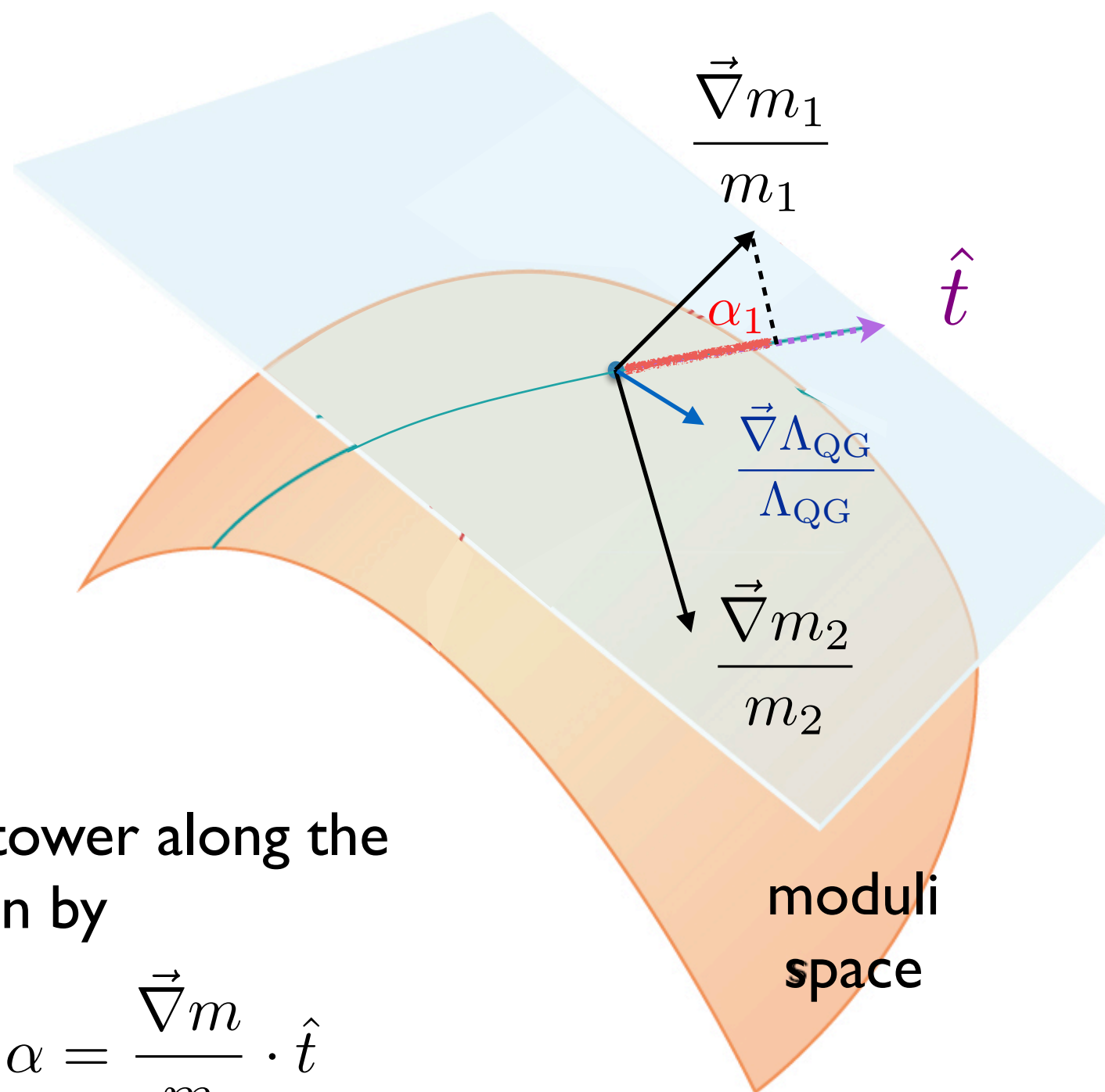
a : labels number of light towers

Species vector: $\frac{\vec{\nabla} \Lambda_{QG}}{\Lambda_{QG}}$

scale at which the local QFT description breaks down

The exponential rate of the tower along the direction \hat{t} is given by

$$m_{\text{tower}} \sim m_0 e^{-\alpha \Delta \phi} \quad \alpha = \frac{\vec{\nabla} m}{m} \cdot \hat{t}$$



Towers in Asymptotically Minkowski Space

So far, all string theory examples satisfy the [Emergent String Conjecture](#), so the leading tower is (in some dual frame) either: [Lee,Lerche,Weigand'19]

- ❖ Tower of Kaluza-Klein states
- ❖ Tower of string oscillator modes of a critical string

Towers in Asymptotically Minkowski Space

So far, all string theory examples satisfy the **Emergent String Conjecture**, so the leading tower is (in some dual frame) either: [Lee,Lerche,Weigand'19]

- ❖ Tower of Kaluza-Klein states
- ❖ Tower of string oscillator modes of a critical string

If so, for every **regular*** infinite distance limit:

$$\frac{\vec{\nabla} m_a}{m_a} \cdot \frac{\vec{\nabla} m_b}{m_b} = \frac{1}{d-2} + \frac{1}{n_a} \delta_{ab} \qquad \frac{\vec{\nabla} m}{m} \cdot \frac{\vec{\nabla} \Lambda_{\text{QG}}}{\Lambda_{\text{QG}}} = \frac{1}{d-2}$$

d = space-time dimension
 n = # of extra dimensions

$$\left| \frac{\vec{\nabla} \Lambda_{\text{QG}}}{\Lambda_{\text{QG}}} \right|^2 = \frac{1}{d-2} - \frac{1}{D-2}$$

Taxonomy rules

[Etheredge,Heidenreich,Rudelius,Ruiz,IV'24]

- *regular:
- 1) Leading tower is non-degenerate
 - 2) The decompactification manifold is asymptotically empty

Towers in Asymptotically Minkowski Space

So far, all string theory examples satisfy the **Emergent String Conjecture**, so the leading tower is (in some dual frame) either: [Lee,Lerche,Weigand'19]

- ❖ Tower of Kaluza-Klein states
- ❖ Tower of string oscillator modes of a critical string

If so, for every **regular*** infinite distance limit: Pattern observed in [Castellano,Ruiz,IV'23]

$$\frac{\vec{\nabla} m_a}{m_a} \cdot \frac{\vec{\nabla} m_b}{m_b} = \frac{1}{d-2} + \frac{1}{n_a} \delta_{ab} \quad \left(\frac{\vec{\nabla} m}{m} \cdot \frac{\vec{\nabla} \Lambda_{\text{QG}}}{\Lambda_{\text{QG}}} = \frac{1}{d-2} \right)$$

d = space-time dimension
 n = # of extra dimensions

$$\left| \frac{\vec{\nabla} \Lambda_{\text{QG}}}{\Lambda_{\text{QG}}} \right|^2 = \frac{1}{d-2} - \frac{1}{D-2}$$

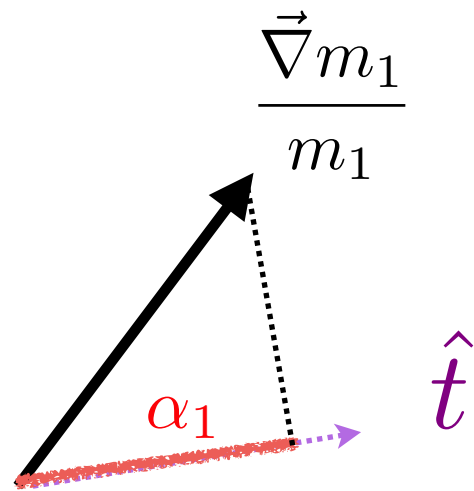
Taxonomy rules

[Etheredge,Heidenreich,Rudelius,Ruiz,IV'24]

- *regular:
- 1) Leading tower is non-degenerate
 - 2) The decompactification manifold is asymptotically empty

Towers in Asymptotically Minkowski Space

The microscopic nature of the tower constrains the lengths of the tower vectors (and therefore, the exponential rate of the masses)



$$m_{\text{tower}} \sim m_0 e^{-\alpha \Delta \phi}$$

$$\alpha = \frac{\vec{\nabla} m}{m} \cdot \hat{t}$$

KK tower: $\left| \frac{\vec{\nabla} m}{m} \right| = \sqrt{\frac{d+n-2}{n(d-2)}}$
(without warping)

string tower: $\left| \frac{\vec{\nabla} m}{m} \right| = \frac{1}{\sqrt{d-2}}$

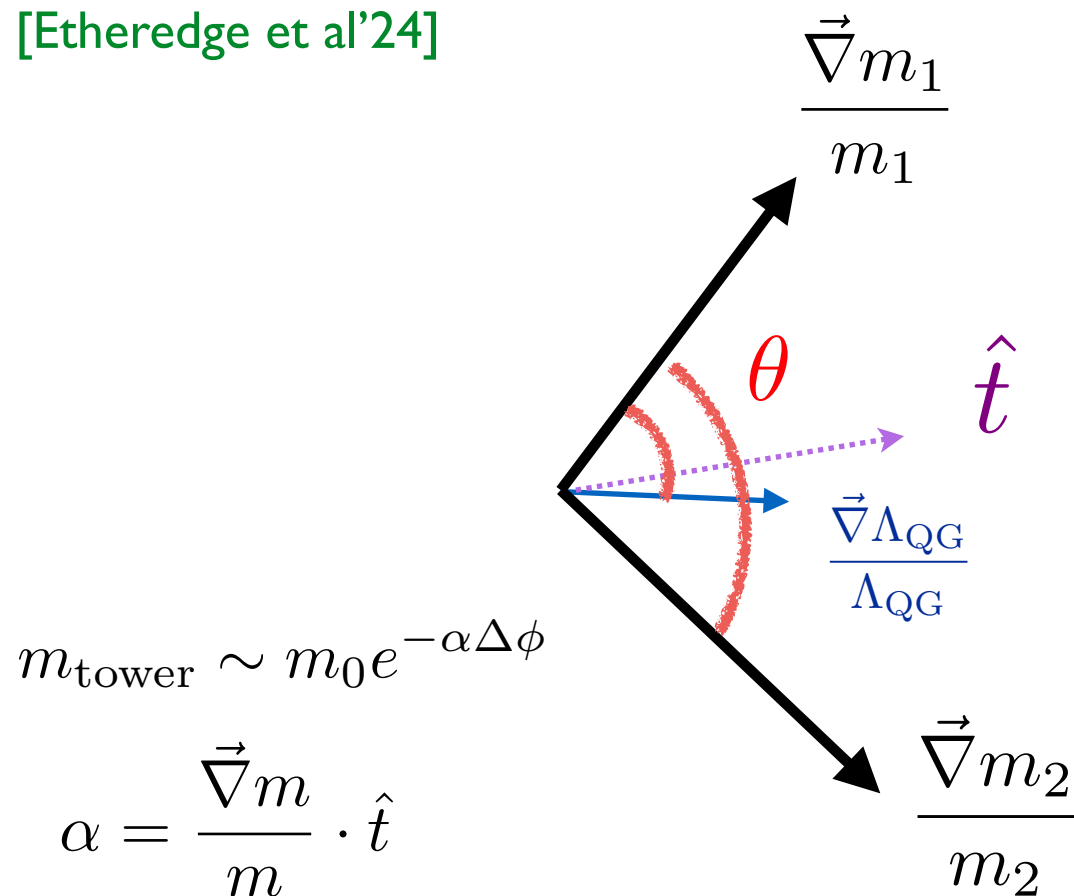
d = space-time dimension

n = number of extra dimensions

Towers in Asymptotically Minkowski Space

The microscopic nature of the tower constrains the lengths of the tower vectors (and therefore, the exponential rate of the masses) but also the angles and the full geometry of their convex hull (frame simplex)

[Etheredge et al'24]



KK tower: $\left| \frac{\vec{\nabla} m}{m} \right| = \sqrt{\frac{d+n-2}{n(d-2)}}$
(without warping)

string tower: $\left| \frac{\vec{\nabla} m}{m} \right| = \frac{1}{\sqrt{d-2}}$

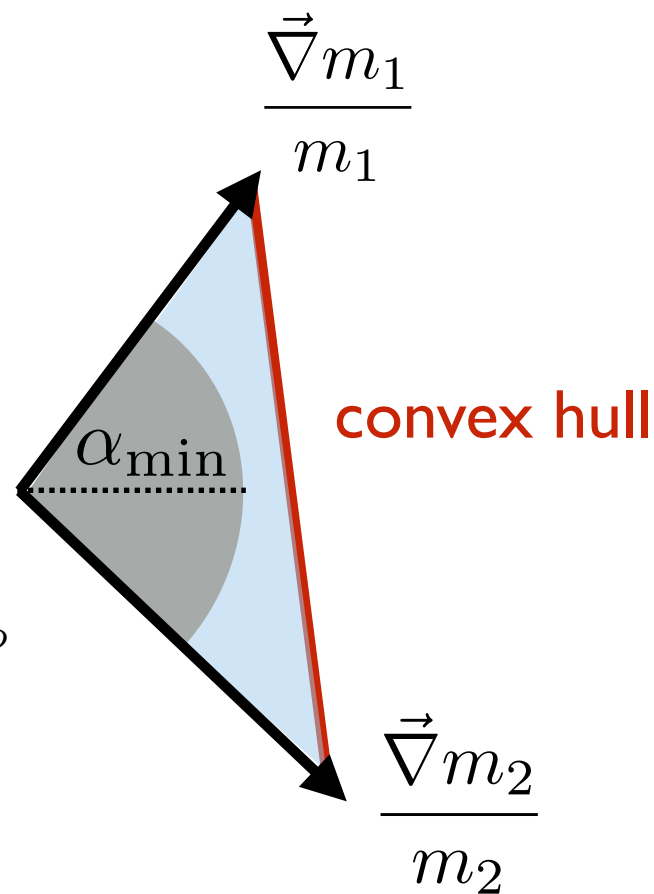
d = space-time dimension

n = number of extra dimensions

Towers in Asymptotically Minkowski Space

The microscopic nature of the tower constrains the lengths of the tower vectors (and therefore, the exponential rate of the masses) but also the angles and the full geometry of their convex hull (frame simplex)

[Etheredge et al'24]



$$m_{\text{tower}} \sim m_0 e^{-\alpha \Delta \phi}$$

$$\alpha = \frac{\vec{\nabla} m}{m} \cdot \hat{t}$$

KK tower: $\left| \frac{\vec{\nabla} m}{m} \right| = \sqrt{\frac{d+n-2}{n(d-2)}}$
(without warping)

string tower: $\left| \frac{\vec{\nabla} m}{m} \right| = \frac{1}{\sqrt{d-2}}$

d = space-time dimension

n = number of extra dimensions

Convex Hull Distance Conjecture with $\alpha_{\min} = \frac{1}{\sqrt{d-2}}$

[Calderon-Infante, Uranga, IV'20]

[Etheredge et al'22]

(i.e., \exists tower with exponential rate $\alpha \geq \alpha_{\min}$ along any geodesic of this plane)

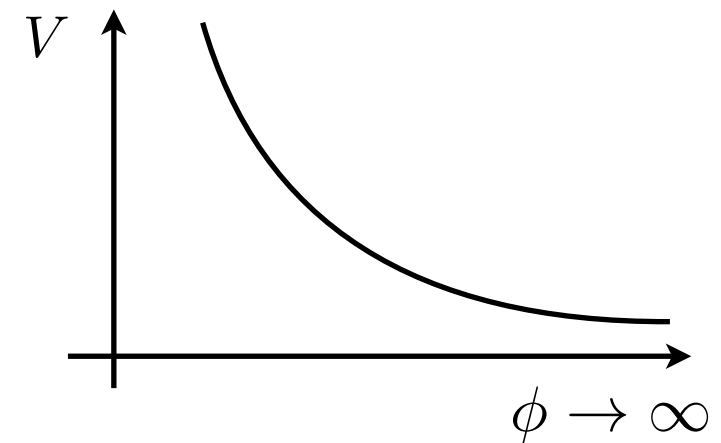
Towers in Asymptotically Minkowski Space

These constraints also hold for asymptotically Minkowski spaces
(even if there is a runaway potential):

We are studying the convex hull of the towers
in 4d $N=1$ string compactifications

Interesting interplay with BPS EFT strings

[Grieco,Ruiz,IV'ongoing] see **Alessandra Grieco's talk**
(Thursday)



Novel cosmological scenarios of transient acceleration using the
towers of states

[Casas,Montero,Ruiz'24] see **Ignacio Ruiz's talk (today)**

2) AdS space



[Calderon-Infante, IV' ongoing]

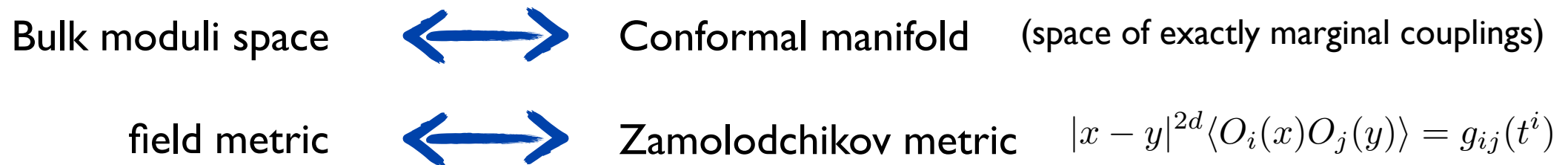
Distance Conjecture in CFTs

Consider AdS_d/CFT_{d-1}

Bulk moduli space \longleftrightarrow Conformal manifold (space of exactly marginal couplings)
field metric \longleftrightarrow Zamolodchikov metric $|x - y|^{2d} \langle O_i(x) O_j(y) \rangle = g_{ij}(t^i)$

Distance Conjecture in CFTs

Consider AdS_d/CFT_{d-1}



Distance conjecture implies:

\exists infinite tower of operators saturating the unitarity bound at every infinite distance limit measured by Zamolodchikov metric, such that

$$\gamma_J \sim e^{-\alpha d(\tau, \tau')} \quad \text{as} \quad d(\tau, \tau') \rightarrow \infty \quad \begin{array}{l} \text{[Perlmutter, Rastelli, Vafa, IV'21]} \\ \text{[Baume, Calderon-Infante'21]} \end{array}$$

anomalous dimension

$$\gamma_J = \Delta - \Delta_{\text{unitarity}}$$

distance measured by Zamolodchikov metric

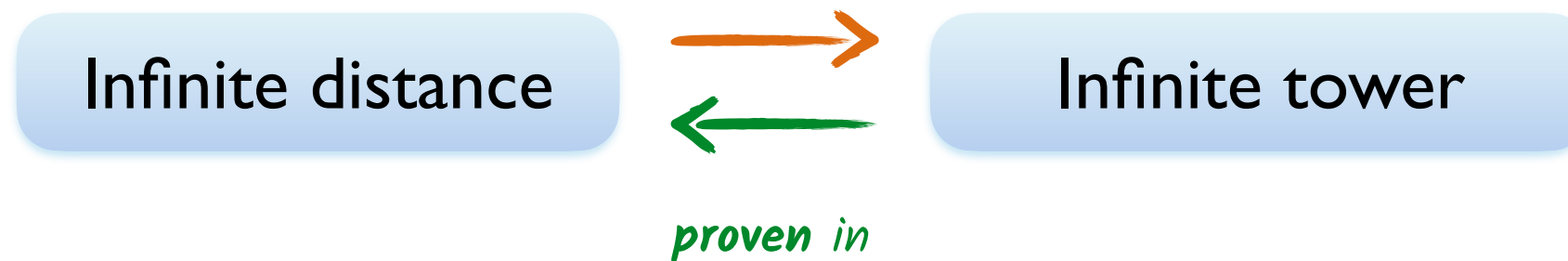
Distance Conjecture in CFTs

It works in all known examples (even beyond holographic CFTs!)



Distance Conjecture in CFTs

It works in all known examples (even beyond holographic CFTs!)

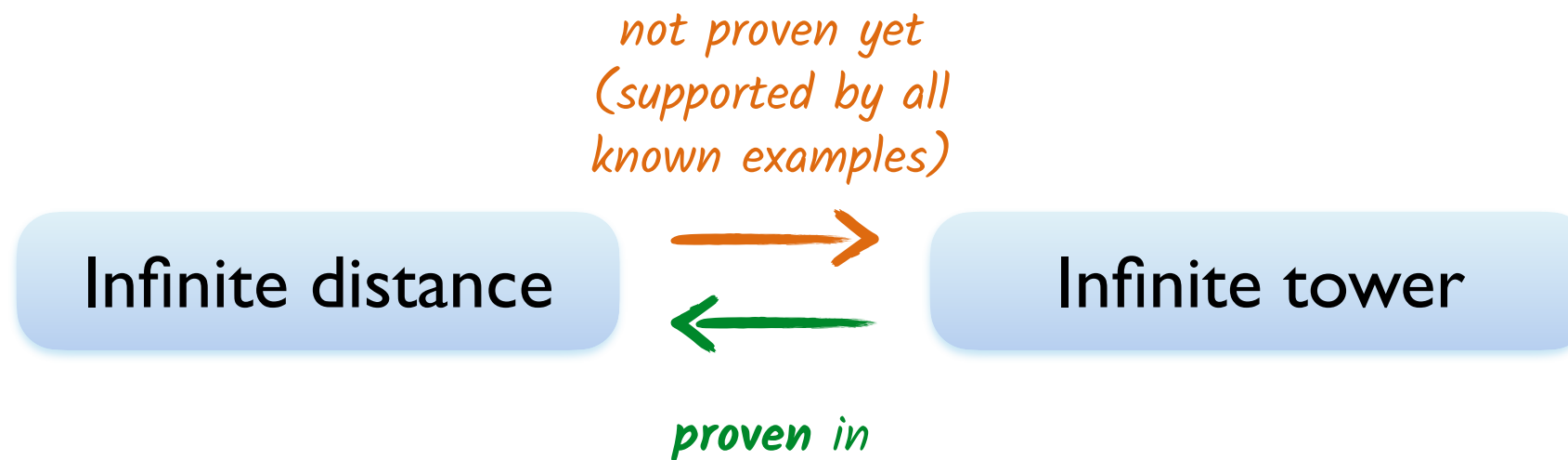


[Baume,Calderon-Infante'23] for higher spin gap in $d>3$

[Ooguri,Wang'24] for scalar gap in $d=3$

Distance Conjecture in CFTs

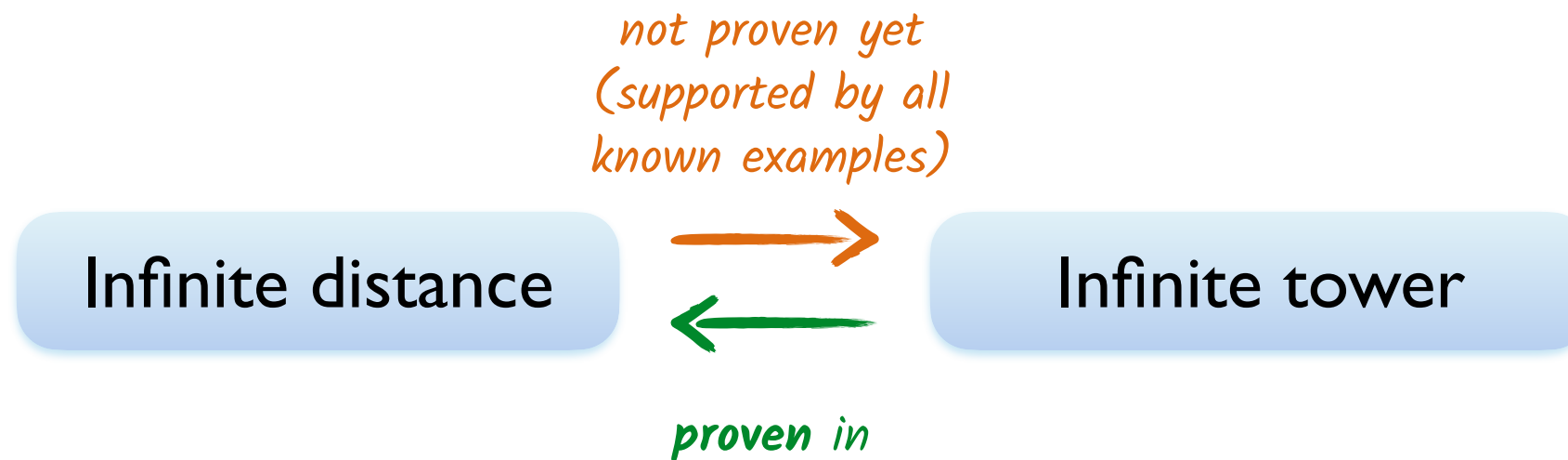
It works in all known examples (even beyond holographic CFTs!)



[Baume,Calderon-Infante'23] for higher spin gap in $d>3$
[Ooguri,Wang'24] for scalar gap in $d=3$

Distance Conjecture in CFTs

It works in all known examples (even beyond holographic CFTs!)



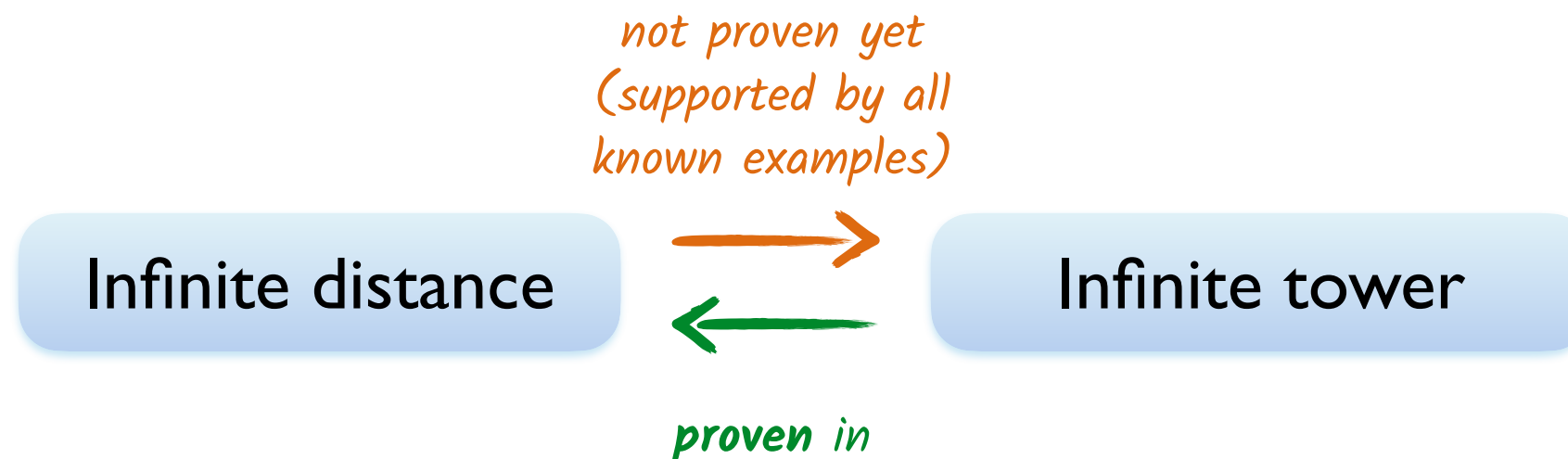
[Baume,Calderon-Infante'23] for higher spin gap in $d>3$
[Ooguri,Wang'24] for scalar gap in $d=3$

What is the nature of the tower?

Is the Emergent String Conjecture still satisfied? (KK or critical string?)

Distance Conjecture in CFTs

It works in all known examples (even beyond holographic CFTs!)



[Baume,Calderon-Infante'23] for higher spin gap in $d>3$
[Ooguri,Wang'24] for scalar gap in $d=3$

What is the nature of the tower?

Is the Emergent String Conjecture still satisfied? (KK or critical string?)

For $d = 3$: Tower of scalar modes [Kontsevich,Soibelman'00] [Acharya,Douglas'06]

For $d > 3$: Tower of higher spin modes “CFT Distance conjecture”

[Perlmutter,Rastelli,Vafa,IV'21] (see also [Baume,Calderon-Infante'21])

Classification of Towers in $d > 3$

Let us focus on 4d SCFT ($\mathcal{N} = 1, \mathcal{N} = 2, \mathcal{N} = 4$) (see though [Bobev et al'23])
(they are the only known examples of $d > 2$ non-compact conformal manifolds)

Classification of Towers in $d > 3$

Let us focus on 4d SCFT ($\mathcal{N} = 1, \mathcal{N} = 2, \mathcal{N} = 4$) (see though [Bobev et al'23])
(they are the only known examples of $d > 2$ non-compact conformal manifolds)

All known infinite distance limits are **weak coupling limits**:

$g_{YM} \rightarrow 0 \rightarrow \text{CFT}_{\text{free}} \times \text{CFT}'$ [Perlmutter, Rastelli, Vafa, IV'21]
(see also [Baume, Calderon-Infante'21])

Classification of Towers in $d > 3$

Let us focus on 4d SCFT ($\mathcal{N} = 1, \mathcal{N} = 2, \mathcal{N} = 4$) (see though [Bobev et al'23])
(they are the only known examples of $d > 2$ non-compact conformal manifolds)

All known infinite distance limits are **weak coupling limits**:

$g_{YM} \rightarrow 0 \rightarrow \text{CFT}_{\text{free}} \times \text{CFT}'$ [Perlmutter, Rastelli, Vafa, IV'21]
(see also [Baume, Calderon-Infante'21])

↓ [Maldacena, Zhiboedov'11]

Higher spin tower

$$m_{\text{tower}} \sim m_0 e^{-\alpha \Delta \phi} \quad \text{with}$$

$$\alpha = \sqrt{\frac{2c}{\dim G}}$$

central charge

gauge group getting free

Classification of Towers in $d > 3$

Let us focus on 4d SCFT ($\mathcal{N} = 1, \mathcal{N} = 2, \mathcal{N} = 4$) (see though [Bobev et al'23])
(they are the only known examples of $d > 2$ non-compact conformal manifolds)

All known infinite distance limits are **weak coupling limits**:

$g_{YM} \rightarrow 0 \rightarrow \text{CFT}_{\text{free}} \times \text{CFT}'$ [Perlmutter, Rastelli, Vafa, IV'21]
(see also [Baume, Calderon-Infante'21])

↓ [Maldacena, Zhiboedov'11]

Higher spin tower

$$m_{\text{tower}} \sim m_0 e^{-\alpha \Delta \phi} \quad \text{with}$$

$$\alpha = \sqrt{\frac{2c}{\dim G}}$$

↖ central charge

↘ gauge group getting free

No pure decompactification limits, always higher spin operators ($J > 2$)

Classification of Towers in $d > 3$

Let us focus on 4d SCFT ($\mathcal{N} = 1, \mathcal{N} = 2, \mathcal{N} = 4$) (see though [Bobev et al'23])
(they are the only known examples of $d > 2$ non-compact conformal manifolds)

All known infinite distance limits are **weak coupling limits**:

$g_{YM} \rightarrow 0 \rightarrow \text{CFT}_{\text{free}} \times \text{CFT}'$ [Perlmutter, Rastelli, Vafa, IV'21]
(see also [Baume, Calderon-Infante'21])

↓ [Maldacena, Zhiboedov'11]

Higher spin tower

$$m_{\text{tower}} \sim m_0 e^{-\alpha \Delta \phi} \quad \text{with}$$

$$\alpha = \sqrt{\frac{2c}{\dim G}}$$

↖ central charge

↘ gauge group getting free

No pure decompactification limits, always higher spin operators ($J > 2$)

Do the higher spin fields always correspond to a critical string becoming tensionless in the bulk?

Classification of Towers in $d > 3$

Exponential rate of the higher spin tower: $\alpha = \sqrt{\frac{2c}{\dim G}}$

Classification of Towers in $d > 3$

Exponential rate of the higher spin tower: $\alpha = \sqrt{\frac{2c}{\dim G}}$

Consider all large N 4d SCFTs with simple gauge factor:

Classification of Towers in $d > 3$

Exponential rate of the higher spin tower: $\alpha = \sqrt{\frac{2c}{\dim G}}$

Consider all large N 4d SCFTs with simple gauge factor:

We get only three values!

1) $\alpha = \frac{1}{\sqrt{2}}$

2) $\alpha = \sqrt{\frac{7}{12}}$

3) $\alpha = \sqrt{\frac{2}{3}}$

Classification of Towers in $d > 3$

Exponential rate of the higher spin tower: $\alpha = \sqrt{\frac{2c}{\dim G}} = \frac{1}{\sqrt{4a/c - 2}}$

Consider all large N 4d SCFTs with simple gauge factor:

We get only three values!

- 1) $\alpha = \frac{1}{\sqrt{2}}$ for $\frac{a}{c} = 1$
- 2) $\alpha = \sqrt{\frac{7}{12}}$ for $\frac{a}{c} = \frac{13}{14}$
- 3) $\alpha = \sqrt{\frac{2}{3}}$ for $\frac{a}{c} = \frac{7}{8}$

Classification of Towers in $d > 3$

Exponential rate of the higher spin tower: $\alpha = \sqrt{\frac{2c}{\dim G}} = \frac{1}{\sqrt{4a/c - 2}}$

Consider all large N 4d SCFTs with simple gauge factor:

We get only three values!

$$1) \quad \alpha = \frac{1}{\sqrt{2}} \quad \text{for } \frac{a}{c} = 1$$

$$2) \quad \alpha = \sqrt{\frac{7}{12}} \quad \text{for } \frac{a}{c} = \frac{13}{14}$$

$$3) \quad \alpha = \sqrt{\frac{2}{3}} \quad \text{for } \frac{a}{c} = \frac{7}{8}$$

see Jose Calderon-Infante's talk

We compute the thermal partition function to show that they have a different Hagedorn-like density of states (different Hagedorn temperature)

$$\rho(E) \sim e^{E/T_H}, \quad T_H = T_H(a/c) = T_H(\alpha) \quad [\text{Calderon-Infante, IV' ongoing}]$$

Classification of Towers in $d > 3$

Exponential rate of the higher spin tower: $\alpha = \sqrt{\frac{2c}{\dim G}} = \frac{1}{\sqrt{4a/c - 2}}$

Consider all large N 4d SCFTs with simple gauge factor:

We get only three values!

- 1) $\alpha = \frac{1}{\sqrt{2}}$ for $\frac{a}{c} = 1$ \rightarrow critical string in all Einstein theories
- 2) $\alpha = \sqrt{\frac{7}{12}}$ for $\frac{a}{c} = \frac{13}{14}$ \rightarrow see Jose Calderon-Infante's talk
- 3) $\alpha = \sqrt{\frac{2}{3}}$ for $\frac{a}{c} = \frac{7}{8}$ \rightarrow non-critical strings in non-Einstein theories

We compute the thermal partition function to show that they have a different Hagedorn-like density of states (different Hagedorn temperature)

$$\rho(E) \sim e^{E/T_H}, \quad T_H = T_H(a/c) = T_H(\alpha) \quad [\text{Calderon-Infante, IV' ongoing}]$$

Classification of Towers in $d > 3$

Exponential rate of the higher spin tower: $\alpha = \sqrt{\frac{2c}{\dim G}} = \frac{1}{\sqrt{4a/c - 2}}$

Consider all large N 4d SCFTs with simple gauge factor:

We get only three values!

1) $\alpha = \frac{1}{\sqrt{2}}$ for $\frac{a}{c} = 1$



critical string in all Einstein theories

2) $\alpha = \sqrt{\frac{7}{12}}$ for $\frac{a}{c} = \frac{13}{14}$

see Jose Calderon-Infante's talk

3) $\alpha = \sqrt{\frac{2}{3}}$ for $\frac{a}{c} = \frac{7}{8}$



non-critical strings
in non-Einstein theories

We compute the thermal partition function to show that they have a different Hagedorn-like density of states (different Hagedorn temperature)

$$\rho(E) \sim e^{E/T_H}, \quad T_H = T_H(a/c) = T_H(\alpha) \quad [\text{Calderon-Infante, IV' ongoing}]$$

Critical string limit

If the bulk dual is Einstein gravity
(so that $a=c$)



the higher spin tower always
comes from the critical string

Exponential rate: $\alpha = \frac{1}{\sqrt{2}}$

Critical string limit

If the bulk dual is Einstein gravity
(so that $a=c$)



the higher spin tower always
comes from the critical string

Exponential rate: $\alpha = \frac{1}{\sqrt{2}}$

How is it compatible with

$$\left| \frac{\vec{\nabla} m}{m} \right| = \frac{1}{\sqrt{d-2}} = \frac{1}{\sqrt{3}} ?$$

Critical string limit

If the bulk dual is Einstein gravity
(so that $a=c$)



the higher spin tower always
comes from the critical string

Exponential rate: $\alpha = \frac{1}{\sqrt{2}}$

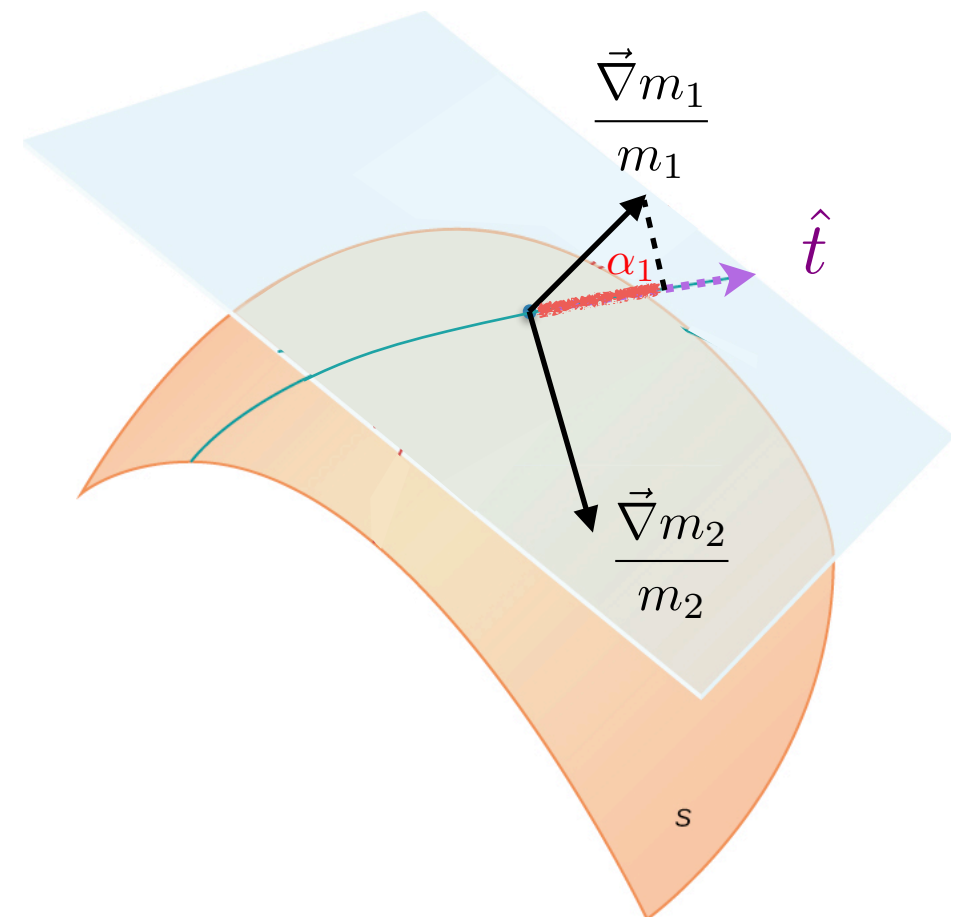
How is it compatible with

$$\left| \frac{\vec{\nabla} m}{m} \right| = \frac{1}{\sqrt{d-2}} = \frac{1}{\sqrt{3}} \quad ?$$

- Quantization of the string changes in AdS

- $\alpha = \frac{\vec{\nabla} m}{m} \cdot \hat{t}$ is the projection on the conformal manifold

but $m(R_{S^5}, g_s) \longleftrightarrow m(N, g_{YM})$



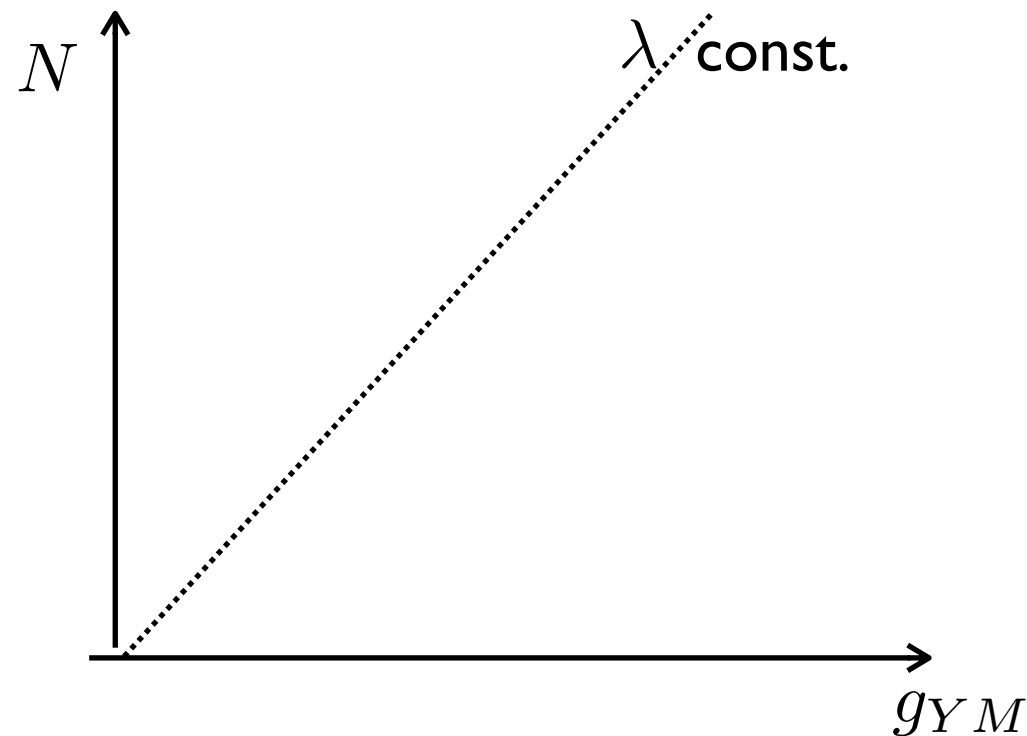
Convex Hull in N=4 SYM

Let us focus on N=4 SYM / $AdS_5 \times S^5$ [Calderon-Infante,IV' ongoing]

$$R = M_{p,5}^{-1} N^{2/3}$$

$$g_s = g_{YM}^2$$

$$\lambda = g_{YM}^2 N$$



Convex Hull in N=4 SYM

Let us focus on N=4 SYM / $AdS_5 \times S^5$ [Calderon-Infante,IV' ongoing]

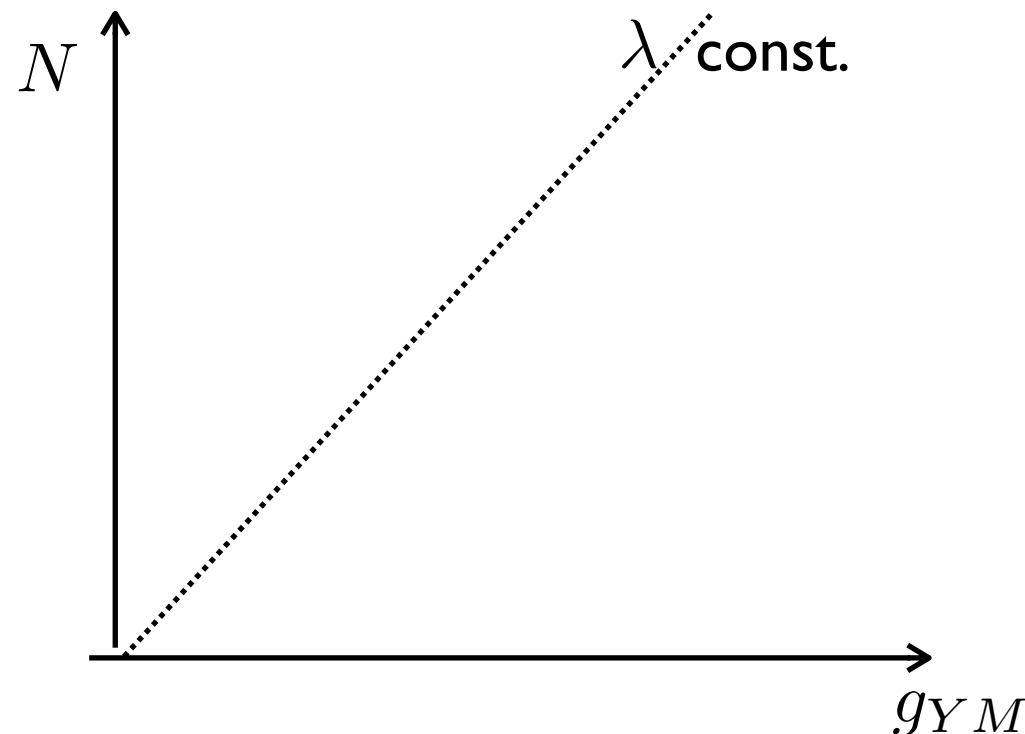
• Field theory pert. regime: $\lambda \ll 1$

• Supergravity regime: $\lambda \gg 1$

$$R = M_{p,5}^{-1} N^{2/3}$$

$$g_s = g_{YM}^2$$

$$\lambda = g_{YM}^2 N$$



Convex Hull in N=4 SYM

Let us focus on N=4 SYM / $AdS_5 \times S^5$ [Calderon-Infante, IV' ongoing]

- Field theory pert. regime: $\lambda \ll 1$

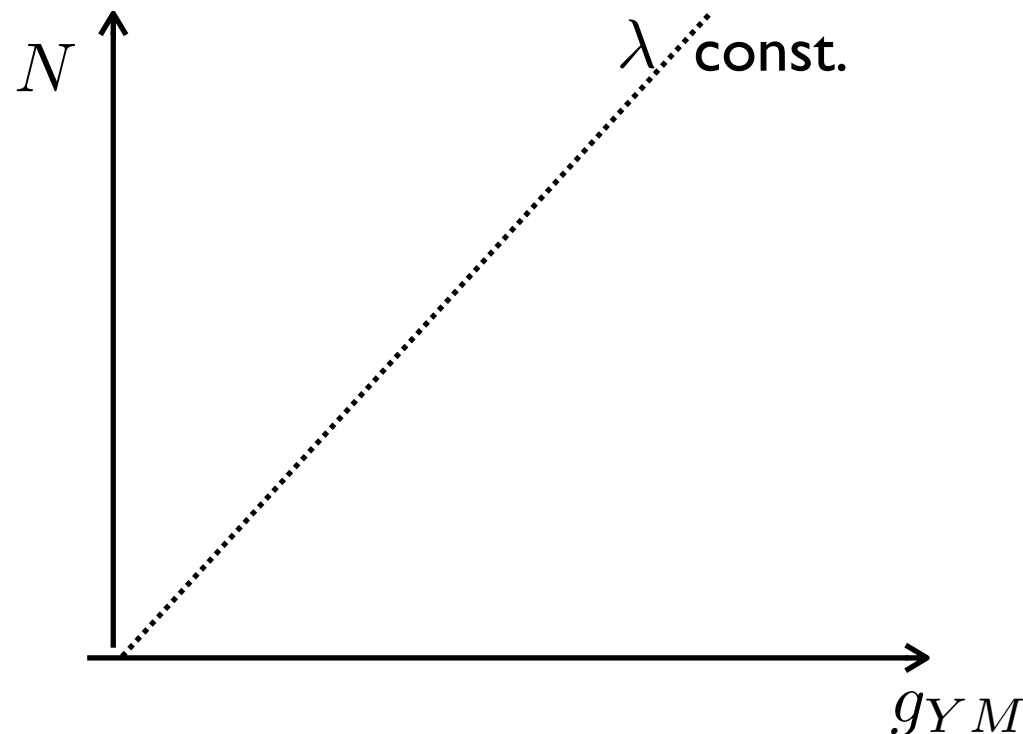
$$\begin{aligned} \gamma &\sim N g_{YM}^2 \\ m &\sim \frac{\sqrt{\gamma}}{N^{2/3}} \longrightarrow \frac{\vec{\nabla} m_{HS}}{m_{HS}} = \left(\frac{1}{\sqrt{2}}, \frac{1}{\sqrt{30}} \right) \\ & \left| \frac{\vec{\nabla} m_{HS}}{m_{HS}} \right| = \sqrt{\frac{8}{15}} \end{aligned}$$

- Supergravity regime: $\lambda \gg 1$

$$R = M_{p,5}^{-1} N^{2/3}$$

$$g_s = g_{YM}^2$$

$$\lambda = g_{YM}^2 N$$



Convex Hull in N=4 SYM

Let us focus on N=4 SYM / $AdS_5 \times S^5$ [Calderon-Infante, IV' ongoing]

- Field theory pert. regime: $\lambda \ll 1$

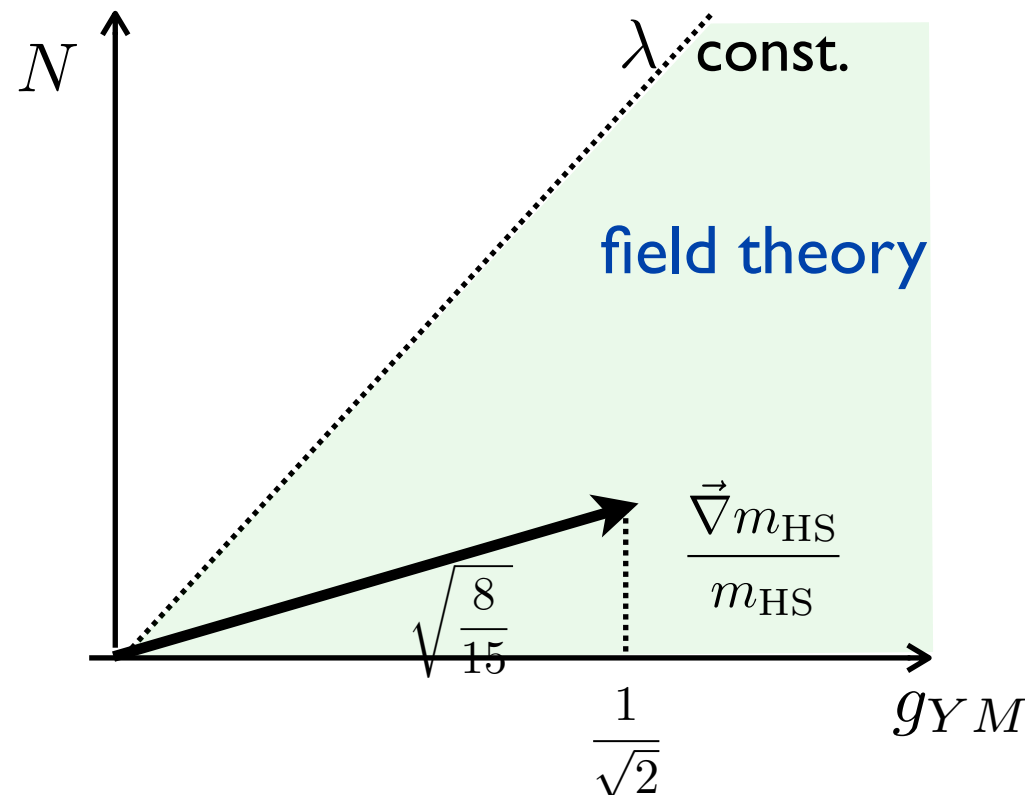
$$\begin{aligned} \gamma &\sim N g_{YM}^2 \\ m &\sim \frac{\sqrt{\gamma}}{N^{2/3}} \longrightarrow \frac{\vec{\nabla} m_{HS}}{m_{HS}} = \left(\frac{1}{\sqrt{2}}, \frac{1}{\sqrt{30}} \right) \\ & \left| \frac{\vec{\nabla} m_{HS}}{m_{HS}} \right| = \sqrt{\frac{8}{15}} \end{aligned}$$

- Supergravity regime: $\lambda \gg 1$

$$R = M_{p,5}^{-1} N^{2/3}$$

$$g_s = g_{YM}^2$$

$$\lambda = g_{YM}^2 N$$



Convex Hull in N=4 SYM

Let us focus on N=4 SYM / $AdS_5 \times S^5$ [Calderon-Infante, IV' ongoing]

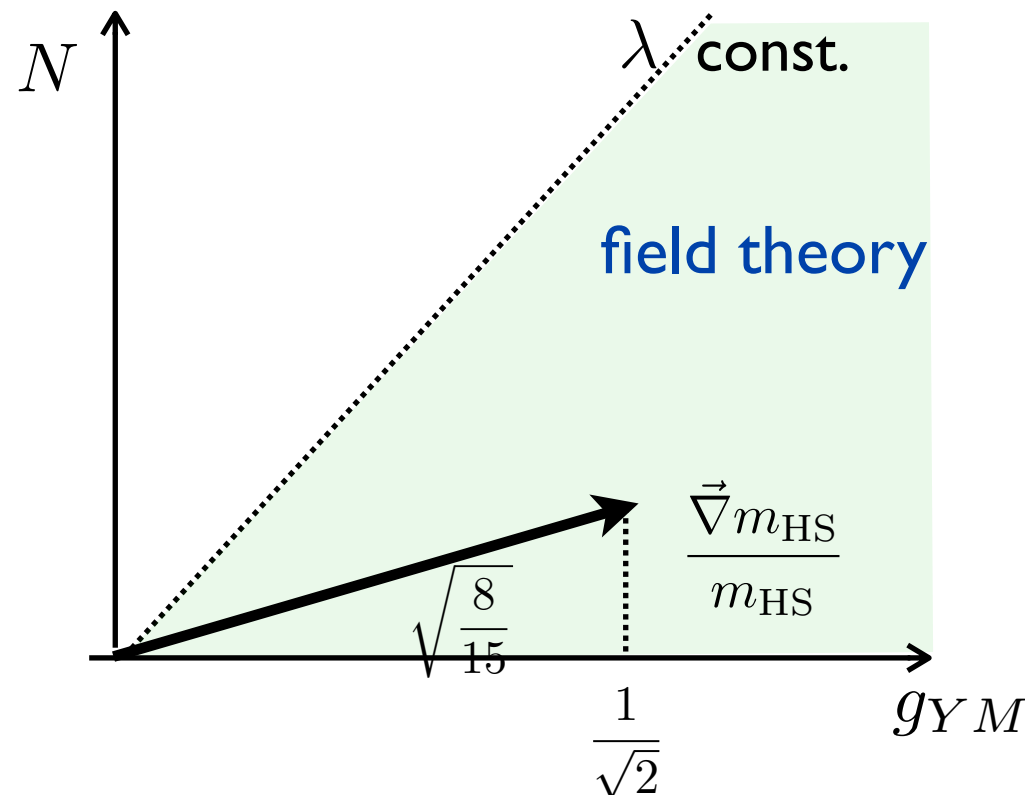
- Field theory pert. regime: $\lambda \ll 1$

$$\begin{aligned} \gamma &\sim N g_{YM}^2 \\ m &\sim \frac{\sqrt{\gamma}}{N^{2/3}} \longrightarrow \frac{\vec{\nabla} m_{HS}}{m_{HS}} = \left(\frac{1}{\sqrt{2}}, \frac{1}{\sqrt{30}} \right) \\ & \left| \frac{\vec{\nabla} m_{HS}}{m_{HS}} \right| = \sqrt{\frac{8}{15}} \end{aligned}$$

- Supergravity regime: $\lambda \gg 1$

$$\begin{aligned} T_s &= M_{p,5}^2 g_s^{1/2} R^{-5/4} \\ m &\sim \sqrt{T_s} \longrightarrow \frac{\vec{\nabla} m_{str}}{m_{str}} = \left(\frac{1}{2\sqrt{2}}, \frac{\sqrt{30}}{12} \right) \\ & \left| \frac{\vec{\nabla} m_{str}}{m_{str}} \right| = \frac{1}{\sqrt{3}} \end{aligned}$$

$$\begin{aligned} R &= M_{p,5}^{-1} N^{2/3} \\ g_s &= g_{YM}^2 \\ \lambda &= g_{YM}^2 N \end{aligned}$$



Convex Hull in N=4 SYM

Let us focus on N=4 SYM / $AdS_5 \times S^5$ [Calderon-Infante, IV' ongoing]

- Field theory pert. regime: $\lambda \ll 1$

$$\gamma \sim N g_{YM}^2$$

$$m \sim \frac{\sqrt{\gamma}}{N^{2/3}} \longrightarrow \frac{\vec{\nabla} m_{HS}}{m_{HS}} = \left(\frac{1}{\sqrt{2}}, \frac{1}{\sqrt{30}} \right)$$

$$\left| \frac{\vec{\nabla} m_{HS}}{m_{HS}} \right| = \sqrt{\frac{8}{15}}$$

- Supergravity regime: $\lambda \gg 1$

$$T_s = M_{p,5}^2 g_s^{1/2} R^{-5/4}$$

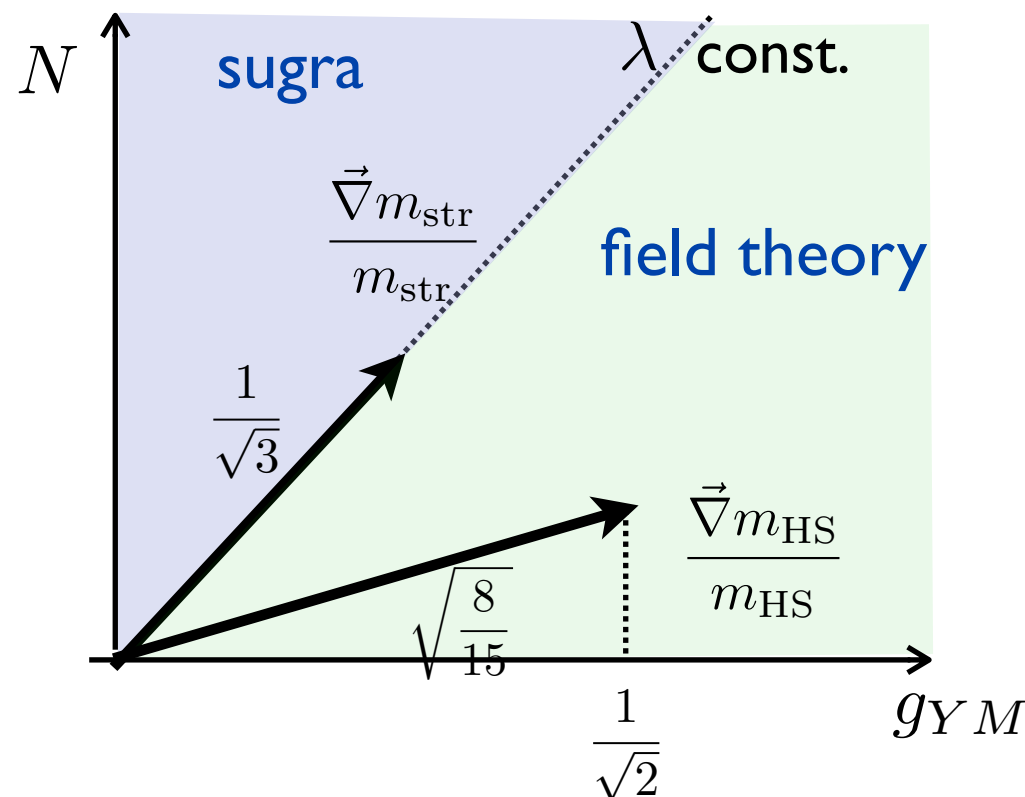
$$m \sim \sqrt{T_s} \longrightarrow \frac{\vec{\nabla} m_{str}}{m_{str}} = \left(\frac{1}{2\sqrt{2}}, \frac{\sqrt{30}}{12} \right)$$

$$\left| \frac{\vec{\nabla} m_{str}}{m_{str}} \right| = \frac{1}{\sqrt{3}}$$

$$R = M_{p,5}^{-1} N^{2/3}$$

$$g_s = g_{YM}^2$$

$$\lambda = g_{YM}^2 N$$



Convex Hull in N=4 SYM

Let us focus on N=4 SYM / $AdS_5 \times S^5$ [Calderon-Infante, IV' ongoing]

- Field theory pert. regime: $\lambda \ll 1$

$$\gamma \sim N g_{YM}^2$$

$$m \sim \frac{\sqrt{\gamma}}{N^{2/3}} \longrightarrow \frac{\vec{\nabla} m_{HS}}{m_{HS}} = \left(\frac{1}{\sqrt{2}}, \frac{1}{\sqrt{30}} \right)$$

$$\left| \frac{\vec{\nabla} m_{HS}}{m_{HS}} \right| = \sqrt{\frac{8}{15}}$$

- Supergravity regime: $\lambda \gg 1$

$$T_s = M_{p,5}^2 g_s^{1/2} R^{-5/4}$$

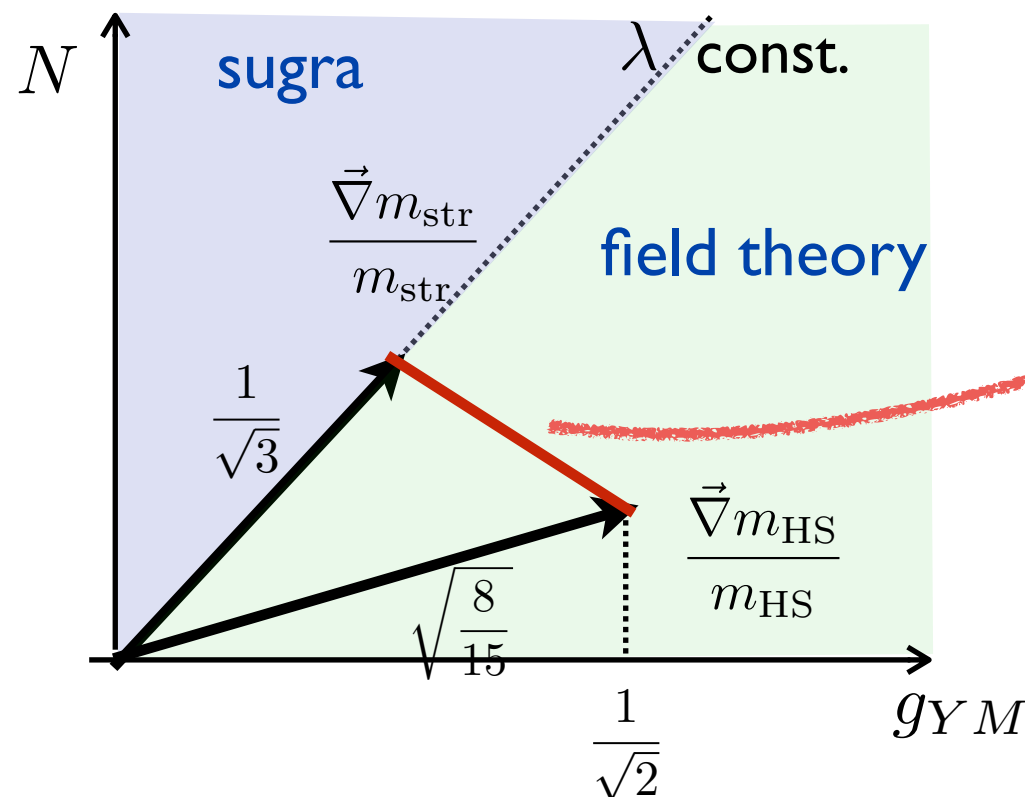
$$m \sim \sqrt{T_s} \longrightarrow \frac{\vec{\nabla} m_{str}}{m_{str}} = \left(\frac{1}{2\sqrt{2}}, \frac{\sqrt{30}}{12} \right)$$

$$\left| \frac{\vec{\nabla} m_{str}}{m_{str}} \right| = \frac{1}{\sqrt{3}}$$

$$R = M_{p,5}^{-1} N^{2/3}$$

$$g_s = g_{YM}^2$$

$$\lambda = g_{YM}^2 N$$



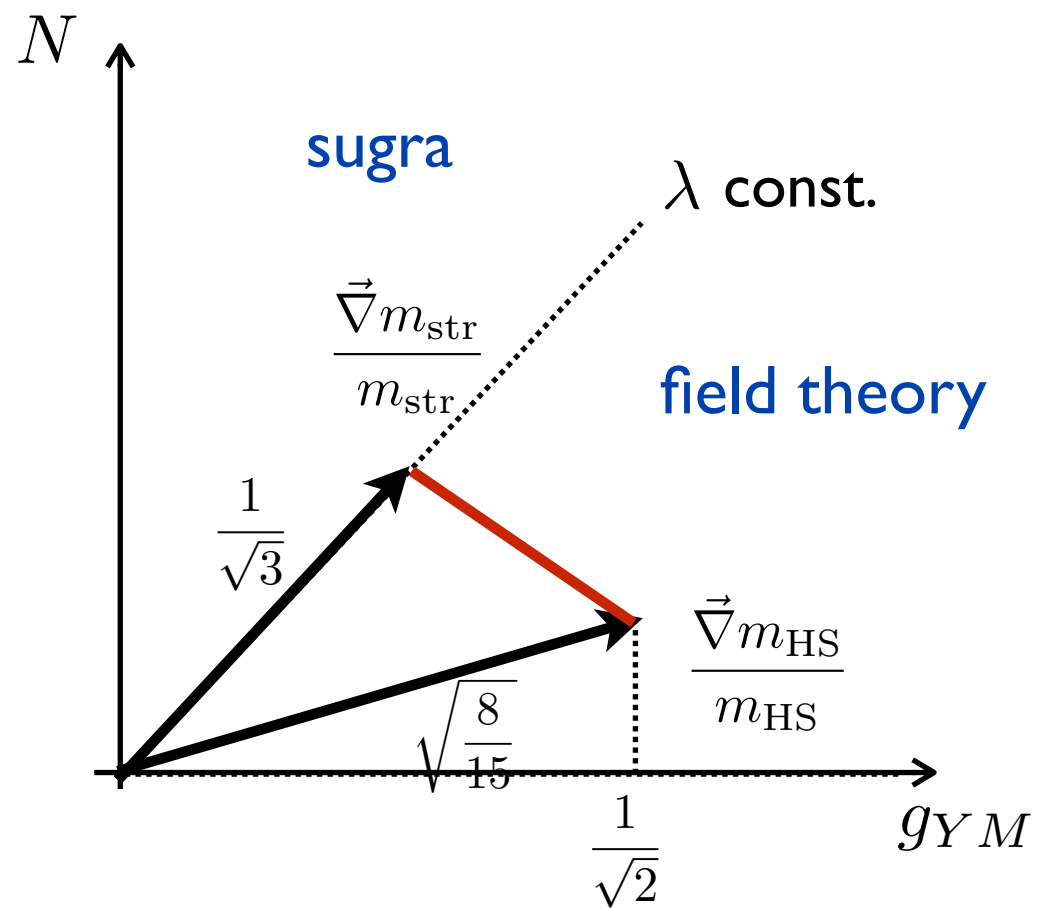
Results from integrability as λ changes

$$m \sim \frac{\sqrt{\gamma_{\text{cusp}}(\lambda)}}{N^{2/3}}$$

see e.g. [Dorigoni, Hatsuda'15]

Convex Hull and Scale Separation

Let us plot the convex hull of all the light towers, including the KK towers

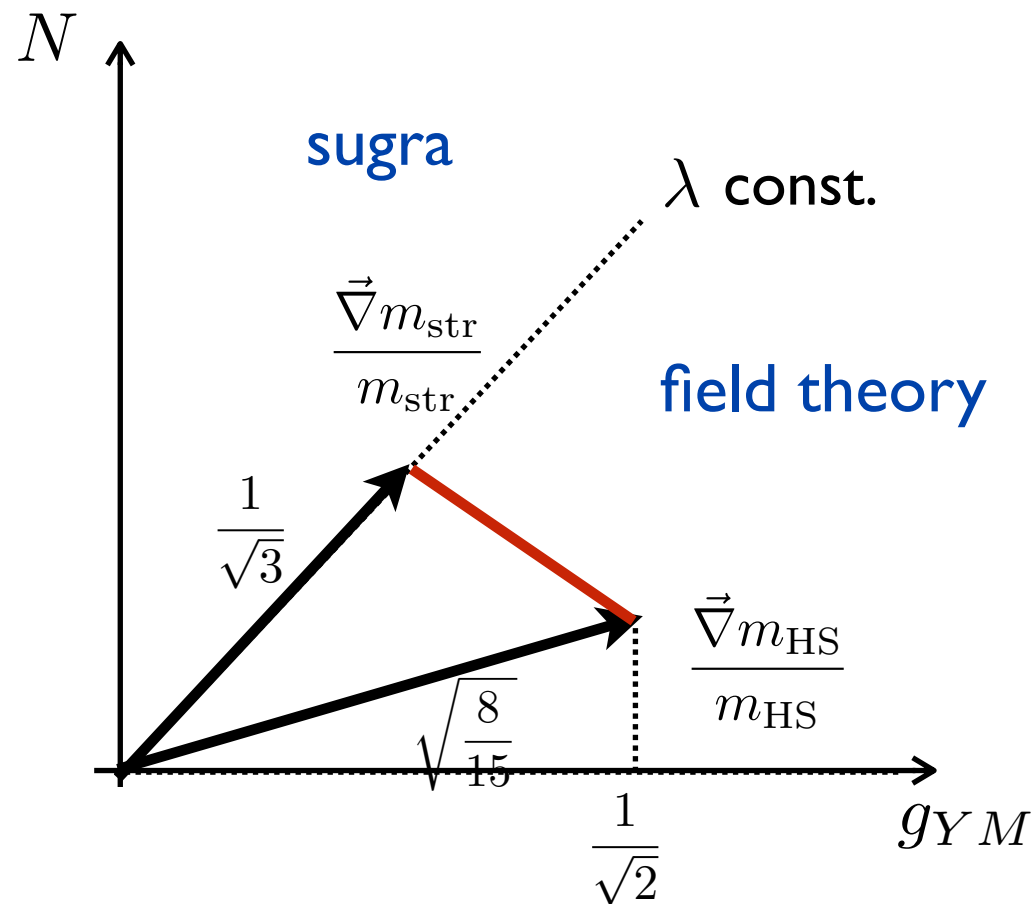


Convex Hull and Scale Separation

Let us plot the convex hull of all the light towers, including the KK towers

$$m_{KK} R_{AdS} \sim \mathcal{O}(1) \longrightarrow m_{KK} \sim N^{-2/3} \longrightarrow \frac{\vec{\nabla} m_{KK}}{m_{KK}} = \left(0, \frac{2\sqrt{30}}{15} \right)$$

$$R_{AdS} \sim R_{S^5}$$

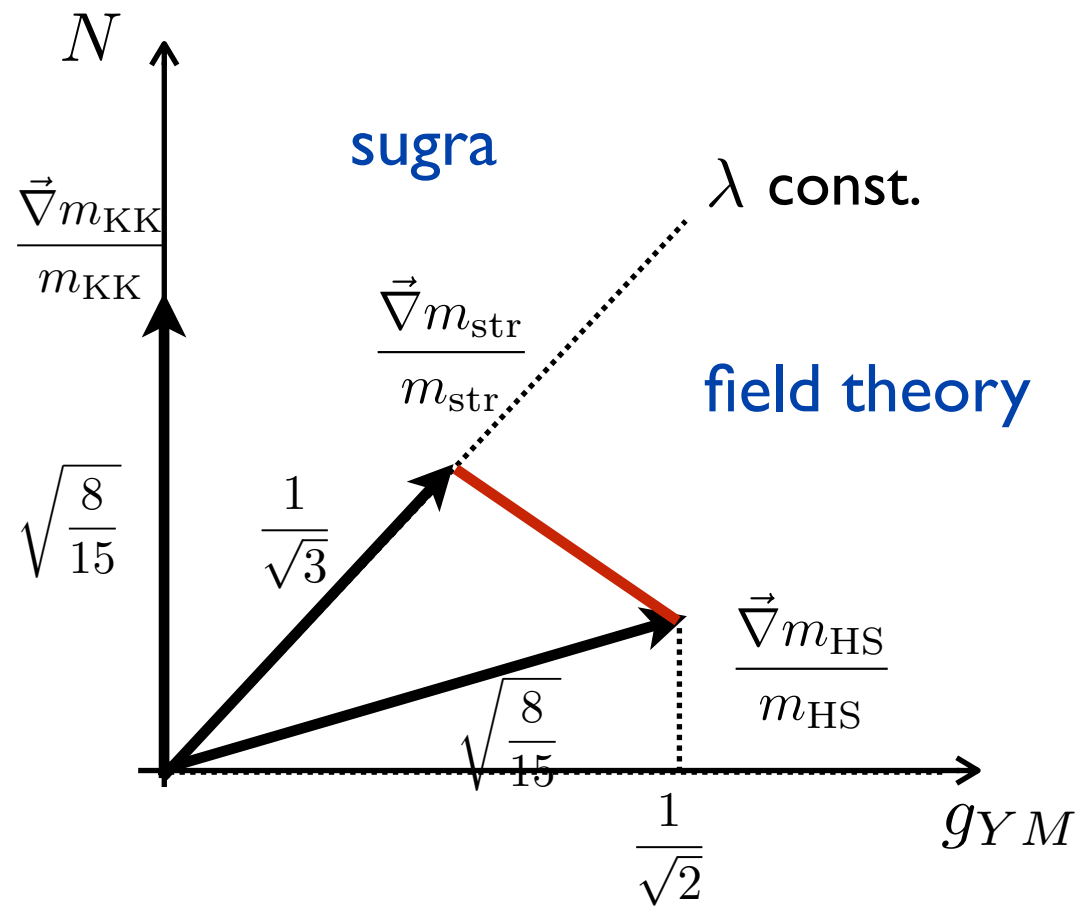


Convex Hull and Scale Separation

Let us plot the convex hull of all the light towers, including the KK towers

$$m_{KK} R_{AdS} \sim \mathcal{O}(1) \longrightarrow m_{KK} \sim N^{-2/3} \longrightarrow \frac{\vec{\nabla} m_{KK}}{m_{KK}} = \left(0, \frac{2\sqrt{30}}{15} \right)$$

$$R_{AdS} \sim R_{S^5}$$

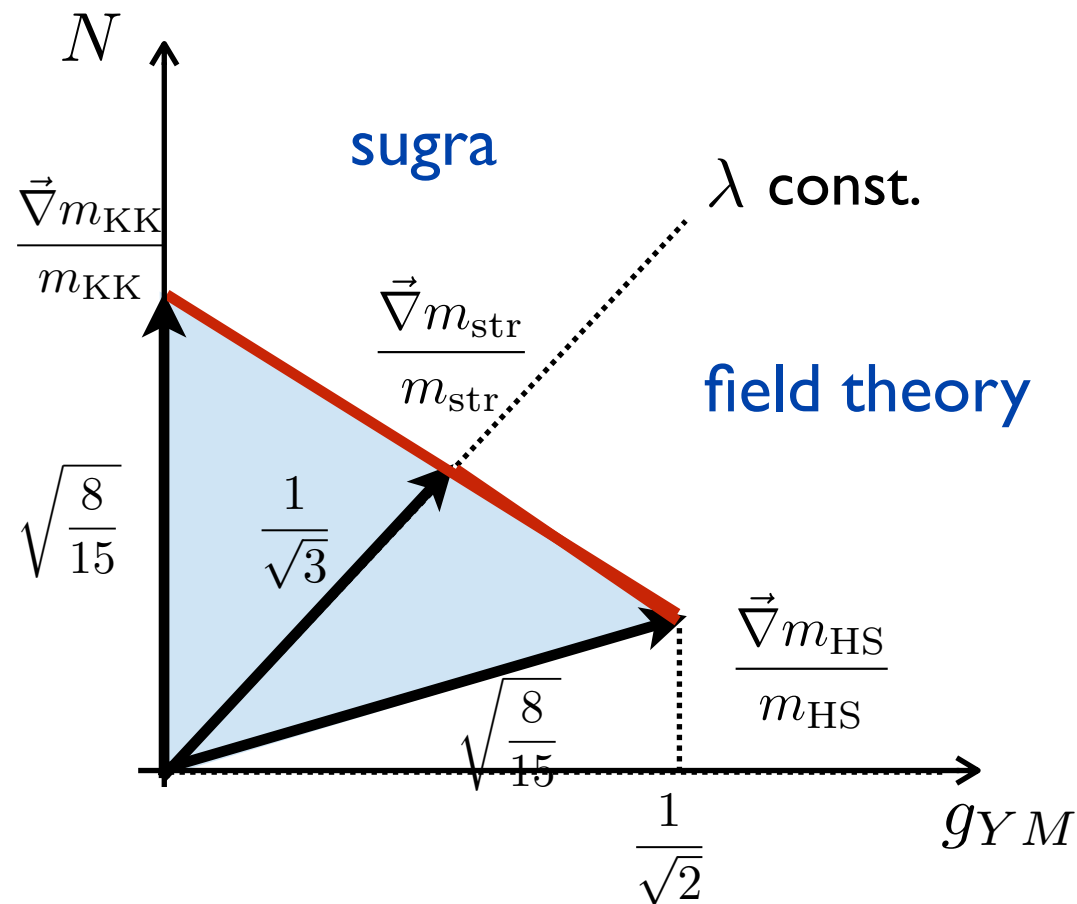


Convex Hull and Scale Separation

Let us plot the convex hull of all the light towers, including the KK towers

$$m_{KK} R_{AdS} \sim \mathcal{O}(1) \longrightarrow m_{KK} \sim N^{-2/3} \longrightarrow \frac{\vec{\nabla} m_{KK}}{m_{KK}} = \left(0, \frac{2\sqrt{30}}{15} \right)$$

$$R_{AdS} \sim R_{S^5}$$



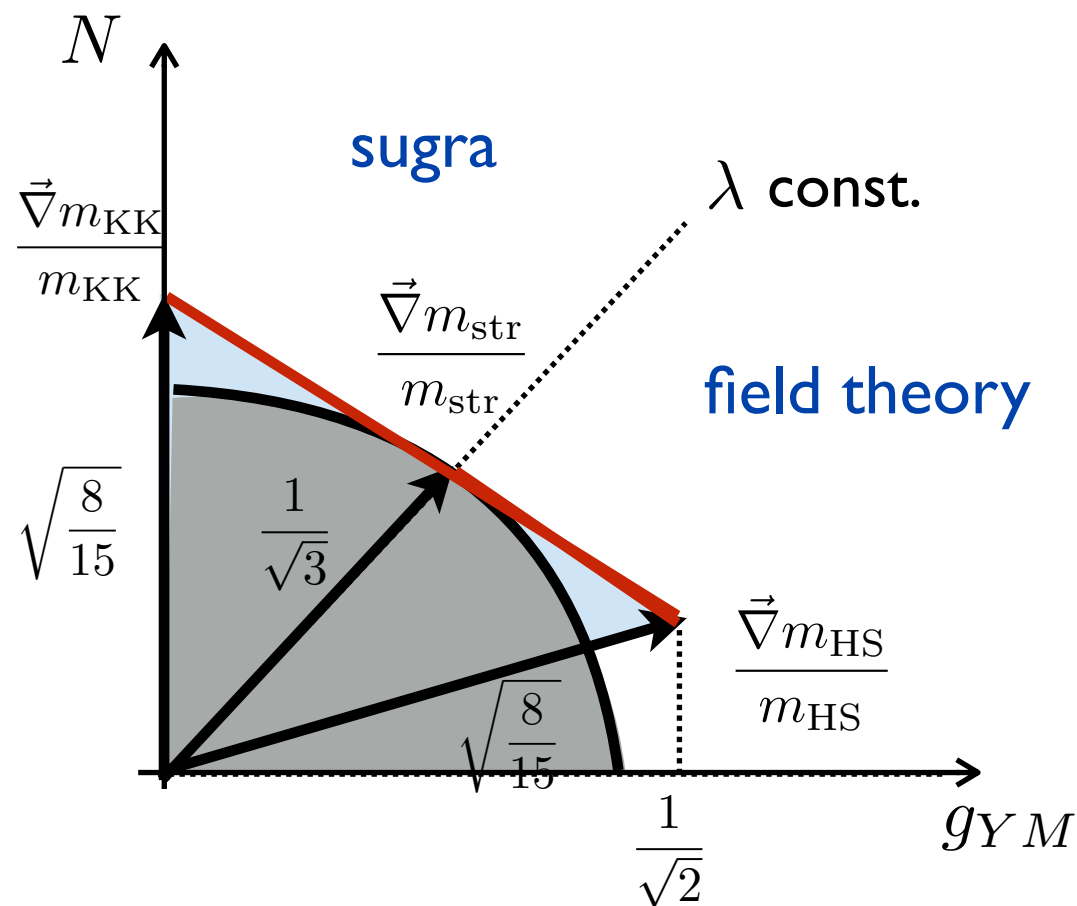
It satisfies taxonomy rules for the towers

Convex Hull and Scale Separation

Let us plot the convex hull of all the light towers, including the KK towers

$$m_{KK} R_{AdS} \sim \mathcal{O}(1) \longrightarrow m_{KK} \sim N^{-2/3} \longrightarrow \frac{\vec{\nabla} m_{KK}}{m_{KK}} = \left(0, \frac{2\sqrt{30}}{15} \right)$$

$$R_{AdS} \sim R_{S^5}$$



It satisfies taxonomy rules for the towers

It also satisfies the CH DC (it includes the ball of radius

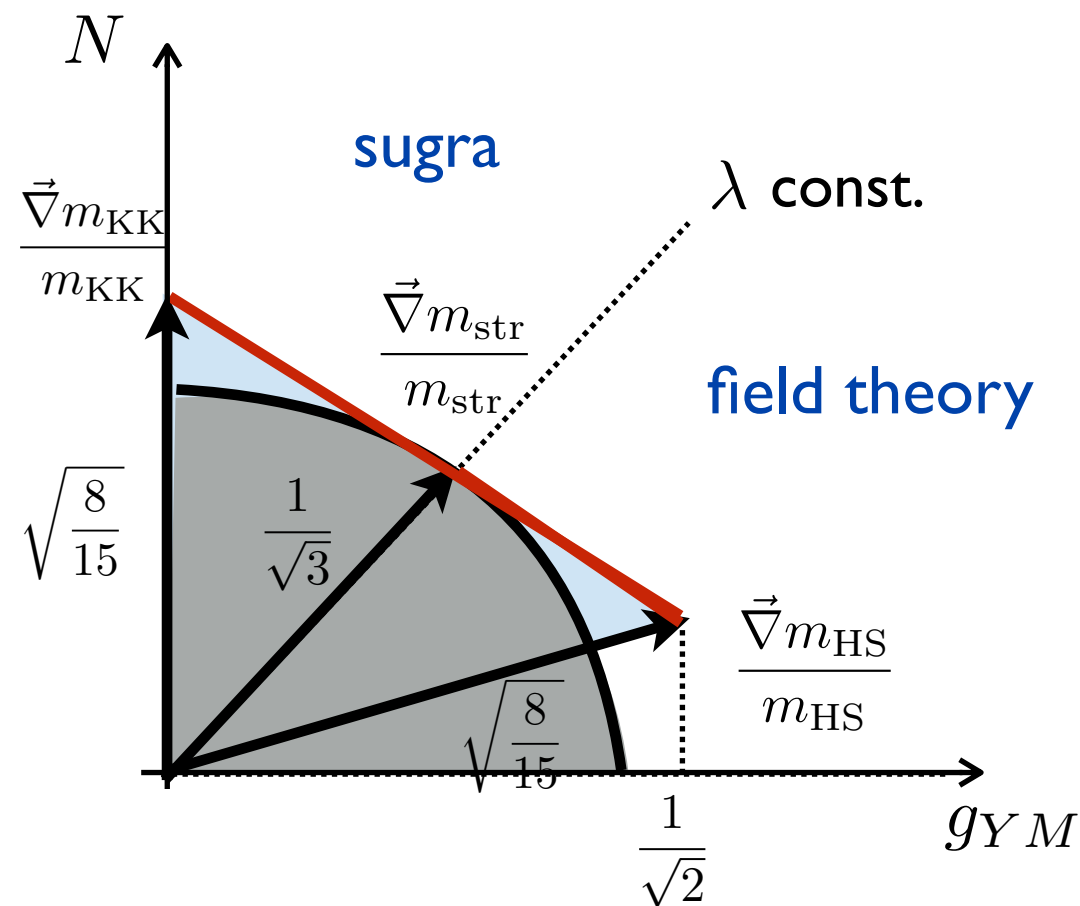
$$\alpha \geq \frac{1}{\sqrt{d-2}} = \frac{1}{\sqrt{3}})$$

Convex Hull and Scale Separation

Let us plot the convex hull of all the light towers, including the KK towers

$$m_{KK} R_{AdS} \sim \mathcal{O}(1) \longrightarrow m_{KK} \sim N^{-2/3} \longrightarrow \frac{\vec{\nabla} m_{KK}}{m_{KK}} = \left(0, \frac{2\sqrt{30}}{15} \right)$$

$$R_{AdS} \sim R_{S^5}$$



It satisfies taxonomy rules for the towers

It also satisfies the CH DC (it includes the ball of radius

$$\alpha \geq \frac{1}{\sqrt{d-2}} = \frac{1}{\sqrt{3}})$$

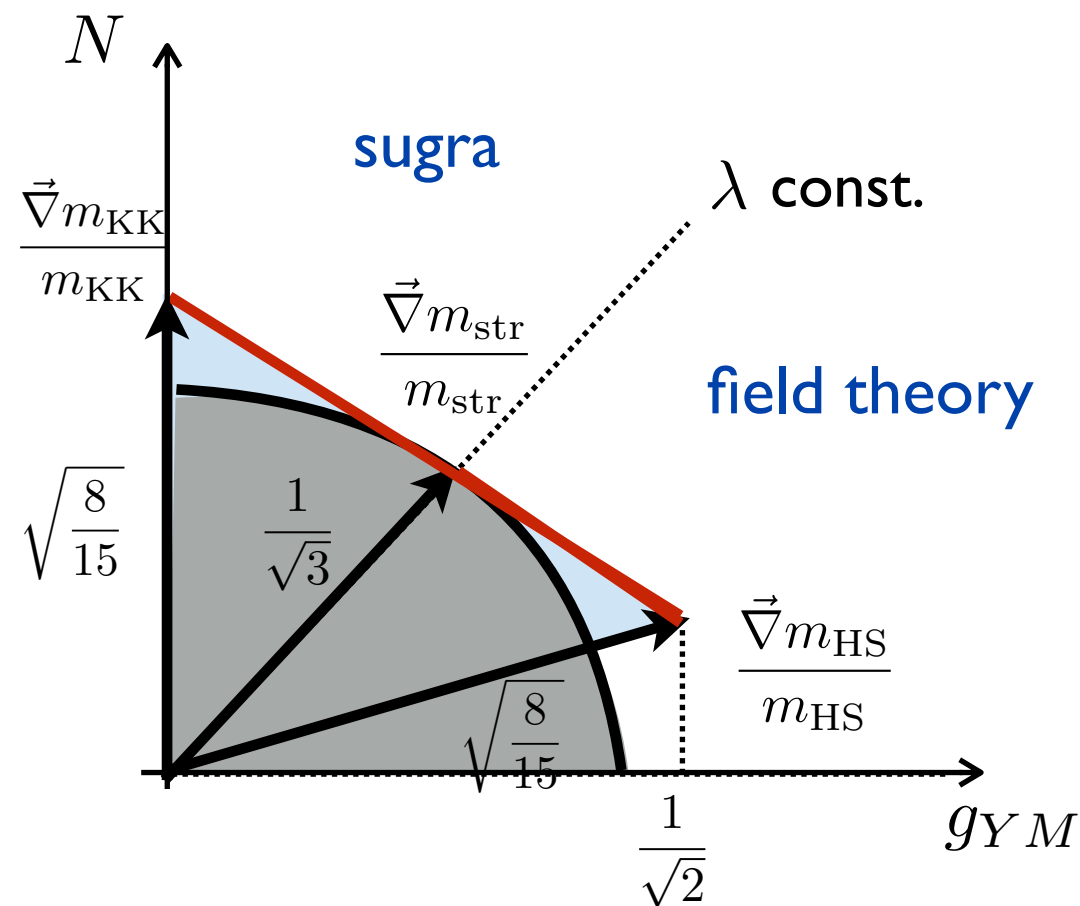
The sharpened bound for the Distance conjecture $\alpha \geq \frac{1}{\sqrt{d-2}}$ is satisfied for every infinite distance direction

Convex Hull and Scale Separation

Let us plot the convex hull of all the light towers, including the KK towers

$$m_{KK} R_{AdS} \sim \mathcal{O}(1) \longrightarrow m_{KK} \sim N^{-2/3} \longrightarrow \frac{\vec{\nabla} m_{KK}}{m_{KK}} = \left(0, \frac{2\sqrt{30}}{15} \right)$$

$$R_{AdS} \sim R_{S^5}$$



It satisfies taxonomy rules for the towers

It also satisfies the CH DC (it includes the ball of radius

$$\alpha \geq \frac{1}{\sqrt{d-2}} = \frac{1}{\sqrt{3}})$$

The sharpened bound for the Distance conjecture $\alpha \geq \frac{1}{\sqrt{d-2}}$ is satisfied for every infinite distance direction

It would be violated if the AdS vacuum was scale-separated $R_{KK} \ll R_{AdS}$

Convex Hull and Scale Separation

The Convex Hull Distance Conjecture with $\alpha \geq \frac{1}{\sqrt{d-2}}$
implies that **holographic AdS spaces with a conformal manifold**
(namely, 5d AdS spaces with 8 or more supercharges)
cannot be scale separated

Convex Hull and Scale Separation

The Convex Hull Distance Conjecture with $\alpha \geq \frac{1}{\sqrt{d-2}}$
implies that **holographic AdS spaces with a conformal manifold**
(namely, 5d AdS spaces with 8 or more supercharges)
cannot be scale separated

Here we are using the moduli space distance

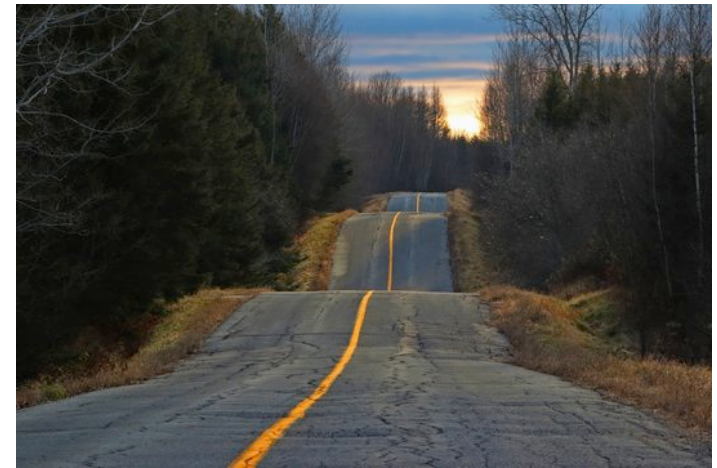
Convex Hull and Scale Separation

The Convex Hull Distance Conjecture with $\alpha \geq \frac{1}{\sqrt{d-2}}$
implies that **holographic AdS spaces with a conformal manifold**
(namely, 5d AdS spaces with 8 or more supercharges)
cannot be scale separated

Here we are using the moduli space distance

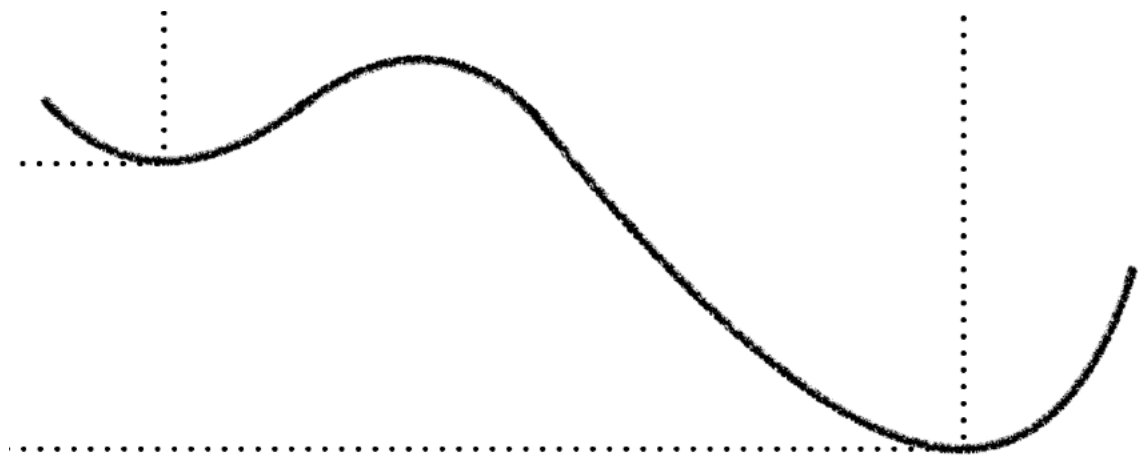
but can we generalize the notion of distance including a scalar potential?

3) Generalized distance with a potential



[Mohseni,Montero,Vafa,IV'ongoing]

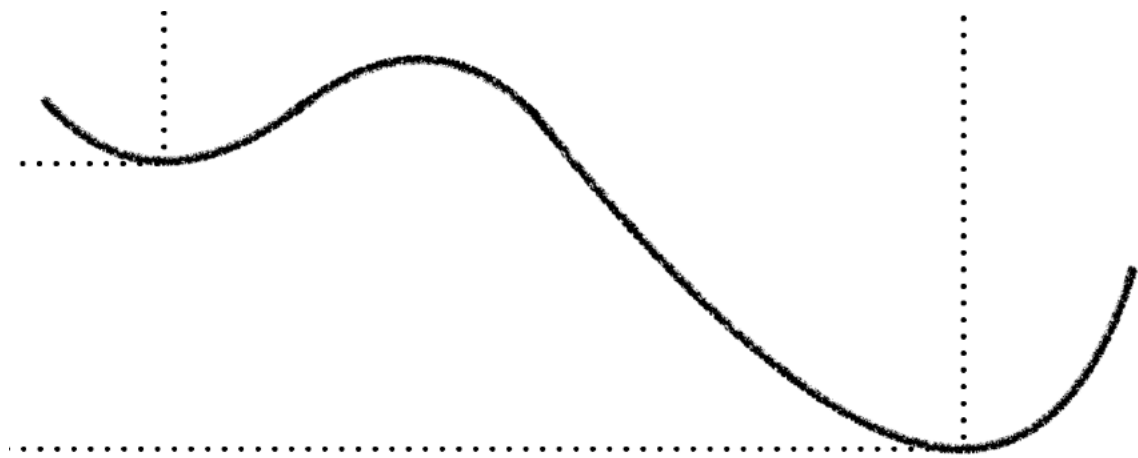
Distance between theories



Very challenging, many attempts...

[Douglas'13][Bachas et al'13][Luest,Palti,Vafa'19]
[De Biasio,Luest et al,20-22][Stout'22]
[Basile,Montella'23][Li,Palti,Petri'23-24]
[Shiu,Tonioni, Van Hemelryck, Van Riet'23-24]
[Tonioni, Van Riet'ongoing]...

Distance between theories

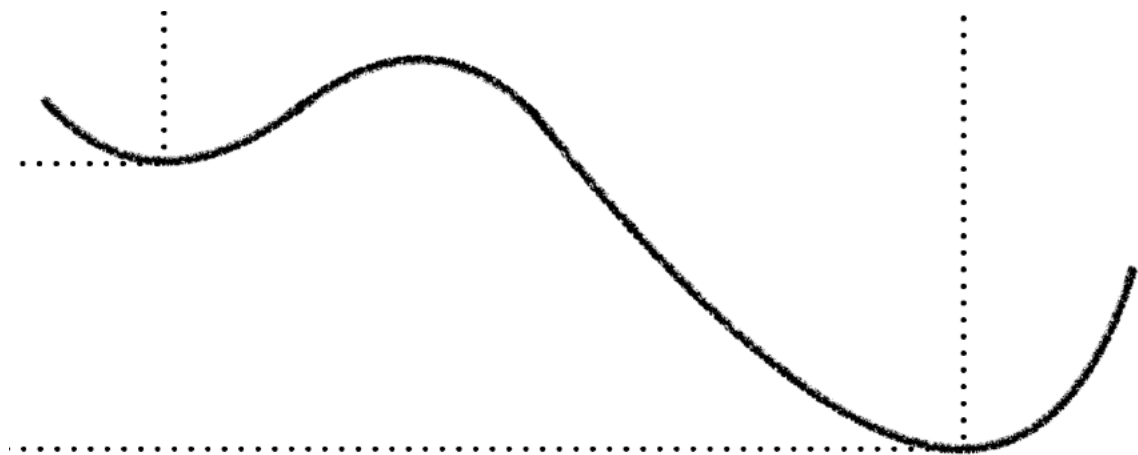


Very challenging, many attempts...

[Douglas'13][Bachas et al'13][Luest,Palti,Vafa'19]
[De Biasio,Luest et al,20-22][Stout'22]
[Basile,Montella'23][Li,Palti,Petri'23-24]
[Shiu,Tonioni, Van Hemelryck, Van Riet'23-24]
[Tonioni, Van Riet'ongoing]...

- Wish list:**
- A notion of metric and geodesics in the space of theories
 - It recovers moduli space distance if $V=0$

Distance between theories



Very challenging, many attempts...

[Douglas'13][Bachas et al'13][Luest,Palti,Vafa'19]
[De Biasio,Luest et al,20-22][Stout'22]
[Basile,Montella'23][Li,Palti,Petri'23-24]
[Shiu,Tonioni, Van Hemelryck, Van Riet'23-24]
[Tonioni, Van Riet'ongoing]...

- Wish list:**
- A notion of metric and geodesics in the space of theories
 - It recovers moduli space distance if $V=0$

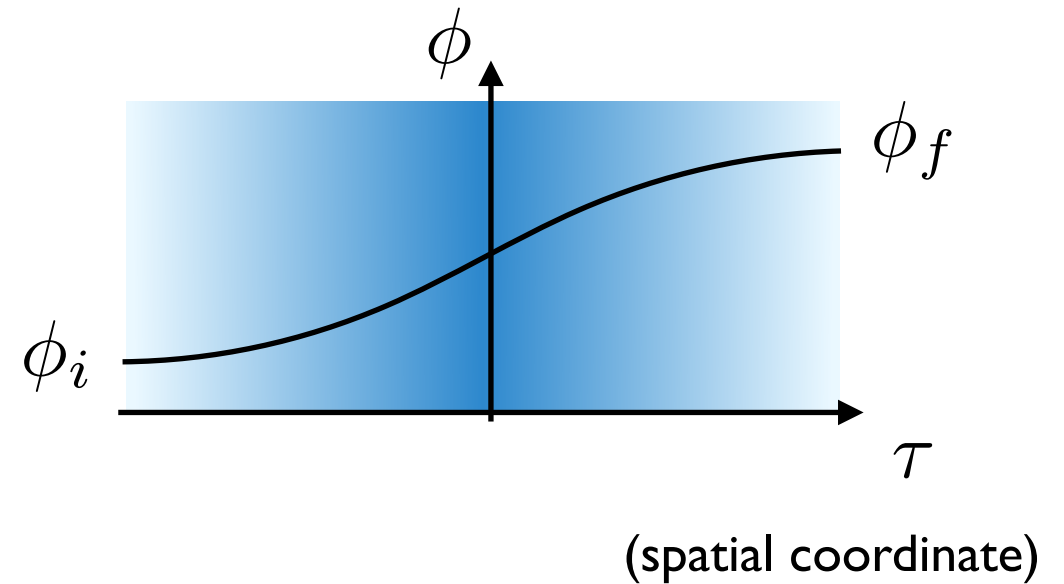
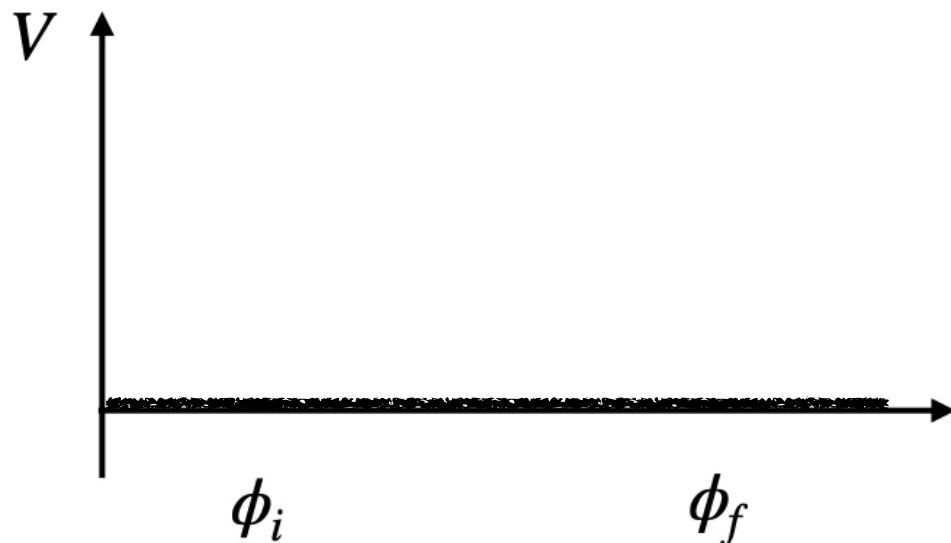
Let us take inspiration on the **cobordism conjecture**: [McNamara,Vafa'19]

Any two d -dim theories are connected by a finite energy domain wall

[Mohseni,Montero,Vafa,IV'ongoing]

Distance between theories

Let us start with $V=0$ (a moduli space)



$$T_{DW} = \int (\partial_\tau \phi)^2 d\tau = \frac{(\Delta\phi)^2}{\Delta\tau} \rightarrow 0$$

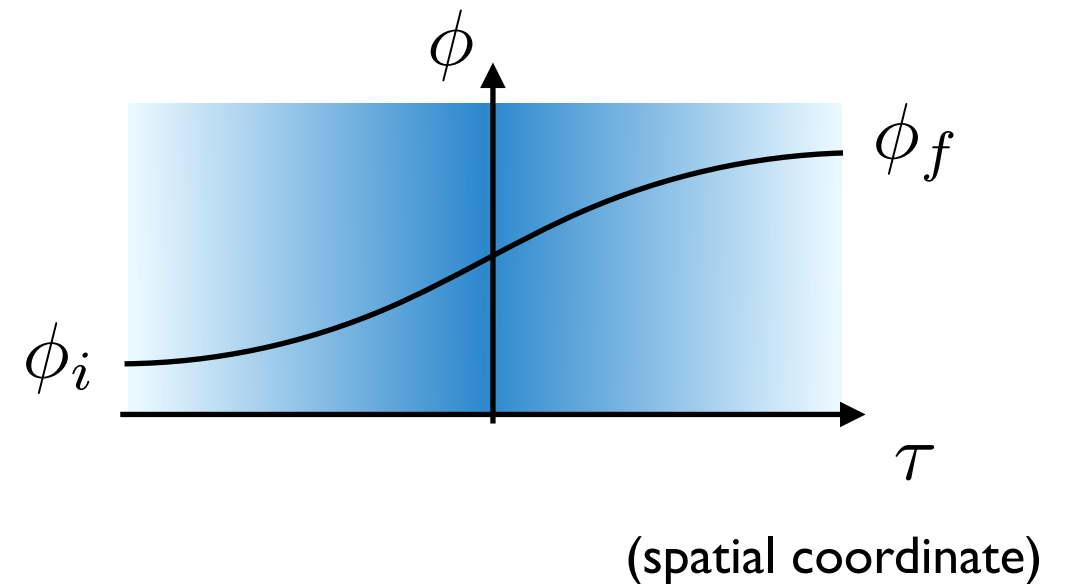
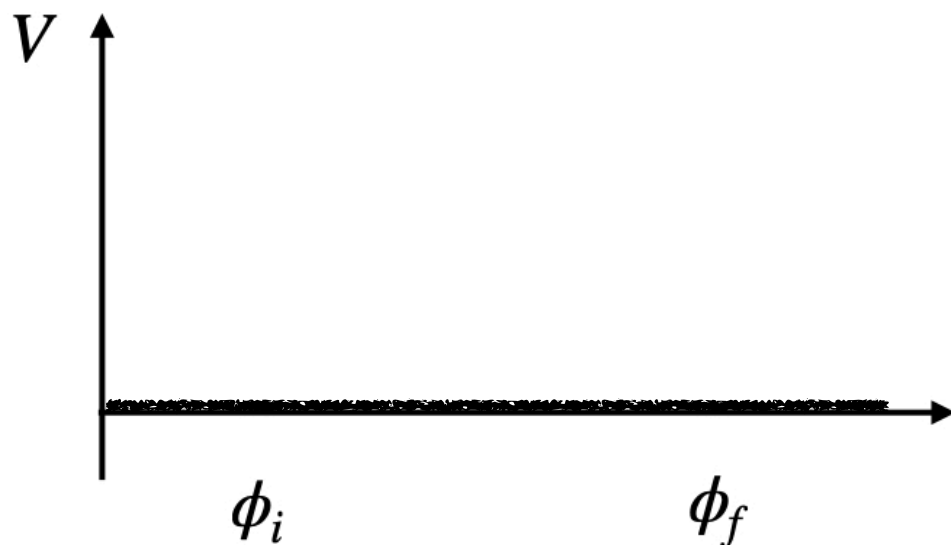
minimized for $\phi(\tau) = \phi_0 + \sqrt{2\rho_E\tau}$

since minimal tension occurs for

$$\Delta\tau \rightarrow \infty \rightarrow \rho_E \rightarrow 0$$

Distance between theories

Let us start with $V=0$ (a moduli space)



$$T_{DW} = \int (\partial_\tau \phi)^2 d\tau = \frac{(\Delta\phi)^2}{\Delta\tau} \rightarrow 0$$

minimized for $\phi(\tau) = \phi_0 + \sqrt{2\rho_E\tau}$

since minimal tension occurs for

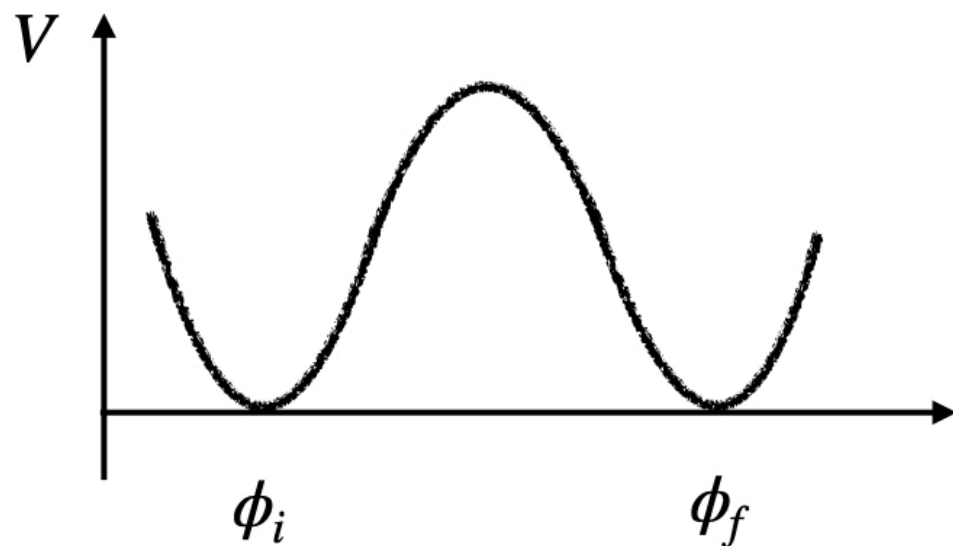
$$\Delta\tau \rightarrow \infty \rightarrow \rho_E \rightarrow 0$$

Define $\Delta \equiv \frac{T_{DW}}{\sqrt{2\rho_E}} = \Delta\phi$

recovers moduli space distance!

Distance between theories

Let us add a potential between two Minkowski vacua (ignoring gravitational effects):

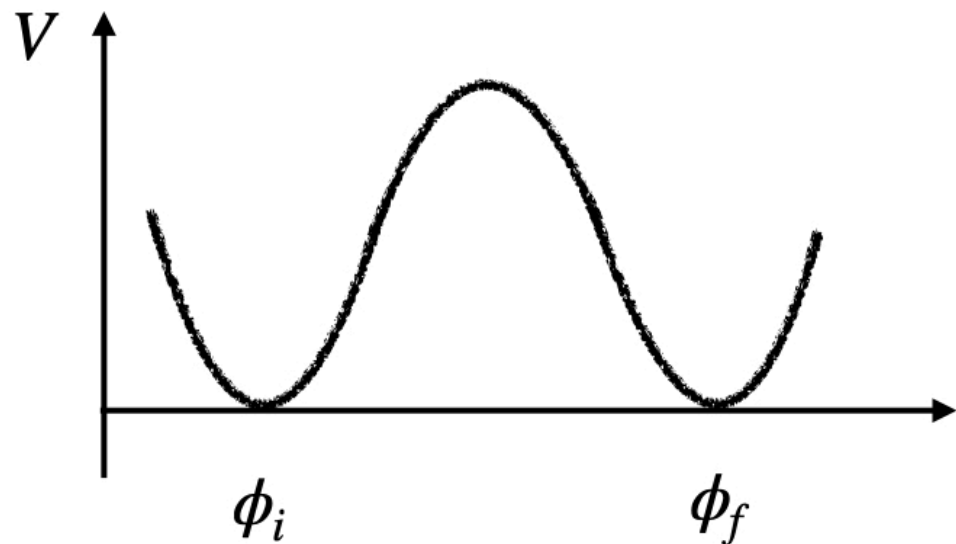


$$\Delta = \frac{T(\rho_E)}{\sqrt{2\rho_E}} = \frac{1}{\sqrt{2\rho_E}} \int_{\phi_i}^{\phi_f} \sqrt{2(V + \rho_E)} d\phi$$

$T(\rho_E)$ is the minimal euclidean action
for a path of energy ρ_E

Distance between theories

Let us add a potential between two Minkowski vacua (ignoring gravitational effects):



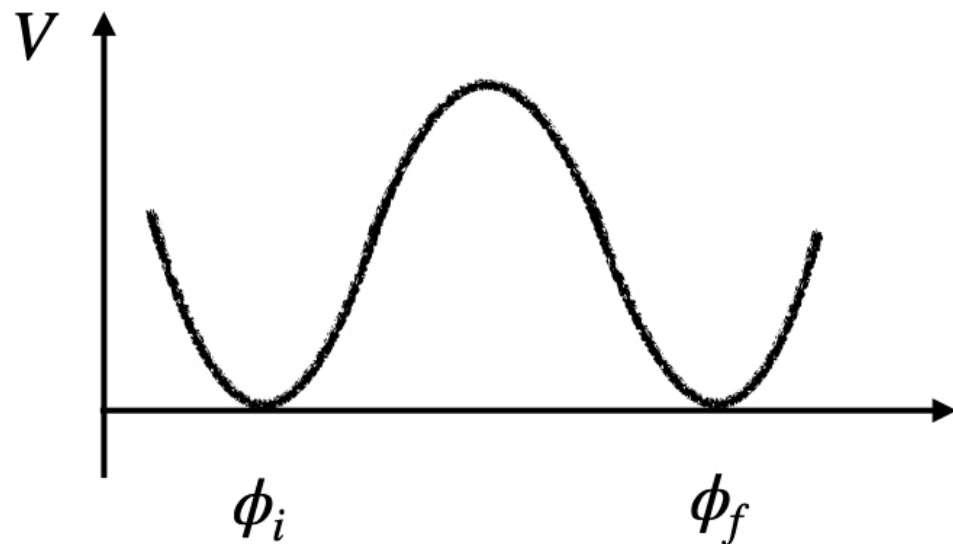
$$\Delta = \frac{T(\rho_E)}{\sqrt{2\rho_E}} = \frac{1}{\sqrt{2\rho_E}} \int_{\phi_i}^{\phi_f} \sqrt{2(V + \rho_E)} d\phi$$

$T(\rho_E)$ is the minimal euclidean action
for a path of energy ρ_E

- This is known as **Mapertuis action principle** $S_M = \int d\tau \left(\frac{1}{2} \dot{\phi}^2 + V \right) + \int \rho_E d\tau$

Distance between theories

Let us add a potential between two Minkowski vacua (ignoring gravitational effects):



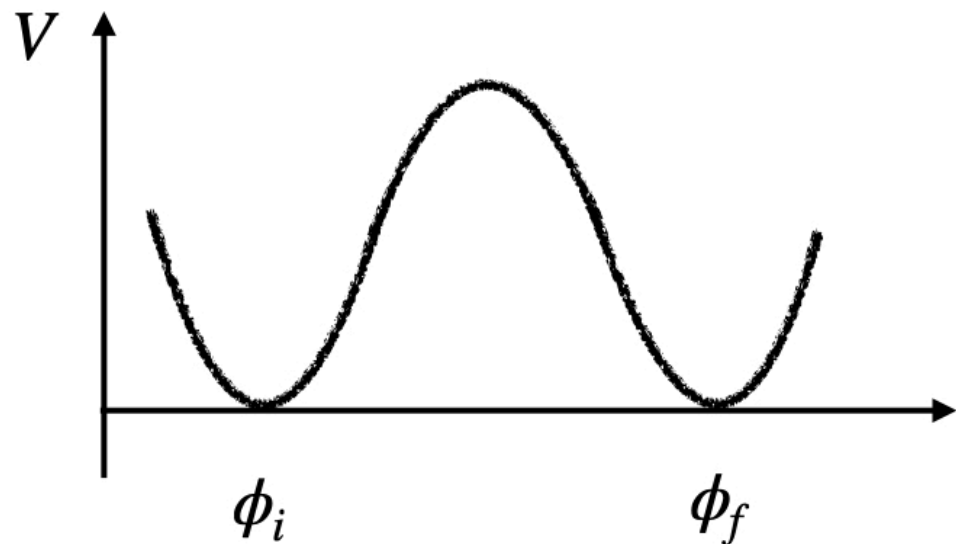
$$\Delta = \frac{T(\rho_E)}{\sqrt{2\rho_E}} = \frac{1}{\sqrt{2\rho_E}} \int_{\phi_i}^{\phi_f} \sqrt{2(V + \rho_E)} d\phi$$

$T(\rho_E)$ is the minimal euclidean action
for a path of energy ρ_E

- This is known as **Mapertuis action principle** $S_M = \int d\tau \left(\frac{1}{2} \dot{\phi}^2 + V \right) + \int \rho_E d\tau$
- It provides a well-behaved notion for a **metric**: $G_{ij} \equiv g_{ij} \left(1 + \frac{V}{\rho_E} \right)$ (Jacobi metric)
(the distance is positive definite, symmetric and satisfies triangular inequality)

Distance between theories

Let us add a potential between two Minkowski vacua (ignoring gravitational effects):



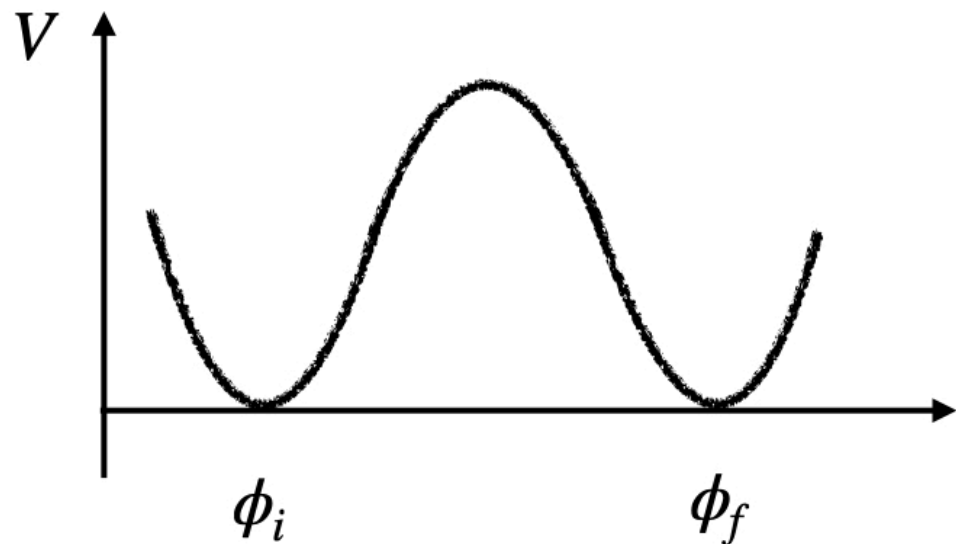
$$\Delta = \frac{T(\rho_E)}{\sqrt{2\rho_E}} = \frac{1}{\sqrt{2\rho_E}} \int_{\phi_i}^{\phi_f} \sqrt{2(V + \rho_E)} d\phi$$

$T(\rho_E)$ is the minimal euclidean action
for a path of energy ρ_E

- This is known as **Mapertuis action principle** $S_M = \int d\tau \left(\frac{1}{2} \dot{\phi}^2 + V \right) + \int \rho_E d\tau$
- It provides a well-behaved notion for a **metric**: $G_{ij} \equiv g_{ij} \left(1 + \frac{V}{\rho_E} \right)$ (Jacobi metric)
(the distance is positive definite, symmetric and satisfies triangular inequality)
- It **recovers moduli space** distance if $V=0$

Distance between theories

Let us add a potential between two Minkowski vacua (ignoring gravitational effects):



$$\Delta = \frac{T(\rho_E)}{\sqrt{2\rho_E}} = \frac{1}{\sqrt{2\rho_E}} \int_{\phi_i}^{\phi_f} \sqrt{2(V + \rho_E)} d\phi$$

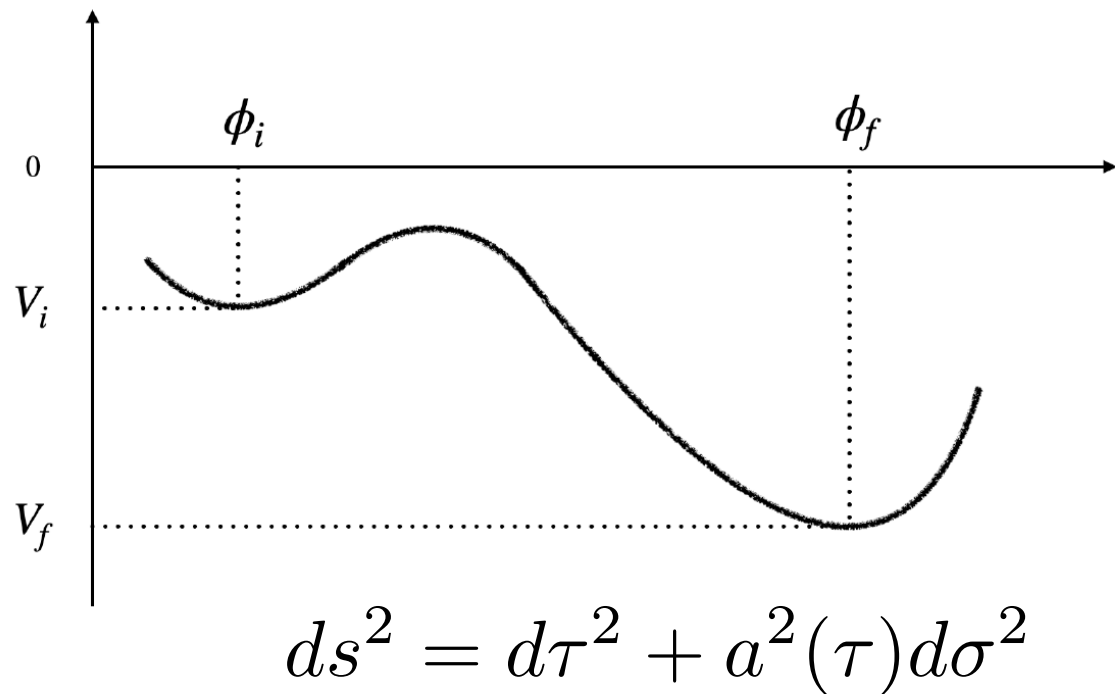
$T(\rho_E)$ is the minimal euclidean action for a path of energy ρ_E

- This is known as **Mapertuis action principle** $S_M = \int d\tau \left(\frac{1}{2} \dot{\phi}^2 + V \right) + \int \rho_E d\tau$
- It provides a well-behaved notion for a **metric**: $G_{ij} \equiv g_{ij} \left(1 + \frac{V}{\rho_E} \right)$ (Jacobi metric)
(the distance is positive definite, symmetric and satisfies triangular inequality)
- It **recovers moduli space** distance if $V=0$
- It **depends on the euclidean energy scale** ρ_E such that:

{	If	$\rho_E \rightarrow \infty$:	$\Delta \rightarrow \Delta\phi$
	If	$\rho_E \rightarrow 0$:	$\Delta \rightarrow \frac{T_{DW}}{\sqrt{2\rho_E}} \rightarrow \infty$

Distance between theories

Let us include gravitational effects and consider e.g. two AdS vacua:

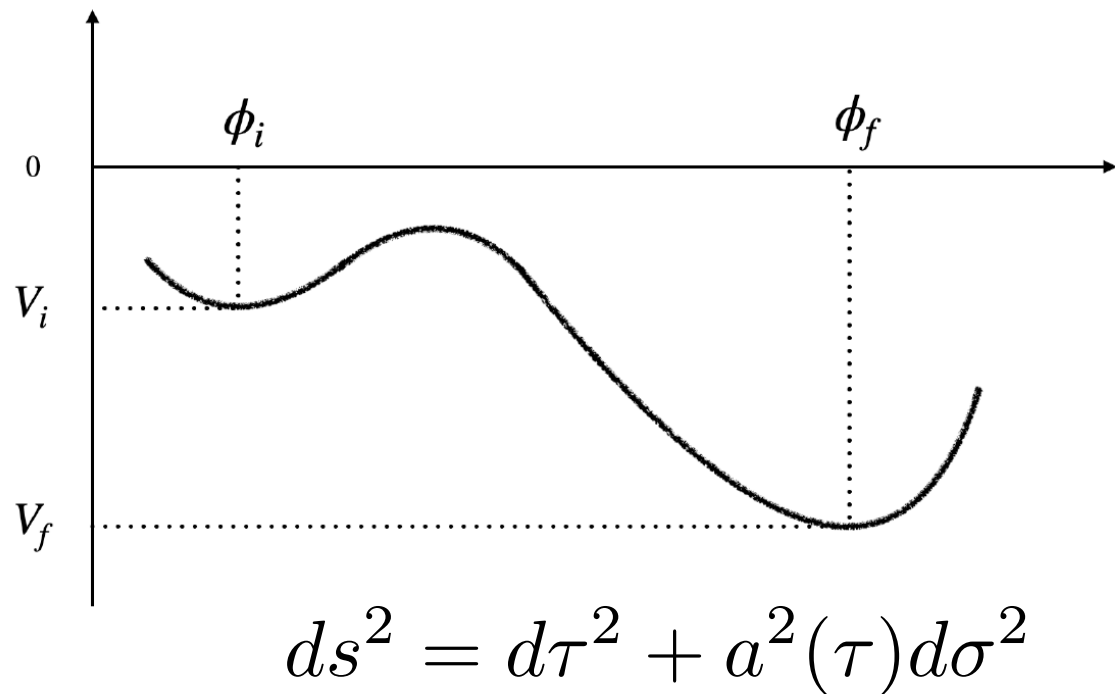


$$\Delta = \int_{\phi_i}^{\phi_f} \sqrt{1 + \frac{V(\phi)}{\rho_E}} d\phi$$

$$\begin{aligned} \frac{1}{2}\dot{\phi}^2 - V &= \rho_E(\phi) \neq \text{const.} \\ &= -\Lambda(\phi) \propto \frac{\dot{a}^2}{a^2} \end{aligned}$$

Distance between theories

Let us include gravitational effects and consider e.g. two AdS vacua:

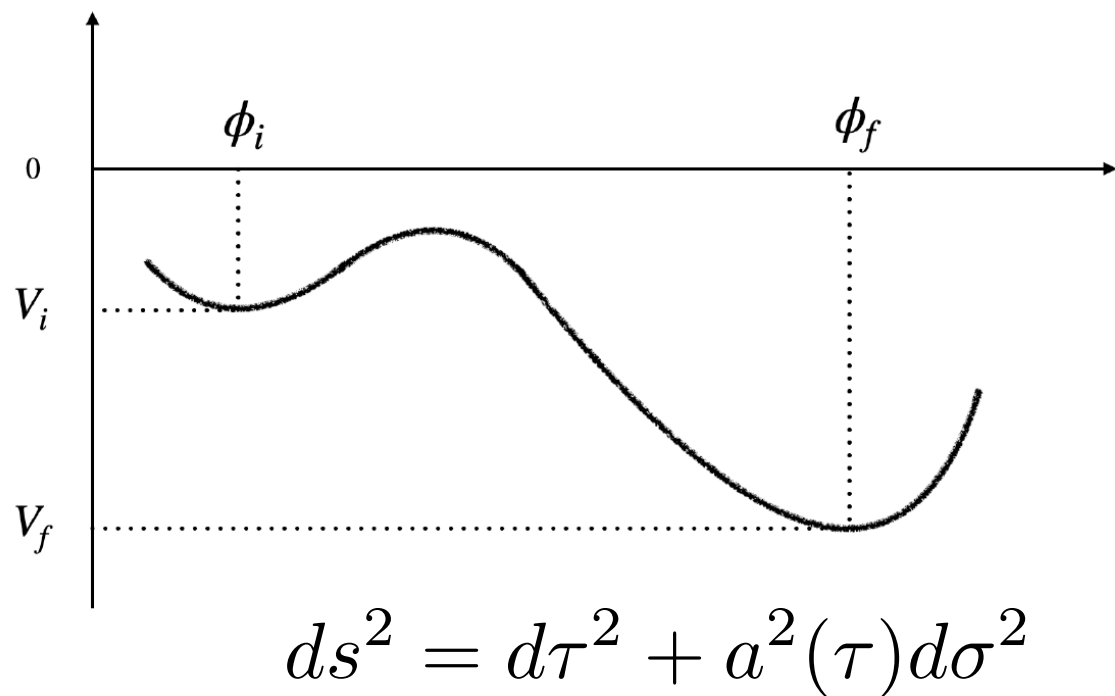


$$\Delta = \int_{\phi_i}^{\phi_f} \sqrt{1 - \frac{V(\phi)}{\Lambda(\phi)}} d\phi$$

$$\begin{aligned} \frac{1}{2}\dot{\phi}^2 - V &= \rho_E(\phi) \neq \text{const.} \\ &= -\Lambda(\phi) \propto \frac{\dot{a}^2}{a^2} \end{aligned}$$

Distance between theories

Let us include gravitational effects and consider e.g. two AdS vacua:



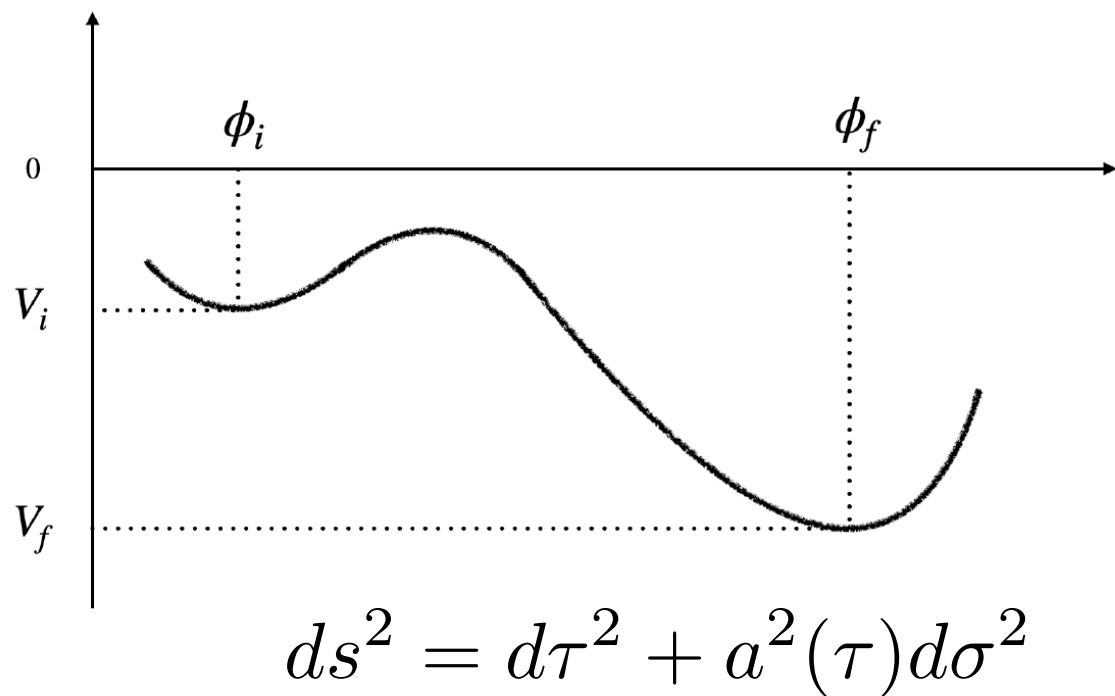
$$\Delta = \int_{\phi_i}^{\phi_f} \sqrt{1 - \frac{V(\phi)}{\Lambda(\phi)}} d\phi$$

$$\begin{aligned} \frac{1}{2}\dot{\phi}^2 - V &= \rho_E(\phi) \neq \text{const.} \\ &= -\Lambda(\phi) \propto \frac{\dot{a}^2}{a^2} \end{aligned}$$

Euclidean Einstein eqs + eom for scalar: $\frac{\Lambda'}{\Lambda} = \sqrt{\alpha(1 - V/\Lambda)}$ with $\Lambda(\phi_i) = \Lambda_i \geq V(\phi_i)$

Distance between theories

Let us include gravitational effects and consider e.g. two AdS vacua:



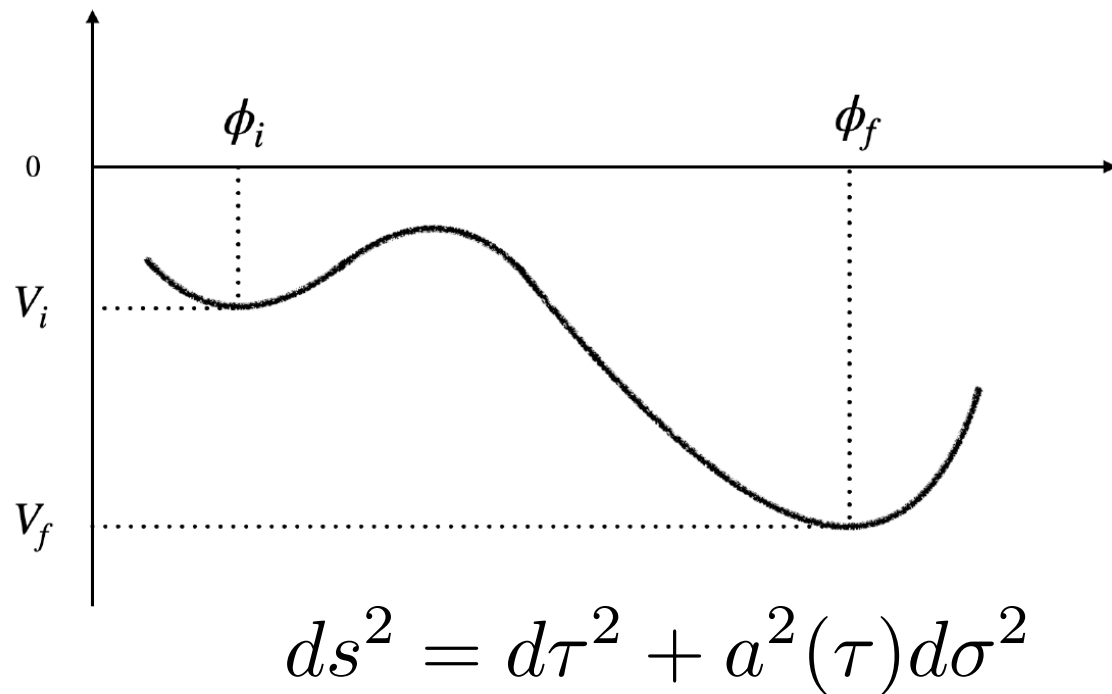
$$\Delta = \int_{\phi_i}^{\phi_f} \sqrt{1 - \frac{V(\phi)}{\Lambda(\phi)}} d\phi = \frac{1}{\sqrt{\alpha}} \log \left(\frac{\Lambda_f}{\Lambda_i} \right)$$

$$\begin{aligned} \frac{1}{2} \dot{\phi}^2 - V &= \rho_E(\phi) \neq \text{const.} & \alpha &= 4 \frac{d-1}{d-2} \\ &= -\Lambda(\phi) \propto \frac{\dot{a}^2}{a^2} \end{aligned}$$

Euclidean Einstein eqs + eom for scalar: $\frac{\Lambda'}{\Lambda} = \sqrt{\alpha(1 - V/\Lambda)}$ with $\Lambda(\phi_i) = \Lambda_i \geq V(\phi_i)$

Distance between theories

Let us include gravitational effects and consider e.g. two AdS vacua:



$$\Delta = \int_{\phi_i}^{\phi_f} \sqrt{1 - \frac{V(\phi)}{\Lambda(\phi)}} d\phi = \frac{1}{\sqrt{\alpha}} \log \left(\frac{\Lambda_f}{\Lambda_i} \right)$$

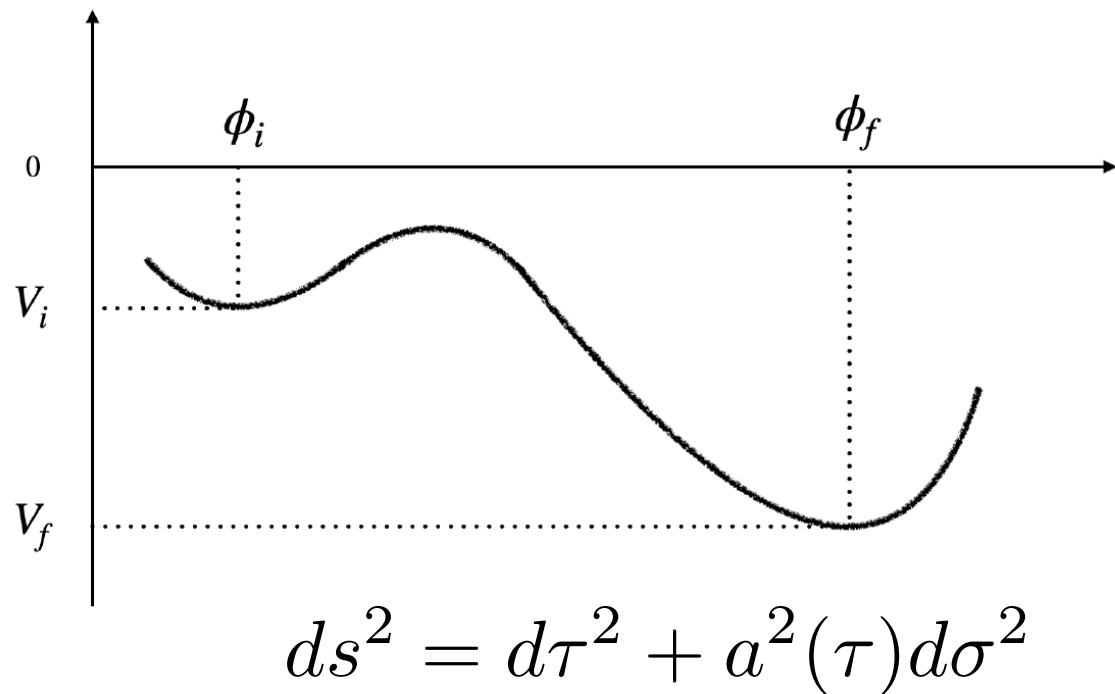
$$\begin{aligned} \frac{1}{2} \dot{\phi}^2 - V &= \rho_E(\phi) \neq \text{const.} & \alpha &= 4 \frac{d-1}{d-2} \\ &= -\Lambda(\phi) \propto \frac{\dot{a}^2}{a^2} \end{aligned}$$

Euclidean Einstein eqs + eom for scalar: $\frac{\Lambda'}{\Lambda} = \sqrt{\alpha(1 - V/\Lambda)}$ with $\Lambda(\phi_i) = \Lambda_i \geq V(\phi_i)$

- It depends on the scale Λ_i
 - If $\Lambda_i \gg V$: $\Delta \rightarrow \Delta\phi$ recovers moduli space distance
 - If $\Lambda_{i,f} = V_{i,f}$: $\Delta \rightarrow \frac{1}{\sqrt{\alpha}} \log \left(\frac{V_f}{V_i} \right)$ recovers AdS distance
- [Luest, Palti, Vafa'19]

Distance between theories

Let us include gravitational effects and consider e.g. two AdS vacua:



$$\Delta = \int_{\phi_i}^{\phi_f} \sqrt{1 - \frac{V(\phi)}{\Lambda(\phi)}} d\phi = \frac{1}{\sqrt{\alpha}} \log \left(\frac{\Lambda_f}{\Lambda_i} \right)$$

$$\begin{aligned} \frac{1}{2} \dot{\phi}^2 - V &= \rho_E(\phi) \neq \text{const.} & \alpha &= 4 \frac{d-1}{d-2} \\ &= -\Lambda(\phi) \propto \frac{\dot{a}^2}{a^2} \end{aligned}$$

Euclidean Einstein eqs + eom for scalar: $\frac{\Lambda'}{\Lambda} = \sqrt{\alpha(1 - V/\Lambda)}$ with $\Lambda(\phi_i) = \Lambda_i \geq V(\phi_i)$

- It depends on the scale Λ_i
 - If $\Lambda_i \gg V$: $\Delta \rightarrow \Delta\phi$ recovers moduli space distance
 - If $\Lambda_{i,f} = V_{i,f}$: $\Delta \rightarrow \frac{1}{\sqrt{\alpha}} \log \left(\frac{V_f}{V_i} \right)$ recovers AdS distance
[Luest, Palti, Vafa'19]

- Not a distance in the mathematical sense

(unless in the SUSY case where there is a DW solution with $\Lambda_{i,f} = V_{i,f}$)

Conclusions

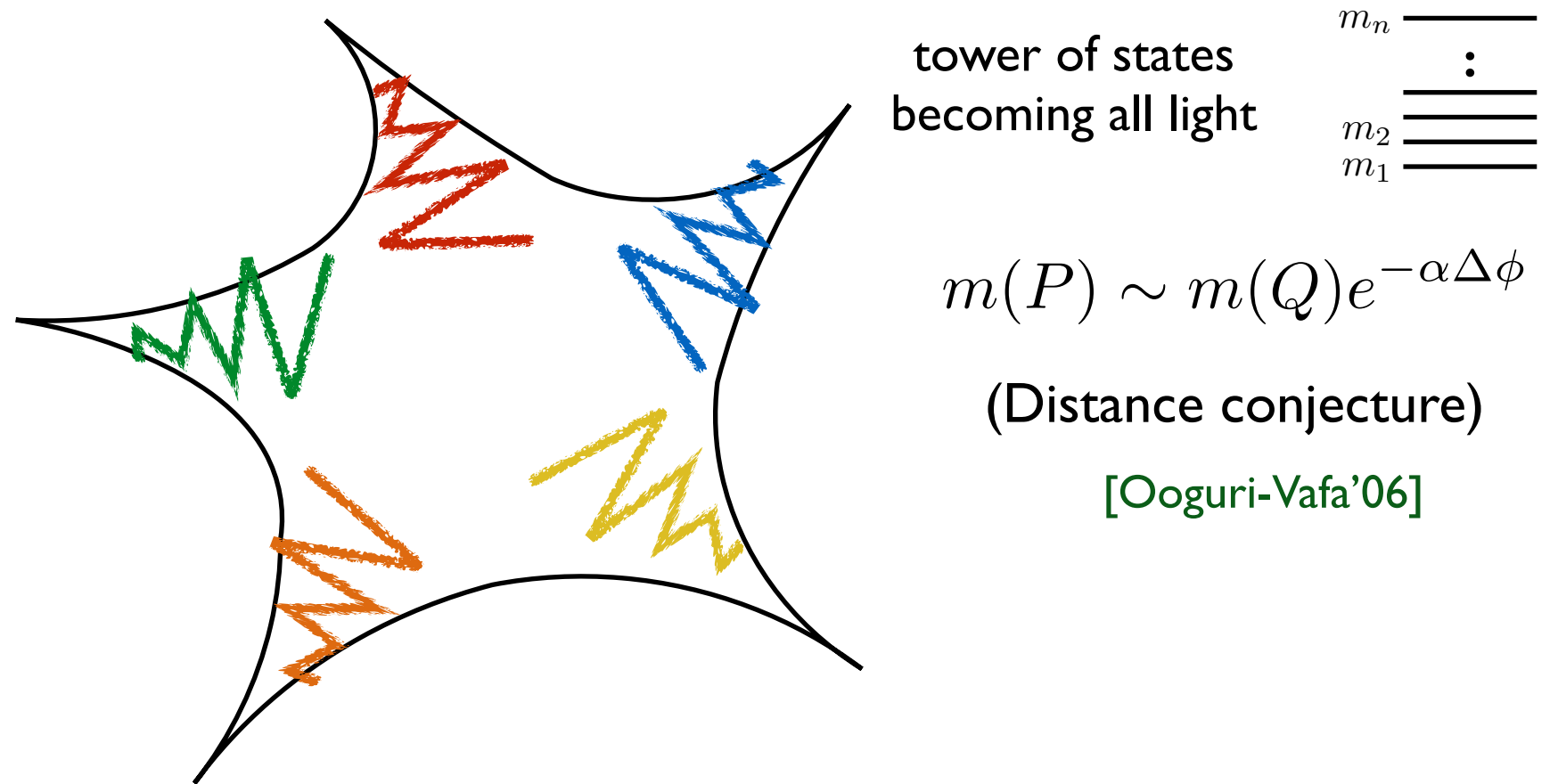
- ❖ The microscopic nature of the tower constrains the value of the exponential rate of the towers and how different limits fit together in moduli space
- ❖ We start a classification of infinite distance limits in the conformal manifold of 4d SCFTs, obtaining non-critical tensionless string limits
- ❖ The Convex Hull Distance Conjecture with $\alpha \geq \frac{1}{\sqrt{d-2}}$ in holographic AdS spaces (with a conformal manifold) if they are not scale separated
- ❖ We propose a notion of distance in the presence of a scalar potential inspired by the cobordism domain wall

Thank you!

back-up slides

Dualities

Different regions of the moduli space described by different perturbative descriptions related by dualities

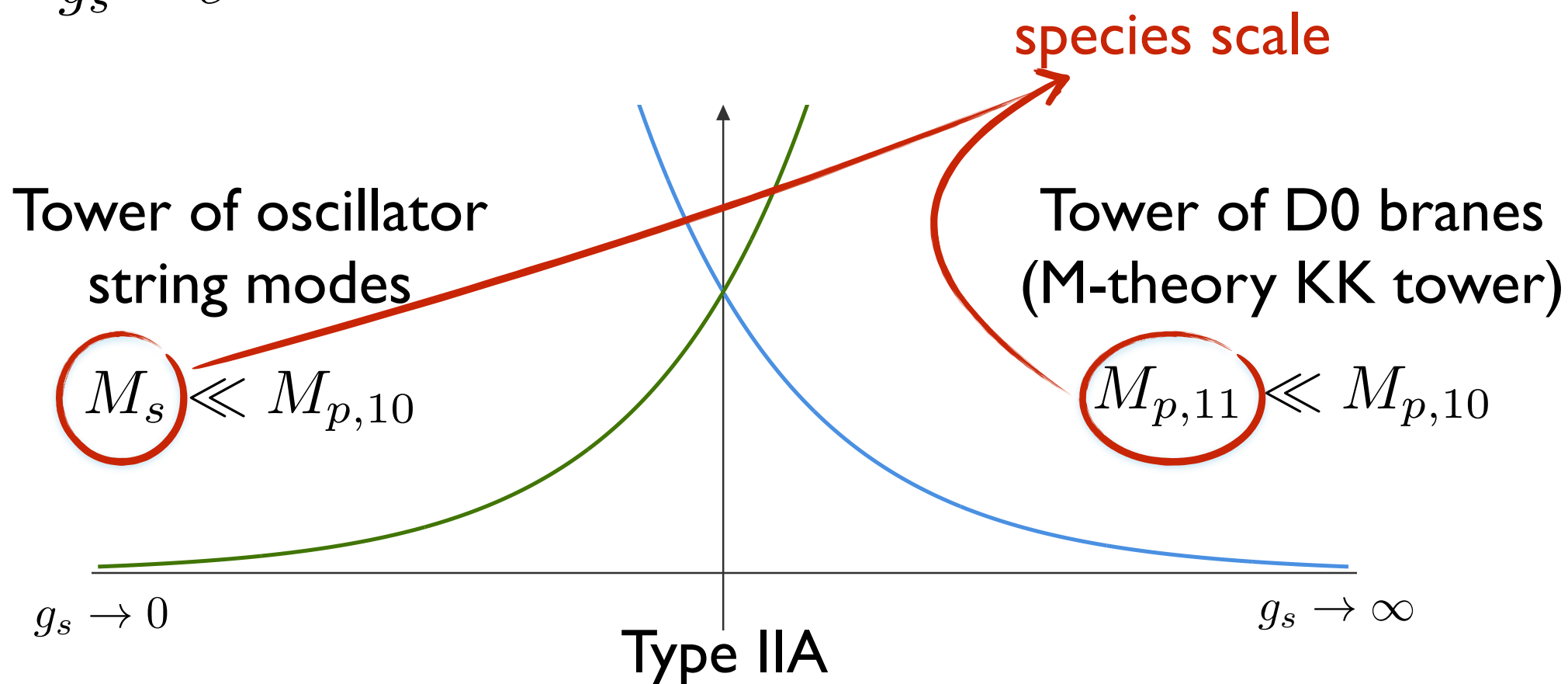


What are the possible descriptions?

How can they be combined? What are the possible dualities?

Example: 10d Type IIA

One dimensional moduli space: $S = M_p^2 \int d^{10}x \sqrt{-g} \left(\frac{1}{2} R - \frac{1}{4} (\partial\phi)^2 + \dots \right)$
 $g_s = e^{-\phi}$

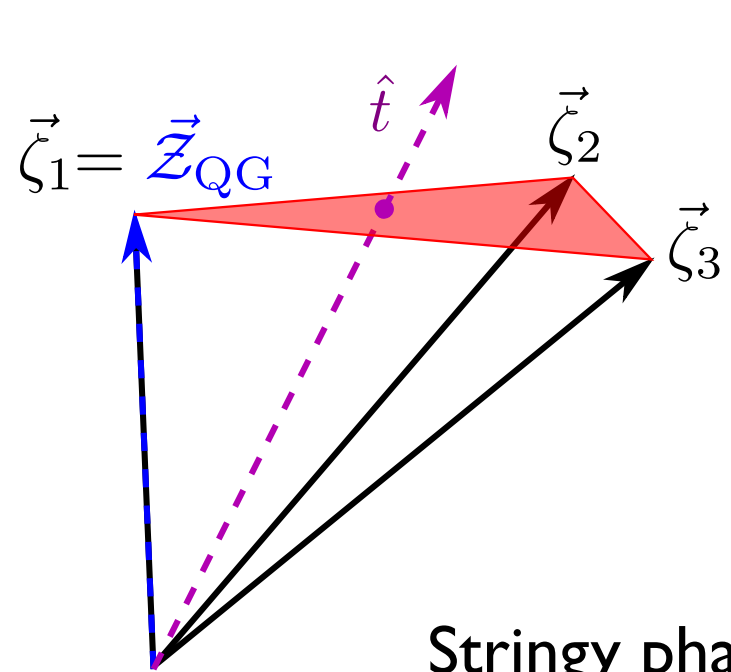


As $g_s \rightarrow 0 \rightarrow M_s = M_{p,10} g_s^{1/4} = M_{p,10} \exp\left(-\frac{1}{\sqrt{8}} \Delta\phi\right)$

As $g_s \rightarrow \infty \rightarrow m_{\text{KK}} = \frac{M_s}{g_s} = M_{p,10} \exp\left(-\frac{3\sqrt{2}}{4} \Delta\phi\right) ; M_{p,11} \sim m_{\text{KK}}^{1/8}$

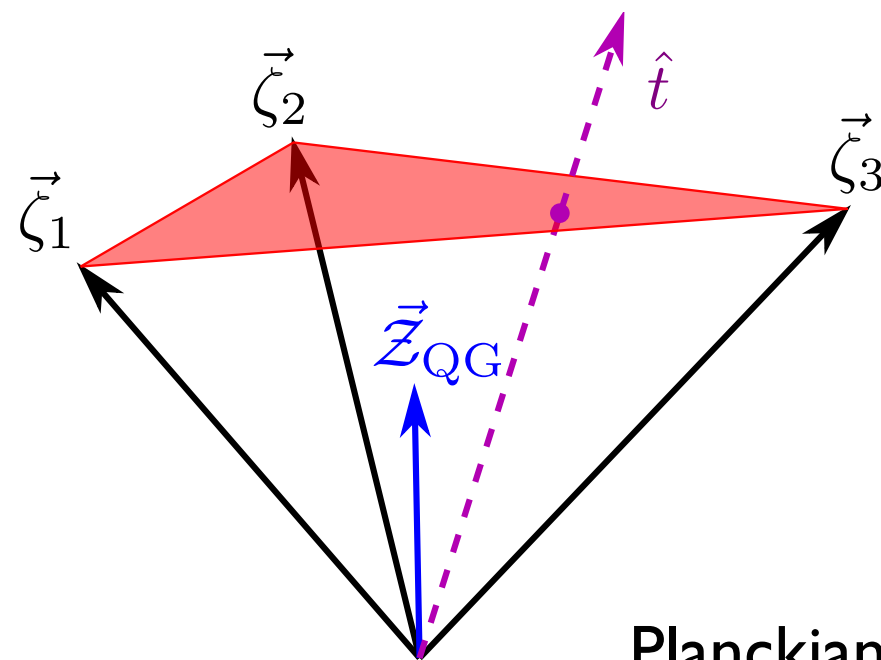
Frame simplex

The towers generate a frame simplex (convex hull of the tower vectors)
to which the species scale is orthogonal



Stringy phase

$$\Lambda_{QG} = M_s$$



Planckian phase

$$\Lambda_{QG} = M_{p,D}$$

$$\vec{\zeta}_i \equiv \frac{\vec{\nabla} m_i}{m_i}$$

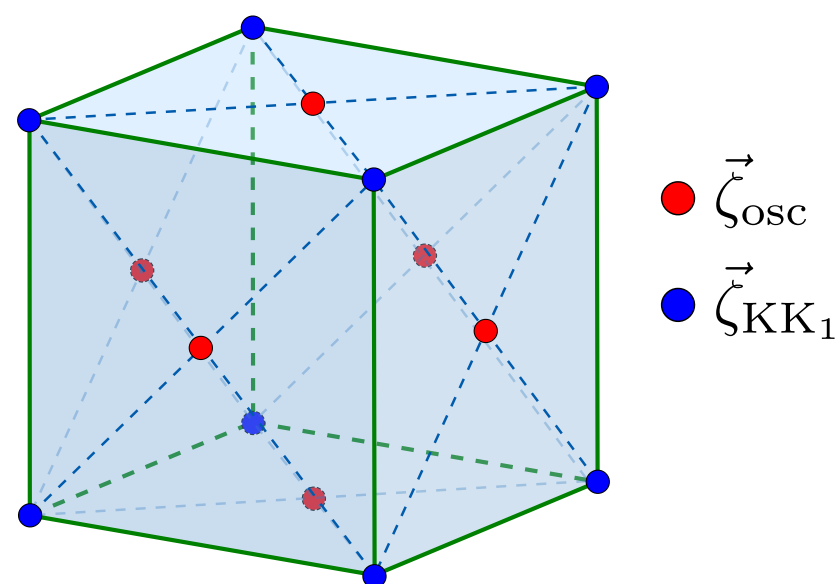
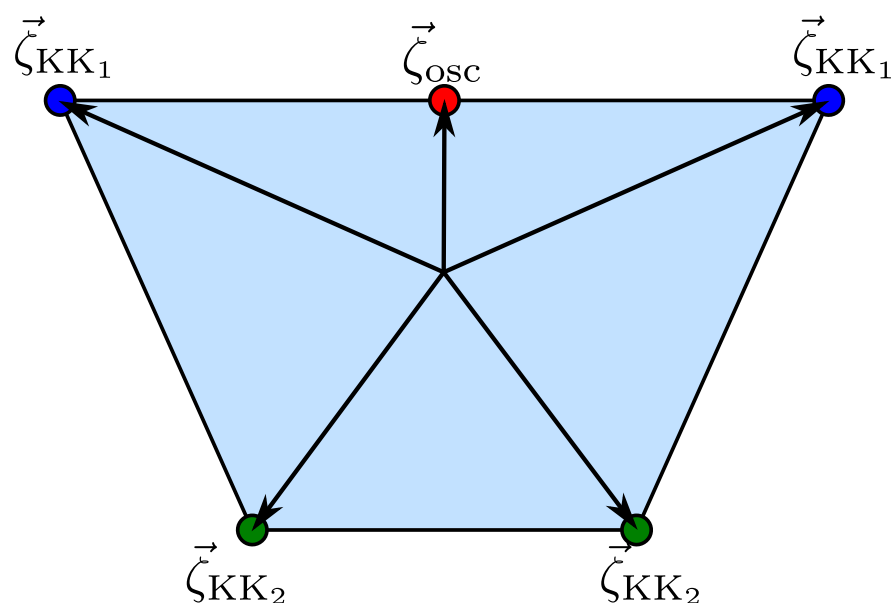
$$\vec{Z}_{QG} \equiv \frac{\vec{\nabla} \Lambda_{QG}}{\Lambda_{QG}}$$

The geometry is rigid under variations of the direction of the infinite distance limit
(as long as we do not break the assumption of regularity and we stay inside the cone)

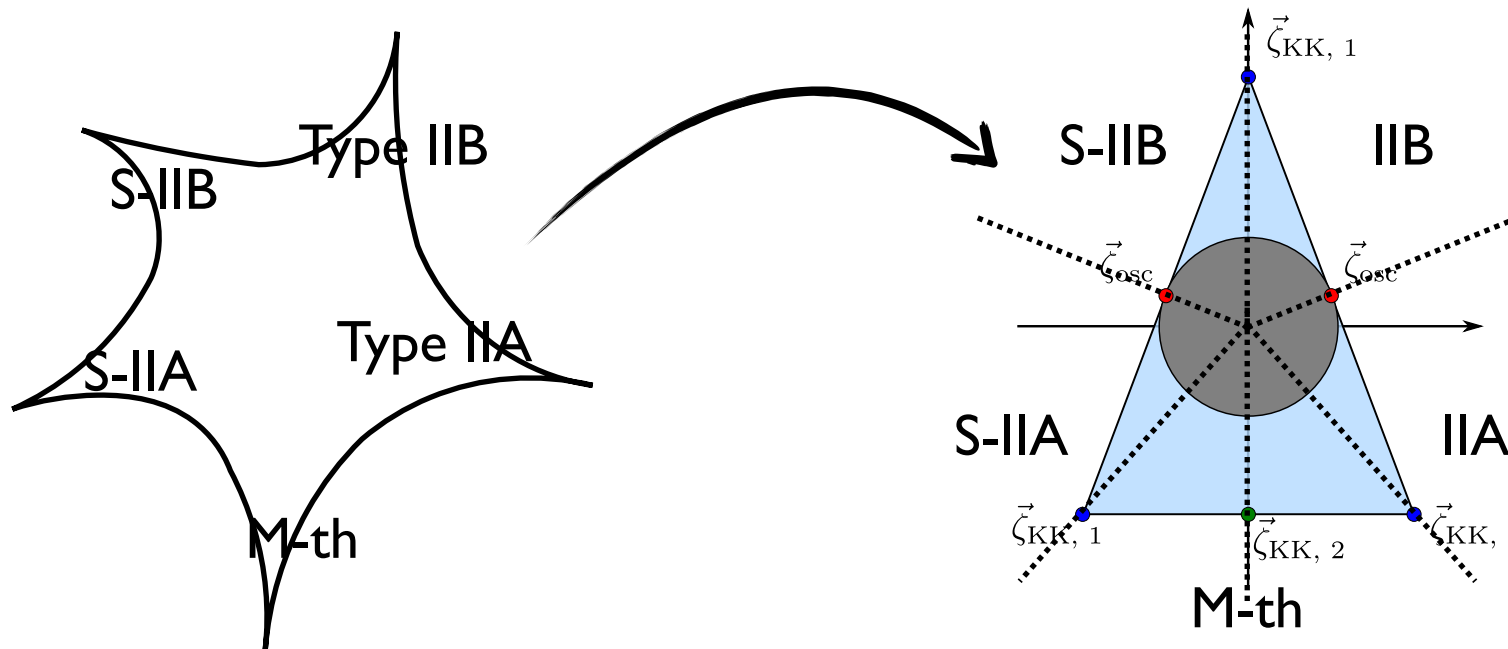
We can associate a frame simplex to each **duality frame**

Tower polytopes

In certain cases (e.g. if there is an asymptotically flat slice of the moduli space) we can **combine duality frames** by gluing individual frame simplices to form a full polytope



For example:
M-theory on T^2



Classification of 2d polytopes

This puts constraints on the structure of dualities of the moduli space

Example:

Classification of
2d polytopes
(with $D_{\max} = 11$)



The sharpened
DC is satisfied

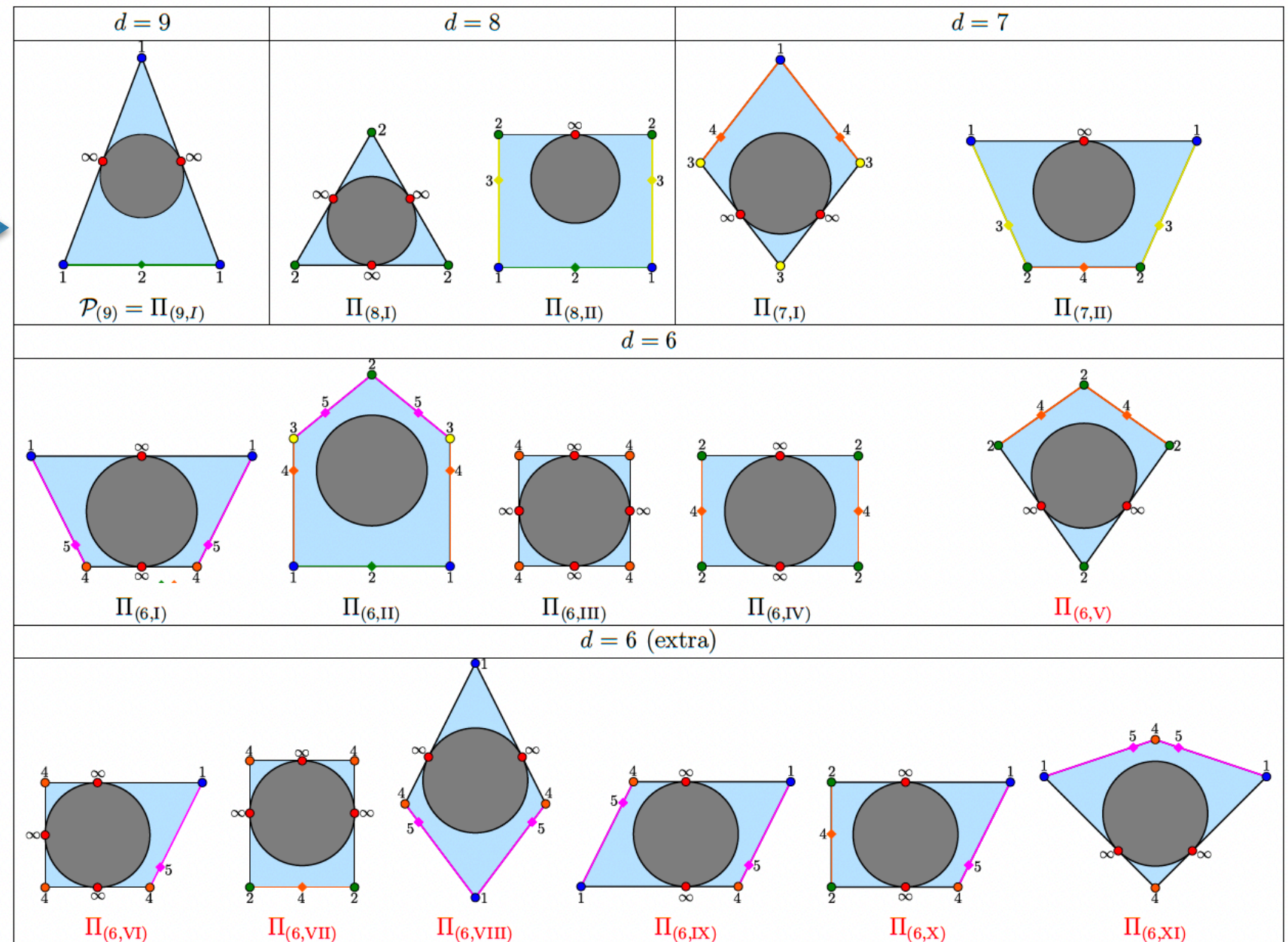
$$\alpha \geq \frac{1}{\sqrt{d-2}}$$

[Etheredge et al'22]

following the
Convex Hull

Distance Conjecture

[Calderon-Infante,Uranga,IV'20]



Many arise in string theory but others are new!

(similar story for species polytopes)


Classification of Infinite Distance Limits in $d > 2$

Let us focus on AdS_{d+1}/CFT_d with $d > 2$ [Perlmutter,Rastelli,Vafa,IV'21]
 (see also [Baume,Calderon-Infante'21])

In the free limit $g_{YM} \rightarrow 0 \longrightarrow \mathcal{O}_\tau = \text{Tr}(F^2 + \dots)$ $\tau = \frac{4\pi i}{g_{YM}^2} + \frac{\theta}{2\pi}$

By perturbation theory:

$$ds^2 = \beta^2 \frac{d\tau d\bar{\tau}}{(\text{Im}\tau)^2} \quad \text{as } \text{Im}\tau \rightarrow \infty \quad \beta^2 = 24 \dim G$$



$$\gamma_J \sim f(J) g_{YM}^2 \sim f(J) \exp\left(-\frac{d(\tau, \tau')}{\beta}\right)$$

If there is a weakly coupled AdS dual, it implies:

Infinite distance limits
at fixed AdS_5 radius



Tower of higher spin fields with an
exponential rate:

Lower bound for α !

$$\alpha = \sqrt{\frac{2c}{\dim G}} \geq \frac{1}{\sqrt{3}} \quad \text{for 4d N=2}$$

$$\geq \frac{1}{2} \quad \text{for 4d N=1}$$

Classification of Infinite Distance Limits in $d > 2$

[Bhardwaj, Tachikawa'13] [Razamat, Sabag, Zafrir'20]

Consider the full classification of 4d SCFTs with large N and simple factor for the gauge group $G = SU(N), USp(2N), SO(N)$

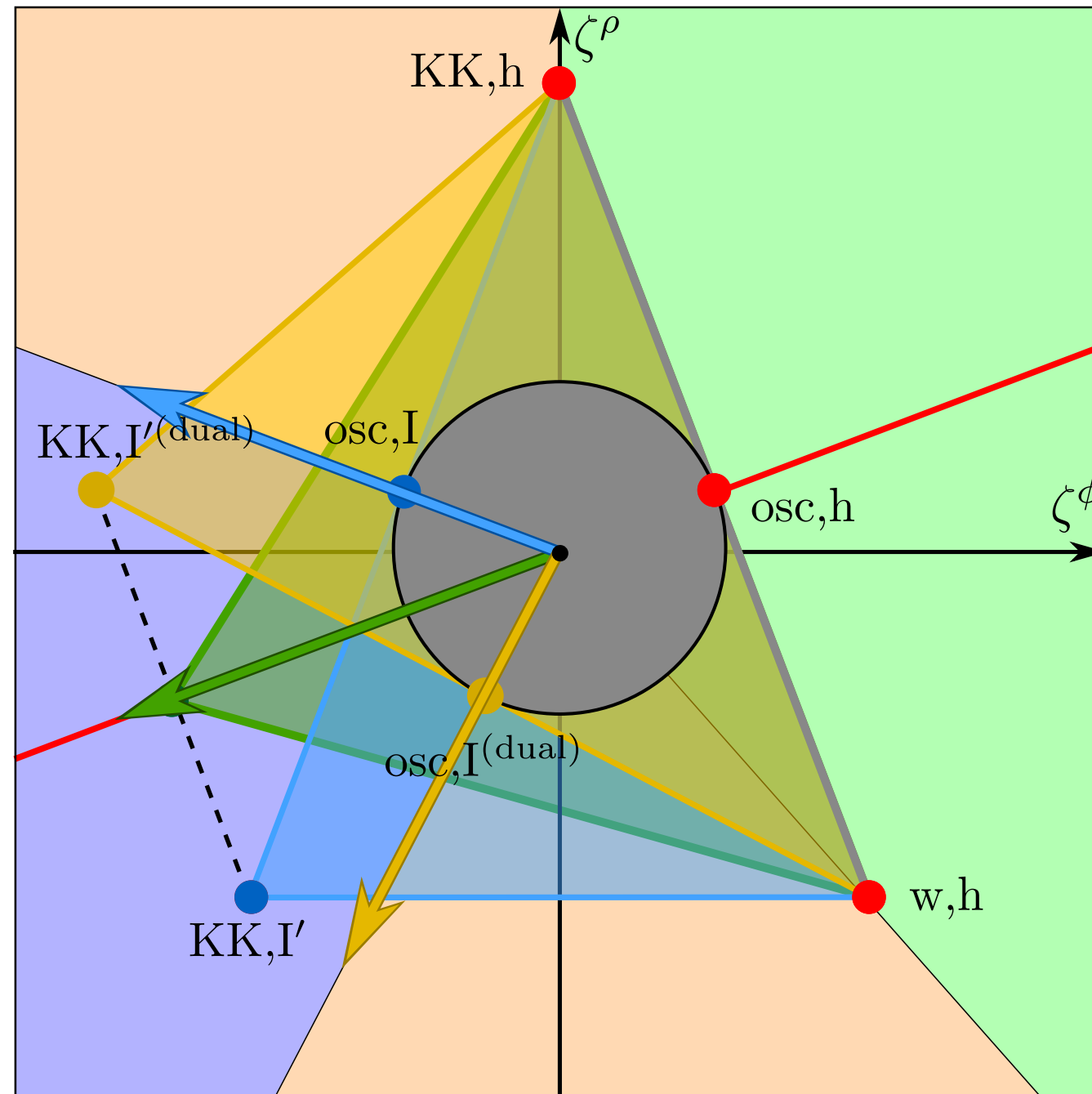
$$\mathcal{N} = 2$$

G	Hypermultiplets	c	α
$SU(N)$	$2N$ fund	$\frac{1}{6}(2N^2 - 1)$	$\sqrt{\frac{2}{3}}$
$SU(N)$	1 asym, $N + 2$ fund	$\frac{1}{24}(7N^2 + 3N - 4)$	$\sqrt{\frac{7}{12}}$
$SU(N)$	2 asym, 4 fund	$\frac{1}{12}(3N^2 + 3N - 2)$	$\frac{1}{\sqrt{2}}$
$SU(N)$	1 asym, $N - 2$ fund	$\frac{1}{24}(7N^2 - 3N - 4)$	$\sqrt{\frac{7}{12}}$
$SU(N)$	1 sym, 1 asym	$\frac{1}{12}(3N^2 - 2)$	$\frac{1}{\sqrt{2}}$
$USp(2N)$	$4N + 4 \frac{1}{2}$ fund	$\frac{1}{6}N(4N + 3)$	$\sqrt{\frac{2}{3}}$
$USp(2N)$	1 asym, 4 fund	$\frac{1}{12}(6N^2 + 9N - 1)$	$\frac{1}{\sqrt{2}}$
$SO(N)$	$N - 2$ vect	$\frac{1}{12}N(2N - 3)$	$\sqrt{\frac{2}{3}}$

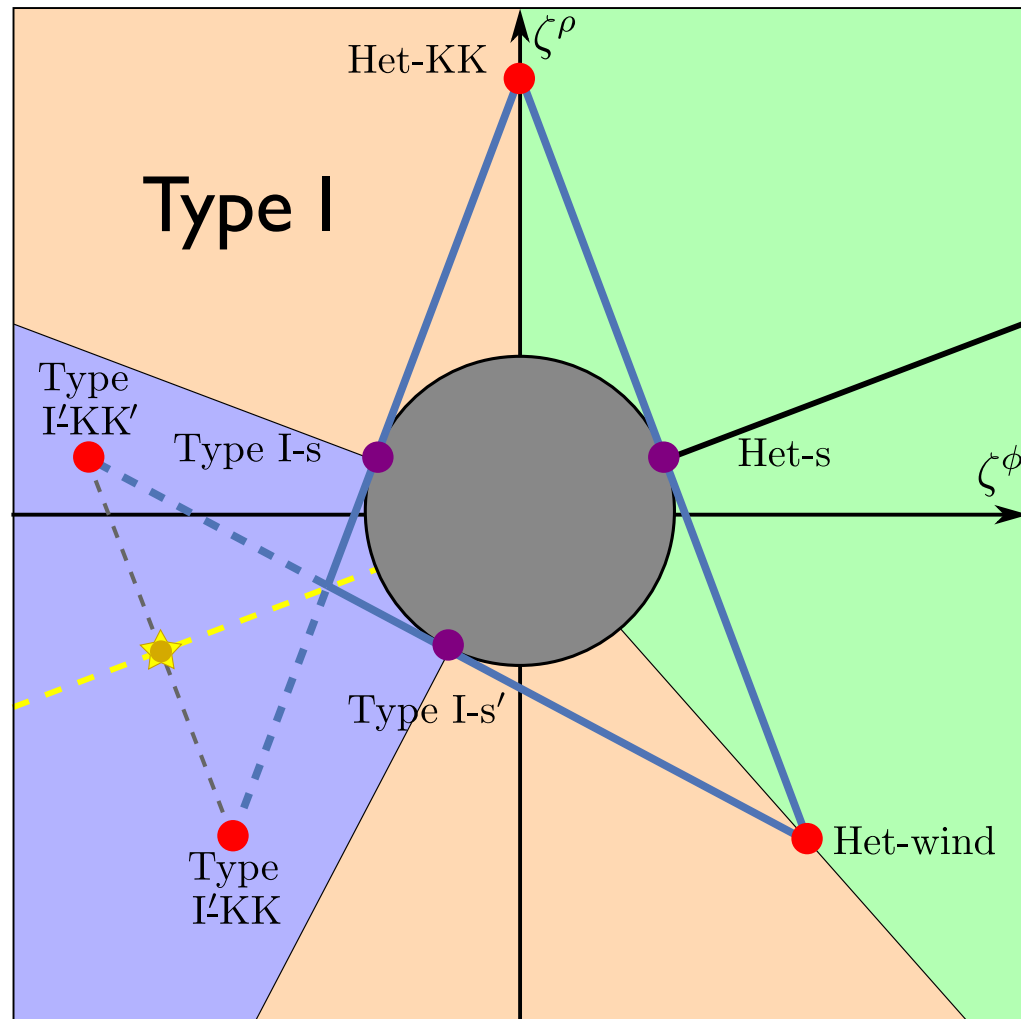
$$\mathcal{N} = 1$$

G	Theory	c	α
$SU(N)$	Table 2, #1	$\frac{1}{24}(7N^2 - 5)$	$\sqrt{\frac{7}{12}}$
$SU(N)$	Table 2, #5	$\frac{1}{24}(6N^2 + 3N - 5)$	$\frac{1}{\sqrt{2}}$
$SU(N)$	Table 3, #4	$\frac{1}{24}(7N^2 - 4)$	$\sqrt{\frac{7}{12}}$
$SU(N)$	Table 5, #4	$\frac{1}{24}(8N^2 - 3)$	$\sqrt{\frac{2}{3}}$
$USp(2N)$	Table 12, #1	$\frac{1}{24}(14N^2 + 15N - 1)$	$\sqrt{\frac{7}{12}}$
$USp(2N)$	Table 13, #9	$\frac{1}{8}(4N^2 + 8N - 1)$	$\frac{1}{\sqrt{2}}$
$USp(2N)$	Table 13, #10	$\frac{1}{24}(14N^2 + 21N - 2)$	$\sqrt{\frac{7}{12}}$
$SO(N)$	Table 18, #1	$\frac{1}{48}(7N^2 - 21N - 4)$	$\sqrt{\frac{7}{12}}$
$SO(N)$	Table 18, #2	$\frac{1}{48}(7N^2 - 15N - 2)$	$\sqrt{\frac{7}{12}}$
$SO(N)$	Table 18, #3	$\frac{1}{24}(4N^2 - 9N - 1)$	$\sqrt{\frac{2}{3}}$

Example of running decompactification



SO(32) slice



E8xE8 slice

